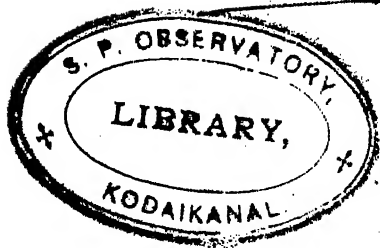


6039





Survey of India.

PROFESSIONAL PAPER—No. 16.

THE
EARTH'S AXES
AND
TRIANGULATION

BY

J. DE GRAAFF HUNTER, M.A.,
MATHEMATICAL ADVISER TO THE SURVEY OF INDIA

PUBLISHED BY ORDER OF THE GOVERNMENT OF INDIA.



Dehra Dun:

PRINTED AT THE OFFICE OF THE TRIGONOMETRICAL SURVEY.

1918.

Price Rs. 4 or 5s. 4d.

TABLE OF CONTENTS.

	PAGE
ERRATA ET ADDENDA	v
INTRODUCTION	vii
CHAPTER I. First method of finding the changes of coordinates of triangulated points due to changes in axes of the terrestrial spheroid and coordinates of origin	1
CHAPTER II. Geodesics on a spheroid	40
CHAPTER III. Changes of coordinates of triangulated points, due to changes in axes of the terrestrial spheroid, calculated along geodesics	53
CHAPTER IV. Geometrical change from one Spheroid of Reference to another ...	69
CHAPTER V. Laplace's Equation and the Choice of a Spheroid of Reference ...	75
CHAPTER VI. Strength and Adjustment of Triangulation. Mechanical Analogy ...	89
CHAPTER VII. Probable errors of triangulation before and after adjustment ...	105
CHAPTER VIII. Numerical values of the probable and actual errors in the Indian triangulation. Note on the solution of linear equations	120
CHAPTER IX. Deflections of the Plumb-line and values of "g" derived from observations of the Survey of India	162
CHAPTER X. Deflections of the Plumb-line and values of "g" derived in Turkistan (Ferghana) by the Russian observations	211
APPENDIX. Various determinations of the Axes of the Earth... ..	217
CHART No. I. Index of G. T. operations	}
CHARTS Nos. II, III, IV, V. Diagrammatic charts of the triangulation of India and Burma.	
CHART No. VI. Gravity and Deflection stations in Turkistan (Ferghana).	

at end

ERRATA ET ADDENDA.

Page

- 3 line 4 from top, column (Old value) *insert* (=6,377,276 metres) *after* 20,922,840·95 feet
 „ „ 5 „ „ „ „ (=6,356,075 metres) „ 20,853,284·03 feet
 13 „ 23 „ *for* known *read* know.
 15 After equation Nos. (18), (19), (20) *add* A
 16 line 7 from top *for* (20) *read* (20) A
 30 For value of u under Long. 82° „ 0·977 „ 0·997
 39 In equation (47) „ $6 \cos 4\lambda$ „ $6 \cos^4 \lambda$
 „ At the bottom *add a footnote* :—In equations (42) to (47) dL , $d\lambda$ are expressed in radians.
 40 At the bottom *add* an alternative proof of (2):—

For the curve whose osculating plane is normal to the spheroid we have for a small element

$$d\lambda = -\frac{ds}{\rho} \cos A$$

$$dA = -\frac{ds}{\nu} \sin A \tan \lambda$$

Hence

$$\cot A \cdot dA = \frac{\rho}{\nu} \tan \lambda \cdot d\lambda = \frac{1-e^2}{1-e^2 \sin^2 \lambda} \cdot \tan \lambda \cdot d\lambda = \frac{1-e^2}{2} \cdot \frac{d\eta}{(1-\eta)(1-e^2\eta)},$$

where $\eta = \sin^2 \lambda$

$$\therefore d \log \sin A = \frac{1}{2} \left(\frac{1}{1-\eta} - \frac{e^2}{1-e^2\eta} \right) d\eta = \frac{1}{2} d \log \frac{1-e^2\eta}{1-\eta} = -\frac{1}{2} d \log \cos^2 \phi = -d \log \cos \phi$$

$$\therefore \sin A \cos \phi = \text{constant on a geodesic.}$$

- 52 line 5 from top *for* ϕ *read* Q
 54 For equation (5) *read*

$$s = -a \sqrt{1-e^2} \left[\chi \left(1 + \frac{1}{4} h^2 - \frac{3}{64} h^4 \right) - \frac{h^2}{8} \sin 2\chi \left(1 - \frac{3}{16} h^2 \right) + \frac{h^4}{32} \cos \chi \sin^3 \chi \right] \chi'$$

- 54 For equation (12) *add a note*—The complete term to next power of e^2 is—

$$-\frac{\delta h^2}{8} \left[(2\chi - \sin 2\chi) \left(1 - \frac{3h^2}{8} \right) + \frac{h^2}{4} \sin 2\chi \sin^2 \chi \right]$$

- 55 In equations (16) and (17) *for* $\sin 2\phi$, $\sin 2\phi'$ *read* $\sin 2\lambda$, $\sin 2\lambda'$
 59 Against $10^\circ \lambda$ in the table in col. L—L' *for* $+v$ *read* $\pm v$
 68 At the end *add a para*—The values found from tables XXIX—XXXIV for the central meridian ($L=77^\circ 40'$ approximately) differ slightly from those found in Chapter I. This is due to an approximation in Chapter I (*vide* footnote on page 4).

ERRATA ET ADDENDA—(Continued).

Page

70	line 6 from top	<i>after</i>	$d\phi = u_1$	<i>insert</i>	(vide p. 54).
84	„ 13 from bottom	<i>for</i>	about	<i>read</i>	above.
96	„ 9 from top	„	$w_r - w_{r-1}$	„	$w_r - w_{r-1} - \eta_{r-1}$
„	„ 13 „ „	„	E_0	„	E
105	„ 13 from bottom	„	10^7	„	10^{-7}
115	„ 7 from bottom, col. 12	„	$+2 \cdot 20$	„	$+2 \cdot 02$
120	„ 17 „ „	„	whence	„	where
148	Against $r=10$, $u=20$, in col. k_n	„	$- \cdot 004$	„	$- \cdot 0041$.
149	In table LXXI against $r=16$, in col. 23	„	$- \cdot 2399$	„	$- \cdot 0399$.

INTRODUCTION.

The preparation of this work has extended over some four years and has been delayed by press of other work resulting mainly from a shortage of officers in the department due to the war. Its final completion has been very hurried as the author has been ordered to Mesopotamia. The end of Chapter VIII has been much abbreviated owing to there not being time to work out a sufficient number of cases to form the basis for a proper discussion. It has been decided however to publish the work at the stage it has reached rather than to wait an indefinite period for its amplification.

The origin of the research was the need which had arisen of converting geodetic results, obtained in India and referred to the now obsolete Everest spheroid, into terms of the best determined spheroid. The work soon showed that certain inconsistencies arose and that multiple values of the changes were found according to the route followed. The reason of this is that the observations of triangulation in India have been *adjusted* to fit the Everest spheroid and will not fit any other without *readjustment*. Had the spheroid of Everest been regarded merely as a reference figure and not, for purposes of triangulation, as identical with the geoid this difficulty would not have arisen: but I believe a similar method of treatment has been followed in all other countries. It would generally be impossible to avoid this treatment, as in the early stages of survey work values of deflections are not usually available. There is little reason why these deflections should not be determined roughly as the triangulation proceeds: and were this done all computations could be made quite correctly, and the deflection results would also be useful. The discrepancies involved, however, are not as large as the probable errors and offer what is more of a computation difficulty than a practical difficulty: and it is believed that the conclusions reached in the first four chapters overcome the difficulty quite satisfactorily for all practical geodetic purposes. The explanation of the discrepancy was by no means discovered at once and it is in Chapter V that the fundamental inconsistency is discussed.

Meanwhile other related subjects come under consideration and in Chapter VI reference is made to some of these. Here a complete analogy between adjustment of errors and small strains in a mechanical framework is shown to exist. This has led to the idea of "strength" of triangulation and gives a less abstract view of the adjustments by the method of least squares than was hitherto available. It enables one to picture in a tangible way the changes necessary, which will also be those most probable. A new method of adjustment of chains of triangulation has been developed, and applied to a number of Indian series.

In this connection a quantity M has been introduced as a criterion of the strength of triangulation. It is based on General Ferrero's quantity " m ", but also takes cognisance of the length of sides and general formation of a series of triangulation. This quantity has been taken out numerically for all the Indian series.

The quantity M permits of probable errors not only of side and azimuth but also of latitude and longitude being expressed at any point of the triangulation. Application has been made to all the circuits of India and the closing errors actually met with are found in good accordance with the theoretical probable errors based on M .

The case of probable errors after adjustment has also been considered. This is much more troublesome and involves very heavy work in the form of solution of numerous equations. The value of the enquiry, however, will be considerable, as an answer can be found to questions such as,

how often should extra base lines or Laplace stations be introduced? It appears that Laplace stations should be just as numerous as extra bases: and that the observation of extra bases alone is only a half measure not likely to improve matters much. It may be roughly compared to closing a traverse circuit for northing and omitting to do so for easting. The improvement due to adjustment is but briefly considered owing to force of circumstances; but the necessary equations have been solved for the cases of the N.W. Quadrilateral of the Indian triangulation and this question will be resumed when it is possible to do so. As has been mentioned the process involves the solution of a large number of simultaneous linear equations; and this has to be done for a large number of values of the absolute term in these equations. It is believed that several novelties have been introduced in this connection which may be of general interest. This question is treated on pages 126-153.

In Chapter IX the results of all the deflection observations are expressed in terms of Helmert's spheroid; the azimuth observations all having been adjusted on the Laplace conditions, none of which had been made use of during the simultaneous adjustment of the triangulation. The quantities given also permit of easy reference to any other spheroid which may at any future time be adopted. Certain observations in Russian Turkistan have also been added, as by means of the recent Indo-Russian connection these can now be stated in terms of the Indian survey. It has also been considered convenient to give a tabular statement of all determinations of g , so as to make a complete statement of gravity, not only its direction, but also its intensity. Full details of the pendulum operations are to be found in Professional Papers Nos. 10, 15.

These quantities are of immediate interest when the question of the form of the geoid and the underlying reasons for that form are considered. A start had been made with this question, which I had hoped to include in this work, but which must now be held over. An approximate method of finding the underground density anomalies by means of Poisson's equations seems possible. This is independent of the usual isostatic hypotheses, and may throw light on the whole question of isostasy.

For convenience of reference a certain number of various determinations of the figure of the earth have been included.

The computations incidental to the preparation of this work have been very heavy. Those of the earlier chapters I to IV have been performed mostly by Babu Mukundananda Acharya and Babu Hem Chandra Banerji, B.A. The solution of equations in Chapter VIII has been entirely carried out by Babu Diwan Chand Nanda, to whom I am especially indebted for his industry and accuracy in a troublesome and monotonous piece of work. The data in Chapter IX have been compiled by Babu Surendranath Mitra, M.R.A.S.

Mr. Sarat Kumar Mukerji has been responsible for the printing of the whole letter press.

DEHRA DUN,
15th Sept. 1917. }

J. DE GRAAFF HUNTER.

CHAPTER I.

First method of finding the changes of coordinates of triangulated points due to changes in axes of the terrestrial spheroid and coordinates of origin.

1. The subject of the first 16 sections of this chapter was printed in abstracted form in 1912 to draw attention to some of the difficulties of the problem, and with the hope that some light might be thrown on it at the Triennial Geodetic Conference held at Hamburg in that year.

The spheroid on which the triangulation of India has been adjusted is now believed to be considerably in error, as from the nature of the case was inevitable. Owing to possible deflection of the plumb line at the origin of the survey at Kalianpur, the values of latitude and azimuth at that point are somewhat in doubt. The problem for solution was to find the changes in latitude, longitude and azimuth of all triangulated points in India due to changes in the adopted values of the axes of the terrestrial spheroid and in the adopted coordinates of the origin of the Survey.

2. The new spheroid which is adopted in the first 16 sections of this chapter is defined by

$$a = \text{semi major axis} = 6378200 \text{ metres}$$

$$e = \text{compression or flattening} = \frac{1}{298.3} = \frac{a-b}{a}.$$

These values are given on page 173 of "*The Figure of the Earth and Isostasy, from Measurements in U.S.A.*" Washington 1909, where they are said to be Dr. Helmert's latest values.

Heretofore the axes used in the Survey of India are those due to Everest, known as "Everest's constants, first set". The numerical values are—

$$a = 20,922,931.80 \text{ feet}$$

$$e = \frac{1}{300.8}$$

All base lines of the Survey of India have been expressed in terms of the Indian ten-foot standard, known as bar *A*. The base lines were not reduced to standard British feet but were

given as some number of times $\frac{A}{10}$ feet. In making use of Everest's constants we have accordingly been taking the semi major axis as

$$20,922,981 \cdot 80 \frac{A}{10} \text{ British feet.}$$

The value of A is given* as $3 \cdot 333,318,86$ Y , Y being the British standard yard.

We accordingly have $\frac{A}{10} = 1 - \cdot 000,004,342$ from which it follows that the semi major axis which has actually been used in India is

$$a = 20,922,840 \cdot 95 \text{ British feet.}$$

Similarly

$$b = 20,853,284 \cdot 03 \quad \text{,, ,,}$$

3. Converting 6,378,200 metres into feet by means of the relation

$$1 \text{ metre} = 39 \cdot 370113 \text{ inches}$$

deduced by Benoit (see "*Rapport du Yard au mètre, Paris 1896*") we get as our new semi major axis $20,925,871 \cdot 23$ British feet, and denoting by δa and δb the corrections which have to be applied to the values used in the Survey of India, we have

$$\delta a = +3030 \cdot 28 \text{ feet} = 923 \cdot 63 \text{ metres}$$

and

$$\delta b = +2436 \cdot 78 \text{ feet} = 742 \cdot 73 \text{ metres}$$

Also since $e^2 = e(2 - e)$ and $\delta e = \frac{1}{298 \cdot 3} - \frac{1}{300 \cdot 8} = \cdot 000,027,86$ it follows that,

$$\delta e^2 = \cdot 000,055,54\dagger$$

where e is the eccentricity.

4. It is now considered that the coordinates of the origin of the survey at Kalianpur require modification. Captain G. P. Lenox Conyngham R. E. observed a group of azimuths and latitudes round Kalianpur. His results gave the mean value reduced to Kalianpur

Latitude $24^\circ 7' 11'' \cdot 57$

Azimuth of Surantal $190^\circ 27' 6'' \cdot 39$.

The values heretofore adopted in the triangulation are,

Latitude $24^\circ 7' 11'' \cdot 26$

Azimuth of Surantal $190^\circ 27' 5'' \cdot 10$.

We have to apply corrections to the origin of $+0'' \cdot 31$ in latitude and $+1'' \cdot 29$ in azimuth. As regards the old value of azimuth a correction of $-1'' \cdot 1$ was applied to the observed value by General Walker in order to make azimuths observed at other parts of the triangulation agree with geodetic values. We are now annulling this by reverting to an observed value of azimuth.

5. Accordingly it is necessary to investigate equations giving the change in coordinates due to the changes of both axes of the spheroid and of the latitude of the origin and of the azimuth of a ray through it, as exhibited in the following table

* Account of the Operations of the G.T. Survey of India Vol. I. p. 28

† This corresponds to $e = \frac{1}{300 \cdot 8}$. Everest's actual value was $\frac{1}{300 \cdot 8017}$ and the corresponding value of δe^2 is $\cdot 000,055,58$.

TABLE I.

	Old value	New value
Longitude of Kalianpur	77° 39' 17"·57
Latitude of Kalianpur ...	24° 7' 11"·26	24° 7' 11"·57
Azimuth at Kalianpur of Surantal ...	190° 27' 5"·10	190° 27' 6"·39
Length of semi major axis ...	20,922,840·95 feet	20,925,871·23 feet (= 6,378,200 metres)
Length of semi minor axis ...	20,853,284·03 feet	20,855,720·81 feet (= 6,356,818, metres)
Compression ...	$\frac{1}{300\cdot8_{017}}$	$\frac{1}{298\cdot3}$

The latest value of longitude* is merely given for convenience of reference. Any change in longitude of origin is of course immediately applicable to the whole of the triangulation by simple addition (or subtraction).

6. The old triangulation was adjusted, that is to say its apparent errors were distributed, by a process following the method of least squares as closely as was thought to be practicable in view of the great number of observed angles involved. Owing to the errors in the chosen values of the axes, the equations which the errors were made to satisfy were not quite correct. In the first place the spherical excesses of the several triangles were computed with uncorrect values of the axes: but, owing to the smallness of these spherical excesses, the change on this account is not appreciable to 0."01—the accuracy to which they were computed. None the less the error on this account being of a systematic kind—always of the same sign—will have had some small effect. With the "circuit equations" the case is less favourable. In following series of triangulation, which embrace much larger areas, the spherical excess becomes much more appreciable, and its value on the new spheroid differs from the old value by about one second in an area of 75 square degrees in Indian latitudes. This difference modifies the circuit equations. It is a smaller error than the errors generated in the triangulation, but is systematic.

The only theoretically accurate course would be to readjust all the triangulation. This would be a very large piece of work, and one object of the present paper is to avoid this labour by putting forward alternative methods, which will give the desired changes, with a departure from strict theoretical accuracy smaller than the errors due to fallible observations. The methods will also be applicable to any further changes that may be found desirable at any subsequent date.

7. The following notation is used

a = semi major axis

b = semi minor axis

e = ellipticity or compression

e = eccentricity

ρ = radius of curvature to meridian = $a(1-e^2)(1-e^2\sin^2\lambda)^{-\frac{3}{2}}$

ν = normal terminated by the minor axis = $a(1-e^2\sin^2\lambda)^{-\frac{1}{2}}$

which is the other principal radius of curvature

* Account of the Operations of the G.T. Survey of India Vol. XVII p. xv.

$$\beta^2 = \nu/\rho$$

λ = latitude

L = longitude

A = azimuth, measured from South by West

u = change in latitude due to change of origin and axes

v = „ longitude „ „ „

w = „ azimuth „ „ „

c = distance between points whose coordinates are λ, L and $\lambda + \Delta \lambda, L + \Delta L$.

As only very small values of c will be considered it is unnecessary to specify whether this distance is measured along a normal plain section or a geodesic line.

8. For small values of c

$$\left. \begin{aligned} \Delta \lambda &= -\frac{c}{\rho} \cos A \\ \Delta L &= -\frac{c}{\nu} \frac{\sin A}{\cos \lambda} \\ \Delta A &= -\frac{c}{\nu} \sin A \tan \lambda \end{aligned} \right\} \dots \dots \dots (1)$$

Differentiating* these equations with respect to A, λ, ρ, ν the corresponding changes of $\Delta \lambda, \Delta L, \Delta A$ are obtained: and remembering that $\delta \Delta \lambda = \delta u$ and $\delta \lambda = u$ etc., we obtain

$$\left. \begin{aligned} \delta u &= \frac{c}{\rho} \cos A \cdot \frac{\delta \rho}{\rho} + \frac{c}{\rho} \sin A \cdot w \\ \delta v &= \frac{c}{\nu} \frac{\sin A}{\cos \lambda} \cdot \frac{\delta \nu}{\nu} - \frac{c}{\nu} \frac{\cos A}{\cos \lambda} \cdot w - \frac{c}{\nu} \frac{\sin A}{\cos^2 \lambda} \sin \lambda \cdot u \\ \delta w &= \frac{c}{\nu} \sin A \tan \lambda \cdot \frac{\delta \nu}{\nu} - \frac{c}{\nu} \cos A \tan \lambda \cdot w - \frac{c}{\nu} \sin A \sec^2 \lambda \cdot u \end{aligned} \right\} \dots \dots \dots (2)$$

These are three simultaneous partial equations from which u, v, w are to be determined. They express the small changes in u, v, w developed along a short (elementary) line in direction of azimuth A . Before they can be integrated it is necessary to define the route along which to travel. It might be supposed at first that the only important matter was the terminal points of the route: but it will be seen later that a different result is found from each route followed. The equations are not integrable in finite terms for all routes, and two special cases are now considered, firstly along a parallel of latitude and secondly along a meridian. These cases correspond to $A = 90^\circ$ and $A = 0$ respectively. A means of dealing with the general case of an oblique curvilinear ray is given later, §13 *et seq.*

9. *Case I, when $A = 90^\circ$.* In the case of a route along a parallel of latitude it is clear that $\frac{c}{\nu \cos \lambda} = dL$ and equations (2) can accordingly be written

$$\left. \begin{aligned} -\frac{du}{dL} &= \frac{\nu}{\rho} \cos \lambda \cdot w \\ -\frac{dv}{dL} &= \frac{\delta \nu}{\nu} - \tan \lambda \cdot u \\ -\frac{dw}{dL} &= \sin \lambda \cdot \frac{\delta \nu}{\nu} - \sec \lambda \cdot u \end{aligned} \right\} \dots \dots \dots (3)$$

* The quantities $\frac{d\rho}{d\lambda}, \frac{d\nu}{d\lambda}$ were neglected as they contain the factor c^2 .

Putting $\beta^2 = \frac{\nu}{\rho}$ it follows at once from (3) that

$$\begin{aligned}\frac{du^2}{dL^2} &= -\frac{\nu}{\rho} \cos \lambda \cdot \frac{dw}{dL} \\ &= \beta^2 \sin \lambda \cos \lambda \cdot \frac{\delta \nu}{\nu} - \beta^2 u\end{aligned}$$

$$\text{or } \frac{d^2 u}{dL^2} + \beta^2 u = \frac{1}{2} \beta^2 \sin 2\lambda \frac{\delta \nu}{\nu} \quad \dots \dots \dots (4)$$

The solution of this is

$$u = \frac{1}{2} \sin 2\lambda \frac{\delta \nu}{\nu} + P \cos (\beta L) + Q \sin (\beta L) \quad \dots \dots \dots (5)$$

where P and Q are constants. Using (3) and differentiating (5) it follows that

$$\begin{aligned}-\beta^2 \cos \lambda \cdot w &= \frac{dw}{dL} = \beta \left\{ -P \sin (\beta L) + Q \cos (\beta L) \right\} \\ w &= \frac{1}{\beta \cos \lambda} \left\{ P \sin (\beta L) - Q \cos (\beta L) \right\} \quad \dots \dots \dots (6)\end{aligned}$$

To determine P and Q put $L = 0$ in (5) and (6)

$$\left. \begin{aligned}u_0 &= \frac{1}{2} \sin 2\lambda \frac{\delta \nu}{\nu} + P \\ w_0 &= -\frac{Q}{\beta \cos \lambda}\end{aligned} \right\} \quad \dots \dots \dots (7)$$

where the suffix zero indicates values at the beginning of the line.

Further, using (3) and (5)

$$-\frac{dv}{dL} = \frac{\delta \nu}{\nu} - \tan \lambda \left\{ \frac{1}{2} \sin 2\lambda \cdot \frac{\delta \nu}{\nu} + P \cos (\beta L) + Q \sin (\beta L) \right\}$$

whence

$$v - v_0 = -\frac{\delta \nu}{\nu} \cdot \cos^2 \lambda L + \frac{\tan \lambda}{\beta} \left\{ P \sin (\beta L) + Q (1 - \cos (\beta L)) \right\} \quad \dots \dots \dots (8)$$

longitude being measured from the starting point.

Expressing these equations in terms of seconds—they are at present in radian units—we write

$$\left. \begin{aligned}R'' &= \frac{1}{2} \sin 2\lambda \cdot \frac{\delta \nu}{\nu} \operatorname{cosec} 1' \\ P'' &= u''_0 - R'' \\ Q'' &= -\beta w''_0 \cos \lambda\end{aligned} \right\} \quad \dots \dots \dots (9)$$

and

$$\left. \begin{aligned}u'' &= R'' + P'' \cos (\beta L) + Q'' \sin (\beta L) \\ v'' &= v''_0 - R'' \cot \lambda \cdot \frac{L^\circ}{57.3} + \frac{\tan \lambda}{\beta} \left\{ P'' \sin (\beta L) + Q'' (1 - \cos (\beta L)) \right\} \\ w'' &= \frac{1}{\beta \cos \lambda} \left\{ P'' \sin (\beta L) - Q'' \cos (\beta L) \right\}\end{aligned} \right\} \quad \dots \dots (10)$$

Since $\nu = a(1 - e^2 \sin^2 \lambda)^{-\frac{1}{2}}$ it follows from logarithmic differentiation that

$$\frac{\delta \nu}{\nu} = \frac{\delta a}{a} + \frac{\sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \frac{\delta e^2}{2}$$

Putting in the values of $a, e, \delta a, \delta e^2$ from §§3, 4 and expanding

$$\frac{\delta \nu}{\nu} = .000,144,83 + .000,027,77 \sin^2 \lambda + .000,000,18 \sin^4 \lambda + \dots \quad (11)$$

10. *Case, II, when $A = 0$.* In the case of a route along a meridian it is clear that $\frac{c}{\rho} = -d\lambda$ and equation (2) can accordingly be written

$$\left. \begin{aligned} \frac{du}{d\lambda} &= -\frac{\delta \rho}{\rho} \\ \frac{dv}{d\lambda} &= \frac{\rho}{\nu} \sec \lambda. w \\ \frac{dw}{d\lambda} &= \frac{\rho}{\nu} \tan \lambda. w \end{aligned} \right\} \dots \dots \dots (12)$$

Differentiating

$$\rho = a(1 - e^2)(1 - e^2 \sin^2 \lambda)^{-\frac{3}{2}} \text{ logarithmically}$$

$$\frac{\delta \rho}{\rho} = \frac{\delta a}{a} - \delta e^2 \left(\frac{1}{1 - e^2} - \frac{3}{2} \cdot \frac{\sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \right)$$

whence, expanding and putting in numerical values

$$\begin{aligned} \frac{\delta \rho}{\rho} &= .000,144,83 - .000,055,54 \left(1.00668 - \frac{3}{2} \sin^2 \lambda - \frac{3}{2} e^2 \sin^4 \lambda \dots \dots \dots \right) \\ &= .000,088,92 + .000,083,31 \sin^2 \lambda + .000,000,55 \sin^4 \lambda \\ &= .000,130,78 - .000,041,93 \cos 2\lambda + .000,000,07 \cos 4\lambda \dots \dots \dots (13) \end{aligned}$$

Integrating the first equation of (12)

$$\begin{aligned} u - u_0 &= - \int \frac{\delta \rho}{\rho} d\lambda \\ &= - .000,130,78 (\lambda - \lambda_0) + \left[.000,020,97 \sin 2\lambda - .000,000,02 \sin 4\lambda \dots \dots \right]_{\lambda_0}^{\lambda} \quad (14) \end{aligned}$$

From last equation of (12)

$$\frac{1}{w} \cdot \frac{dw}{d\lambda} = \frac{\rho}{\nu} \tan \lambda = \frac{1 - e^2}{1 - e^2 \sin^2 \lambda} \cdot \tan \lambda$$

Put $y = \sin^2 \lambda$ and $dy = 2 \sin \lambda \cos \lambda d\lambda$

Then

$$\begin{aligned} d \log w &= \frac{1 - e^2}{1 - e^2 y} \tan \lambda \cdot \frac{dy}{2 \sin \lambda \cos \lambda} \\ &= \frac{1}{2} \cdot \frac{1 - e^2}{1 - e^2 y} \cdot \frac{dy}{1 - y} \\ &= \frac{1}{2} \left(\frac{1}{1 - y} - \frac{e^2}{1 - e^2 y} \right) dy \end{aligned}$$

Integrating

$$\log w = -\frac{1}{2} \log (1 - y) + \frac{1}{2} (1 - e^2 y) + \text{constant}$$

$$w = K \sqrt{\frac{1 - e^2 y}{1 - y}} \quad \text{where } K \text{ is a constant}$$

$$w = \frac{K \sqrt{1-e^2 \sin^2 \lambda}}{\cos \lambda} = \frac{K a}{\nu \cos \lambda}$$

$$\therefore w = w_0 \cdot \frac{\nu_0 \cos \lambda_0}{\nu \cos \lambda} \quad (15)$$

Again $\frac{dv}{d\lambda} = \frac{1-e^2}{1-e^2 \sin^2 \lambda} \sec \lambda \cdot K \frac{\sqrt{1-e^2 \sin^2 \lambda}}{\cos \lambda}$

$$\therefore v = K (1-e^2) \int \frac{d\lambda}{\cos^2 \lambda \sqrt{1-e^2 \sin^2 \lambda}}$$

Putting $x = \sin \lambda$ then $dx = \cos \lambda d\lambda$ and

$$v = K (1-e^2) \int \frac{dx}{(1-x^2)^{\frac{3}{2}} \sqrt{1-e^2 x^2}}$$

Now $\frac{1}{(1-x^2)^{\frac{3}{2}} \sqrt{1-e^2 x^2}} = \frac{1}{(1-x^2)^{\frac{3}{2}} \sqrt{1-e^2 + e^2 (1-x^2)}}$

$$= \frac{1}{\sqrt{1-e^2}} \cdot \frac{1}{(1-x^2)^{\frac{3}{2}}} \left\{ 1 - \frac{1}{2} k (1-x^2) + \frac{3}{8} k^2 (1-x^2)^2 \dots \right\} \quad (16)$$

where $k = \frac{e^2}{1-e^2} = 0.006,682,2$

$$\therefore v = K \sqrt{1-e^2} \int \left\{ \frac{dx}{(1-x^2)^{\frac{3}{2}}} - \frac{1}{2} k \cdot \frac{1}{\sqrt{1-x^2}} + \frac{3}{8} k^2 x^2 \sqrt{1-x^2} \dots \right\} dx$$

$$= K \sqrt{1-e^2} \left[\frac{x}{\sqrt{1-x^2}} - \frac{1}{2} k \sin^{-1} x + \frac{3}{8} k^2 \left(\frac{1}{2} x \sqrt{1-x^2} + \frac{1}{2} \sin^{-1} x \right) \dots \right]$$

$$= K \sqrt{1-e^2} \left[\tan \lambda - \frac{1}{2} k \lambda + \frac{3}{16} k^2 (\sin \lambda \cos \lambda + \lambda) \dots \right]$$

Collecting results and expressing them in terms of seconds we get

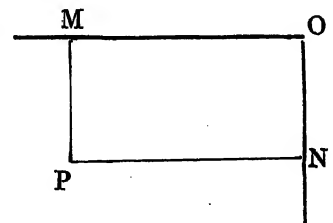
$$\left. \begin{aligned} u'' - w_0'' &= -0.000,130,78 (\lambda'' - \lambda_0'') + 4.325 (\sin 2\lambda - \sin 2\lambda_0) - 0.0035 (\sin 4\lambda - \sin 4\lambda_0) \\ v'' - v_0'' &= w_0'' \cdot \frac{\nu_0 \cos \lambda_0}{a} \sqrt{1-e^2} \left[\tan \lambda - 0.003,332,7 \lambda'' \sin 1'' + 0.000,004,2 \sin 2\lambda \right]_{\lambda_0}^{\lambda} \\ w'' &= w_0'' \cdot \frac{\nu_0 \cos \lambda_0}{\nu \cos \lambda} \end{aligned} \right\} \quad (17)$$

The value of $\log \sqrt{1-e^2}$ is $\bar{1}.9985538$.

11. With the equations (10) and (17) we can now deduce the values of u, v, w for any point P. Starting from the origin O, we may compute along the parallel OM and find the values at M. Using these as initial values we can then proceed along the meridian MP and get values for P.

Or we may first proceed along meridian ON and then along the parallel NP.

The values arrived at by the two routes are not identical. This is inevitable. The discrepancy in azimuth is the change in spherical excess on the given area from the old to the new spheroid. We shall proceed to consider the discrepancies which occur. To study these



discrepancies the values of the changes were computed to five places of decimals. In the first place it was considered convenient to take as origin for this computation the point whose latitude and longitude were $24^\circ, 78^\circ$ on the old spheroid. The double values of u, v, w , for this point differ by a small amount and in view of what follows the following mean value of w was taken:—

$$w = \frac{\frac{w_x}{x} + \frac{w_y}{y}}{\frac{1}{x} + \frac{1}{y}} = \frac{yw_x + xw_y}{x+y}$$

The suffix x indicates the value found by route ONP and the suffix y " " " " " " " " OMP and $x = \text{PN}$, $y = \text{PM}$, but x and y are always treated as positive. The values of u_x, u_y and v_x, v_y were practically identical.

Starting from this origin by means of our equations the values exhibited in the following three tables are obtained:—

TABLE II.

LATITUDE (u'').

Lat.	Long.	58°	63°	68°	73°	78°	83°	88°	93°	98°
34°	u_x	-2.07723	-2.65069	-3.09166	-3.39675	-3.56361				
	u_y	-2.50858	-2.89432	-3.20019	-3.42385					
	$u_x - u_y$	+0.43135	+0.24363	+0.10853	+0.02710		0.00000			
29°	u_x	-0.27462	-0.75738	-1.13306	-1.39881	-1.55258				
	u_y	-0.49755	-0.88329	-1.18916	-1.41282					
	$u_x - u_y$	+0.22293	+0.12591	+0.05610	+0.01401		0.00000			
24°	u_x	+1.40241	+1.01667	+0.71080	+0.48714	+0.34738	+0.29261	+0.32323	+0.43903	+0.63911
	u_y									
	$u_x - u_y$									
19°	u_x	+2.97072	+2.65699	+2.45452	+2.27510	+2.15009	+2.08045	+2.06672	+2.10900	+2.20697
	u_y	+3.20512	+2.81988	+2.51851	+2.28985		+2.09582	+2.12594	+2.24174	+2.44182
	$u_x - u_y$	-0.23440	-0.16289	-0.06399	-0.01475		-0.01487	-0.05922	-0.13274	-0.23485
14°	u_x	+4.44968	+4.27179	+4.11549	+3.98199	+3.87299	+3.78724	+3.72751	+3.69353	+3.68557
	u_y	+4.92732	+4.54158	+4.23571	+4.01205		+3.81752	+3.84814	+3.96394	+4.16402
	$u_x - u_y$	-0.47764	-0.26979	-0.12022	-0.03006		-0.03028	-0.12063	-0.27041	-0.47845

TABLE III.

LONGITUDE (v).

		58°	63°	68°	73°	78°	83°	88°	93°	98°
34°	v_x	+11.82474	+8.92142	+5.98821	+3.03302	+0.06397				
	v_y	+11.55396	+8.70715	+5.84003	+2.95734					
	$v_x - v_y$	+0.27078	+0.21427	+0.14818	+0.07568	0.00000				
29°	v_x	+11.03493	+8.28493	+5.51411	+2.72776	-0.06874				
	v_y	+10.97521	+8.23535	+5.47878	+2.70942					
	$v_x - v_y$	+0.05977	+0.04958	+0.03533	+0.01834	0.00000				
24°	v_x	+10.45012	+7.80729	+5.15103	+2.48448	-0.18914	-2.86654	-5.54441	-8.21943	-10.88833
	v_y									
	$v_x - v_y$	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000
19°	v_x	+10.03969	+7.46470	+4.83195	+2.29303	-0.30049	-2.89693	-5.49462	-8.09189	-10.68705
	v_y	+9.96451	+7.41142	+4.84792	+2.27645		-2.88086	-5.46060	-8.08863	-10.61192
	$v_x - v_y$	+0.07518	+0.05328	+0.03406	+0.01658	0.00000	-0.01657	-0.03402	-0.05826	-0.07513
14°	v_x	+9.78031	+7.23867	+4.69380	+2.14557	-0.40531	-2.95831	-5.51288	-8.06847	-10.62452
	v_y	+9.50738	+7.03876	+4.56250	+2.08063		-2.89337	-5.38170	-7.86843	-10.35171
	$v_x - v_y$	+0.27293	+0.20011	+0.13121	+0.06494	+0.00000	-0.06494	-0.13118	-0.20004	-0.27281

TABLE IV.

AZIMUTH (w).

		58°	63°	68°	73°	78°	83°	88°	93°	98°
34°	w_x	+8.78306	+6.98775	+5.13898	+3.25090	+1.33803				
	w_y	+5.83531	+4.75699	+3.64230	+2.49973					
	$w_x - w_y$	+2.94775	+2.23076	+1.49668	+0.75117	0.00000				
29°	w_x	+6.97539	+5.60208	+4.18593	+2.73775	+1.26862				
	w_y	+5.53262	+4.51023	+3.45336	+2.37006					
	$w_x - w_y$	+1.44277	+1.09185	+0.73257	+0.36769	0.00000				
24°	w_x	+5.29810	+4.31905	+3.30698	+2.26960	+1.21485	+0.15080	-0.91440	-1.97260	-3.01574
	w_y									
	$w_x - w_y$	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000	0.00000
19°	w_x	+3.71867	+3.11339	+2.48428	+1.83619	+1.17400	+0.50282	-0.17217	-0.84588	-1.51311
	w_y	+5.11995	+4.17883	+3.19579	+2.19329		+0.14573	-0.88365	-1.90627	-2.91483
	$w_x - w_y$	-1.40128	-1.06044	-0.71151	-0.35710	0.00000	+0.35709	+0.71148	+1.06039	+1.40122
14°	w_x	+2.20928	+1.96355	+1.70277	+1.42397	+1.14420	+0.85068	+0.55065	+0.24640	-0.05974
	w_y	+4.99000	+4.06789	+3.11467	+2.13762		+0.14203	-0.86122	-1.85789	-2.84036
	$w_x - w_y$	-2.78072	-2.10434	-1.41190	-0.70865	0.00000	+0.70865	+1.41187	+2.10429	+2.78062

12. Denote the distances PN, PM (in any linear unit, not in angular units) by x and y then.

$$x = (L - L_0) \nu \cos \lambda$$

$$y = \int_{\lambda_0}^{\lambda} \rho d\lambda.$$

By inspection of the numbers shown in tables II, III, IV the following equations are found to be approximately true:

$$\left. \begin{aligned} u_x - u_y &= Ax^2y \\ v_x - v_y &= Bxy^2 \\ w_x - w_y &= Cxy \end{aligned} \right\} \dots \dots \dots (18)$$

where A, B, C are quantities varying slightly with the latitude, but which may be treated as constants with their mean values over any area with which we shall need to deal. The last equation simply expresses that the closing error in azimuth is equal to the change in spherical excess.

Now $u_x - u_y$ is what we will call the "closing error in latitude" in proceeding round the circuit OMPN; $v_x - v_y$ and $w_x - w_y$ being corresponding quantities for longitude and azimuth. Over any elementary area

$$\left. \begin{aligned} dU &= d(u_x - u_y) = 2Ax dx dy \\ dV &= d(v_x - v_y) = 2By dx dy \\ dW &= d(w_x - w_y) = C dx dy \end{aligned} \right\} \dots \dots \dots (19)$$

By integrating over any area the closing error of the circuit enclosing that area is found.

To find the values of u, v, w then which would be obtained by proceeding along any route it is only necessary to find the values of u_x, v_x, w_x (or u_y, v_y, w_y) and apply the closing error with the correct sign. Integrating (18) it follows for moderate areas,

$$\left. \begin{aligned} U &= 2A \bar{x}a \\ V &= 2B \bar{y}a \\ W &= Ca \end{aligned} \right\} \dots \dots \dots (20)$$

where a is the area of the circuit and \bar{x}, \bar{y} are the coordinates of its centre of gravity. We say for moderate areas because the coordinates x and y are curvilinear: but for the areas we shall require to apply the formulæ to, x and y may be treated as rectilinear coordinates.

13. By means of the above equations it is possible to find the result of the change of axes and origin as computed along any line of any curvature or along any route whatever, by computing first along a parallel and then along a meridian (or in the reverse order) and then applying the "closing errors" of the circuit formed by the line in question and the parallel and meridian.

This, then, would solve the problem as far as solitary lines were concerned. When we come to a network of lines the case is different, for several values of the changes which occur at a point can be found corresponding to the several possible routes by which the point can be reached. In view of the fact that most of the triangulation of India is along meridian or parallel (see triangulation chart at end), the following procedure is suggested:—

- (1). Select central meridian and parallel for India (Burma will be dealt with separately). The selected meridian is 78° and the selected parallel $24^\circ N$.

(2). Assume the values of u, v, w found by the formulæ on these lines, which we will call axes, to be correct. We have then to distribute the closing errors in PM and PN. (see fig. §11).

(3). If PM is a meridional series the computations fixing the *length* PM depend only in a small measure on the size of the earth's axes. The way in which these axes have come in is through the spherical excess. In nearly all triangulation in the Survey of India, the spherical excess is such a small quantity that the change of axes proposed will not appreciably affect it (to $0''.01$). There is reason then for assuming that the length PM is correct. In the same way the length PN may be regarded as correct. If then u_y and v_x are taken for the changes in coordinates of P there should be no error to the first order: and as the values of u, v are so small the second order quantities may surely be neglected.

(4). This process would hold for the corners of circuits formed by meridian and longitudinal series, though some modification would be more correct for oblique series. In the Indian triangulation meridional and longitudinal series are the rule. Oblique series occur practically only along the coast of the Bay of Bengal and along the first range of the Himalayas. (See index chart of triangulation at end). As far as latitude and longitude are concerned we should not be committing much error in accepting the values of latitude and longitude, u_y and v_x .

(5). Now consider the azimuth change. This can be found from the change in position of two contiguous points. If we take two points originally on the same latitude whose changes are u_y and $u_y + \frac{\delta u_y}{\delta L} dL$ the azimuth change on the line joining them is

$$\rho \frac{\delta u_y}{\delta L} dL \frac{1}{v \cos \lambda \cdot dL} = \frac{\rho}{v \cos \lambda} \cdot \frac{\delta u_y}{\delta L} = w_y$$

whereas the azimuth change deduced from two points originally on the same longitude is

$$\frac{v \cos \lambda}{\rho} \frac{\delta v_x}{\delta \lambda} = w_x$$

It has been seen that the azimuth closing errors is $C \cdot xy$ where $OM = x$ $ON = y$. and C is a quantity which varies slightly with the latitude. Treating C as a constant and equal to its mean value over the area in consideration is permissible. This will be satisfied if the azimuth error is put into the lines PN and PM to amount proportional to their lengths.

This gives

$$\frac{\frac{w_x}{y} + \frac{w_x}{x}}{\frac{1}{y} + \frac{1}{x}} = \frac{xw_y + yw_x}{x + y}$$

as the best correction to the azimuth at P.

(6). The difference $w_y - w_x$ is not to be regarded as an error contained in the value of the angle $M P N$. Its effects is to alter the curvature of the lines $P M$ and $P N$.

(7). In the case then of a gridiron system of meridional and longitudinal series at regular intervals all of equal weight, it seems that the best values we could assign to the changes are $u_y, v_x, \frac{xw_y + yw_x}{x + y}$. In this case the next step in correcting the triangulation would be to find the changes of intermediate points on the series as follows:—

$$\frac{u_y}{v_x} \text{ for change in } \begin{array}{l} \text{latitude} \\ \text{longitude} \end{array} \text{ along a } \begin{array}{l} \text{meridian} \\ \text{parallel} \end{array}$$

and for the other coordinate and azimuth to simply interpolate between the terminal values.

(8). A difficulty arises when the actual points of the triangulation series are considered. For if the above formula for two adjacent points A and B is used, the difference of coordinates will not exactly give the correct change of azimuth. Adopting the rule of computing the azimuth from the coordinates, a different azimuth change on a ray going east from that found on a ray going south is arrived at. That is to say the actual change of $w_x - w_y$ would be forced into the single angle formed by these rays.

To avoid this it appears better to take the azimuth change to be $\frac{xw_y + yw_x}{x + y}$ and the change in coordinates of *one point only* to be given by u_y, v_x , and compute the coordinates of second adjacent point from this with the corrected azimuth (*i. e.* old azimuth $+ \frac{xw_y + yw_x}{x + y}$) and the old value of the distance c .

The computation alluded to in (8) may be performed with tables such as are given in the "Auxiliary Tables"* prepared for the new values of the axes: or we may at once deduce the changes in the position of the second point B by differentiation of the equations for $\Delta\lambda, \Delta L$ and ΔA .

The equations are:—

$$\Delta\lambda = -\frac{c}{\rho} \cos A - \frac{1}{2} \frac{c^3}{\rho^3} \sin^3 A \tan \lambda = \delta_1 \lambda + \delta_3 \lambda$$

$$\Delta L = -\frac{c}{\nu} \frac{\sin A}{\cos \lambda} + \frac{1}{2} \frac{c^3}{\nu^3} \frac{\sin 2 A \tan \lambda}{\cos \lambda} = \delta_1 L + \delta_3 L$$

$$\Delta A = -\frac{c}{\nu} \sin A \tan \lambda + \frac{1}{4} \frac{c^3}{\nu^3} (1 + 2 \tan^2 \lambda) = \delta_1 A + \delta_3 A$$

being equations (1) carried to an extra term in consideration of the larger value of c now contemplated.

Differentiating we have at once

$$\delta \Delta \lambda = \delta_1 \lambda \left(-\frac{\delta \rho}{\rho} - \tan A \cdot w \right) + \delta_3 \lambda \left(-\frac{\delta \rho}{\rho} - \frac{\delta \nu}{\nu} + 2 \cot A \cdot w + \frac{2}{\sin 2\lambda} u \right)$$

$$\delta \Delta L = \delta_1 L \left(-\frac{\delta \nu}{\nu} + \cot A \cdot w + \tan \lambda u \right) + \delta_3 L \left\{ -\frac{2\delta \nu}{\nu} + 2 \cot 2A \cdot w + (\cot \lambda + 2 \tan \lambda) u \right\}$$

* Auxiliary Tables of the Survey of India, Dehra Dūn, 1906.

$$\delta\Delta A = \delta_1 A \left(-\frac{\delta\nu}{\nu} + \cot A.w + \frac{2}{\sin 2\lambda} u \right) + \delta_2 A \left(-\frac{2\delta\nu}{\nu} + \frac{4 \tan \lambda \sec^2 \lambda}{1 + 2 \tan^2 \lambda} u + 2 \cot 2A.w \right)$$

In above u, v, w are the values found for one end of the base A : the values for the other end B are then $u + \delta\Delta\lambda, v + \delta\Delta L, w + \delta\Delta A$.

A third method is to reach B by proceeding first along the parallel AC through A and then down the meridian CB through B , by means of the formulæ (or tables) already given: and then by applying the closing error of the area ACB .

It appears, then, that the expressions $u_y, v_x, \frac{xw_y + yw_x}{x + y}$ may be taken to represent the changes in latitude, longitude and azimuth respectively of any point in India (excluding Burma) with the restriction that adjacent points must be treated differently, the changes for the second point being deduced by one of the three methods just explained. On this basis the results may be given in convenient tabular form. They will represent the changes with accuracy considerably greater than the accuracy with which the points can be considered to be fixed in space by triangulation.

14. These values are believed to be satisfactory for all the purposes for which they can be used. As far as map producing goes the discrepancies are negligible. For geodetic purposes we require to know the absolute corrections to latitudes, longitudes and azimuths of a base where a junction is to be made with another survey—such as the Russian survey, or the Burma survey. We can do this as described in § 13 for one end of the base and then compute the coordinates of the other end of the base from a knowledge of its length. In the case of Burma the triangulation has not yet been adjusted. It will perhaps be adjusted with the new values of the axes and made to fit on to the most eastern series of the North-East Quadrilateral, *viz.*, the Shillong Meridional Series, after this has been corrected for change of axes.

We also wish to know corrections to triangulated latitudes or azimuths at stations where these quantities have also been observed astronomically, so as to know the actual plumb-line deflections. As regards latitude we have uncertainty of perhaps $0''.1$ on account of axes change after leaving the central latitude by 10° , *i. e.* one part in 360,000 which is of the order of accuracy of our base-lines in India. The error generated in the triangulation must eventually be greater than this. The same argument holds as regards the azimuth, where the uncertainty of change due to change of axes, and due to error generated in triangulation are necessarily larger numbers when expressed in seconds of arc than occur in the latitude. The astronomic observations for azimuth are less precise, considered from point of view of plumb-line deflection, than the latitude observations. Apart from these considerations an error in plumb-line deflection in latitude of $0''.1$ is of little account. In India we have plumb-line deflections of over $50''$ and, at least at present, tenths of second are too minute to be taken account of in any discussion of deflections.

15. It seems then that the method sketched above is sufficiently precise for the geodetic uses to which the results can be put, and higher accuracy could not be applied with advantage to the results of triangulation. The method of § 13 is applicable to points which can be reached by either route (meridian or parallel) without the route departing out of the region of triangulation. Thus while it applies to all the triangulation in India which has been adjusted, it could not be fairly applied without modification to Burma, for this would imply the existence of triangulation across the Bay of Bengal. As the Burma triangulation remains to be adjusted, this does not matter and it will only be necessary to apply the method as far as the Shillong Meridional Series, which can be done very satisfactorily, the more so as our selected central latitude crosses this series.

16. The Survey of India was asked in 1912 by the Siamese Survey Department to furnish the best possible values of the coordinates of Bangkok. The way in which this has been done will serve as a good illustration of the method of using the closing error to determine the changes which occur along a route which is neither meridional nor longitudinal. As far as longitude 90° the route may be taken to follow the central parallel, latitude 24° (*see triangulation chart at end*). From there it proceeds along the Burma Coast Series down to latitude $13^\circ 45'$, and thence to Bangkok along latitude $13^\circ 45'$. In this case then we first compute along parallel 24° up to longitude 98° : we then proceed along meridian 98° down to latitude $13^\circ 45'$. The result at this point is found from tables II, III, IV by extrapolation to be

$$u_y = + 4.248$$

$$v_y = -10.339$$

$$w_y = - 2.837$$

Now treating longitude 98° as axis from which x is measured, we evaluate the closing errors over the area between the Coast series and latitude 24° and meridian 98° and get

$$\Sigma Ax^2y = U = +.017$$

$$\Sigma Bxy^2 = V = -.026$$

$$\Sigma Cxy = W = +.426$$

Hence the changes at latitude $13^\circ 45'$, longitude 98° , as determined by the route following the Burma Coast Series, are $u_y + U$, $v_y + V$, $w_y + W$.

One further correction remains. The Bangkok Series which emanates from this point is expressed in "preliminary terms"—it was computed from preliminary values of the side from which it emanates. Later values of this side, found after the Coast Series had been computed from the preliminary value of the side, require the following changes to be applied to the beginning of the Bangkok Series, viz.

in	latitude	...	$-1''.80$
	longitude	...	$-0''.17$
	azimuth	...	$+5''.50$

Combining these we arrive at the changes to be made at latitude $13^\circ 45'$, longitude 98° .

	Latitude	Longitude	Azimuth
By parallel and meridian route
	$+4.248$	-10.339	-2.837
Correction to bring into terms of Burma Coast Series route
	$+0.017$	-0.026	$+0.426$
Correction from "preliminary terms"
	-1.80	-0.17	$+5.50$
Total	$+2.47$	-10.54	$+3.09$

With these initial values by computing along parallel $13^\circ 45'$ up to longitude $100^\circ 33' 3''.5$ the old value of the longitude of Phukhao Thong Station* in Bangkok we get the changes

$$u = + 2''.34$$

$$v = - 11''.82$$

$$w = + 2''.9$$

To bring into Greenwich terms the farther correction $- 2' 27''.18$ is required to the longitude, the final corrections being

$$\begin{aligned} &+ 2''.34 \text{ in latitude} \\ &- 2' 39''.00 \text{ in longitude} \\ &+ 2''.9 \text{ in azimuth} \end{aligned}$$

* Phukhao Thong Station is the most easterly triangulated point shown on the triangulation chart.

Owing to an unfortunate confusion of the quantities u, v, w , with u, v, w , the following corrections were wrongly supplied to the Royal Survey Department Siam in 1912

$$\begin{array}{ll} +1''\cdot78 & \text{in latitude} \\ -2'39''\cdot32 & \text{longitude} \\ +5''\cdot7 & \text{azimuth} \end{array}$$

17. So far the particular case of definite numerical values of $\frac{\delta a}{a}, \delta e^2, u_0, w_0$ has been considered. It is desirable to put the solution in a form in which the results of any desired change can be calculated rapidly.

Let u_1, u_2, u_3, u_4 be the changes in latitude respectively due to $\delta a = 1000$ metres, $\delta b = 1000$ metres, $u_0 = 1''$, $w_0 = 1''$ with corresponding notation for v and w . Since the quantities involved are small it is clear that

$$u = \delta a \cdot u_1 + \delta b \cdot u_2 + u_0 \cdot u_3 + w_0 \cdot u_4 \dots \dots \dots (18)$$

with similar equations for v and w : where $\delta a, \delta b$ are expressed in kilometers and u_0, w_0 in seconds. If values of u_1, u_2, u_3, u_4 are tabulated, equation (18) will enable the quantities u, v, w to be evaluated for any desired case. Of the quantities $\delta a, \delta b, \delta e^2$ any two may be regarded as independent, the third being determined from the result of differentiating $b = a \sqrt{1-e^2}$ logarithmically.

$$i. e. \quad \frac{\delta a}{a} - \frac{\delta b}{b} = \frac{1}{2} \frac{\delta e^2}{1-e^2} \dots \dots \dots (19)$$

When $\delta a = \delta b = 1000$ metres = $39370\cdot113$ inches = $3280\cdot843$ feet

$$\frac{1}{2} \frac{\delta a}{a} \operatorname{cosec} 1'' = 16\cdot1718 \quad \text{of which the logarithm is } 1\cdot2087596$$

$$\frac{1}{2} \frac{\delta b}{b} \operatorname{cosec} 1'' = 16\cdot2258 \quad ,, \quad ,, \quad ,, \quad 1\cdot2102058$$

Also

$$\begin{aligned} \frac{\delta v}{v} &= \frac{\delta a}{a} + \frac{\sin^2 \lambda}{1-e^2 \sin^2 \lambda} \cdot \frac{\delta e^2}{2} \\ &= \frac{\delta a}{a} \left(1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) - \frac{\delta b}{b} \cdot \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \dots \dots \dots \text{by (19)} \end{aligned}$$

Hence equation (9) may be written, omitting the dashes

$$\left. \begin{aligned} R &= \sin 2\lambda \left\{ 16\cdot1718 \left(1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) \delta a - 16\cdot2258 \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \delta b \right\} \\ P &= u_0 - R \\ Q &= w_0 \beta \cos \lambda \end{aligned} \right\} \dots \dots \dots (20)$$

where $\delta a, \delta b$ are expressed in kilometres and u_0 and w_0 in seconds.

18. *Along a parallel of latitude*

Case I, when $\delta a = 1, \delta b = 0, u_0 = 0, w_0 = 0$

$$\text{By (20) } \left. \begin{aligned} R &= 16\cdot1718 \sin 2\lambda \left(1 + \frac{(1-e^2) \sin^2 \lambda}{1-e^2 \sin^2 \lambda} \right) = -P \\ Q &= 0 \end{aligned} \right\} \dots \dots \dots (21)$$

and (10) becomes

$$\left. \begin{aligned} u &= R(1 - \cos(\beta L)) \\ v - v_0 &= -R \left(\frac{L^2 \cot \lambda}{57.3} + \frac{1}{\beta} \tan \lambda \sin(\beta L) \right) \\ w &= -\frac{R}{\beta} \sec \lambda \sin(\beta L) \end{aligned} \right\} \quad \dots (22)$$

where L is expressed in degrees.

Case II, when $\delta a = 0$, $\delta b = 1$, $u_0 = 0$, $w_0 = 0$

From (20)

$$R = -16.2258 \sin 2\lambda \cdot \frac{(1 - e^2) \sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \quad \dots (23)$$

and equations (22) hold for this case also.

Case III, when $\delta a = 0$, $\delta b = 0$, $u_0 = 1$, $w_0 = 0$

In this case $\frac{\delta v}{v} = 0$ so that

$$\left. \begin{aligned} R &= 0 \\ P &= u_0 = 1 \\ Q &= 0 \end{aligned} \right\} \quad \dots (24)$$

and

$$\left. \begin{aligned} u &= \cos(\beta L) \\ v - v_0 &= \frac{1}{\beta} \tan \lambda \sin(\beta L) \\ w &= \frac{1}{\beta} \sec \lambda \sin(\beta L) \end{aligned} \right\} \quad \dots (25)$$

Case IV, when $\delta a = 0$, $\delta b = 0$, $u_0 = 0$, $w_0 = 1$

$$\left. \begin{aligned} R &= P = 0 \\ Q &= -\beta \cos \lambda \end{aligned} \right\} \quad \dots (26)$$

and

$$\left. \begin{aligned} u &= -\beta \cos \lambda \sin(\beta L) \\ v - v_0 &= -\sin \lambda (1 - \cos(\beta L)) \\ w &= \cos(\beta L) \end{aligned} \right\} \quad \dots (27)$$

19. From (10) and (19)

$$\begin{aligned} \frac{\delta \rho}{\rho} &= \frac{\delta a}{a} - 2 \left(\frac{\delta a}{a} - \frac{\delta b}{b} \right) \left(1 - \frac{e^2}{2} \frac{(1 - e^2) \sin^2 \lambda}{1 - e^2 \sin^2 \lambda} \right) \\ &= \frac{\delta a}{a} - 2 \left(\frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 1 - \frac{e^2}{2} (1 - e^2) \sin^2 \lambda (1 + e^2 \sin^2 \lambda + \dots) \right\} \\ &= \frac{\delta a}{a} - 2 \left(\frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 1 - \frac{e^2}{2} (1 - e^2) \left(\frac{1 - \cos 2\lambda}{2} + \frac{e^2}{8} (3 - 4 \cos 2\lambda + \cos 4\lambda) \dots \right) \right\} \end{aligned}$$

Hence from (12)

$$\begin{aligned} u - u_0 &= - \int_{\lambda_0}^{\lambda} \frac{\delta \rho}{\rho} d\lambda \\ &= -\frac{\delta a}{a} (\lambda - \lambda_0) + 2 \left(\frac{\delta a}{a} - \frac{\delta b}{b} \right) \left\{ 0.2518 (\lambda - \lambda_0) + 0.3750 (\sin 2\lambda - \sin 2\lambda_0) - 0.0003 (\sin 4\lambda - \sin 4\lambda_0) \right\} \end{aligned}$$

Expressing in seconds and introducing δa , δb expressed in kilometres and λ, λ_0 in degrees this becomes

$$u - u_0 = \delta a \left\{ -0.2807 (\lambda^\circ - \lambda_0^\circ) + 24.26 (\sin 2\lambda + \sin 2\lambda_0) - 0.019 (\sin 4\lambda - \sin 4\lambda_0) \right\} \\ + \delta b \left\{ -0.2847 (\lambda^\circ - \lambda_0^\circ) - 24.34 (\sin 2\lambda - \sin 2\lambda_0) + 0.019 (\sin 4\lambda - \sin 4\lambda_0) \right\}. \quad (28)$$

20. Along a meridian

Case I, when $\delta a = 1$, $\delta b = 0$, $u_0 = 0$, $w_0 = 0$

By (28) and (17)

$$\left. \begin{aligned} u &= -0.2807(\lambda^\circ - \lambda_0^\circ) + 24.26 (\sin 2\lambda - \sin 2\lambda_0) - 0.02 (\sin 4\lambda - \sin 4\lambda_0) \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \quad (29)$$

Case II, when $\delta a = 0$, $\delta b = 1$, $u_0 = 0$, $w_0 = 0$

By (28) and (17)

$$\left. \begin{aligned} u &= -0.2847(\lambda^\circ - \lambda_0^\circ) - 24.34 (\sin 2\lambda - \sin 2\lambda_0) + 0.02 (\sin 4\lambda - \sin 4\lambda_0) \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \quad (30)$$

Case III, when $\delta a = 0$, $\delta b = 0$, $u_0 = 1$, $w_0 = 0$

Then

$$\left. \begin{aligned} u &= u_0 = 1 \\ v - v_0 &= 0 \\ w &= 0 \end{aligned} \right\} \quad (31)$$

Case IV, when $\delta a = 0$, $\delta b = 0$, $u_0 = 0$, $w_0 = 1$

Then

$$\left. \begin{aligned} u - u_0 &= 0 \\ v - v_0 &= \frac{v_0 \cos \lambda_0}{a} \cdot \sqrt{1 - e^2} \left[\tan \lambda - 0.000,058,2 \lambda^\circ + 0.000,004,2 \sin 2\lambda \right]_{\lambda_0}^{\lambda} \\ w &= \frac{v_0 \cos \lambda_0}{v \cos \lambda} \end{aligned} \right\} \quad (32)$$

21. With the equations of §18, 20 changes can be computed at any point for any case. There are two possible routes. As an example consider Case I. We can first proceed along a parallel to the appropriate longitude. In proceeding thence along a meridian we have to apply Cases I, III and IV (because the values given by (22) now become initial values), and so find the values u_y, v_y, w_y : secondly we can proceed along a meridian and afterwards along a parallel and so find the second set of values u_x, v_x, w_x .

These computations have been made and the results for every degree square corner, so far as concerns the Indian triangulation, are exhibited in the following tables.

Values of $u_x - u_y$ in seconds.

Case I. $-\delta a = 1$ km.

Long.	Lat.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°		
86° 85 84 83 82 81 80 79 28 27 26 25	86°	0.102	0.086	0.070	0.057	0.044	0.034	0.024	0.016	0.010	0.005	0.002	0.000	0.000	0.001	0.004	0.008	0.014														
	85	0.092	0.077	0.063	0.051	0.040	0.030	0.022	0.015	0.009	0.005	0.002	0.000	0.000	0.001	0.004	0.008	0.013	0.019	0.025	0.032	0.040	0.049	0.059	0.070							
	84	0.082	0.068	0.056	0.045	0.035	0.027	0.019	0.013	0.008	0.004	0.002	0.000	0.000	0.001	0.003	0.007	0.011	0.016	0.021	0.027	0.034	0.042	0.051	0.060							
	83	0.072	0.060	0.049	0.040	0.031	0.024	0.017	0.012	0.007	0.004	0.002	0.000	0.000	0.001	0.003	0.006	0.010	0.013	0.019	0.025	0.032	0.040	0.049	0.059	0.070						
	82	0.063	0.052	0.043	0.035	0.027	0.020	0.015	0.010	0.006	0.003	0.001	0.000	0.000	0.001	0.003	0.005	0.009	0.013	0.019	0.025	0.032	0.040	0.049	0.059	0.070						
	81	0.054	0.045	0.037	0.029	0.023	0.017	0.013	0.009	0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.007	0.011	0.016	0.021	0.027	0.034	0.042	0.051	0.060						
	80	0.045	0.037	0.031	0.025	0.019	0.015	0.011	0.007	0.004	0.002	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.013	0.019	0.025	0.032	0.040	0.049	0.059	0.070					
	79	0.036	0.030	0.025	0.020	0.015	0.012	0.009	0.006	0.004	0.002	0.001	0.000	0.000	0.000	0.000	0.001	0.003	0.005	0.008	0.011	0.014	0.019	0.023	0.028	0.034	0.040					
	28	0.028	0.023	0.019	0.015	0.012	0.009	0.007	0.005	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.006	0.008	0.011	0.014	0.018	0.022	0.026	0.031	0.036	0.041	0.045	
	27	0.020	0.017	0.014	0.011	0.009	0.007	0.005	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.006	0.008	0.011	0.014	0.018	0.022	0.026	0.031	0.036	0.041	0.045	
26	0.012	0.011	0.009	0.007	0.006	0.004	0.003	0.002	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.007	0.008	0.010	0.012	0.014	0.017	0.019	0.022	0.025		
25	0.006	0.005	0.004	0.003	0.003	0.002	0.002	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.002	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011		
24 23 22 21 20 19 18 17 16 15 14 13 12 11 10 9 8	24	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.005	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.002		
	23	0.007	0.006	0.005	0.004	0.003	0.002	0.002	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.002	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.011	0.012	0.014	0.016	
	22	0.013	0.011	0.009	0.007	0.005	0.004	0.003	0.002	0.001	0.001	0.000	0.000	0.000	0.000	0.001	0.001	0.002	0.003	0.004	0.005	0.006	0.008	0.010	0.012	0.014	0.017	0.019	0.022	0.025	0.028	
	21	0.018	0.015	0.012	0.010	0.008	0.006	0.004	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.001	0.001	0.002	0.003	0.004	0.006	0.007	0.009	0.012	0.014	0.017	0.020	0.024	0.027	0.031	0.035	
	20	0.025	0.019	0.016	0.013	0.010	0.008	0.005	0.004	0.003	0.002	0.001	0.000	0.000	0.000	0.001	0.002	0.003	0.005	0.007	0.009	0.012	0.015	0.018	0.022	0.026	0.030	0.035	0.040	0.045	0.050	
	19					0.012	0.009	0.007	0.005	0.003	0.002	0.001	0.000	0.000	0.000	0.001	0.002	0.004	0.006	0.008	0.011	0.014	0.017	0.021	0.025	0.029	0.034	0.039	0.044	0.049	0.054	
	18					0.014	0.011	0.008	0.005	0.003	0.002	0.001	0.000	0.000	0.000	0.001	0.003	0.005	0.007	0.010	0.013	0.016	0.020	0.024	0.028	0.033	0.037	0.042	0.047	0.052	0.057	
	17					0.016	0.012	0.009	0.006	0.004	0.003	0.002	0.001	0.000	0.000	0.001	0.004	0.006	0.009	0.012	0.015	0.019	0.023	0.027	0.032	0.036	0.041	0.046	0.051	0.056	0.061	
	16								0.008			0.004	0.002	0.001	0.000	0.000	0.001	0.003	0.005	0.008	0.011	0.014	0.018	0.022	0.026	0.030	0.035	0.039	0.044	0.049	0.054	0.059
	15						0.017	0.013				0.005	0.003	0.001	0.000	0.000	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060
14										0.004	0.002	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
13										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
12										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
11										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
10										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
9										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	
8										0.005	0.003	0.001	0.000	0.000	0.001	0.002	0.004	0.006	0.009	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.041	0.045	0.050	0.055	0.060	

Case I, $-\delta a = 1$ km.

Lat.	P o s i t i v e												N e g a t i v e																	
	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
36°	0.128	0.118	0.107	0.097	0.086	0.075	0.064	0.053	0.042	0.030	0.019	0.008	0.004	0.015	0.027	0.038	0.050	0.063	0.077	0.092	0.107	0.122	0.137	0.152	0.167	0.182	0.197	0.212	0.227	0.242
35	0.104	0.095	0.087	0.078	0.070	0.061	0.052	0.043	0.034	0.025	0.015	0.003	0.003	0.012	0.022	0.033	0.040	0.049	0.059	0.069	0.079	0.089	0.099	0.109	0.119	0.129	0.139	0.149	0.159	0.169
34	0.082	0.076	0.069	0.062	0.056	0.049	0.041	0.034	0.027	0.020	0.012	0.005	0.005	0.010	0.017	0.025	0.032	0.039	0.047	0.055	0.063	0.071	0.079	0.087	0.095	0.103	0.111	0.119	0.127	0.135
33	0.064	0.059	0.054	0.049	0.043	0.038	0.032	0.027	0.021	0.015	0.010	0.004	0.004	0.008	0.014	0.020	0.026	0.032	0.038	0.044	0.050	0.056	0.062	0.068	0.074	0.080	0.086	0.092	0.098	0.104
32	0.049	0.045	0.041	0.037	0.033	0.029	0.025	0.020	0.016	0.012	0.007	0.003	0.003	0.006	0.010	0.015	0.019	0.023	0.028	0.033	0.038	0.043	0.048	0.053	0.058	0.063	0.068	0.073	0.078	0.083
31	0.035	0.033	0.030	0.027	0.024	0.021	0.018	0.015	0.012	0.009	0.005	0.002	0.002	0.004	0.008	0.012	0.016	0.020	0.024	0.028	0.032	0.036	0.040	0.044	0.048	0.052	0.056	0.060	0.064	0.068
30	0.025	0.023	0.021	0.019	0.017	0.015	0.013	0.010	0.008	0.006	0.004	0.002	0.002	0.003	0.005	0.008	0.010	0.012	0.014	0.016	0.018	0.020	0.022	0.024	0.026	0.028	0.030	0.032	0.034	0.036
29	0.016	0.015	0.014	0.013	0.011	0.010	0.008	0.007	0.005	0.004	0.002	0.001	0.001	0.002	0.004	0.005	0.006	0.008	0.009	0.011	0.012	0.013	0.014	0.016	0.017	0.018	0.019	0.020	0.021	0.022
28	0.010	0.009	0.008	0.007	0.006	0.005	0.004	0.003	0.003	0.002	0.001	0.001	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011	0.012	0.013	0.014	0.015	0.016	0.017
27	0.005	0.005	0.004	0.004	0.003	0.003	0.003	0.002	0.002	0.001	0.001	0.000	0.000	0.001	0.001	0.002	0.002	0.003	0.003	0.004	0.004	0.005	0.005	0.006	0.006	0.007	0.007	0.008	0.008	0.009
26	0.002	0.002	0.002	0.002	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001
25	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
24	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000
23	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001
22	0.004	0.003	0.003	0.003	0.003	0.002	0.002	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.002	0.002	0.002	0.003	0.003	0.003	0.003	0.004	0.004	0.005	0.005	0.005
21	0.007	0.006	0.005	0.005	0.004	0.004	0.003	0.003	0.002	0.002	0.001	0.000	0.000	0.001	0.001	0.001	0.002	0.003	0.004	0.004	0.005	0.005	0.006	0.006	0.007	0.007	0.008	0.008	0.009	0.010
20	0.011	0.010	0.009	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.001	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.008	0.009	0.010	0.011	0.012	0.013	0.014	0.015	0.016	0.016
19					0.010	0.009	0.008	0.003	0.005	0.008	0.003	0.003	0.001	0.001	0.002	0.003	0.005	0.006	0.007	0.009	0.010	0.011	0.011							
18				0.014	0.012	0.010	0.009	0.009	0.007	0.005	0.003	0.001	0.001	0.002	0.003	0.004	0.006	0.007	0.009	0.010	0.012	0.014	0.015							
17				0.018	0.016	0.014	0.011	0.009	0.008	0.006	0.004	0.001	0.001	0.002	0.003	0.004	0.006	0.008	0.010	0.012	0.015	0.018	0.020							
16				0.023	0.020	0.017	0.014	0.011	0.009	0.008	0.005	0.002	0.001	0.002	0.003	0.005	0.008	0.010	0.013	0.016	0.020	0.025	0.028							
15											0.013	0.010	0.008	0.002	0.001	0.005	0.009	0.012	0.016											
14											0.016	0.012	0.007	0.003	0.001	0.006	0.010	0.015	0.019											
13											0.019	0.013	0.008	0.003	0.002	0.007	0.012	0.017	0.022											
12											0.021	0.015	0.010	0.004	0.002	0.008	0.014	0.019	0.025											
11											0.024	0.017	0.011	0.004	0.002	0.009	0.015	0.022	0.029											
10											0.027	0.020	0.012	0.005	0.002	0.010	0.017	0.025	0.032											
9											0.030	0.022	0.014	0.005	0.003	0.011	0.019	0.027	0.035											
8											0.033	0.024	0.015	0.006	0.003	0.012	0.021	0.030	0.039											

Values of w_x in seconds.

TABLE IX.

Case I. — $\delta a = 1$ km.

Long.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
Lat.	P o s i t i v e																N e g a t i v e													
36°	4.760	4.847	3.945	3.538	3.132	2.725	2.317	1.909	1.499	1.089	0.679	0.269	0.148	0.553	0.953	1.372	1.781													
35°	4.567	4.150	3.792	3.402	3.019	2.620	2.228	1.835	1.441	1.047	0.653	0.253	0.136	0.531	0.935	1.319	1.713													
34°	4.384	4.021	3.648	3.278	2.897	2.531	2.148	1.867	1.487	1.098	0.628	0.249	0.131	0.511	0.920	1.289	1.648													
33°	4.229	3.871	3.511	3.150	2.789	2.428	2.063	1.699	1.335	0.970	0.605	0.239	0.126	0.492	0.857	1.222	1.586													
32°	4.073	3.727	3.381	3.034	2.685	2.336	1.986	1.638	1.285	0.931	0.582	0.230	0.121	0.473	0.825	1.176	1.527	1.878	2.228	2.577	2.928	3.274	3.620	3.968	4.310					
31°	3.925	3.582	3.258	2.928	2.598	2.251	1.914	1.577	1.239	0.900	0.561	0.222	0.117	0.456	0.795	1.131	1.472	1.809	2.146	2.488	2.819	3.151	3.488	3.821	4.163					
30°	3.784	3.461	3.143	2.819	2.495	2.171	1.846	1.520	1.194	0.868	0.541	0.214	0.113	0.440	0.767	1.088	1.419	1.745	2.070	2.396	2.719	3.042	3.364	3.685	4.005					
29°	3.653	3.342	3.032	2.720	2.408	2.095	1.781	1.467	1.153	0.837	0.522	0.207	0.100	0.425	0.740	1.055	1.370	1.681	1.998	2.311	2.624	2.936	3.246	3.556	3.865					
28°	3.528	3.229	2.920	2.608	2.306	2.024	1.721	1.417	1.113	0.820	0.504	0.200	0.105	0.410	0.715	1.019	1.323	1.627	1.930	2.232	2.534	2.836	3.138	3.435	3.738					
27°	3.410	3.121	2.831	2.540	2.319	1.957	1.654	1.370	1.078	0.782	0.488	0.193	0.102	0.397	0.691	0.985	1.279	1.573	1.866	2.168	2.460	2.741	3.032	3.321	3.609	3.896	4.182	4.467	4.751	
26°	3.301	3.021	2.740	2.450	2.176	1.898	1.610	1.328	1.042	0.757	0.472	0.187	0.099	0.384	0.669	0.954	1.238	1.523	1.806	2.089	2.371	2.653	2.934	3.214	3.493	3.771	4.047	4.323	4.598	
25°	3.198	2.927	2.655	2.382	2.109	1.835	1.560	1.285	1.009	0.733	0.457	0.181	0.095	0.372	0.615	0.924	1.200	1.475	1.750	2.024	2.298	2.571	2.843	3.114	3.385	3.654	3.922	4.189	4.455	
24°	3.102	2.839	2.570	2.311	2.046	1.780	1.513	1.246	0.979	0.711	0.444	0.176	0.093	0.361	0.623	0.886	1.164	1.431	1.697	1.963	2.229	2.494	2.758	3.021	3.283	3.544	3.805	4.061	4.322	
23°	3.013	2.757	2.501	2.244	1.987	1.729	1.470	1.210	0.951	0.691	0.431	0.170	0.090	0.350	0.610	0.870	1.130	1.390	1.649	1.907	2.165	2.423	2.678	2.931	3.189	3.442	3.695	3.946	4.197	
22°	2.930	2.683	2.433	2.183	1.933	1.681	1.429	1.177	0.925	0.672	0.419	0.166	0.087	0.341	0.594	0.847	1.099	1.351	1.603	1.855	2.106	2.356	2.605	2.854	3.101	3.348	3.593	3.838	4.082	
21°	2.854	2.613	2.370	2.126	1.882	1.637	1.392	1.147	0.901	0.651	0.406	0.161	0.085	0.332	0.578	0.821	1.070	1.316	1.561	1.806	2.050	2.294	2.537	2.779	3.021	3.261	3.500	3.738	3.976	
20°	2.784	2.543	2.311	2.074	1.836	1.597	1.358	1.118	0.879	0.638	0.398	0.153	0.083	0.324	0.564	0.801	1.044	1.281	1.523	1.762	2.000	2.238	2.475	2.711	2.946	3.181	3.414	3.646	3.878	
19°																														
18°																														
17°																														
16°																														
15°																														
14°																														
13°																														
12°																														
11°																														
10°																														
9°																														
8°																														

Case I. — $\delta\alpha = 1$ km.

Long. Lat.	63°	67°	63°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°
86°	1.839	1.134	1.028	0.923	0.817	0.711	0.605	0.498	0.391	0.284	0.177	0.070	0.037	0.144	0.251	0.358	0.435												
85	1.100	1.007	0.913	0.820	0.726	0.631	0.537	0.443	0.347	0.253	0.157	0.064	0.083	0.128	0.233	0.318	0.418												
84	0.968	0.838	0.804	0.731	0.638	0.545	0.472	0.330	0.305	0.232	0.133	0.035	0.020	0.113	0.193	0.280	0.323												
83	0.843	0.771	0.692	0.637	0.535	0.453	0.411	0.338	0.266	0.193	0.120	0.048	0.025	0.098	0.171	0.243	0.316												
82	0.723	0.661	0.601	0.539	0.477	0.415	0.353	0.290	0.238	0.168	0.103	0.041	0.022	0.081	0.147	0.209	0.271												
81	0.611	0.559	0.507	0.455	0.403	0.340	0.298	0.245	0.193	0.140	0.097	0.035	0.018	0.071	0.134	0.173	0.239												
80	0.504	0.461	0.419	0.376	0.332	0.287	0.243	0.203	0.159	0.116	0.072	0.022	0.015	0.059	0.102	0.146	0.193												
79	0.404	0.377	0.335	0.301	0.266	0.232	0.197	0.163	0.128	0.093	0.058	0.023	0.012	0.047	0.082	0.117	0.153												
78	0.310	0.284	0.257	0.231	0.204	0.173	0.151	0.125	0.098	0.071	0.044	0.018	0.009	0.033	0.063	0.090	0.116												
77	0.223	0.203	0.184	0.165	0.146	0.127	0.108	0.083	0.070	0.051	0.032	0.013	0.007	0.026	0.045	0.064	0.083												
76	0.140	0.128	0.116	0.104	0.092	0.083	0.068	0.053	0.044	0.032	0.020	0.008	0.004	0.016	0.028	0.040	0.052												
75	0.063	0.057	0.052	0.047	0.041	0.036	0.031	0.025	0.020	0.015	0.009	0.004	0.002	0.007	0.013	0.018	0.014												
74	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.003	0.002	0.001	0.001	0.000	0.001	0.002	0.003	0.003												
24	0.074	0.068	0.063	0.055	0.049	0.043	0.037	0.030	0.023	0.017	0.011	0.004	0.002	0.009	0.015	0.023	0.028												
23	0.135	0.123	0.112	0.100	0.089	0.077	0.066	0.054	0.043	0.031	0.019	0.008	0.001	0.016	0.037	0.059	0.051												
22	0.190	0.174	0.158	0.143	0.126	0.109	0.093	0.076	0.030	0.014	0.027	0.011	0.003	0.023	0.038	0.055	0.071												
21	0.241	0.220	0.200	0.179	0.159	0.138	0.117	0.097	0.076	0.055	0.035	0.014	0.007	0.029	0.049	0.070	0.090												
20																													
19																													
18																													
17																													
16																													
15																													
14																													
13																													
12																													
11																													
10																													
9																													
8																													

TABLE XI.

Values of u_z in seconds.Case II.— $\delta\delta = 1$ km.

Long.	Lat.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°		
36°	0	8-819	8-830	8-839	8-848	8-856	8-862	8-868	8-873	8-877	8-880	8-883	8-885	8-888	8-893	8-898	8-900	8-903	8-907	8-912	8-918	8-924	8-930	8-937	8-943	8-950	8-957	8-963	8-970	8-977	8-984	
35	A	7-769	7-772	7-781	7-788	7-796	7-802	7-807	7-811	7-815	7-818	7-821	7-823	7-825	7-828	7-830	7-833	7-835	7-838	7-841	7-844	7-847	7-850	7-853	7-856	7-859	7-862	7-865	7-868	7-871	7-874	
34		7-178	7-186	7-194	7-201	7-208	7-213	7-218	7-222	7-225	7-228	7-230	7-232	7-234	7-236	7-238	7-240	7-242	7-244	7-246	7-248	7-250	7-252	7-254	7-256	7-258	7-260	7-262	7-264	7-266	7-268	
33		6-566	6-574	6-581	6-587	6-593	6-598	6-602	6-605	6-608	6-610	6-612	6-613	6-615	6-617	6-619	6-621	6-623	6-625	6-627	6-629	6-631	6-633	6-635	6-637	6-639	6-641	6-643	6-645	6-647	6-649	
32		5-929	5-935	5-941	5-947	5-951	5-955	5-959	5-962	5-964	5-967	5-969	5-971	5-973	5-975	5-977	5-979	5-981	5-983	5-985	5-987	5-989	5-991	5-993	5-995	5-997	5-999	6-001	6-003	6-005	6-007	
31		5-286	5-291	5-295	5-299	5-303	5-307	5-310	5-313	5-316	5-319	5-322	5-325	5-328	5-331	5-334	5-337	5-340	5-343	5-346	5-349	5-352	5-355	5-358	5-361	5-364	5-367	5-370	5-373	5-376	5-379	
30		4-577	4-580	4-584	4-587	4-590	4-593	4-596	4-598	4-599	4-600	4-601	4-602	4-603	4-604	4-605	4-606	4-607	4-608	4-609	4-610	4-611	4-612	4-613	4-614	4-615	4-616	4-617	4-618	4-619	4-620	
29		3-849	3-856	3-867	3-869	3-871	3-872	3-873	3-874	3-875	3-876	3-877	3-878	3-879	3-880	3-881	3-882	3-883	3-884	3-885	3-886	3-887	3-888	3-889	3-890	3-891	3-892	3-893	3-894	3-895	3-896	
28		3-126	3-126	3-126	3-127	3-127	3-127	3-128	3-128	3-128	3-128	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	3-129	
27		2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	2-363	
26		1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	1-578	
25		0-771	0-768	0-761	0-757	0-753	0-750	0-747	0-743	0-740	0-737	0-734	0-731	0-728	0-725	0-722	0-719	0-716	0-713	0-710	0-707	0-704	0-701	0-698	0-695	0-692	0-689	0-686	0-683	0-680	0-677	
24		0-059	0-063	0-072	0-078	0-083	0-088	0-092	0-095	0-098	0-100	0-102	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	0-103	
23		0-910	0-919	0-927	0-935	0-942	0-948	0-953	0-957	0-961	0-963	0-965	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	0-966	
22		1-781	1-792	1-803	1-813	1-821	1-828	1-835	1-840	1-845	1-848	1-850	1-851	1-852	1-853	1-854	1-855	1-856	1-857	1-858	1-859	1-860	1-861	1-862	1-863	1-864	1-865	1-866	1-867	1-868	1-869	
21		2-672	2-686	2-699	2-710	2-720	2-729	2-737	2-744	2-749	2-753	2-756	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	2-757	
20		3-552	3-568	3-584	3-597	3-609	3-620	3-630	3-637	3-643	3-648	3-651	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	3-653	
19		4-437	4-459	4-480	4-500	4-517	4-532	4-547	4-560	4-571	4-580	4-588	4-593	4-596	4-598	4-599	4-600	4-601	4-602	4-603	4-604	4-605	4-606	4-607	4-608	4-609	4-610	4-611	4-612	4-613	4-614	
18		5-323	5-347	5-369	5-389	5-407	5-423	5-438	5-451	5-462	5-471	5-478	5-483	5-486	5-488	5-489	5-490	5-491	5-492	5-493	5-494	5-495	5-496	5-497	5-498	5-499	5-500	5-501	5-502	5-503	5-504	
17		6-206	6-231	6-253	6-273	6-290	6-305	6-318	6-329	6-338	6-345	6-350	6-353	6-355	6-356	6-357	6-358	6-359	6-360	6-361	6-362	6-363	6-364	6-365	6-366	6-367	6-368	6-369	6-370	6-371	6-372	
16		7-094	7-112	7-127	7-139	7-149	7-157	7-163	7-168	7-171	7-173	7-175	7-176	7-177	7-178	7-179	7-180	7-181	7-182	7-183	7-184	7-185	7-186	7-187	7-188	7-189	7-190	7-191	7-192	7-193	7-194	
15		7-981	8-000	8-016	8-029	8-040	8-049	8-056	8-062	8-066	8-069	8-071	8-073	8-074	8-075	8-076	8-077	8-078	8-079	8-080	8-081	8-082	8-083	8-084	8-085	8-086	8-087	8-088	8-089	8-090	8-091	
14		8-868	8-888	8-904	8-917	8-928	8-937	8-944	8-949	8-953	8-956	8-958	8-960	8-961	8-962	8-963	8-964	8-965	8-966	8-967	8-968	8-969	8-970	8-971	8-972	8-973	8-974	8-975	8-976	8-977	8-978	8-979
13		9-755	9-775	9-791	9-804	9-815	9-824	9-831	9-836	9-840	9-843	9-845	9-847	9-848	9-849	9-850	9-851	9-852	9-853	9-854	9-855	9-856	9-857	9-858	9-859	9-860	9-861	9-862	9-863	9-864	9-865	9-866
12		10-642	10-662	10-678	10-691	10-702	10-711	10-718	10-723	10-726	10-728	10-729	10-730	10-731	10-732	10-733	10-734	10-735	10-736	10-737	10-738	10-739	10-740	10-741	10-742	10-743	10-744	10-745	10-746	10-747	10-748	10-749
11		11-529	11-549	11-565	11-578	11-589	11-598	11-605	11-610	11-613	11-615	11-616	11-617	11-618	11-619	11-620	11-621	11-622	11-623	11-624	11-625	11-626	11-627	11-628	11-629	11-630	11-631	11-632	11-633	11-634	11-635	11-636
10		12-416	12-436	12-452	12-465	12-476	12-485	12-492	12-497	12-500	12-502	12-503	12-504	12-505	12-506	12-507	12-508	12-509	12-510	12-511	12-512	12-513	12-514	12-515	12-516	12-517	12-518	12-519	12-520	12-521	12-522	12-523
9		13-303	13-323	13-339	13-352	13-363	13-372	13-379	13-384	13-388	13-391	13-393	13-395	13-396	13-397	13-398	13-399	13-400	13-401	13-402	13-403	13-404	13-405	13-406	13-407	13-408	13-409	13-410	13-411	13-412	13-413	13-414
8		14-190	14-210	14-226	14-239	14-250	14-259	14-266	14-271	14-274	14-276	14-277	14-278	14-279	14-280	14-281	14-282	14-283	14-284	14-285	14-286	14-287	14-288	14-289	14-290	14-291	14-292	14-293	14-294	14-295	14-296	14-297

Case II.— $\delta b = 1$ km.

Long. Lat.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°
36°																													
35°	0.105	0.088	0.072	0.058	0.045	0.034	0.025	0.017	0.010	0.005	0.002	0.000	0.000	0.001	0.004	0.009	0.015												
34°	0.100	0.081	0.069	0.055	0.043	0.033	0.024	0.016	0.010	0.005	0.002	0.000	0.000	0.001	0.004	0.008	0.014												
33°	0.095	0.079	0.065	0.052	0.041	0.031	0.022	0.015	0.009	0.006	0.002	0.000	0.000	0.001	0.004	0.008	0.013												
32°	0.088	0.074	0.060	0.046	0.038	0.029	0.021	0.014	0.008	0.005	0.002	0.000	0.000	0.001	0.004	0.007	0.012												
31°	0.081	0.068	0.056	0.045	0.035	0.028	0.019	0.013	0.003	0.004	0.002	0.000	0.000	0.001	0.003	0.006	0.010	0.015	0.022	0.029	0.037	0.047	0.053	0.064	0.077	0.091	0.108	0.126	0.146
30°	0.073	0.061	0.050	0.041	0.032	0.024	0.017	0.012	0.007	0.004	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.014	0.019	0.025	0.033	0.041	0.051	0.061	0.072	0.083	0.099	0.119	0.146
29°	0.065	0.054	0.044	0.036	0.028	0.021	0.015	0.010	0.006	0.003	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.014	0.019	0.025	0.033	0.041	0.051	0.061	0.072	0.083	0.099	0.119	0.146
28°	0.055	0.046	0.038	0.031	0.024	0.018	0.013	0.009	0.005	0.003	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.012	0.016	0.022	0.028	0.036	0.044	0.053	0.062	0.073	0.088	0.108	0.136
27°	0.045	0.038	0.031	0.025	0.020	0.015	0.011	0.007	0.004	0.002	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.010	0.013	0.018	0.023	0.029	0.036	0.043	0.051	0.061	0.073	0.088	0.108
26°	0.035	0.029	0.024	0.019	0.015	0.011	0.008	0.006	0.003	0.002	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.007	0.010	0.014	0.018	0.022	0.027	0.033	0.039	0.046	0.055	0.066	0.083
25°	0.024	0.020	0.016	0.013	0.010	0.008	0.005	0.004	0.002	0.001	0.000	0.000	0.000	0.001	0.003	0.006	0.009	0.005	0.007	0.009	0.012	0.015	0.018	0.022	0.026	0.030	0.035	0.040	0.046
24°	0.013	0.010	0.008	0.006	0.005	0.004	0.003	0.002	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.003	0.004	0.006	0.007	0.008	0.011	0.013	0.015	0.017	0.019	0.022
23°	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.001	0.001	0.001	0.001	0.002	0.002	0.003	0.003	0.003
22°	0.015	0.013	0.010	0.008	0.006	0.005	0.004	0.002	0.002	0.001	0.000	0.000	0.000	0.000	0.001	0.003	0.006	0.009	0.012	0.015	0.019	0.023	0.028	0.033	0.038	0.044	0.051	0.057	0.063
21°	0.029	0.025	0.020	0.016	0.013	0.010	0.007	0.005	0.003	0.002	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.007	0.010	0.013	0.016	0.020	0.025	0.029	0.033	0.038	0.044	0.051	0.057
20°	0.044	0.037	0.030	0.024	0.020	0.016	0.013	0.009	0.006	0.004	0.002	0.001	0.000	0.001	0.003	0.006	0.009	0.006	0.009	0.012	0.015	0.019	0.023	0.028	0.033	0.038	0.044	0.051	0.057
19°	0.060	0.050	0.041	0.033	0.026	0.020	0.014	0.010	0.006	0.003	0.001	0.000	0.000	0.001	0.003	0.006	0.009	0.006	0.009	0.012	0.015	0.019	0.023	0.028	0.033	0.038	0.044	0.051	0.057
18°																													
17°																													
16°																													
15°																													
14°																													
13°																													
12°																													
11°																													
10°																													
9°																													
8°																													

Case II. — $\delta b = 1$ km.

Long. Lat.	P o s i t i v e										N e g a t i v e																			
	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
86°	0.178	0.139	0.144	0.130	0.115	0.100	0.085	0.070	0.055	0.040	0.025	0.010	0.005	0.021	0.036	0.051	0.068													
85	0.148	0.128	0.128	0.110	0.098	0.085	0.078	0.060	0.047	0.034	0.021	0.008	0.005	0.017	0.030	0.043	0.056													
84	0.121	0.111	0.108	0.093	0.081	0.071	0.061	0.050	0.039	0.028	0.018	0.007	0.004	0.014	0.025	0.036	0.047													
83	0.088	0.081	0.083	0.075	0.067	0.058	0.050	0.041	0.032	0.023	0.015	0.006	0.003	0.013	0.021	0.029	0.038													
82	0.079	0.072	0.068	0.060	0.053	0.046	0.040	0.033	0.026	0.019	0.013	0.006	0.002	0.009	0.016	0.023	0.030													
81	0.061	0.056	0.051	0.046	0.041	0.036	0.030	0.025	0.020	0.014	0.009	0.004	0.002	0.007	0.013	0.018	0.023													
80	0.045	0.041	0.038	0.034	0.030	0.026	0.023	0.019	0.015	0.011	0.007	0.003	0.001	0.005	0.009	0.013	0.017													
79	0.031	0.029	0.026	0.024	0.021	0.018	0.016	0.013	0.010	0.008	0.005	0.002	0.001	0.004	0.007	0.009	0.013													
78	0.020	0.019	0.017	0.015	0.014	0.012	0.010	0.008	0.007	0.005	0.003	0.001	0.000	0.003	0.004	0.006	0.008													
77	0.011	0.010	0.008	0.009	0.008	0.007	0.006	0.006	0.005	0.004	0.003	0.002	0.001	0.000	0.003	0.003	0.004													
76	0.005	0.005	0.004	0.004	0.003	0.003	0.003	0.003	0.002	0.001	0.001	0.000	0.000	0.000	0.001	0.001	0.002													
75	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.001													
74	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000													
73	0.002	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000													
72	0.008	0.005	0.005	0.004	0.004	0.003	0.003	0.003	0.003	0.002	0.001	0.000	0.000	0.001	0.001	0.002	0.003													
71	0.013	0.013	0.011	0.009	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.000	0.002	0.003	0.004	0.005													
70	0.023	0.021	0.019	0.017	0.015	0.013	0.011	0.009	0.007	0.005	0.003	0.001	0.000	0.003	0.005	0.006	0.008													
19					0.028	0.020	0.017	0.014	0.011	0.008	0.005	0.002	0.001	0.004	0.007	0.010	0.013													
18					0.033	0.029	0.025	0.020	0.016	0.012	0.007	0.003	0.000	0.008	0.010	0.015	0.019													
17					0.046	0.040	0.034	0.028	0.023	0.016	0.010	0.004	0.000	0.008	0.014	0.020	0.026													
16					0.080	0.068	0.054	0.037	0.029	0.021	0.013	0.005	0.000	0.008	0.011	0.018	0.026													
15							0.057	0.047	0.037	0.027	0.017	0.007	0.003	0.003	0.013	0.023	0.030													
14							0.045	0.035	0.025	0.015	0.005	0.003	0.004	0.017	0.029	0.042	0.054													
13							0.035	0.020	0.010	0.005	0.002	0.010	0.005	0.005	0.021	0.036	0.048													
12							0.068	0.045	0.020	0.010	0.005	0.012	0.006	0.006	0.025	0.043	0.059													
11							0.078	0.037	0.015	0.008	0.005	0.014	0.008	0.008	0.029	0.050	0.068													
10							0.091	0.066	0.041	0.021	0.013	0.016	0.009	0.009	0.034	0.059	0.079													
9							0.108	0.077	0.048	0.027	0.017	0.023	0.010	0.009	0.045	0.077	0.106													
8							0.131	0.088	0.055	0.031	0.023	0.028	0.011	0.011	0.045	0.078	0.144													

TABLE XVI.
Values of $w_x - w_y$ in seconds.
Case II.— $\delta\delta = 1$ km.

Long.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
Lat.	86°	85°	84°	83°	82°	81°	80°	79°	78°	77°	76°	75°	74°	73°	72°	71°	70°	69°	68°	67°	66°	65°	64°	63°	62°	61°	60°	59°	58°	
0	1.889	1.103	1.064	0.945	0.837	0.728	0.619	0.510	0.401	0.291	0.181	0.072	0	0.068	0.148	0.237	0.327	0.416	0.509	0.598	0.688	0.747	0.826	0.905	0.984					
1	1.108	1.082	0.980	0.888	0.798	0.684	0.568	0.459	0.349	0.238	0.127	0.017	0.087	0.176	0.265	0.354	0.442	0.530	0.618	0.705	0.792	0.879	0.966	1.053	1.140	1.227	1.314	1.401	1.488	1.575
2	1.110	1.016	0.922	0.827	0.728	0.627	0.524	0.419	0.314	0.208	0.102	0.000	0.088	0.176	0.264	0.352	0.440	0.528	0.615	0.702	0.789	0.876	0.963	1.050	1.137	1.224	1.311	1.398	1.485	1.572
3	1.028	0.938	0.849	0.762	0.674	0.587	0.498	0.411	0.323	0.234	0.146	0.058	0.068	0.156	0.244	0.332	0.420	0.508	0.595	0.682	0.769	0.856	0.943	1.030	1.117	1.204	1.291	1.378	1.465	1.552
4	0.949	0.861	0.772	0.682	0.613	0.538	0.453	0.373	0.293	0.213	0.133	0.053	0.063	0.151	0.239	0.327	0.415	0.502	0.589	0.676	0.763	0.850	0.937	1.024	1.111	1.198	1.285	1.372	1.459	1.546
5	0.867	0.780	0.690	0.619	0.549	0.477	0.406	0.334	0.262	0.190	0.119	0.047	0.057	0.145	0.233	0.321	0.409	0.496	0.583	0.670	0.757	0.844	0.931	1.018	1.105	1.192	1.279	1.366	1.453	1.540
6	0.788	0.696	0.603	0.511	0.417	0.324	0.232	0.139	0.046	0.053	0.160	0.247	0.334	0.421	0.508	0.595	0.682	0.769	0.856	0.943	1.030	1.117	1.204	1.291	1.378	1.465	1.552	1.639	1.726	1.813
7	0.718	0.624	0.531	0.439	0.346	0.254	0.161	0.068	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817
8	0.651	0.556	0.463	0.370	0.277	0.184	0.091	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817
9	0.588	0.493	0.400	0.307	0.214	0.121	0.028	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817
10	0.525	0.430	0.337	0.244	0.151	0.058	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904
11	0.462	0.367	0.274	0.181	0.088	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991
12	0.400	0.305	0.212	0.119	0.026	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991
13	0.337	0.242	0.149	0.056	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078
14	0.274	0.179	0.086	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165
15	0.212	0.117	0.024	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165
16	0.149	0.054	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252
17	0.086	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252	2.339	2.426
18	0.024	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252	2.339	2.426
19	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252	2.339	2.426	2.513
20	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252	2.339	2.426
21	0.000	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252	2.339
22	0.000	0.000	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165	2.252
23	0.000	0.000	0.000	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078	2.165
24	0.000	0.000	0.000	0.000	0.000	0.000	0.057	0.164	0.251	0.338	0.425	0.512	0.599	0.686	0.773	0.860	0.947	1.034	1.121	1.208	1.295	1.382	1.469	1.556	1.643	1.730	1.817	1.904	1.991	2.078

TABLE XVII.

Values of μ and ν in seconds.Case III.— $\mu_0 = 1''$.

Long.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	
Lat.	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	50°	
Values of μ (Positive)																														
Values of ν																														
P o s i t i v e																														
86°	0.147	0.134	0.122	0.109	0.097	0.084	0.073	0.063	0.054	0.046	0.039	0.033	0.028	0.024	0.020	0.017	0.015	0.013	0.011	0.009	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.000	
85	0.143	0.130	0.118	0.105	0.093	0.081	0.070	0.060	0.051	0.043	0.036	0.030	0.025	0.021	0.017	0.014	0.012	0.010	0.008	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.000	0.000	0.000	
84	0.139	0.126	0.114	0.101	0.089	0.077	0.066	0.056	0.047	0.039	0.032	0.026	0.021	0.017	0.014	0.011	0.009	0.007	0.006	0.005	0.004	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.000	
83	0.135	0.122	0.110	0.097	0.085	0.073	0.062	0.052	0.043	0.035	0.028	0.022	0.017	0.013	0.010	0.008	0.006	0.005	0.004	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
82	0.131	0.118	0.106	0.093	0.081	0.069	0.058	0.048	0.039	0.031	0.024	0.018	0.013	0.009	0.006	0.004	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
81	0.127	0.114	0.102	0.089	0.077	0.065	0.054	0.044	0.035	0.027	0.020	0.014	0.009	0.005	0.003	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
80	0.123	0.110	0.098	0.085	0.073	0.061	0.050	0.040	0.031	0.023	0.016	0.010	0.005	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
79	0.119	0.106	0.094	0.081	0.069	0.057	0.046	0.036	0.027	0.019	0.012	0.006	0.002	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
78	0.115	0.102	0.090	0.077	0.065	0.053	0.042	0.032	0.023	0.015	0.008	0.003	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
77	0.111	0.098	0.086	0.073	0.061	0.049	0.038	0.028	0.019	0.011	0.004	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
76	0.107	0.094	0.082	0.069	0.057	0.045	0.034	0.024	0.015	0.007	0.002	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
75	0.103	0.090	0.078	0.065	0.053	0.041	0.030	0.020	0.011	0.003	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
74	0.099	0.086	0.074	0.061	0.049	0.037	0.026	0.016	0.007	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
73	0.095	0.082	0.070	0.057	0.045	0.033	0.022	0.012	0.003	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
72	0.091	0.078	0.066	0.053	0.041	0.029	0.018	0.008	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
71	0.087	0.074	0.062	0.049	0.037	0.025	0.014	0.005	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
70	0.083	0.070	0.058	0.045	0.033	0.021	0.010	0.001	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
69	0.079	0.066	0.054	0.041	0.029	0.017	0.006	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
68	0.075	0.062	0.050	0.037	0.025	0.013	0.002	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
67	0.071	0.058	0.046	0.033	0.021	0.009	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	
66	0.067	0.054	0.042	0.029	0.017	0.005	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	

P

o

s

i

t

i

v

e

e

e

e

e

e

e

e

e

e

e

e

Values of w in seconds.

TABLE XVIII.

Case III. $-\alpha_0 = 1''$

Long.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°
Lat.	N													P															
	e													o															
86°	0.280	0.289	0.297	0.188	0.165	0.148	0.132	0.100	0.079	0.057	0.036	0.014	0.007	0.029	0.051	0.073	0.094	0.110	0.130	0.151	0.171	0.191	0.212	0.233	0.253	0.273	0.293	0.313	0.333
85	0.247	0.258	0.265	0.154	0.133	0.118	0.100	0.069	0.048	0.027	0.005	0.014	0.007	0.029	0.051	0.071	0.088	0.109	0.129	0.149	0.169	0.189	0.210	0.230	0.250	0.270	0.290	0.310	0.330
84	0.244	0.253	0.260	0.153	0.131	0.114	0.103	0.068	0.047	0.026	0.005	0.014	0.007	0.028	0.049	0.070	0.081	0.108	0.128	0.148	0.168	0.188	0.208	0.228	0.248	0.268	0.288	0.308	0.328
83	0.241	0.251	0.258	0.150	0.128	0.111	0.097	0.065	0.044	0.023	0.004	0.014	0.007	0.028	0.049	0.070	0.080	0.107	0.127	0.147	0.167	0.187	0.207	0.227	0.247	0.267	0.287	0.307	0.327
82	0.238	0.248	0.255	0.147	0.125	0.108	0.093	0.061	0.040	0.019	0.004	0.014	0.007	0.028	0.049	0.069	0.079	0.106	0.126	0.146	0.166	0.186	0.206	0.226	0.246	0.266	0.286	0.306	0.326
81	0.236	0.246	0.253	0.146	0.124	0.107	0.092	0.060	0.039	0.018	0.004	0.013	0.007	0.027	0.048	0.068	0.078	0.105	0.125	0.145	0.165	0.185	0.205	0.225	0.245	0.265	0.285	0.305	0.325
80	0.233	0.243	0.250	0.143	0.121	0.104	0.089	0.057	0.036	0.015	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
79	0.231	0.241	0.248	0.141	0.119	0.102	0.087	0.055	0.034	0.013	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
78	0.229	0.239	0.246	0.139	0.117	0.100	0.085	0.053	0.032	0.011	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
77	0.227	0.237	0.244	0.137	0.115	0.098	0.083	0.051	0.030	0.009	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
76	0.225	0.235	0.242	0.135	0.113	0.096	0.081	0.049	0.028	0.007	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
75	0.223	0.233	0.240	0.133	0.111	0.094	0.079	0.047	0.026	0.005	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
74	0.221	0.231	0.238	0.131	0.109	0.092	0.077	0.045	0.024	0.003	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
73	0.219	0.229	0.236	0.129	0.107	0.090	0.075	0.043	0.022	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
72	0.217	0.227	0.234	0.127	0.105	0.088	0.073	0.041	0.020	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
71	0.215	0.225	0.232	0.125	0.103	0.086	0.071	0.039	0.018	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
70	0.213	0.223	0.230	0.123	0.101	0.084	0.069	0.037	0.016	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
69	0.211	0.221	0.228	0.121	0.099	0.082	0.067	0.035	0.014	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
68	0.209	0.219	0.226	0.119	0.097	0.080	0.065	0.033	0.012	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
67	0.207	0.217	0.224	0.117	0.095	0.078	0.063	0.031	0.010	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324
66	0.205	0.215	0.222	0.115	0.093	0.076	0.061	0.029	0.008	0.001	0.003	0.013	0.007	0.027	0.047	0.067	0.077	0.104	0.124	0.144	0.164	0.184	0.204	0.224	0.244	0.264	0.284	0.304	0.324

Values of u and v in seconds.

TABLE XIX.

Case IV.— $\omega_0 = 1''$.

Long.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°							
Lat.	P o s i t i v e																N e g a t i v e																			
36°-8'	Values of <i>u</i>																Values of <i>v</i>																			
	0.185	0.170	0.154	0.138	0.123	0.106	0.090	0.074	0.058	0.043	0.027	0.011	0.005	0.003	0.004	0.070	0.086	0.101	0.117	0.133	0.149	0.165	0.181	0.196	0.213	0.227	0.243	0.258								
86°	0.299	0.293	0.284	0.276	0.268	0.259	0.249	0.247	0.243	0.238	0.233	0.228	0.223	0.218	0.213	0.208	0.203	0.198	0.193	0.188	0.183	0.178	0.173	0.168	0.163	0.158	0.153	0.148								
85°	0.216	0.210	0.200	0.192	0.184	0.175	0.165	0.163	0.159	0.154	0.149	0.144	0.139	0.134	0.129	0.124	0.119	0.114	0.109	0.104	0.099	0.094	0.089	0.084	0.079	0.074	0.069	0.067								
84°	0.168	0.162	0.152	0.144	0.136	0.127	0.117	0.115	0.111	0.106	0.101	0.096	0.091	0.086	0.081	0.076	0.071	0.066	0.061	0.056	0.051	0.046	0.041	0.036	0.031	0.026	0.021	0.019								
83°	0.171	0.175	0.176	0.178	0.179	0.180	0.181	0.181	0.182	0.183	0.183	0.184	0.184	0.185	0.185	0.186	0.186	0.187	0.187	0.188	0.188	0.189	0.189	0.190	0.190	0.191	0.191	0.192								
82°	0.149	0.151	0.153	0.154	0.156	0.157	0.158	0.159	0.160	0.161	0.161	0.162	0.163	0.163	0.164	0.165	0.165	0.166	0.167	0.167	0.168	0.169	0.169	0.170	0.170	0.171	0.171	0.172								
81°	0.128	0.130	0.131	0.132	0.133	0.134	0.135	0.136	0.137	0.138	0.138	0.139	0.140	0.140	0.141	0.142	0.142	0.143	0.143	0.144	0.144	0.145	0.145	0.146	0.146	0.147	0.147	0.148								
80°	0.107	0.108	0.110	0.112	0.113	0.114	0.115	0.116	0.117	0.117	0.118	0.118	0.119	0.120	0.120	0.121	0.121	0.122	0.122	0.123	0.123	0.124	0.124	0.125	0.125	0.126	0.126	0.127								
79°	0.086	0.088	0.090	0.092	0.093	0.094	0.095	0.096	0.097	0.098	0.097	0.097	0.098	0.098	0.099	0.099	0.100	0.100	0.101	0.101	0.102	0.102	0.103	0.103	0.104	0.104	0.105	0.105								
78°	0.066	0.068	0.069	0.071	0.072	0.073	0.074	0.075	0.076	0.077	0.077	0.078	0.078	0.079	0.079	0.080	0.080	0.081	0.081	0.082	0.082	0.083	0.083	0.084	0.084	0.085	0.085	0.087								
77°	0.046	0.048	0.049	0.051	0.052	0.053	0.054	0.055	0.056	0.057	0.057	0.058	0.059	0.059	0.060	0.061	0.061	0.062	0.062	0.063	0.063	0.064	0.064	0.065	0.065	0.066	0.066	0.067								
76°	0.027	0.029	0.030	0.031	0.032	0.033	0.034	0.035	0.035	0.036	0.037	0.037	0.038	0.038	0.039	0.040	0.040	0.041	0.041	0.042	0.042	0.043	0.043	0.044	0.044	0.045	0.045	0.047								
75°	0.008	0.010	0.011	0.012	0.013	0.014	0.015	0.015	0.016	0.017	0.017	0.018	0.018	0.019	0.020	0.020	0.021	0.021	0.022	0.022	0.023	0.023	0.024	0.024	0.025	0.025	0.026	0.027								
74°	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014	0.014	0.015	0.015	0.016	0.016	0.017	0.017	0.018								
73°	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014	0.014	0.015	0.015	0.016								
72°	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014	0.014	0.015	0.016								
71°	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014	0.014	0.015								
70°	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014	0.015								
69°	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.013	0.014								
68°	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013	0.014								
67°	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.012	0.013								
66°	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.000	0.001	0.002	0.003	0.004	0.005	0.006	0.007	0.007	0.008	0.009	0.009	0.010	0.010	0.011	0.011	0.012	0.013								

TABLE XX.

Values of w in seconds.

Case IV. — $w_0 = 1''$.

Long.	Lat.	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°
36°	1-104	1-108	1-112	1-115	1-117	1-120	1-122	1-124	1-126	1-127	1-128	1-129	1-130	1-131	1-132	1-133	1-134	1-135	1-136	1-137	1-138	1-139	1-140	1-141	1-142	1-143	1-144	1-145	1-146	1-147
35	1-081	1-086	1-088	1-091	1-094	1-096	1-098	1-101	1-103	1-105	1-107	1-109	1-110	1-111	1-112	1-113	1-114	1-115	1-116	1-117	1-118	1-119	1-120	1-121	1-122	1-123	1-124	1-125	1-126	1-127
34	1-077	1-081	1-085	1-088	1-091	1-093	1-095	1-097	1-098	1-099	1-100	1-101	1-102	1-103	1-104	1-105	1-106	1-107	1-108	1-109	1-110	1-111	1-112	1-113	1-114	1-115	1-116	1-117	1-118	1-119
33	1-065	1-069	1-073	1-075	1-078	1-080	1-082	1-084	1-085	1-086	1-087	1-088	1-089	1-090	1-091	1-092	1-093	1-094	1-095	1-096	1-097	1-098	1-099	1-100	1-101	1-102	1-103	1-104	1-105	1-106
32	1-054	1-057	1-060	1-063	1-066	1-068	1-070	1-072	1-074	1-075	1-076	1-077	1-078	1-079	1-080	1-081	1-082	1-083	1-084	1-085	1-086	1-087	1-088	1-089	1-090	1-091	1-092	1-093	1-094	1-095
31	1-049	1-046	1-049	1-052	1-055	1-057	1-059	1-061	1-063	1-064	1-065	1-066	1-067	1-068	1-069	1-070	1-071	1-072	1-073	1-074	1-075	1-076	1-077	1-078	1-079	1-080	1-081	1-082	1-083	1-084
30	1-038	1-035	1-038	1-041	1-044	1-046	1-048	1-050	1-051	1-052	1-053	1-054	1-055	1-056	1-057	1-058	1-059	1-060	1-061	1-062	1-063	1-064	1-065	1-066	1-067	1-068	1-069	1-070	1-071	1-072
29	1-023	1-025	1-028	1-031	1-034	1-036	1-038	1-040	1-041	1-042	1-043	1-044	1-045	1-046	1-047	1-048	1-049	1-050	1-051	1-052	1-053	1-054	1-055	1-056	1-057	1-058	1-059	1-060	1-061	1-062
28	1-013	1-016	1-019	1-022	1-024	1-026	1-028	1-030	1-031	1-032	1-033	1-034	1-035	1-036	1-037	1-038	1-039	1-040	1-041	1-042	1-043	1-044	1-045	1-046	1-047	1-048	1-049	1-050	1-051	1-052
27	1-003	1-007	1-010	1-013	1-015	1-017	1-019	1-021	1-023	1-024	1-025	1-026	1-027	1-028	1-029	1-030	1-031	1-032	1-033	1-034	1-035	1-036	1-037	1-038	1-039	1-040	1-041	1-042	1-043	1-044
26	0-994	0-998	1-001	1-004	1-006	1-008	1-011	1-013	1-015	1-016	1-017	1-018	1-019	1-020	1-021	1-022	1-023	1-024	1-025	1-026	1-027	1-028	1-029	1-030	1-031	1-032	1-033	1-034	1-035	1-036
25	0-988	0-990	0-993	0-996	0-998	1-000	1-002	1-004	1-006	1-007	1-008	1-009	1-010	1-011	1-012	1-013	1-014	1-015	1-016	1-017	1-018	1-019	1-020	1-021	1-022	1-023	1-024	1-025	1-026	1-027
24	0-979	0-982	0-985	0-988	0-990	0-992	0-994	0-996	0-997	0-998	0-999	1-000	1-001	1-002	1-003	1-004	1-005	1-006	1-007	1-008	1-009	1-010	1-011	1-012	1-013	1-014	1-015	1-016	1-017	1-018
23	0-971	0-974	0-977	0-980	0-983	0-985	0-987	0-989	0-990	0-991	0-992	0-993	0-994	0-995	0-996	0-997	0-998	0-999	1-000	1-001	1-002	1-003	1-004	1-005	1-006	1-007	1-008	1-009	1-010	1-011
22	0-964	0-968	0-970	0-973	0-976	0-978	0-980	0-981	0-982	0-983	0-984	0-985	0-986	0-987	0-988	0-989	0-990	0-991	0-992	0-993	0-994	0-995	0-996	0-997	0-998	0-999	1-000	1-001	1-002	1-003
21	0-958	0-961	0-964	0-967	0-969	0-971	0-973	0-975	0-977	0-978	0-979	0-980	0-981	0-982	0-983	0-984	0-985	0-986	0-987	0-988	0-989	0-990	0-991	0-992	0-993	0-994	0-995	0-996	0-997	0-998
20	0-951	0-955	0-958	0-960	0-963	0-965	0-967	0-969	0-970	0-971	0-972	0-973	0-974	0-975	0-976	0-977	0-978	0-979	0-980	0-981	0-982	0-983	0-984	0-985	0-986	0-987	0-988	0-989	0-990	0-991
19		0-957	0-959	0-961	0-962	0-964	0-965	0-966	0-967	0-968	0-969	0-970	0-971	0-972	0-973	0-974	0-975	0-976	0-977	0-978	0-979	0-980	0-981	0-982	0-983	0-984	0-985	0-986	0-987	0-988
18		0-951	0-953	0-955	0-957	0-958	0-959	0-960	0-961	0-962	0-963	0-964	0-965	0-966	0-967	0-968	0-969	0-970	0-971	0-972	0-973	0-974	0-975	0-976	0-977	0-978	0-979	0-980	0-981	0-982
17		0-946	0-948	0-950	0-952	0-953	0-954	0-955	0-956	0-957	0-958	0-959	0-960	0-961	0-962	0-963	0-964	0-965	0-966	0-967	0-968	0-969	0-970	0-971	0-972	0-973	0-974	0-975	0-976	0-977
16		0-941	0-943	0-945	0-947	0-948	0-949	0-950	0-951	0-952	0-953	0-954	0-955	0-956	0-957	0-958	0-959	0-960	0-961	0-962	0-963	0-964	0-965	0-966	0-967	0-968	0-969	0-970	0-971	0-972
15		0-937	0-939	0-941	0-943	0-944	0-945	0-946	0-947	0-948	0-949	0-950	0-951	0-952	0-953	0-954	0-955	0-956	0-957	0-958	0-959	0-960	0-961	0-962	0-963	0-964	0-965	0-966	0-967	0-968
14		0-933	0-935	0-936	0-938	0-939	0-940	0-941	0-942	0-943	0-944	0-945	0-946	0-947	0-948	0-949	0-950	0-951	0-952	0-953	0-954	0-955	0-956	0-957	0-958	0-959	0-960	0-961	0-962	0-963
13		0-929	0-931	0-932	0-934	0-935	0-936	0-937	0-938	0-939	0-940	0-941	0-942	0-943	0-944	0-945	0-946	0-947	0-948	0-949	0-950	0-951	0-952	0-953	0-954	0-955	0-956	0-957	0-958	0-959
12		0-925																												
11																														
10																														
9																														
8																														

The closing errors.

22. The tables just given exhibit the "closing errors" or differences between u, v, w and u_y, v_y, w_y respectively. The formulæ for these differences $u - u_y$ etc. for the four cases will now be considered separately and approximate expressions found for them.

Case I, when $\delta a = 1, \delta b = 0, u_0 = 0, w_0 = 0$

Along the parallel OM the changes at M are given by equations (21) and (22) in which suffix zero may be added to β, R, λ to indicate that it applies to latitude λ_0 . Apply the changes at M to the case of a meridian: it is necessary to consider cases I, III, IV of § 20 and the following equations are deduced:

$$\left. \begin{aligned} u_y &= \left[-0.2807\lambda + 24.26\sin 2\lambda - 0.02\sin 4\lambda \right]_{\lambda_0}^{\lambda} + R_0(1 - (\cos \beta_0 L)) \\ v_y - v_0 &= -R_0 \left(\frac{L^0 \cot \lambda_0}{57.3} + \frac{1}{\beta_0} \tan \lambda_0 \sin (\beta_0 L) \right) \\ &\quad - \frac{R_0 v_0}{\beta_0 a} \sqrt{1 - e^2} \sin (\beta_0 L) \left[\tan \lambda - 0.000,058, 2\lambda^0 + 0.000,004, 2\sin 2\lambda \right]_{\lambda_0}^{\lambda} \dots (33) \\ w_y &= -\frac{R_0 v_0}{\beta_0 \nu} \sec \lambda \sin (\beta_0 L) \dots \dots \dots \\ R_0 &= 16.1718 \sin 2\lambda_0 \left(1 + \frac{\sin^3 \lambda_0 (1 - e^2)}{1 - e^2 \sin^2 \lambda_0} \right) \dots \dots \dots \end{aligned} \right\}$$

Next proceeding along ON, the changes at N are given by (29), and applying these initial values to the parallel NP the following equations are formed by consideration of cases I and III of § 18

$$\left. \begin{aligned} u_n &= R(1 - \cos (\beta L)) + \cos (\beta L) \left[-0.2807\lambda^0 + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ v_n - v_0 &= -R \left(\frac{L^0 \cot \lambda}{57.3} + \frac{1}{\beta} \tan \lambda \sin (\beta L) \right) \\ &\quad + \frac{1}{\beta} \tan \lambda \sin (\beta L) \left[-0.2807\lambda^0 + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ w_n &= -\frac{R}{\beta} \sec \lambda \sin (\beta L) + \frac{1}{\beta} \sec \lambda \sin (\beta L) \left[-0.2807\lambda^0 + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \end{aligned} \right\} (34)$$

From (33) and (34) it follows that

$$\left. \begin{aligned} u_n - u_y &= \left[R(1 - \cos (\beta L)) \right]_{\lambda_0}^{\lambda} + (1 - \cos (\beta L)) \left[0.2807\lambda^0 - 24.26 \sin 2\lambda + 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ v_n - v_y &= - \left[R \left(\frac{L^0 \cot \lambda}{57.3} + \frac{1}{\beta} \tan \lambda \sin (\beta L) \right) \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{1}{\beta} \tan \lambda \sin (\beta L) \left[-0.280 \lambda L^0 + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{R_0 v_0}{\beta_0 a} \sqrt{1 - e^2} \sin (\beta_0 L) \left[\tan \lambda - 0.000,058, 2\lambda^0 + 0.000,004, 2\sin 2\lambda \right]_{\lambda_0}^{\lambda} \\ w_n - w_y &= -\frac{R}{\beta} \sec \lambda \sin (\beta L) + \frac{1}{\beta} \sec \lambda \sin (\beta L) \left[-0.2807\lambda^0 + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &\quad + \frac{R_0 v_0}{\beta_0 \nu} \sec \lambda \sin (\beta_0 L) \end{aligned} \right\} (35)$$

23. Equations (35) may be simplified and written in approximate form if terms depending on ϵ^2 are neglected. The closing errors will still be expressed with sufficient accuracy. Then β becomes unity and " a " may be substituted for ν . Denote $\lambda - \lambda_0$ by θ . In what follows u , etc. are expressed in seconds and λ , L , θ are expressed in radians except when the degree mark is added—thus λ° , and $\lambda^\circ/\lambda = 57.3$. The successive terms of (35) are taken one by one and reduced. \doteq denotes approximate equality.

Case I, when $\delta a = 1$, $\delta b = 0$, $u_0 = 0$, $w_0 = 0$

$$\text{Here } \frac{R}{16.17} \doteq \sin 2\lambda (1 + \sin^2 \lambda) = \frac{3}{4} \sin 2\lambda - \frac{1}{4} \sin 4\lambda$$

$$\begin{aligned} u_x - u_y &\doteq (1 - \cos L) \left[16.17 \left(\frac{3}{4} \sin 2\lambda - \frac{1}{4} \sin 4\lambda \right) + 16.08\lambda - 24.26 \sin 2\lambda + 0.02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ &= (1 - \cos L) \left[16.08\lambda - 4.02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ &= 16.1 (1 - \cos L) \left\{ \theta - \frac{1}{4} \sin 2\theta \cos 2(\lambda + \lambda_0) \right\} \\ &= .281 \theta^\circ (1 - \cos L) \left\{ 1 - \frac{\sin 2\theta}{2\theta} \cos 2(\lambda + \lambda_0) \right\} \dots \dots \dots (36) \end{aligned}$$

$$\frac{R}{16.17} \cot \lambda = 2 \cos^2 \lambda (1 + \sin^2 \lambda) = \frac{5}{4} + \cos 2\lambda - \frac{1}{4} \cos 4\lambda$$

$$\frac{R}{16.17} \tan \lambda = 2 \sin^2 \lambda (1 + \sin^2 \lambda) = \frac{7}{4} - 2 \cos 2\lambda + \frac{1}{4} \cos 4\lambda$$

$$\begin{aligned} \left[-R \left(\frac{L^\circ \cot \lambda}{57.3} + \tan \lambda \sin L \right) \right]_{\lambda_0}^\lambda &= -16.17 \frac{L^\circ}{57.3} \left[\frac{5}{4} + \cos 2\lambda - \frac{1}{4} \cos 4\lambda + \frac{\sin L}{L} \left(\frac{7}{4} - 2 \cos 2\lambda + \frac{1}{4} \cos 4\lambda \right) \right]_{\lambda_0}^\lambda \\ &= +.2828 L^\circ \left\{ 2 \left(1 - 2 \frac{\sin L}{L} \right) \sin \theta \sin (\lambda + \lambda_0) - \frac{1}{2} \left(1 - \frac{\sin L}{L} \right) \sin 2\theta \sin 2(\lambda + \lambda_0) \right\} \\ &= -0.0098 L^\circ \theta^\circ \left\{ \left(2 \frac{\sin L}{L} - 1 \right) \frac{\sin \theta}{\theta} \sin (\lambda + \lambda_0) + \frac{1}{2} \left(1 - \frac{\sin L}{L} \right) \frac{\sin 2\theta}{2\theta} \sin 2(\lambda + \lambda_0) \right\} \end{aligned}$$

$$\begin{aligned} \frac{1}{\beta} \tan \lambda \sin (\beta L) &\left[-0.2807 \lambda^\circ + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^\lambda \\ &\doteq \theta^\circ \tan \lambda \sin L \left\{ -0.2807 + .8469 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) - .0014 \frac{\sin 2\theta}{2\theta} \cos 2(\lambda + \lambda_0) \right\} \\ &\doteq -0.0049 L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ 1 - 3 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) \right\} \end{aligned}$$

$$\begin{aligned} \frac{R_0 \nu_0}{\beta_0 a} \sqrt{1 - \epsilon^2} \sin (\beta_0 L) &\left[\tan \lambda - 0.000,058, 2\lambda^\circ + 0.000,004, 2 \sin 2\lambda \right]_{\lambda_0}^\lambda \\ &\doteq 16.17 \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \sin L \left\{ \frac{\sin \theta}{\cos \lambda \cos \lambda_0} - 0.000,058, 2\theta^\circ + 0.000,008, 4 \cos (\lambda + \lambda_0) \sin \theta \right\} \\ &\doteq +0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \left\{ \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0.00384 \right\} \end{aligned}$$

Hence

$$\begin{aligned}
 v_x - v_y &= +0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ 2 \left(\frac{L}{\sin L} - 2 \right) \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} + \left(1 - \frac{L}{\sin L} \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{\theta} \right. \\
 &\quad \left. + \left(-1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right) \tan \lambda \right. \\
 &\quad \left. + \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \left(\sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0.00334 \right) \right\} \\
 &= 0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left[-\tan \lambda - 0.00334 \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \right. \\
 &\quad \left. + \frac{\sin \theta}{\theta} \left\{ 2 \left(\frac{L}{\sin L} - 2 \right) \sin (\lambda + \lambda_0) + 3 \cos (\lambda + \lambda_0) \tan \lambda + 2 \sin \lambda_0 (1 + \sin^2 \lambda_0) \sec \lambda \right\} \right. \\
 &\quad \left. + \left(1 - \frac{L}{\sin L} \right) \sin 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right]
 \end{aligned}$$

Now $\frac{\sin \theta}{\theta} \doteq 1$ and $\frac{L}{\sin L} \doteq 1$ the error being about .01 when θ or $L = 15^\circ$: hence $\frac{\sin \theta}{\theta}$ or $\frac{\sin 2 \theta}{2 \theta}$ may be treated as unity when multiplied by $\frac{L}{\sin L} - 1$. It follows that

$$\begin{aligned}
 v_x - v_y &\doteq +0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \tan \lambda \left(\cos (\lambda + \lambda_0) - 1 \right) - 0.00334 \sin 2 \lambda_0 (1 + \sin^2 \lambda_0) \right. \\
 &\quad \left. + 2 \sec \lambda \sin^3 \lambda_0 + \left(\frac{L}{\sin L} - 1 \right) \left(2 \sin (\lambda + \lambda_0) - \sin 2 (\lambda + \lambda_0) \right) \right\} \quad (37)
 \end{aligned}$$

$$\begin{aligned}
 -\frac{R}{\beta} \sec \lambda \sin (\beta L) + \frac{R_0 v_0}{\beta_0 v} \sec \lambda \sin (\beta_0 L) &\doteq -\sec \lambda \sin L (R - R_0) \\
 &= -0.2823 L^\circ \frac{\sin L}{L} \sec \lambda \left[\frac{3}{4} \sin 2 \lambda - \frac{1}{4} \sin 4 \lambda \right]_{\lambda_0}^\lambda \\
 &= -0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} - \cos 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right\} \\
 \frac{1}{\beta} \sec \lambda \sin (\beta L) &\left[-0.2807 \lambda^\circ + 24.26 \sin 2 \lambda - 0.02 \sin 4 \lambda \right]_{\lambda_0}^\lambda \\
 &\doteq \sec \lambda \sin L \left\{ -0.2807 \theta^\circ + 48.52 \cos (\lambda + \lambda_0) \sin \theta \right\} \\
 &= -0.0049 L^\circ \theta^\circ \sec \lambda \frac{\sin L}{L} \left\{ 1 - 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
 \end{aligned}$$

Hence

$$w_x - w_y \doteq -0.0049 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ 1 - \cos 2 (\lambda + \lambda_0) \frac{\sin 2 \theta}{2 \theta} \right\} \quad (38)$$

24. Case II, when $\delta a = 0$, $\delta b = 1$, $u_0 = 0$, $w_0 = 0$. Here

$$\begin{aligned}
 R &= -16.2258 \frac{(1 - e^2) \sin^3 \lambda}{1 - e^2 \sin^2 \lambda} \sin 2 \lambda \\
 &\doteq -8.11 (\sin 2 \lambda - \frac{1}{4} \sin 4 \lambda)
 \end{aligned}$$

Equations (35) hold for this case if we use the above value of R and change the quantity $-0.2807 \lambda + 24.26 \sin 2 \lambda - 0.02 \sin 4 \lambda$ into $-0.2847 \lambda - 24.34 \sin 2 \lambda + 0.02 \sin 4 \lambda$.

Then

$$\begin{aligned} u_x - u_y &= (1 - \cos L) \left[-8.11 (\sin 2\lambda - \frac{1}{2} \sin 4\lambda) + 16.3\lambda + 24.26 \sin 2\lambda - 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &= (1 - \cos L) \left[16.3\lambda + 16.1 \sin 2\lambda + 4.04 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &= .283 \theta^\circ (1 - \cos L) \left\{ 1 + 2 \frac{\sin \theta}{\theta} \cos (\lambda + \lambda_0) + \frac{\sin 2\theta}{2\theta} \cos 2(\lambda + \lambda_0) \right\} \quad (39) \end{aligned}$$

$$\frac{R_1}{16.28} \cot \lambda = -2 \sin^3 \lambda \cos \lambda \cot \lambda = -2 \sin^3 \lambda \cos^2 \lambda = -\frac{1}{4} (1 - \cos 4\lambda)$$

$$\frac{R_2}{16.28} \tan \lambda = -2 \sin^3 \lambda \cos \lambda \tan \lambda = -2 \sin^4 \lambda = -\frac{1}{4} (3 - 4 \cos 2\lambda + \cos 4\lambda)$$

$$\begin{aligned} \therefore \left[-R \left(\frac{L^\circ \cot \lambda}{57.8} + \tan \lambda \sin L \right) \right]_{\lambda_0}^{\lambda} &= \frac{16.28}{4} \cdot \frac{L^\circ}{57.8} \left[1 - \cos 4\lambda + \frac{\sin L}{L} (3 - 4 \cos 2\lambda + \cos 4\lambda) \right] \\ &= 0.0049 L^\circ \theta^\circ \left\{ \left(1 - \frac{\sin L}{L} \right) \sin 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} + 2 \frac{\sin L}{L} \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\} \end{aligned}$$

$$\begin{aligned} \frac{1}{\beta} \tan \lambda \sin \beta L &\left[-0.2847 \lambda^\circ - 24.84 \sin 2\lambda + 0.02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\ &= \tan \lambda \sin L \theta^\circ \left\{ -0.2847 - 0.8497 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} + 0.0014 \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\ &= -L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ +0.00497 + 0.0148 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} - 0.00002 \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\ &= -0.0050 L^\circ \theta^\circ \tan \lambda \frac{\sin L}{L} \left\{ 1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\} \\ \frac{R_0 \nu_0}{\beta_0 a} \sqrt{1 - e^2} \sin (\beta_0 L) &\left[\tan \lambda - 0.000,058, 2\lambda^\circ + 0.000,004, 2 \sin 2\lambda \right]_{\lambda_0}^{\lambda} \\ &= -16.23 \sin 2\lambda_0 \sin^3 \lambda_0 \sin L \left\{ \frac{\sin \theta}{\cos \lambda \cos \lambda_0} - 0.000058 2\theta^\circ + 0.0000084 \cos (\lambda + \lambda_0) \sin \theta \right\} \\ &= -0.0050 L^\circ \theta^\circ \frac{\sin L}{L} \sin 2\lambda_0 \sin^3 \lambda_0 \left\{ \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} - 0.00334 \right\} \end{aligned}$$

Hence

$$\begin{aligned} u_x - v_y &= 0.0050 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \left(\frac{L}{\sin L} - 1 \right) \sin 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} + 2 \sin (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right. \\ &\quad \left. - \left(1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right) \tan \lambda \right. \\ &\quad \left. - \sin 2\lambda_0 \sin^3 \lambda_0 \left(-0.00334 + \sec \lambda \sec \lambda_0 \frac{\sin \theta}{\theta} \right) \right\} \\ &= 0.0050 L^\circ \theta^\circ \frac{\sin L}{L} \left\{ \frac{\sin \theta}{\theta} \left(\sin 2\lambda_0 \cos \lambda_0 \sec \lambda - \cos (\lambda + \lambda_0) \tan \lambda \right) - \tan \lambda \right. \\ &\quad \left. + \left(\frac{L}{\sin L} - 1 \right) \sin 2(\lambda + \lambda_0) + 0.00334 \sin 2\lambda_0 \sin^3 \lambda_0 \right\} \quad (40) \end{aligned}$$

$$\begin{aligned}
-\frac{R}{\beta} \sec \lambda \sin (\beta L) + \frac{R_0 \nu_0}{\beta_0 \nu} \sec \lambda \sin (\beta_0 L) &= -\sec \lambda \sin L (R - R_0) \\
&= +16 \cdot 23 \sec \lambda \sin L \left[\frac{1}{2} \sin 2\lambda - \frac{1}{4} \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
&= 0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} - \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \\
\frac{1}{\beta} \sec \lambda \sin (\beta L) &\left[-0 \cdot 2847\lambda - 24 \cdot 34 \sin 2\lambda + 0 \cdot 02 \sin 4\lambda \right]_{\lambda_0}^{\lambda} \\
&= \sec \lambda \sin L \left\{ -0 \cdot 2847\theta^\circ - 48 \cdot 68 \cos (\lambda + \lambda_0) \cdot \sin \theta \right\} \\
&= -0 \cdot 0050 L^\circ \theta^\circ \sec \lambda \frac{\sin L}{L} \left\{ 1 + 3 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} \right\}
\end{aligned}$$

Hence

$$w_x - w_y = -0 \cdot 0050 L^\circ \theta^\circ \frac{\sin L}{L} \sec \lambda \left\{ 1 + 2 \cos (\lambda + \lambda_0) \frac{\sin \theta}{\theta} + \cos 2(\lambda + \lambda_0) \frac{\sin 2\theta}{2\theta} \right\} \quad (41)$$

25. The closing errors for cases III and IV have also been considered and are practically zero. This is at once evident also from the computed values of $u_x u_y$ etc. which agree to at least 0·001 of a second. It is otherwise clear that there would be no closing error on a sphere caused by moving the origin: and accordingly the effect on a spheroid must vanish with e^2 and accordingly have e^2 as a factor. In considering closing errors then it is only necessary to take cases I and II into account, and this may be done by means of equations (36) to (41). The form of these equations explains how the closing errors found in tables II, III, IV approximately satisfied the empirical relations (18). The relations would not have been equally satisfactory for case I and case II considered independently.

In the case of Indian triangulation θ° only exceeds 8° for values of L° between -7° and -1° and is greater than -8° for values of L° between -5° and $+3^\circ$: so that we can always consider one of the quantities, θ° or L° , numerically less than 8° . Closing errors for the elementary area $dL d\lambda$ are now deduced from the equations (36) to (41). In what follows L is treated as identical with $\sin L$.

Putting U_1 for $u_x - u_y$ in case I, U_2 for $u_x - u_y$ in case II etc. we have, omitting small terms

$$dU_1 = 16 \cdot 1 \sin L (1 - \cos 4\lambda) dL d\lambda \quad (42)$$

$$dU_2 = 16 \cdot 1 \sin L (1 + 2 \cos 2\lambda + \cos 4\lambda) dL d\lambda \quad (43)$$

$$\begin{aligned}
dV_1 &= -16 \cdot 17 \cos L \left\{ -2 \sin 2\lambda + \sin 4\lambda + 4 \sin 2\lambda - \sin 4\lambda \right\} dL d\lambda \\
&\quad + \cos L \left\{ \sec^2 \lambda (-16 \cdot 08\lambda + 24 \cdot 26 \sin 2\lambda) + \tan \lambda (-16 \cdot 08 + 48 \cdot 52 \cos 2\lambda) \right\} dL d\lambda \\
&\quad + 16 \cdot 17 \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \cos L \left\{ \sec^2 \lambda - 0 \cdot 00334 \right\} dL d\lambda
\end{aligned}$$

$$= 16 \cdot 1 \cos L \left\{ \sec^2 \lambda \left(-\lambda + \frac{2}{3} \sin 2\lambda + \sin 2\lambda_0 (1 + \sin^2 \lambda_0) \right) + \tan \lambda (3 \cos 2\lambda - 1) - 2 \sin 2\lambda \right\} dL d\lambda$$

$$= 16 \cdot 1 \cos L \left\{ \sec^2 \lambda (0 \cdot 871 - \lambda + \frac{2}{3} \sin 2\lambda) + \sin 2\lambda + 4 \tan \lambda \right\} dL d\lambda$$

$$= 16 \cdot 1 \cos L \left\{ \sec^2 \lambda (0 \cdot 871 - \lambda) + \sin 2\lambda - \tan \lambda \right\} dL d\lambda \quad (44)$$

$$\begin{aligned}
dV_2 &= 16 \cdot 23 \cos L \left\{ \sin 4\lambda + 2 \sin 2\lambda - \sin 4\lambda \right\} dL d\lambda \\
&\quad + \cos L \left\{ \sec^2 \lambda (-16 \cdot 3\lambda - 24 \cdot 34 \sin 2\lambda) + \tan \lambda (-16 \cdot 3 - 48 \cdot 68 \cos 2\lambda) \right\} dL d\lambda \\
&\quad - 16 \cdot 23 \sin 2\lambda_0 \sin^3 \lambda_0 \cos L \sec^2 \lambda dL d\lambda \\
&= 16 \cdot 2 \cos L \left\{ 2 \sin 2\lambda + \sec^2 \lambda \left(-\lambda - \frac{3}{2} \sin 2\lambda - \sin 2\lambda_0 \sin^3 \lambda_0 \right) + \tan \lambda (-1 - 3 \cos 2\lambda) \right\} dL d\lambda \\
&= -16 \cdot 2 \cos L \left\{ \sec^2 \lambda (0 \cdot 125 + \lambda + \frac{3}{2} \sin 2\lambda) + 2 \tan \lambda - \sin 2\lambda \right\} dL d\lambda \quad \dots \dots \dots (45) \\
&= -16 \cdot 2 \cos L \left\{ \sec^2 \lambda (0 \cdot 125 + \lambda) + 5 \tan \lambda - \sin 2\lambda \right\} dL d\lambda
\end{aligned}$$

Putting $\lambda_0 = 24^\circ 7' 11'' \cdot 26$

$$\begin{aligned}
dW_1 &= -32 \cdot 34 \cos L \cos \lambda (1 + 3 \sin^2 \lambda) dL d\lambda \\
&\quad + \cos L \left\{ \sin \lambda \sec^2 \lambda (-16 \cdot 1\lambda + 24 \cdot 26 \sin 2\lambda) + \sec \lambda (-16 \cdot 1 + 48 \cdot 5 \cos 2\lambda) \right\} dL d\lambda \\
&= -16 \cdot 1 \cos L \left\{ 2 \cos \lambda (1 + 3 \sin^2 \lambda) + \sin \lambda \sec^2 \lambda (\lambda - \frac{3}{2} \sin 2\lambda) + \sec \lambda (1 - 3 \cos 2\lambda) \right\} dL d\lambda \\
&= -16 \cdot 1 \cos L \left\{ \lambda \tan \lambda + 1 + 5 \cos^2 \lambda - 6 \cos^4 \lambda \right\} \sec \lambda dL d\lambda \quad \dots \dots \dots (46)
\end{aligned}$$

$$\begin{aligned}
dW_2 &= 16 \cdot 2 \cos L \times 6 \cos \lambda \sin^2 \lambda dL d\lambda \\
&\quad + \cos L \left\{ \sin \lambda \sec^2 \lambda (-16 \cdot 3\lambda - 24 \cdot 34 \sin 2\lambda) + \sec \lambda (-16 \cdot 3 - 48 \cdot 68 \cos 2\lambda) \right\} dL d\lambda \\
&= -16 \cdot 2 \cos L \left\{ -6 \cos \lambda \sin^2 \lambda + \sec \lambda (\lambda \tan \lambda + \frac{3}{2} \tan \lambda \sin 2\lambda + 1 + 3 \cos 2\lambda) \right\} dL d\lambda \\
&= -16 \cdot 2 \cos L \sec \lambda \left\{ \lambda \tan \lambda + 1 - 3 \cos^2 \lambda + 6 \cos 4\lambda \right\} dL d\lambda \quad \dots \dots \dots (47)
\end{aligned}$$

By means of equations (42) to (47) it is possible to find the changes u, v, w at P as computed by any route. For

$$u = u_y + \int dU$$

the integration being taken over the area between the desired route (upper limit) and the central parallel and the meridian through P : and u_y being the quantity to be found by properly combining the four cases. This obviously does not get rid of the multiple values obtainable for u, v, w according to the route followed, but if any route has special advantages, results of following it become available. One such route is the geodesic, or the shortest path between any point and the origin. There is something to be said in favour of following this route, and the subject of the geodesic is accordingly considered in some detail in the following chapter, where a direct method of finding the quantities u, v, w along a geodesic is also made use of.

In concluding this chapter it may be pointed out that the equations (42) to (47) enable the differences of the values of u, v, w to be rapidly estimated. This makes it clear at once how far it is a matter of importance to strictly adhere to any route that may be selected; for the difference in values that will be found by any two routes is the closing error.

CHAPTER II.

Geodesics on a Spheroid.

1. It is now necessary to develop some properties and relations of a geodesic in order that the changes of coordinates due to change of axes may be computed along geodesics. A fundamental relation of a geodesic on a conicoid is*

$$pD = \text{constant} \quad \dots \dots \dots (1)$$

where p is the perpendicular from the centre on the tangent plane at a point and D is the semidiameter of the quadric parallel to the tangent to the curve at the same point. In the case of a spheroid there is symmetry about the polar axis. In the figure ZOX is the equatorial plane and YCY' any meridian. Let P be any point on a geodesic: then the plane through O parallel to the tangent plane at P is the plane DOX , where OD is the diameter conjugate to OP . If ϕ is the eccentric angle of P so that the coordinates of P are $O, a \cos \phi, b \sin \phi$ then

$$Od^2 = a^2 \sin^2 \phi + b^2 \cos^2 \phi$$

For a geodesic proceeding from P in azimuth A , the semidiameter parallel to it is OQ where $\angle DOQ = 180^\circ - A$ and so

$$\begin{aligned} D &= \left(\frac{\sin^2 A}{a^2} + \frac{\cos^2 A}{a^2 \sin^2 \phi + b^2 \cos^2 \phi} \right)^{-\frac{1}{2}} \\ &= \left(\frac{a^2 (a^2 \sin^2 \phi + b^2 \cos^2 \phi)}{a^2 + \sin^2 A \cos^2 \phi (b^2 - a^2)} \right)^{\frac{1}{2}} \end{aligned}$$

Also,

$$p^2 = \frac{1}{\frac{\cos^2 \phi}{a^2} + \frac{\sin^2 \phi}{b^2}}$$

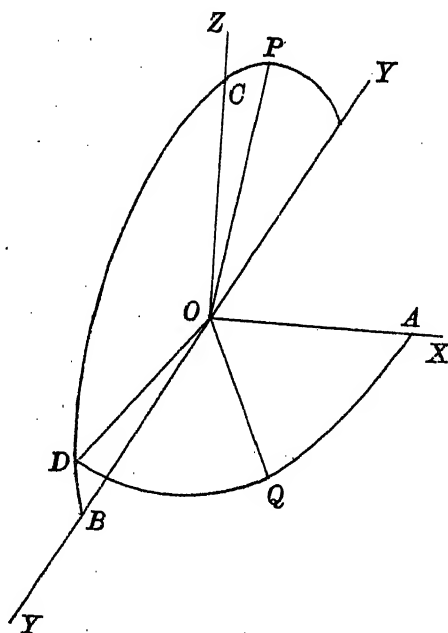
hence

$$p^2 D^2 = \frac{a^4 b^2}{a^2 + \sin^2 A \cos^2 \phi (b^2 - a^2)}$$

which is constant along a geodesic. It follows that

$$\sin A \cos \phi = \text{constant} = k = \sin A_0 \quad \dots \dots \dots (2)$$

along any geodesic, A_0 being the azimuth of the geodesic on crossing the equator.



* *Geometry of Three Dimensions* by George Salmon, 3rd edition, § 397.

2. Take two consecutive points on a spheroid and let A be the azimuth of the elementary line joining them. Then if the latitudes and longitudes are λ, L ; $\lambda + d\lambda, L + dL$ it follows that

$$\tan A = \frac{\nu \cos \lambda \, dL}{\rho \, d\lambda} \quad \dots \dots \dots (8)$$

where ρ is the radius of curvature of the meridian and ν is the normal terminated by the minor axis.

The relation between the latitude λ and the eccentric angle or "reduced latitude" ϕ is

$$\tan \phi = \frac{b}{a} \tan \lambda = \sqrt{1-e^2} \tan \lambda = (1-e) \tan \lambda \quad \dots \dots \dots (4)$$

Differentiating (2) logarithmically

$$\frac{dA}{d\phi} = \tan A \tan \phi \quad \dots \dots \dots (5)$$

and by (4)

$$\frac{d\phi}{d\lambda} = \frac{b}{a} \frac{\cos^2 \phi}{\cos^2 \lambda} \quad \dots \dots \dots (6)$$

Multiplying (5) and (6)

$$\begin{aligned} \frac{dA}{d\lambda} &= \frac{b}{a} \tan A \tan \phi \frac{\cos^2 \phi}{\cos^2 \lambda} = \frac{b^2}{a^2} \cdot \frac{\tan A \tan \lambda}{\cos^2 \lambda (1 + \frac{b^2}{a^2} \tan^2 \lambda)} \\ &= \frac{b^2}{a^2} \cdot \frac{\tan A \tan \lambda}{1 - e^2 \sin^2 \lambda} = \frac{\rho}{\nu} \tan A \tan \lambda \quad \dots \dots \dots (7) \end{aligned}$$

Equation (3) may be written

$$L = \int \frac{\rho}{\nu \cos \lambda} \cdot \tan A \, d\lambda \quad \dots \dots \dots (8)$$

and the integration of this will now be performed.

$$\cot^2 A = \operatorname{cosec}^2 A - 1 = \frac{\cos^2 \phi}{k^2} - 1 \quad \dots \dots \text{by (2)}$$

$$= \frac{1}{k^2 \left(1 + \frac{b^2}{a^2} \tan^2 \lambda\right)} - 1 = \frac{\cos^2 \lambda}{k^2 (1 - e^2 \sin^2 \lambda)} - 1$$

$$\therefore \tan A = \pm \frac{k \sqrt{1 - e^2 \sin^2 \lambda}}{\sqrt{1 - k^2 - (1 - k^2 e^2) \sin^2 \lambda}} = \pm \frac{k}{\sqrt{1 - k^2}} \cdot \frac{\sqrt{1 - e^2 \sin^2 \lambda}}{\sqrt{1 - a^2 \sin^2 \lambda}}$$

where $a^2 = \frac{1 - k^2 e^2}{1 - k^2}$ and the + sign is taken for 1st and 4th quadrants and the minus sign for the 2nd and 3rd quadrants.

Hence by (8)

$$L = \pm \frac{k(1 - e^2)}{\sqrt{1 - k^2}} \int \frac{1}{\sqrt{1 - e^2 \sin^2 \lambda} \sqrt{1 - a^2 \sin^2 \lambda}} \frac{d\lambda}{\cos \lambda}$$

$$\text{since } \frac{\rho}{\nu} = \frac{1 - e^2}{1 - e^2 \sin^2 \lambda}$$

Put x for $\sin \lambda$: then

$$L = \pm \frac{k(1 - e^2)}{\sqrt{1 - k^2}} \int \frac{dx}{\sqrt{1 - e^2 x^2} \sqrt{1 - a^2 x^2} (1 - x^2)} \quad \dots \dots \dots (9)$$

This is an elliptic integral which cannot be integrated exactly: but it may be developed in a series of integrable terms as follows.

Put $1-x^2 = y^2$, then

$$\begin{aligned} \frac{1}{\sqrt{1-e^2x^2}} \cdot \frac{1}{1-x^2} &= \frac{1}{y^2 \sqrt{1-e^2+e^2y^2}} = \frac{1}{\sqrt{1-e^2}} \cdot \frac{1}{y^2 \sqrt{1+\beta^2y^2}} \\ &= \frac{1}{\sqrt{1-e^2}} \left\{ \frac{1}{y^2} - \frac{1}{2} \beta^2 + \frac{1 \cdot 3}{2^2 \cdot 2} \beta^4 y^2 \dots \right\} \dots \quad (10) \end{aligned}$$

where $\beta^2 = e^2/(1-e^2)$: hence

$$\int \frac{dx}{\sqrt{1-e^2x^2} \sqrt{1-a^2x^2} (1-x^2)} = \frac{1}{\sqrt{1-e^2}} \int \frac{dx}{\sqrt{1-a^2x^2}} \left\{ \frac{1}{1-x^2} - \frac{1}{2} \beta^2 + \frac{3}{8} \beta^4 (1-x^2) \dots \right\} \quad (11)$$

Finally put $\sin \theta = ax = a \sin \lambda \dots \dots \dots (12)$

Then

$$\int \frac{dx}{\sqrt{1-a^2x^2}} = \frac{1}{a} \int \frac{\cos \theta d\theta}{\cos \theta} = \frac{\theta}{a} \dots \dots \dots (13)$$

$$\begin{aligned} \int \frac{dx}{(1-x^2) \sqrt{1-a^2x^2}} &= \frac{1}{a} \int \frac{d\theta}{1-\frac{\sin^2 \theta}{a^2}} = a \int \frac{d \tan \theta}{a^2 + a^2 - 1 \tan^2 \theta} \\ &= \frac{1}{\sqrt{a^2-1}} \tan^{-1} \left(\frac{\sqrt{a^2-1}}{a} \tan \theta \right) \dots \dots \dots (14) \end{aligned}$$

The remaining terms of (11) may be dealt with by the formula of reduction (17) now deduced.

$$\begin{aligned} u_n &= \int \frac{x^n dx}{\sqrt{1-a^2x^2}} = \frac{1}{a^2} \int \frac{x^{n-2} (1-\overline{1-a^2x^2}) dx}{\sqrt{1-a^2x^2}} \\ &= \frac{1}{a^2} \cdot u_{n-2} - \frac{1}{a^2} \int x^{n-2} \sqrt{1-a^2x^2} dx \dots \dots \dots (15) \end{aligned}$$

Integrating by parts

$$\begin{aligned} u_n &= -\frac{1}{a^2} \int x^{n-1} d \sqrt{1-a^2x^2} \\ &= -\frac{1}{a^2} x^{n-1} \sqrt{1-a^2x^2} + \frac{n-1}{a^2} \int x^{n-2} \sqrt{1-a^2x^2} dx \dots (16) \end{aligned}$$

Multiplying (15) by $(n-1)$ and adding to (16)

$$nu_n = \frac{n-1}{a^2} \cdot u_{n-2} - \frac{1}{a^2} \cdot x^{n-1} \sqrt{1-a^2x^2} \dots \dots \dots (17)$$

Hence

$$\int \frac{1-x^2}{\sqrt{1-a^2x^2}} dx = \frac{\theta}{a} \left(1 - \frac{1}{2a^2} \right) + \frac{1}{2a^2} x \sqrt{1-a^2x^2} \dots \dots \dots (18)$$

and from (9), (11), (13), (14) and (18)

$$\begin{aligned} L-L' &= \pm \frac{k \sqrt{1-e^2}}{\sqrt{1-k^2}} \left[\frac{1}{\sqrt{a^2-1}} \tan^{-1} \left(\frac{\sqrt{a^2-1}}{a} \tan \theta \right) - \frac{1}{2} \beta^2 \cdot \frac{\theta}{a} \right. \\ &\quad \left. + \frac{3}{8} \beta^4 \left\{ \frac{\theta}{a} \left(1 - \frac{1}{2a^2} \right) + \frac{1}{2a^2} \sin \lambda \cos \theta \right\} \dots \right] \quad (19) \end{aligned}$$

in which

$$\left. \begin{aligned} \theta &= \sin^{-1}(a \sin \lambda) \\ a^2 &= \frac{1-k^2 e^2}{1-k^2} \\ \beta^2 &= \frac{e^2}{1-e^2} \end{aligned} \right\} \dots \dots \dots (20)$$

Hence

$$\begin{aligned} \sin \theta &= \sqrt{\frac{1-k^2 e^2}{1-k^2}} \sin \lambda \\ \tan \theta &= \frac{\sqrt{1-k^2 e^2} \sin \lambda}{\sqrt{1-k^2 - (1-k^2 e^2) \sin^2 \lambda}} = \frac{\sqrt{1-k^2 e^2} \tan \lambda}{\sqrt{1-k^2 - k^2 \tan^2 \lambda (1-e^2)}} \\ &= \sqrt{\frac{1-k^2 e^2}{1-e^2}} \frac{\tan \phi}{\sqrt{1-k^2 \sec^2 \phi}} \quad \text{by (4)} \\ &= \pm \sqrt{\frac{1-k^2 e^2}{1-e^2}} \tan \phi \sec A \quad \dots \dots \dots (21) \end{aligned}$$

$$\therefore \frac{\sqrt{a^2-1}}{a} \tan \theta = \pm k \tan \phi \sec A = \pm \tan A \sin \phi \quad \dots \dots \dots (22)$$

Since θ is given by (20) we may always arrange that it shall be in 1st quadrant and the sign in (22) must be taken accordingly.

Put

$$\tan \psi = \pm \tan A \sin \phi \quad \dots \dots \dots (23)$$

ψ being always in the first quadrant.

Then (19) may be written

$$L-L' = \pm \left[\pm \psi - \frac{e^2}{2} (1 + \frac{1}{1+k^2} \frac{e^2}{8} + \dots) k \theta + \frac{3e^4}{16} (1 + \dots) k \sqrt{1-k^2} \sin \lambda \cos \theta + \dots \right] \quad \dots \dots (24)$$

Now by (2) it follows that

$$\tan A \sin \phi = \frac{\sin A_0 \sec \phi}{\sqrt{1 - \sin^2 A_0 \sec^2 \phi}} \sin \phi = \frac{\sin A_0 \tan \phi}{\sqrt{\cos^2 A_0 - \sin^2 A_0 \tan^2 \phi}} = \frac{\tan A_0 \tan \phi}{\sqrt{1 - \tan^2 A_0 \tan^2 \phi}}$$

Hence

$$\psi = \pm \tan^{-1} (\tan A \sin \phi) = \pm \sin^{-1} (\tan A_0 \tan \phi) \quad \dots \dots \dots (25)$$

Neglecting terms involving e^4 (23) may be written

$$[L] = \pm \left[\pm \psi - \frac{ke^2}{2} \theta \right] \quad \dots \dots \dots (26)$$

where ψ and θ are both in first quadrant and are defined by (25) and (20) respectively. This result is correct to nearest second for the terrestrial spheroid.

The rules for the double sign outside bracket are + 1st and 4th quadrant of azimuth
- 2nd and 3rd quadrant "

and for double signs before ψ + 1st and 2nd quadrant "
- 3rd and 4th quadrant "

The quantity θ may also be found from (21) which may be written

$$\tan \theta = \frac{1}{k} \sqrt{\frac{1-k^2 e^2}{1-e^2}} \cdot \tan \psi \quad \dots \quad (27)$$

3. To solve (26) for A_0 take the first approximation to A_0 , namely A_1 such that

$$\left[\sin^{-1} (\tan A_1 \tan \phi) \right]_{\phi'}^{\phi} = L - L' \quad \dots \quad (28)$$

and for brevity put $\tan A_1 \tan \phi = x$, $\tan A_1 \tan \phi' = y$ and $L - L' = \theta$

Then (26) becomes

$$\begin{aligned} \sin^{-1} x - \sin^{-1} y &= \theta \quad \dots \quad (29) \\ \therefore x \sqrt{1-y^2} - y \sqrt{1-x^2} &= \sin \theta \end{aligned}$$

Squaring and transposing

$$x^2 + y^2 - 2x^2 y^2 - \sin^2 \theta = 2xy \sqrt{(1-x^2)(1-y^2)}$$

Squaring again

$$\begin{aligned} \therefore \sin^4 \theta - 2 \sin^2 \theta (x^2 + y^2 - 2x^2 y^2) + (x^2 - y^2)^2 &= 0 \\ \text{or } (x^2 - y^2)^2 + 4x^2 y^2 \sin^2 \theta - 2(x^2 + y^2) \sin^2 \theta + \sin^4 \theta &= 0 \end{aligned}$$

and putting this into factors

$$(x^2 + y^2 + 2xy \cos \theta - \sin^2 \theta) (x^2 + y^2 - 2xy \cos \theta - \sin^2 \theta) = 0$$

Substituting for x and y it follows that

$$\tan^2 A_1 \left(\tan^2 \phi + \tan^2 \phi' \pm 2 \tan \phi \tan \phi' \cos (L - L') \right) = \sin^2 (L - L') \quad \dots \quad (30)$$

The double signs have been introduced by the process of squaring and it is necessary to return to (28) to decide which signs give the required solution.

First supposing $\tan \phi$ and $\tan \phi'$ of the same sign and $\phi > \phi'$: then from (28) changing the sign of $\tan \phi'$ diminishes the value of $\tan A_1$: hence by (30) we see that the lower sign must be taken. The same is true if $\tan \phi$ and $\tan \phi'$ are of opposite sign, and as ϕ and ϕ' are interchangeable this shows that the lower sign in (29) must always be taken.

Again if $\phi > \phi'$ the sign of $\tan A$ is the same as that of $L - L'$, and if $\phi < \phi'$ the sign of $\tan A$ is opposite to that of $L - L'$.

Hence we may write (30)

$$\tan A_1 = \pm \sqrt{\frac{\sin (L - L')}{\tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos L - L'}} \quad \dots \quad (31)$$

the upper or lower sign being taken according as $\phi >$ or $< \phi'$.

Also $\tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos L - L'$

$$= (1 - e^2) \left\{ \tan^2 \lambda + \tan^2 \lambda' - 2 \tan \lambda \tan \lambda' \cos L - L' \right\} \quad \text{by (4)}$$

$$= (1 - e^2) (\tan \lambda - \tan \lambda')^2 \left\{ 1 + \frac{4 \tan \lambda \tan \lambda'}{(\tan \lambda - \tan \lambda')^2} \sin^2 \frac{L - L'}{2} \right\}$$

$$= (1 - e^2) \left(\frac{\sin (\lambda - \lambda')}{\cos \lambda \cos \lambda'} \right)^2 \sec^2 \omega$$

where

$$\tan^2 \omega = \sin^2 \frac{L-L'}{2} \cdot \frac{\sin 2\lambda \sin 2\lambda'}{\sin^2 (\lambda - \lambda')}$$

so that finally

$$\left. \begin{aligned} \tan A_1 &= + \frac{\sin (L-L') \cos \omega \cos \lambda \cos \lambda'}{\sqrt{1-e^2} \sin (\lambda - \lambda')} \\ \tan \omega &= \frac{\sin \frac{1}{2} (L-L')}{\sin (\lambda - \lambda')} \sin 2\lambda \sin 2\lambda' \end{aligned} \right\} \dots \dots \dots (32)$$

where

and $\cos \omega$ is taken positive.

4. Denote by ${}_1A$ the approximate value of A which corresponds to A_1 which is an approximate value of A_0 .

Then

$$\left. \begin{aligned} \sin {}_1A \cos \phi &= \sin A_1 \\ {}_1A + \delta_1 A &= A \end{aligned} \right\} \dots \dots \dots (33)$$

and

Hence

$$\begin{aligned} \tan A_1 &= \frac{\sin {}_1A \cos \phi}{\sqrt{1 - \sin^2 {}_1A \cos^2 \phi}} = \frac{1}{\sqrt{\sec^2 \phi \operatorname{cosec}^2 {}_1A - 1}} \\ &= \frac{1}{\sqrt{\sec^2 \phi \cot^2 {}_1A + \tan^2 \phi}} \end{aligned}$$

By (30)

$$\sin^2 (L-L') (\tan^2 \phi + \sec^2 \phi \cot^2 {}_1A) = \tan^2 \phi + \tan^2 \phi' - 2 \tan \phi \tan \phi' \cos (L-L')$$

$$\cot^2 {}_1A \sec^2 \phi \sin^2 (L-L') = \left\{ \tan \phi \cos (L-L') - \tan \phi' \right\}^2$$

$$\therefore \tan_1 A = + \frac{\sec \phi \sin (L-L')}{\tan \phi \cos (L-L') - \tan \phi'} \dots \dots \dots (34)$$

Similarly

$$\tan_1 A' = - \frac{\sec \phi' \sin (L-L')}{\tan \phi' \cos (L-L') - \tan \phi} \dots \dots \dots (35)$$

By differentiating (33) logarithmically with regard to A_1 and ${}_1A$ we get the relation between δA_1 and $\delta_1 A$ as follows:

$$\cot {}_1A \delta_1 A = \cot A_1 \delta A_1 \dots \dots \dots (36)$$

Equations (34) and (35) correspond to the ordinary equations of spherical trigonometry to which they reduce if the eccentric angles ϕ, ϕ' are replaced by latitudes λ, λ' .

5. Suppose next that

$$A_0 = A_1 + \delta A_1 \dots \dots \dots (37)$$

where δA_1 gives a second approximation to A_0 . Then by (26) neglecting terms in e^4 , it follows that

$$\begin{aligned} \frac{e^2 k}{2} (\theta - \theta') &= \pm \delta [\psi] = \delta \left[\sin^{-1} (\tan A_1 \tan \phi) \right]_{\phi'}^{\phi} \\ &= \delta A_1 \left[\frac{\sec^2 A_1 \tan \phi}{\sqrt{1 - \tan^2 A_1 \tan^2 \phi}} \right]_{\phi'}^{\phi} \dots \dots \dots (38) \end{aligned}$$

With notation of (20)

$$\sin \theta = a \sin \lambda = \sqrt{\frac{1-e^2 \sin^2 A_1}{1-\sin^2 A_1}} \sin \lambda = \sec A_1 \sin \lambda \sqrt{1-e^2 \sin^2 A_1}$$

since $k = \sin A_0 = \sin A_1$

$$\begin{aligned} \therefore \cos^2 \theta &= 1 - \sin^2 \lambda \sec^2 A_1 (1 - e^2 \sin^2 A_1) \\ &= \cos^2 \lambda - (1 - e^2) \tan^2 A_1 \sin^2 \lambda \end{aligned}$$

$$\therefore \frac{\sec^2 A_1 \tan \phi}{\sqrt{1 - \tan^2 A_1 \tan^2 \phi}} = \frac{\sqrt{1 - e^2} \sec^2 A_1 \tan \lambda}{\sqrt{1 - (1 - e^2) \tan^2 A_1 \tan^2 \lambda}} = \sqrt{1 - e^2} \sec^2 A_1 \sin \lambda \sec \theta$$

\therefore by (38)

$$\delta A_1 = \frac{e^2 \sin A_1 \cos^2 A_1}{2\sqrt{1-e^2}} \frac{(\theta - \theta')}{\left[\sin \lambda \sec \theta \right]_{\lambda'}^{\lambda}}$$

Now

$$\sin \lambda \sec \theta = \frac{1}{a} \tan \theta$$

$$\therefore \delta A_1 = \frac{e^2 \sin 2 A_1}{4} \sqrt{\frac{1 - e^2 \sin^2 A_1}{1 - e^2}} \cdot \frac{\theta - \theta'}{\tan \theta - \tan \theta'} \left. \vphantom{\frac{\theta - \theta'}{\tan \theta - \tan \theta'}} \right\} \dots \dots \dots (39)$$

in which

$$\theta = \sin^{-1} \left(\sin \lambda \sec A_1 \sqrt{1 - e^2 \sin^2 A_1} \right)$$

For computation

$$\sin \theta = \sin \phi \sec A_1 \left\{ 1 + \frac{e^2}{2} (\cos^2 \phi - \sin^2 A_1) \right\} \quad \text{since } \sin \lambda = \sin \phi \sqrt{1 - e^2 \cos^2 \phi}$$

$$= \sin \phi \sec A_1 \left(1 + \frac{e^2}{2} \cos^2 \phi \cos^2 A_1 \right) \quad \text{since } \sin A_1 = \sin A \cos \phi$$

Let $\sin \theta_1 = \sin \phi \sec A_1$ and $\theta = \theta_1 + \delta \theta$

$$\tan \theta_1 = \frac{\sin \phi}{\sqrt{\cos^2 A_1 - \sin^2 \phi}} = \frac{\sin \phi}{\sqrt{\cos^2 \phi - \sin^2 A \cos^2 \phi}} = \tan \phi \sec A$$

and $\delta \theta = \frac{e^2}{2} \cos^2 \phi \cos^2 A \tan \theta_1 = \frac{e^2}{4} \sin 2 \phi \cos A.$

With a given value of the larger quantity θ , θ' say θ it is clear that $\frac{\theta - \theta'}{\tan \theta - \tan \theta'}$ is greater the smaller the value of θ' ; its maximum value accordingly is $\frac{\theta}{\tan \theta}$ and the maximum value of this quantity occurs when $\theta = 0$, when it becomes unity. It is quite clear then that δA_1 cannot exceed $\frac{1}{2} e^2 \sin 2 A$, i. e. $6' \cdot \sin 2 A_1$ in the case of the terrestrial spheroid where $e^2 = \frac{1}{150}$.

6. If ds is the length of an elementary line

$$\rho d\lambda = -ds \cos A$$

$$\therefore s = -\int \rho \sec A d\lambda$$

On a geodesic

$$\cos A = \sqrt{1 - \sin^2 A_0 \sec^2 \phi} = \cos A_0 \sqrt{1 - \tan^2 A_0 \tan^2 \phi}$$

$$\rho = \frac{a(1-e^2)}{(1-e^2 \sin^2 \lambda)^{\frac{3}{2}}} ; \quad \tan \lambda = \frac{\tan \phi}{\sqrt{1-e^2}}$$

$$\sin \lambda = \frac{\sin \phi}{\sqrt{1-e^2 \cos^2 \phi}} ; \quad 1-e^2 \sin^2 \lambda = \frac{1-e^2}{1-e^2 \cos^2 \phi} ; \quad \cos \lambda = \cos \phi \sqrt{\frac{1-e^2}{1-e^2 \cos^2 \phi}}$$

whence

$$\rho = \frac{a}{\sqrt{1-e^2}} (1-e^2 \cos^2 \phi)^{\frac{3}{2}}$$

Also $\frac{d\lambda}{\sin \lambda \cos \lambda} = \frac{d\phi}{\sin \phi \cos \phi}$

$$d\lambda = d\phi \frac{\sqrt{1-e^2}}{1-e^2 \cos^2 \phi}$$

$$\begin{aligned} \therefore s &= - \int \frac{a(1-e^2 \cos^2 \phi)^{\frac{3}{2}}}{\sqrt{1-e^2}} \cdot \frac{\sec A_0}{\sqrt{1-\tan^2 A_0 \tan^2 \phi}} \cdot \frac{\sqrt{1-e^2}}{1-e^2 \cos^2 \phi} \cdot d\phi \\ &= -a \sec A_0 \int \sqrt{\frac{1-e^2 \cos^2 \phi}{1-\tan^2 A_0 \tan^2 \phi}} \cdot d\phi \end{aligned}$$

Put $\sin \phi = x \quad d\phi = \frac{dx}{\sqrt{1-x^2}}$

$$\begin{aligned} s &= -a \sec A_0 \int \sqrt{\frac{1-e^2(1-x^2)}{1-\tan^2 A_0 x^2/(1-x^2)}} \cdot \frac{dx}{\sqrt{1-x^2}} \\ &= -a \int \sqrt{\frac{1-e^2(1-x^2)}{\cos^2 A_0 - x^2}} \cdot dx = -a \sqrt{1-e^2} \int \frac{\sqrt{1+\beta^2 x^2}}{\sqrt{\cos^2 A_0 - x^2}} \cdot dx \end{aligned}$$

where $\beta^2 = \frac{e^2}{1-e^2}$

Put then $x = \cos A_0 \sin \chi = \sin \phi$

$$\begin{aligned} s &= -a \sqrt{1-e^2} \int_{\chi'}^{\chi} \frac{\sqrt{1+\beta^2 \cos^2 A_0 \sin^2 \chi}}{\cos A_0 \cos \chi} \cdot \cos A_0 \cos \chi d\chi \\ &= -a \sqrt{1-e^2} \int_{\chi'}^{\chi} \sqrt{1+h^2 \sin^2 \chi} d\chi \quad \text{where } h^2 = \beta^2 \cos^2 A_0 = \frac{e^2 \cos^2 A_0}{1-e^2} \\ &= -a \sqrt{1-e^2} \left[\chi + \frac{1}{2} h^2 \int \sin^2 \chi d\chi - \frac{1}{8} h^4 \int \sin^4 \chi d\chi \dots \right]_{\chi'}^{\chi} \end{aligned}$$

Now $\int \sin^n \chi d\chi = -\frac{1}{n} \cos \chi \sin^{n-1} \chi + \frac{(n-1)}{n} \int \sin^{n-2} \chi d\chi$

$$\therefore \int \sin^2 \chi d\chi = -\frac{1}{2} \cos \chi \sin \chi + \frac{\chi}{2}$$

$$\begin{aligned}
 \int \sin^4 \chi d\chi &= -\frac{1}{4} \cos \chi \sin^3 \chi - \frac{3}{4} \int \sin^2 \chi d\chi \\
 &= -\frac{1}{4} \cos \chi \sin^3 \chi - \frac{3}{8} \cos \chi \sin \chi + \frac{3}{8} \chi \\
 &\quad \text{etc.} \\
 \therefore s &= -a \sqrt{1-e^2} \left[\chi \left(1 + \frac{1}{4} h^2 - \frac{3}{64} h^4 \dots \right) - \frac{h^2}{4} \cos \chi \sin \chi \left(1 - \frac{3}{16} h^2 \dots \right) \right. \\
 &\quad \left. + \frac{h^4}{32} \cos \chi \sin^3 \chi (1 \dots) - \dots \right] \frac{\chi}{\chi} \dots (40)
 \end{aligned}$$

where

$$\left. \begin{aligned} \sin \chi &= \frac{\sin \phi}{\cos A_0} \\ h^2 &= \frac{e^2 \cos^2 A_0}{1-e^2} \end{aligned} \right\} \dots \dots \dots (41)$$

Otherwise

$$\text{since } \sin A \cos \phi = \sin A_0$$

$$\sin \chi = \frac{\sin \phi \sin A_0}{\sin A \cos \phi \cos A_0} = \tan \phi \tan A_0 \operatorname{cosec} A$$

Also

$$\tan^2 A_0 = \frac{\sin^2 A \cos^2 \phi}{1 - \sin^2 A \cos^2 \phi} = \frac{\sin^2 A}{\tan^2 \phi + \cos^2 A}$$

and

$$\sin \chi = \pm \sqrt{\frac{\tan^2 \phi}{\tan^2 \phi + \cos^2 A}}$$

$$\tan \chi = \pm \tan \phi \sec A \dots \dots \dots (42)$$

7. To facilitate reductions, tables are now given enabling the conversion from λ to ϕ and *vice versa* to be easily performed. They have been computed from formula $\lambda - \phi = \frac{1}{2} \cdot \frac{e \sin^2 \lambda}{1 - e \sin^2 \lambda}$ which is readily deducible from (4).

Table XXI.

λ	$\phi - \lambda$	λ	$\phi - \lambda$	λ	$\phi - \lambda$	λ	$\phi - \lambda$
°	' "	°	' "	°	' "	°	' "
0	-0 0'0	10	-1 57'3	20	-3 40'5	30	-4 57'1
1	-0 12'0	11	-2 8'4	21	-3 49'5	31	-5 3'0
2	-0 23'8	12	-2 19'4	22	-3 58'2	32	-5 8'5
3	-0 35'8	13	-2 30'3	23	-4 6'7	33	-5 13'5
4	-0 47'7	14	-2 41'0	24	-4 14'9	34	-5 18'2
5	-0 59'5	15	-2 51'4	25	-4 22'8	35	-5 22'5
6	-1 11'3	16	-3 1'8	26	-4 30'4	36	-5 26'4
7	-1 23'0	17	-3 11'9	27	-4 37'6	37	-5 29'9
8	-1 34'5	18	-3 21'6	28	-4 44'5	38	-5 33'1
9	-1 45'9	19	-3 31'2	29	-4 50'9	39	-5 35'8
10	-1 57'3	20	-3 40'5	30	-4 57'1	40	-5 38'2

Table XXII.

ϕ	$\lambda - \phi$	ϕ	$\lambda - \phi$	ϕ	$\lambda - \phi$	ϕ	$\lambda - \phi$
°	' "	°	' "	°	' "	°	' "
0	+0 0'0	10	+1 57'7	20	+3 41'1	30	+4 57'6
1	+0 12'0	11	+2 8'8	21	+3 50'1	31	+5 3'5
2	+0 23'9	12	+2 19'8	22	+3 58'8	32	+5 8'9
3	+0 35'9	13	+2 30'7	23	+4 7'3	33	+5 13'9
4	+0 47'9	14	+2 41'5	24	+4 15'5	34	+5 18'6
5	+0 59'7	15	+2 51'9	25	+4 23'4	35	+5 22'8
6	+1 11'5	16	+3 2'3	26	+4 30'9	36	+5 26'7
7	+1 23'3	17	+3 12'4	27	+4 38'1	37	+5 30'2
8	+1 34'8	18	+3 22'1	28	+4 45'0	38	+5 33'3
9	+1 46'2	19	+3 31'7	29	+4 51'4	39	+5 36'0
10	+1 57'7	20	+3 41'1	30	+4 57'6	40	+5 38'4

Values of A_0 , k , s together with certain of the quantities by means of which they are computed are now given in tabular form for geodesics passing through the origin and points L , ϕ , for values of $L'-L$ differing by 4° from 0 to 24° and for values of ϕ from 10° to 38° . It is clear that for longitudes east of the origin the value of A is $360 -$ (its value in table); and that for s there is no change.

TABLE XXIII.

ϕ	$L'-L$	0°	4°	8°	12°	16°	20°	24°
38°	$180^\circ - A_1$	$0^\circ 0' 0''$	$11^\circ 40' 32''.2$	$21^\circ 58' 5''.6$	$30^\circ 12' 43''.8$	$36^\circ 27' 5''.0$	$41^\circ 2' 45''.3$	$44^\circ 23' 53''.7$
	$-\delta A_1$	$0 0$	$1 37.2$	$2 41.6$	$3 6.7$	$3 2.2$	$2 48.9$	$2 27.1$
	$180^\circ - A_0$	$0 0 0$	$11 42 9$	$22 0 47$	$30 15 51$	$36 30 7$	$41 5 34$	$44 26 21$
	$\log k$	$-\infty$	1.8071363	1.5738218	1.7024180	1.7744081	1.8177511	1.8451916
	k^2	0.006682	0.006407	0.005743	0.004985	0.004318	0.003795	0.003407
	s/b	0.243714	0.250921	0.271891	0.302418	0.341099	0.385090	0.432724
34°	$180^\circ - A_1$	$0 0 0$	$16 46 26.0$	$30 1 52.5$	$39 8 30.0$	$45 6 17.2$	$48 59 53.3$	$51 34 18.2$
	$-\delta A_1$	$0 0$	$2 14.0$	$3 21.5$	$3 20.0$	$2 53.0$	$2 21.7$	$1 52.7$
	$180^\circ - A_0$	$0 0 0$	$16 48 40$	$30 5 14$	$39 11 50$	$45 9 10$	$49 2 15$	$51 36 6$
	$\log k$	$-\infty$	1.4612245	1.7001131	1.8007114	1.8506405	1.8780268	1.8941560
	k^2	0.006682	0.006123	0.005003	0.004013	0.003323	0.002872	0.002578
	s/b	0.173820	0.184246	0.212464	0.252555	0.299755	0.351166	0.405139
30°	$180^\circ - A_1$	$0 0 0$	$27 11 14.3$	$48 2 45.7$	$50 56 59.8$	$55 2 12.4$	$57 17 10.9$	$58 34 48.1$
	$-\delta A_1$	$0 0$	$3 24.5$	$3 27.8$	$2 36.8$	$1 51.6$	$1 18.3$	$0 54.9$
	$180^\circ - A_0$	$0 0 0$	$27 14 39$	$48 6 14$	$50 59 37$	$55 4 4$	$57 18 29$	$58 35 43$
	$\log k$	$-\infty$	1.6606593	1.8346251	1.8904627	1.9137238	1.9250992	1.9312075
	k^2	0.006682	0.005282	0.003562	0.002647	0.002191	0.001949	0.001814
	s/b	0.103942	0.121206	0.162319	0.213923	0.270030	0.328293	0.387702
26°	$180^\circ - A_1$	$0 0 0$	$52 54 33.6$	$60 59 16.4$	$62 59 19.7$	$63 41 3.3$	$63 56 27.7$	
	$-\delta A_1$	$0 0$	$2 46.9$	$1 7.7$	$0 32.1$	$0 16.4$	$0 6.5$	
	$180^\circ - A_0$	$0 0 0$	$52 57 21$	$61 0 24$	$62 59 52$	$63 41 20$	$63 56 34$	
	$\log k$	$-\infty$	1.9020954	1.9418474	1.9498721	1.9525018	1.9534436	
	k^2	0.006682	0.002425	0.001570	0.001378	0.001318	0.001289	
	s/b	0.034077	0.072033	0.131409	0.193359	0.255990	0.318858	
22°	$180^\circ - A_1$	$0 0 0$	$53 31 3.3$	$62 24 14.3$	$64 41 45.0$	$65 31 8.4$	$65 50 39.9$	$65 56 48.9$
	$-\delta A_1$	$0 0$	$3 5.1$	$1 19.0$	$0 38.7$	$0 20.4$	$0 9.4$	$0 2.1$
	$180^\circ - A_0$	$0 0 0$	$53 34 8$	$62 25 33$	$64 42 24$	$65 31 29$	$65 50 49$	$65 56 51$
	$\log k$	$-\infty$	1.9055652	1.9476361	1.9562317	1.9591081	1.9602121	1.9605528
	k^2	0.006682	0.002357	0.001432	0.001220	0.001147	0.001119	0.001110
	s/b	0.035775	0.073722	0.133777	0.196608	0.260182	0.324032	0.387973
18°	$180^\circ - A_1$	$0 0 0$	$29 19 11.0$	$46 25 6.2$	$55 5 13.3$	$59 41 41.7$	$62 18 45.2$	$63 52 30.2$
	$-\delta A_1$	$0 0$	$4 2.5$	$4 7.9$	$3 11.5$	$2 21.1$	$1 43.1$	$1 15.0$
	$180^\circ - A_0$	$0 0 0$	$29 23 14$	$46 29 14$	$55 8 25$	$59 44 3$	$62 20 28$	$63 53 45$
	$\log k$	$-\infty$	1.6908226	1.8604705	1.9141068	1.9363608	1.9473002	1.9532744
	k^2	0.006682	0.005073	0.003168	0.002183	0.001697	0.001440	0.001294
	s/b	0.105616	0.124192	0.168012	0.222619	0.281789	0.343184	0.405645
14°	$180^\circ - A_1$	$0 0 0$	$19 22 52.8$	$34 31 14.1$	$44 51 50.8$	$51 41 49.6$	$56 15 27.3$	$59 21 59.3$
	$-\delta A_1$	$0 0$	$3 8.3$	$4 27.9$	$4 27.1$	$3 54.8$	$3 17.9$	$2 43.8$
	$180^\circ - A_0$	$0 0 0$	$19 26 1$	$34 35 42$	$44 56 18$	$51 45 44$	$56 18 45$	$59 24 43$
	$\log k$	$-\infty$	1.5220722	1.7541789	1.8490169	1.8951186	1.9201629	1.9349266
	k^2	0.006682	0.005943	0.004528	0.003348	0.002560	0.002056	0.001730
	s/b	0.157964	0.187488	0.219676	0.264754	0.317253	0.374042	0.433406
10°	$180^\circ - A_1$	$0 0 0$	$14 27 13.6$	$27 2 2.7$	$36 58 9.7$	$44 26 58.7$	$50 0 9.0$	$54 8 7.8$
	$-\delta A_1$	$0 0$	$2 29.4$	$4 5.5$	$4 41.4$	$4 38.9$	$4 18.7$	$3 51.3$
	$180^\circ - A_0$	$0 0 0$	$14 29 43$	$27 6 8$	$37 2 51$	$44 31 38$	$50 4 28$	$54 11 59$
	$\log k$	$-\infty$	1.3984612	1.6585649	1.7799407	1.8458708	1.8847263	1.9090537
	k^2	0.006682	0.006264	0.005295	0.004257	0.003396	0.002752	0.002287
	s/b	0.245270	0.254192	0.279254	0.316635	0.362531	0.414102	0.469464

TABLE XXIV.

ϕ	$L'-L$	4°	8°	12°	16°	20°	24°	$L'-L$	4°	8°	12°	16°	20°	24°
38°	ψ	9 18 48.0	18 24 47.6	27 7 19.7	35 19 17.2	42 57 7.6	50 0 29.7	ψ'	5 18 13.3	10 23 38.1	15 5 35.6	19 16 57.9	22 54 14.4	25 57 2.0
	θ	39 2 46	41 41 34	45 32 6	50 2 52	54 49 40	58 36 57	θ'	24 39 44	26 8 25	28 12 42	30 30 56	32 46 56	34 50 54
	$180^\circ - A$	14 54 57	28 24 7	39 45 84	49 0 53	56 31 21	62 41 9	$180^\circ - A'$	24 35 53.5	26 4 32.5	28 9 8.0	30 27 41.7	32 43 58.1	34 48 9.4
34°	ψ	11 45 31.0	23 0 15.0	33 22 16.3	42 41 47.9	50 58 58.9	58 19 40.8	ψ'	7 44 54.5	14 59 2.1	21 20 26.9	26 39 22.0	30 55 57.1	34 16 3.1
	θ	35 49 35	40 19 53	46 14 27	52 30 24	58 34 31	64 13 30	θ'	25 15 46	28 9 24	31 46 34	35 20 40	38 28 29	41 2 20
	$180^\circ - A$	20 25 3	37 12 26	49 40 9	58 46 54	65 37 29	70 57 58	$180^\circ - A'$	25 11 43.3	28 5 50.3	31 43 29.4	35 17 59.0	38 26 4.7	41 0 8.1
30°	ψ	17 17 39.5	32 42 25.8	45 27 48.4	55 45 11.5	64 6 17.8	71 1 40.0	ψ'	13 17 1.7	24 41 9.8	33 25 54.1	39 42 39.5	44 3 9.4	46 57 53.6
	θ	34 17 29	43 16 20	52 38 4	60 51 33	67 47 47	73 39 40	θ'	27 20 35	33 53 30	40 23 14	45 24 11	49 0 29	51 28 28
	$180^\circ - A$	34 13 16.3	43 13 17.5	52 35 54.0	60 49 57.2	67 46 36.8	73 38 49.5	$180^\circ - A'$	27 16 53.2	33 55 39.9	40 20 59.8	45 22 18.5	48 58 49.2	51 26 56.8
26°	ψ	40 15 21.3	61 39 31.8	73 9 49.7	80 31 41.5	85 55 33.0		ψ'	36 14 41.9	53 38 12.2	61 7 51.0	64 29 4.6	65 52 17.5	
	θ	49 43 33	64 45 39	74 55 13	81 30 7	86 20 26		θ'	42 36 4	57 14 20	63 51 15	68 51 1	68 5 36	
	$180^\circ - A$	46 41 28.7	64 44 35.1	74 54 37.8	81 29 39.5	86 20 19.5		$180^\circ - A'$	42 33 59.6	57 13 6.2	63 50 18.6	68 50 9.7	68 4 50.3	
22°	ψ	33 11 16.7	50 41 8.8	58 45 26.4	62 34 8.8	64 17 12.1	64 51 13.0	ψ'	37 11 57.8	56 42 30.0	70 47 27.9	78 36 50.4	84 20 35.0	88 55 15.0
	θ	39 8 33	54 2 38	61 16 15	64 43 37	68 17 29	66 48 31	θ'	43 21 54	61 42 9	73 31 30	79 37 11	84 50 18	89 0 54
	$180^\circ - A$	39 6 34.5	54 1 38.4	61 15 21.6	64 42 51.2	68 16 46.1	66 47 49.1	$180^\circ - A'$	43 19 53.1	61 41 7.8	73 30 54.2	79 36 49.8	84 50 7.7	89 0 52.3
18°	ψ	10 32 37.6	20 0 48.5	27 48 16.9	33 50 3.8	38 18 48.4	41 32 18.6	ψ'	14 33 19.3	28 2 12.2	39 50 22.2	40 52 51.1	58 22 16.3	65 36 27.9
	θ	20 49 13	26 42 15	33 45 32	37 50 18	41 45 18	44 37 51	θ'	27 56 41	36 19 59	45 30 30	53 58 35	61 24 15	67 51 24
	$180^\circ - A$	20 46 19.6	26 40 3.8	33 43 39.5	37 48 63.7	41 44 7.0	44 36 44.2	$180^\circ - A'$	27 53 5.0	36 17 23.8	45 28 38.0	53 57 11.3	61 23 12.4	67 50 37.4
14°	ψ	5 2 47.9	9 54 8.1	14 24 21.7	18 36 45.7	21 57 51.5	24 56 51.9	ψ'	9 3 30.7	17 55 33.6	26 26 30.3	34 29 36.7	42 1 25.3	49 1 7.3
	θ	14 54 25	17 7 37	20 0 55	23 2 6	25 52 56	28 24 27	θ'	25 40 9	39 43 42	51 11 35	41 12 57	47 18 42	53 14 3
	$180^\circ - A$	14 51 52.8	17 5 23.5	19 59 3.3	23 0 32.1	25 51 33.0	28 23 11.6	$180^\circ - A'$	25 36 10.8	39 40 21.6	51 8 52.9	41 10 46.4	47 16 56.3	53 12 35.0
10°	ψ	2 36 46.0	5 10 38.7	7 38 55.8	9 59 16.4	12 9 46.7	14 9 4.2	ψ'	6 37 29.5	11 12 5.9	19 41 6.8	26 2 11.0	33 13 24.1	38 13 24.9
	θ	10 21 50	11 16 40	12 35 31	14 7 14	15 43 7	17 17 14	θ'	24 57 38	27 15 20	30 45 21	34 54 30	39 27 20	44 11 39
	$180^\circ - A$	10 19 59.4	11 14 55.3	12 33 58.2	14 5 51.2	15 41 52.9	17 16 7.5	$180^\circ - A'$	24 53 31.7	27 14 38.1	30 42 8.1	34 51 45.9	39 25 1.4	44 9 31.7

For the case $L' - L = 0$ it is clear that

$$\psi = \psi' = 0; \quad \theta = \lambda; \quad \theta' = \lambda'; \quad \chi = \pm \phi; \quad \chi' = \pm \phi'$$

8. It may be of interest to find the expression for the azimuthal angle of a vertical plane at the origin which passes through any given point on the earth, so that the difference of this and the geodesic may be studied.

The spheroid may be expressed

$$\frac{x^2 + y^2}{a^2} + \frac{z^2}{b^2} = 1 \quad (43)$$

and P and Q two points on surface

$$P \quad a \cos \phi, 0, b \sin \phi$$

$$Q \quad a \cos \phi' \cos L, a \cos \phi' \sin L, b \sin \phi'$$

Tangent plane at P is

$$\frac{x \cos \phi}{a} + \frac{z \sin \phi}{b} = 1 \quad (44)$$

The vertical plane at P which passes through Q also is

$$lx + my + nz = 1$$

subject to the conditions

$$la \cos \phi + nb \sin \phi = 1$$

since P is in it

$$la \cos \phi' \cos L + ma \cos \phi' \sin L + nb \sin \phi' = 1$$

since Q is in it

$$\frac{l \cos \phi}{a} + \frac{n \sin \phi}{b} = 0$$

since it is perpendicular to (44)

$$\therefore \frac{la \cos \phi}{a^3} = \frac{nb \sin \phi}{-b^3} = \frac{la \cos \phi + nb \sin \phi}{a^3 - b^3} = \frac{1}{a^3 - b^3} \quad \dots \dots \dots (45)$$

Also

$$\begin{aligned} \frac{1}{a^3 - b^3} &= \frac{la}{a^3 \sec \phi} = -\frac{nb}{b^3 \operatorname{cosec} \phi} = \frac{la \cos \phi' \cos L + nb \sin \phi' - 1}{a^3 \sec \phi \cos \phi' \cos L - b^3 \operatorname{cosec} \phi \sin \phi' - a^3 + b^3} \\ &= \frac{ma \cos \phi' \sin L}{a^3 \left(1 - \frac{\cos \phi' \cos L}{\cos \phi}\right) - b^3 \left(1 - \frac{\sin \phi'}{\sin \phi}\right)} \quad \dots \dots \dots (46) \end{aligned}$$

The azimuthal angle of Q from P as determined by the vertical plane through P is the angle this plane makes with ZOX or otherwise it is the angle between the normal to this plane, whose direction cosines are proportional to lmn , and the axis OY .

The angle accordingly is $\cos^{-1} \frac{m}{\sqrt{l^2 + m^2 + n^2}} = \cot^{-1} \frac{m}{\sqrt{l^2 + n^2}} = \psi$ say

$$\therefore \tan \psi = \frac{\sqrt{l^2 + n^2}}{m} \quad \dots \dots \dots (47)$$

Now

$$\frac{l}{a \sec \phi} = \frac{n}{-b \operatorname{cosec} \phi} = \frac{1}{a^3 - b^3} = \frac{\sqrt{l^2 + n^2}}{\sqrt{a^3 \sec^3 \phi + b^3 \operatorname{cosec}^3 \phi}}$$

$$\begin{aligned} \therefore \sqrt{l^2 + n^2} &= \frac{a}{\sin \phi \cos \phi} \sqrt{\sin^2 \phi + (1 - e^2) \cos^2 \phi} \cdot \frac{1}{a^3 e} \\ &= \frac{\sqrt{1 - e^2 \cos^2 \phi}}{ae^3 \sin \phi \cos \phi} \quad \dots \dots \dots (48) \end{aligned}$$

and by (46)

$$\begin{aligned} m &= \frac{1 - \frac{\cos \phi' \cos L}{\cos \phi} - (1 - e^2) \left(1 - \frac{\sin \phi'}{\sin \phi}\right)}{ae^3 \cos \phi' \sin L} \\ &= \left\{ \frac{\sin \phi'}{\sin \phi} - \frac{\cos \phi' \cos L}{\cos \phi} + e^2 \left(1 - \frac{\sin \phi'}{\sin \phi}\right) \right\} / ae^3 \cos \phi' \sin L \end{aligned}$$

$$\begin{aligned} \therefore \tan \psi &= \frac{\sqrt{1 - e^2 \cos^2 \phi}}{ae^3 \sin \phi \cos \phi} \cdot \frac{ae^3 \cos \phi' \sin L}{\frac{\sin \phi'}{\sin \phi} - \frac{\cos \phi' \cos L}{\cos \phi} + e^2 \left(1 - \frac{\sin \phi'}{\sin \phi}\right)} \\ &= \frac{2 \cos \phi' \sin L \sqrt{1 - e^2 \cos^2 \phi}}{\sin 2\phi \left\{ \frac{\sin \phi'}{\sin \phi} - \frac{\cos \phi' \cos L}{\cos \phi} + e^2 \left(1 - \frac{\sin \phi'}{\sin \phi}\right) \right\}} \quad \dots \dots (49) \end{aligned}$$

Substitute for ϕ in terms of λ by means of (4)

$$\tan \phi = \frac{1}{\sqrt{1 - e^2}} \tan \lambda; \quad \frac{\sin \phi}{\sqrt{1 - e^2} \sin \lambda} = \frac{\cos \phi}{\cos \lambda} = \frac{1}{\sqrt{1 - e^2 \sin^2 \lambda}}$$

Hence the value of $\tan \psi$ is

$$\frac{\frac{\cos \lambda' \sin L}{\sqrt{1-e^2 \sin^2 \lambda}} \sqrt{\frac{1-e^2}{1-e^2 \sin^2 \lambda}}}{\frac{(\sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L) \sqrt{1-e^2}}{\sqrt{(1-e^2 \sin^2 \lambda)(1-e^2 \sin^2 \lambda')}} + \frac{e^2 \cos \lambda \sqrt{1-e^2}}{\sqrt{1-e^2 \sin^2 \lambda}} \left(\frac{\sin \lambda}{\sqrt{1-e^2 \sin^2 \lambda}} - \frac{\sin \lambda'}{\sqrt{1-e^2 \sin^2 \lambda'}} \right)}$$

from which it follows that

$$\tan \psi = \cos \lambda' \sin L / \left\{ \sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L + e^2 \cos \lambda \left(\sin \lambda \sqrt{\frac{1-e^2 \sin^2 \lambda'}{1-e^2 \sin^2 \lambda}} - \sin \lambda' \right) \right\}. \quad (50)$$

ψ being the azimuthal angle of ϕ from P .

The case of a sphere is found by putting $e=0$, when (50) becomes

$$\tan \psi_0 = \frac{\cos \lambda' \sin L}{\sin \lambda' \cos \lambda - \cos \lambda' \sin \lambda \cos L} \quad \dots \dots \dots (51)$$

which is the ordinary formula.

$$\text{Let } \psi = \psi_0 + \delta\psi: \text{ then } \cot \overline{\psi_0 + \delta\psi} - \cot \psi_0 = -\frac{\sin \delta\psi}{\sin \psi \sin \psi_0}$$

$$-\delta\psi = e^2 \frac{\sin^2 \psi_0 \cos \lambda}{\sin L \cos \lambda'} \left\{ \sin \lambda \sqrt{\frac{1-e^2 \sin^2 \lambda'}{1-e^2 \sin^2 \lambda}} - \sin \lambda' \right\} \quad \dots \dots \dots (52)$$

This formula gives the correction to be applied to the azimuthal angle, found from the formula for a sphere, to obtain the spheroidal azimuth.

CHAPTER III.

Changes of coordinates of triangulated points, due to changes in axes of the terrestrial spheroid, calculated along geodesics.

1. Consider a geodesic on any surface, and let $A B C$ be three consecutive points on it. Then $A B C$ is the osculating plane at B and from the fundamental property of the geodesic it contains the normal to the surface at B . This shows that measured from B the azimuths of A and C differ by two right angles. It is possible then to describe a geodesic on a surface of unknown form by fulfilling this condition: and, to take a practical case, a traverse along a geodesic can be observed on the earth without knowing its figure if access to a level surface is possible. It follows that if there is a geodesic on one surface which has been selected as representing the earth and it is desired to change to another surface, the geodesic on the first surface will, on transfer to the second surface, remain a geodesic. This property makes it possible to differentiate along a geodesic with respect to the constants of the first surface and so to find relations between changes in these constants and the quantities defining the position of points.

This fact will now be made use of in connection with the equations of Chapter II for the case of a spheroid. Now it has been shown in Chapter I that the effect of slightly changing the latitude and azimuth at the origin may be computed, and that the result is practically independent of the route followed—values of u_x and u_y etc. being identical to nearer than 0.001 of a second: and the resulting changes of latitude, longitude and azimuth for unit changes at the origin are given in tables XVII—XX. These values then are equally applicable to a geodesic and so it is only necessary to consider the effects of changes in a and b .

It is however found convenient not to alter the constant k of the geodesic; and since this is equal to $\sin A \cos \phi$, this is given effect to by not changing the values of A and ϕ at the origin. It is moreover more convenient in dealing with geodesics on a spheroid to make use of the eccentric angle, or reduced latitude ϕ , in place of the latitude λ . Now the relation between λ and ϕ involves e^2 and so if ϕ is unchanged at the origin while e^2 undergoes a change it follows that λ must also change at the origin. In the solution which follows the effect of this origin change of λ occurs: but as its amount can be found from tables XVII, XVIII there is no difficulty in removing it.

2. For convenience of reference equations (26), (25), (27), (2), (40), (42), (41) of Chapter II are repeated.

$$[L] = \pm \left[\pm \psi - \frac{ke^2}{2} \theta \right] \dots \dots \dots (1)$$

$$\psi = \pm \tan^{-1} (\tan A \sin \phi) = \pm \sin^{-1} (\tan A_0 \tan \phi) \dots \dots \dots (2)$$

$$\tan \theta = \frac{1}{k} \sqrt{\frac{1-k^2e^2}{1-e^2}} \tan \psi \dots \dots \dots (3)$$

$$k = \sin A \cos \phi = \sin A_0 \dots \dots \dots (4)$$

$$s = -a \sqrt{1-e^2} \left[\chi \left(1 + \frac{1}{4} k^2 \dots \right) - \frac{k^2}{8} \sin^2 \chi (1 - \dots) + \dots \right] \frac{\chi}{\chi'} \quad (5)$$

$$\tan \chi = \pm \tan \phi \sec A = \pm \frac{1}{k} \tan \psi \dots \dots \dots (6)$$

$$k^2 = \frac{e^2 \cos^2 A_0}{1-e^2} \dots \dots \dots (7)$$

$$\sin \chi = \sin \phi \sec A_0 \dots \dots \dots (8)$$

The signs occurring in (1) and (2) are to be determined as explained in Chapter II. The sign of χ is determined by (8). It is best to consider a and e^2 as independent variables and b as dependent on these. At the end there is no difficulty in passing to the case of a and b considered as independent and e^2 as the dependent variable.

3. Suppose then that a and e^2 are changed, while the azimuth and reduced latitude ϕ of the origin remain unchanged. Values at the origin will be denoted by dashes.

Differentiating (4) and keeping k constant

$$\delta k = 0 = \cos A \cos \phi \delta A - \sin A \sin \phi \delta \phi$$

Changes of *reduced latitude*, longitude and azimuth, in keeping with the notation of Chapter I, will be denoted by u_1, v, w : and the above equation may be written

$$w = \tan A \tan \phi. u_1 \dots \dots \dots (9)$$

Differentiating (1) and remembering that $v' = 0$

$$\pm v = \left[\delta \psi \right] - \frac{ke^2}{2} \left[\delta \theta \right] - \frac{k \delta e^2}{2} \left[\theta \right] \dots \dots \dots (10)$$

By differentiating logarithmically $\tan \psi = \pm \tan A \sin \phi$ which is the same as (2): and by (3) and (6)

$$\frac{\delta \psi}{\sin \psi \cos \psi} = \frac{w}{\sin A \cos A} + u_1 \cot \phi = \frac{\delta \chi}{\sin \chi \cos \chi} = \frac{\delta \theta}{\sin \theta \cos \theta} - (1-k^2) \frac{\delta e^2}{2} \equiv x \dots \dots (11)$$

any of these expressions being denoted by x . This quantity x vanishes at the origin.

Finally differentiate (5) keeping s constant, replacing $\delta \chi$ by means of (11) and $a\sqrt{1-e^2}$ by b

$$\frac{s}{b} \left(-\frac{\delta a}{a} + \frac{\delta e^2}{2(1-e^2)} \right) = -\frac{x}{2} \sin 2\chi \left\{ \left(1 + \frac{k^2}{4} + \dots \right) - \frac{k^2}{4} \cos 2\chi (1 \dots) \dots \right\} - \frac{\delta k^2}{8} \left[2\chi (1 \dots) - \sin 2\chi (\dots) \right]$$

$$\div -\frac{x}{2} \sin 2\chi \left\{ 1 + \frac{k^2}{4} (1 - \cos 2\chi) \right\} - \frac{\delta k^2}{8} [2\chi - \sin 2\chi] \dots \dots \dots (12)$$

where

$$\delta k^2 = \frac{\cos^2 A_0}{(1-e^2)^2} \delta e^2 \dots \dots \dots (13)$$

Equations (12) and (13) serve to determine x in terms of $\frac{\delta a}{a}$ and δe^2 . For the other quantities from (9) and (11) it follows that

$$w = \tan A \tan \phi \cdot u_1 = \frac{x \tan A}{\sec^2 A + \cot^2 \phi} \dots \dots \dots (14)$$

and from (10) and (11)

$$\pm v = \frac{x}{2} \sin 2\psi - \frac{k e^2}{4} \left\{ \left[\sin 2\theta \right] (1 - k^2) \frac{\delta e^2}{2} + x \sin 2\theta \right\} - \frac{k \delta e^2}{2} [\theta] \dots \dots \dots (15)$$

In all the above equations square brackets indicate that the quantity enclosed has to be taken between limits.

4. It remains to give the relation between u and u_1
From (4) of Chapter II

$$\tan \phi = \sqrt{1 - e^2} \tan \lambda$$

Differentiating this logarithmically

$$\frac{u_1}{\sin \phi \cos \phi} = \frac{u}{\sin \lambda \cos \lambda} - \frac{1}{2} \frac{\delta e^2}{1 - e^2}$$

or

$$u = \frac{\sin 2\lambda}{\sin 2\phi} u_1 + \frac{1}{4} \frac{\delta e^2}{1 - e^2} \sin 2\lambda$$

Now

$$\begin{aligned} \frac{\sin 2\lambda}{\sin 2\phi} &= \frac{\sqrt{1 - e^2}}{1 - e^2 \cos^2 \phi} = 1 + \frac{e^2}{2} (2 \cos^2 \phi - 1) \\ &= 1 + \frac{e^2}{2} \cos 2\phi \end{aligned}$$

so that

$$u = \left(1 + \frac{e^2}{2} \cos 2\phi \right) u_1 + \frac{1}{4} \frac{\delta e^2}{1 - e^2} \sin 2\phi \dots \dots \dots (16)$$

At the origin $u_1' = 0$: hence

$$u' = \frac{1}{4} \frac{\delta e^2}{1 - e^2} \sin 2\phi' \dots \dots \dots (17)$$

5. It may be noticed that by this method the changes u, v, w appear to be found without any integration, whereas in Chapter I simultaneous differential equations occurred which had to be solved. The case under consideration is a particular case of the general equations (2) of Chapter I. The decision to follow a geodesic introduces a relation by which these equations can be reduced to total differential equations: and the integration of these equations would lead to the same results as may be obtained from equations (12) to (17). The same results will be seen to be obtainable by application to values of u_x, v_x, w_x of the appropriate closing errors. For the case now under consideration the equations formed in Chapter II give the results of integration: and so no further integration is necessary.

6. In making use of the equations (12) to (17) two cases are considered in which

(i) $\delta a = 1 \text{ km.}$ and $\delta e^2 = 0$

(ii) $\delta a = 0$ and $\delta e^2 = .0001$

The first of these corresponds to a combination of cases I and II of Chapter I, while the second corresponds to a combination of cases II and III. This arrangement simplifies computation and there is no difficulty in deriving cases I and II when the computations are complete.

As no azimuthal change is being made at the origin it is clear that there is symmetry about a central meridian. In Chapter II values of $\psi, \theta, \chi, A, \frac{\delta}{b}, h^2, k$ (*vide* tables XXIII, XXIV) have already been given for every 4° of ϕ from 10° to 38° and for longitude differences of 4° from 4° to 24° . With the help of these the values of u, v, w exhibited in the following two tables have been found. A double sign is prefixed to v and w and of these the upper or lower is to be taken according as the point is west or east of the origin. The results are given to three places of decimals as found by the computations: but the last figure is liable to error, which is not sufficiently large to be practically important for the present purpose.

TABLE XXV.

$$\delta a = 1 \text{ km.}, (\delta e^2 = 0) \quad \delta b = \frac{\delta}{a} = .9967 \text{ km.}$$

ϕ	$L-L'$	4	8	12	16	20	24
38	u_1	- 7.834	- 7.710	- 7.510	- 7.227	- 6.863	- 6.415
	u	- 7.840	- 7.716	- 7.516	- 7.233	- 6.868	- 6.420
	$\pm v$	+ 2.644	+ 5.280	+ 7.914	+ 10.532	+ 13.143	+ 15.733
	$\pm w$	+ 1.630	+ 3.258	+ 4.882	+ 6.499	+ 8.106	+ 9.703
34	u_1	- 5.577	- 5.468	- 5.281	- 5.019	- 4.683	- 4.270
	u	- 5.584	- 5.475	- 5.288	- 5.025	- 4.689	- 4.275
	$\pm v$	+ 2.498	+ 4.996	+ 7.489	+ 9.967	+ 12.437	+ 14.894
	$\pm w$	+ 1.400	+ 2.799	+ 4.195	+ 5.586	+ 6.971	+ 8.348
30	u_1	- 3.325	- 3.222	- 3.052	- 2.812	- 2.502	- 2.123
	u	- 3.331	- 3.227	- 3.057	- 2.817	- 2.506	- 2.127
	$\pm v$	+ 2.384	+ 4.768	+ 7.143	+ 9.516	+ 11.871	+ 14.228
	$\pm w$	+ 1.195	+ 2.390	+ 3.581	+ 4.769	+ 5.951	+ 7.128
26	u_1	- 1.071	- 0.977	- 0.820	- 0.603	- 0.322	+ 0.024
	u	- 1.073	- 0.979	- 0.822	- 0.604	- 0.323	+ 0.024
	$\pm v$	+ 2.294	+ 4.586	+ 6.873	+ 9.152	+ 11.435	+ 13.718
	$\pm w$	+ 1.009	+ 2.016	+ 3.019	+ 4.024	+ 5.025	+ 6.024
22	u_1	+ 1.185	+ 1.268	+ 1.409	+ 1.606	+ 1.860	+ 2.172
	u	+ 1.188	+ 1.271	+ 1.412	+ 1.610	+ 1.864	+ 2.177
	$\pm v$	+ 2.225	+ 4.447	+ 6.666	+ 8.881	+ 11.088	+ 13.291
	$\pm w$	+ 0.836	+ 1.671	+ 2.504	+ 3.338	+ 4.164	+ 4.991
18	u_1	+ 3.439	+ 3.513	+ 3.639	+ 3.814	+ 4.044	+ 4.319
	u	+ 3.448	+ 3.522	+ 3.649	+ 3.824	+ 4.055	+ 4.331
	$\pm v$	+ 2.172	+ 4.343	+ 6.512	+ 8.674	+ 10.834	+ 12.981
	$\pm w$	+ 0.673	+ 1.346	+ 2.018	+ 2.688	+ 3.360	+ 4.024
14	u_1	+ 5.695	+ 5.760	+ 5.869	+ 6.024	+ 6.223	+ 6.467
	u	+ 5.712	+ 5.777	+ 5.886	+ 6.042	+ 6.241	+ 6.486
	$\pm v$	+ 2.135	+ 4.271	+ 6.401	+ 8.536	+ 10.660	+ 12.777
	$\pm w$	+ 0.519	+ 1.037	+ 1.554	+ 2.070	+ 2.587	+ 3.100
10	u_1	+ 7.952	+ 8.008	+ 8.102	+ 8.232	+ 8.404	+ 8.616
	u	+ 7.977	+ 8.033	+ 8.127	+ 8.258	+ 8.430	+ 8.643
	$\pm v$	+ 2.114	+ 4.229	+ 6.343	+ 8.451	+ 10.559	+ 12.662
	$\pm w$	+ 0.369	+ 0.737	+ 1.105	+ 1.472	+ 1.840	+ 2.206

TABLE XXVI.

$$\delta a = 0; (\delta e^2 = 0.0001); \delta b = -\frac{a^2 \delta e^2}{2b} = -0.3200 \text{ km.}; u_0 = 3''.872.$$

ϕ	$L-L'$	4°	8°	12°	16°	20°	24°
38°	u_1	+ 1.837	+ 1.807	+ 1.757	+ 1.689	+ 1.600	+ 1.498
	u	+ 6.881	+ 6.850	+ 6.800	+ 6.738	+ 6.643	+ 6.536
	$\pm v$	- 0.094	- 0.187	- 0.279	- 0.370	- 0.455	- 0.540
	$\pm w$	- 0.882	- 0.763	- 1.142	- 1.519	- 1.890	- 2.258
34°	u_1	+ 1.364	+ 1.336	+ 1.289	+ 1.223	+ 1.139	+ 1.037
	u	+ 6.185	+ 6.157	+ 6.111	+ 6.045	+ 5.960	+ 5.857
	$\pm v$	- 0.060	- 0.121	- 0.182	- 0.240	- 0.295	- 0.346
	$\pm w$	- 0.342	- 0.684	- 1.024	- 1.362	- 1.696	- 2.027
30°	u_1	+ 0.845	+ 0.818	+ 0.775	+ 0.713	+ 0.633	+ 0.536
	u	+ 5.360	+ 5.324	+ 5.280	+ 5.218	+ 5.138	+ 5.041
	$\pm v$	- 0.033	- 0.067	- 0.100	- 0.131	- 0.160	- 0.186
	$\pm w$	- 0.304	- 0.607	- 0.909	- 1.208	- 1.506	- 1.800
26°	u_1	+ 0.232	+ 0.257	+ 0.216	+ 0.158	+ 0.084	- 0.007
	u	+ 4.331	+ 4.357	+ 4.315	+ 4.258	+ 4.184	+ 4.096
	$\pm v$	- 0.011	- 0.023	- 0.033	- 0.041	- 0.049	- 0.056
	$\pm w$	- 0.265	- 0.530	- 0.794	- 1.057	- 1.317	- 1.579
22°	u_1	- 0.321	- 0.344	- 0.382	- 0.435	- 0.504	- 0.587
	u	+ 3.294	+ 3.271	+ 3.233	+ 3.180	+ 3.111	+ 3.027
	$\pm v$	+ 0.008	+ 0.015	+ 0.023	+ 0.032	+ 0.043	+ 0.055
	$\pm w$	- 0.227	- 0.454	- 0.679	- 0.904	- 1.127	- 1.349
18°	u_1	- 0.960	- 0.980	- 1.015	- 1.063	- 1.126	- 1.201
	u	+ 2.098	+ 2.078	+ 2.043	+ 1.995	+ 1.932	+ 1.856
	$\pm v$	+ 0.024	+ 0.046	+ 0.068	+ 0.092	+ 0.116	+ 0.143
	$\pm w$	- 0.188	- 0.375	- 0.563	- 0.749	- 0.935	- 1.119
14°	u_1	- 1.628	- 1.646	- 1.677	- 1.720	- 1.775	- 1.844
	u	+ 0.812	+ 0.794	+ 0.762	+ 0.719	+ 0.664	+ 0.596
	$\pm v$	+ 0.036	+ 0.069	+ 0.102	+ 0.134	+ 0.169	+ 0.205
	$\pm w$	- 0.148	- 0.296	- 0.444	- 0.591	- 0.738	- 0.884
10°	u_1	- 2.322	- 2.337	- 2.364	- 2.401	- 2.450	- 2.508
	u	- 0.547	- 0.562	- 0.590	- 0.637	- 0.676	- 0.734
	$\pm v$	+ 0.042	+ 0.083	+ 0.122	+ 0.162	+ 0.202	+ 0.244
	$\pm w$	- 0.108	- 0.215	- 0.322	- 0.429	- 0.536	- 0.642

By the help of tables XVII and XVIII of Chapter I and their extensions (*vide* tables XXXV, XXXVI below) the effect of u_0 is eliminated: the figures in these tables give by interpolation values of u for each point in table XXVI, and these multiplied by -3.872 are applied to the corresponding figures in XXVI, leaving residuals due to case $\delta b = -0.3200$. The residuals are multiplied by $-\frac{1}{32}$ and the result is the case II,— $\delta b = 1, \delta a = 0$. This case is then multiplied by -0.9967 and applied to table XXV, leaving as a residual case I,— $\delta a = 1, \delta b = 0$. From these tables values corresponding to even degrees of λ are interpolated and the results are exhibited in tables XXVII, XXVIII.

Case I.— $\delta a = 1$ km.Values of u_j
in seconds.

TABLE XXIX.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.
Lat.																							Lat.
40°																							40°
39°																							39°
38°																							38°
37°																							37°
36°																							36°
35°																							35°
34°																							34°
33°																							33°
32°																							32°
31°																							31°
30°																							30°
29°																							29°
28°																							28°
27°																							27°
26°																							26°
25°																							25°
24°																							24°
23°																							23°
22°																							22°
21°																							21°
20°																							20°
19°																							19°
18°																							18°
17°																							17°
16°																							16°
15°																							15°
14°																							14°
13°																							13°
12°																							12°
11°																							11°
10°																							10°
9°																							9°
8°																							8°
7°																							7°
6°																							6°
5°																							5°
4°																							4°
3°																							3°
2°																							2°
1°																							1°
0°																							0°

The sign is + above the horizontal
dividing line and - below it.The sign is + above the horizontal
dividing line and - below it.

Values of v_g
in seconds.

TABLE XXX.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.	
Lat.	P o s i t i v e																			N e g a t i v e		Lat.		
40°													4.348	3.581	2.815	2.048	1.280						40°	
39													4.298	3.530	2.781	2.022	1.282						39	
38													4.246	3.498	2.748	1.997	1.245						38	
37													4.201	3.458	2.716	1.973	1.230						37	
36								7.083	6.352	5.620	4.887	4.154	3.420	2.685	1.951	1.216	0.481	0.254	0.080	1.723	2.458		36	
35								7.007	6.284	5.560	4.834	4.100	3.383	2.656	1.930	1.203	0.476	0.251	0.078	1.704	2.431		35	
34								6.933	6.218	5.502	4.785	4.067	3.348	2.629	1.911	1.192	0.471	0.248	0.068	1.687	2.406		34	
33								6.863	6.155	5.446	4.736	4.025	3.314	2.602	1.891	1.179	0.466	0.246	0.058	1.670	2.382		33	
32								6.794	6.093	5.391	4.688	3.985	3.281	2.578	1.873	1.167	0.462	0.243	0.040	1.654	2.359		32	
31	12.384 11.267	11.688 11.577	10.992 10.887	10.293 10.194	9.595 9.502	8.895 8.809	8.195 8.116	7.494 7.422	6.728 6.034	5.339	4.643	3.948	3.251	2.554	1.855	1.157	0.458	0.241	0.040	1.639	2.337		31	
30	12.154	11.471	10.787	10.100	9.414	8.727	8.040	7.353	6.665	5.877	5.288	4.600	3.911	3.221	2.530	1.838	1.146	0.453	0.239	0.031	1.624	2.316	30	
29	12.044	11.366	10.689	10.007	9.327	8.646	7.965	7.284	6.603	5.821	5.239	4.557	3.874	3.190	2.506	1.821	1.135	0.449	0.237	0.023	1.608	2.294	29	
28	11.938	11.267	10.594	9.919	9.244	8.568	7.893	7.218	6.544	5.868	5.192	4.515	3.838	3.160	2.483	1.804	1.124	0.445	0.234	0.014	1.593	2.272	28	
27	11.837	11.170	10.503	9.833	9.164	8.494	7.824	7.156	6.487	5.817	5.147	4.476	3.806	3.133	2.461	1.788	1.114	0.441	0.232	0.006	1.579	2.252	27	
26	11.740	11.078	10.416	9.751	9.088	8.424	7.750	7.096	6.433	5.769	5.104	4.439	3.773	3.106	2.440	1.773	1.105	0.438	0.230	0.008	1.566	2.238	26	
25	11.646	10.990	10.332	9.674	9.016	8.357	7.696	7.030	6.381	5.722	5.062	4.403	3.743	3.081	2.420	1.758	1.096	0.434	0.228	0.001	1.553	2.215	25	
24								6.985	6.331	5.677	5.022	4.367	3.713	3.058	2.402	1.745	1.088	0.430	0.227	0.894	1.542	2.190	24	
23								6.933	6.284	5.634	4.984	4.335	3.686	3.036	2.385	1.733	1.081	0.427	0.225	0.878	1.531	2.183	23	
22								6.883	6.238	5.593	4.948	4.303	3.650	3.014	2.368	1.721	1.072	0.424	0.224	0.872	1.520	2.168	22	
21								6.834	6.194	5.554	4.913	4.274	3.634	2.994	2.352	1.700	1.066	0.422	0.222	0.866	1.510	2.153	21	
20															3.010	2.674	2.337	1.698	1.058	0.419	0.220	0.860	2.138	20
19															3.586	2.954	2.321	1.686	1.052	0.416	0.219	0.854	2.124	19
18															3.562	2.934	2.305	1.675	1.044	0.413	0.218	0.849	2.110	18
17															3.539	2.915	2.290	1.664	1.038	0.411	0.216	0.843	2.096	17
16															3.516	2.895	2.274	1.652	1.030	0.408	0.215	0.837	2.082	16
15															3.493	2.875	2.258	1.641	1.023	0.405	0.213	0.831	2.067	15
14															3.470	2.856	2.243	1.630	1.016	0.402	0.212	0.826	2.053	14
13															3.448	2.838	2.228	1.619	1.010	0.400	0.210	0.820	2.039	13
12															3.426	2.820	2.214	1.608	1.003	0.397	0.209	0.815	2.026	12
11															3.404	2.802	2.200	1.598	0.996	0.394	0.208	0.810	2.013	11
10															3.382	2.784	2.186	1.588	0.990	0.392	0.206	0.804	2.000	10
9															3.360	2.766	2.172	1.578	0.984	0.390	0.205	0.799	1.989	9
8															3.338	2.748	2.158	1.568	0.978	0.388	0.203	0.794	1.984	8
Long.	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°	Long.	
Lat.	N e g a t i v e																					Lat.		
30°	2.316	3.007	3.697	4.387	5.076	5.764	6.452	7.140	7.827	8.514	9.201	0.888	10.574	11.260	11.943	12.625	13.306	13.988	14.660	15.340	16.020	16.700	30°	
29	2.294	2.978	3.662	4.345	5.028	5.710	6.392	7.073	7.754	8.435	9.116	0.797	10.477	11.157	11.834	12.511	13.187	13.863	14.538	15.214	15.888	16.562	29	
28	2.272	2.950	3.628	4.308	4.983	5.659	6.334	7.010	7.684	8.359	9.034	0.710	10.385	11.050	11.731	12.403	13.073	13.744	14.413	15.083	15.753	16.421	28	
27	2.252	2.925	3.597	4.268	4.939	5.610	6.279	6.940	7.617	8.286	8.956	0.626	10.295	10.964	11.630	12.297	12.963	13.628	14.292	14.957	15.621	16.285	27	
26	2.233	2.900	3.566	4.232	4.898	5.563	6.227	6.891	7.554	8.217	8.891	0.546	10.210	10.873	11.534	12.198	12.857	13.518	14.177	14.836	15.495	16.153	26	
25	2.215	2.876	3.537	4.198	4.865	5.517	6.176	6.835	7.494	8.152	8.811	0.470	10.128	10.788	11.442	12.098	12.753	13.409	14.064	14.719	15.373	16.028	25	
24	2.199	2.855	3.510	4.165	4.819	5.474	6.129	6.783	7.437	8.091	8.746	0.400	10.054	10.705	11.357	12.008	12.659	13.308	13.956	14.603	15.251	15.900	24	
23	2.183	2.834	3.484	4.134	4.783	5.432	6.082	6.732	7.381	8.031	8.670	0.327	9.977	10.625	11.272	11.919	12.563	13.207	13.850	14.494	15.137	15.780	23	
22	2.168	2.814	3.459	4.104	4.748	5.393	6.038	6.684	7.329	7.973	8.617	0.256	9.900	10.545	11.190	11.831	12.472	13.110	13.747	14.385	15.024	15.663	22	
21	2.153	2.795	3.435	4.075	4.715	5.355	5.996	6.637			8.554	0.186	9.827	10.468	11.107	11.744	12.380	13.016	13.649	14.282	14.916	15.550	21	
20	2.138	2.776	3.412	4.048	4.683	5.319	5.955	6.591			8.495	0.117	0.754	10.390	11.024	11.658	12.291	12.923	13.551	14.181	14.812	15.443	20	
19	2.124	2.758	3.390	4.021	4.653						0.681			10.315	10.946	11.574	12.204	12.829	13.454	14.082	14.708	15.336	19	
18	2.110	2.740	3.368	3.995	4.623						0.600			10.237	10.866	11.493	12.116	12.738	13.361	13.983	14.602	15.227	18	
17	2.096	2.721	3.345	3.969	4.593						0.540			10.157	10.788	11.408	12.020	12.647	13.268	13.886	14.503	15.123	17	
16	2.082	2.703	3.323	3.944	4.563									0.471	10.091	10.708	11.327	11.945	12.569	13.176	13.791	14.404	15.018	16
15															11.860	12.471	13.085	13.698	14.308	14.917			15	
14															11.775	12.384	12.993	13.604	14.200	14.815			14	
13																							13	
12																							12	
11																							11	
10																							10	
9																							9	
8																							8	

Values of w_g
in seconds.

TABLE XX XI.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.	
Lat.	P o s i t i v e																			N e g a t i v e			Lat.	
40°														2-261 1-854	1-457 1-080 0-683								40°	
39°														2-187 1-801	1-415 1-029 0-648								39°	
38°														2-125 1-750	1-374 0-909 0-623								38°	
37°														2-065 1-700	1-835 0-970 0-606								37°	
36°																							36°	
35°																							35°	
34°																							34°	
33°																							33°	
32°																							32°	
31°																							31°	
30°	5-587	5-277 4-965 4-658	4-338 4-023 3-708	3-393 3-078 2-762	3-157 2-833	2-807	2-480 2-165 1-854	2-138 1-823 1-516	1-808 0-870 0-543	0-233	0-123 0-478	0-883 1-188											30°	
29°	5-453	5-151 4-847 4-541	4-233 3-926 3-618	3-811 3-504 3-196	3-485 3-179	2-857	2-540 2-225 1-910	2-202 1-887 1-572	1-958 0-774 0-482	0-227	0-119 0-465	0-811 1-156											29°	
28°																								28°
27°																								27°
26°																								26°
25°																								25°
24°																								24°
23°																								23°
22°																								22°
21°																								21°
20°																								20°
19°																								19°
18°																								18°
17°																								17°
16°																								16°
15°																								15°
14°																								14°
13°																								13°
12°																								12°
11°																								11°
10°																								10°
9°																								9°
8°																								8°
Long.	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°	Long.	
Lat.	N e g a t i v e																					Lat.		
30°	1-019 1-823 1-628	1-933 2-237 2-540	2-841 3-141 3-441	3-741	4-042 4-342 4-641	4-938 5-234 5-528	5-822 6-118 6-404	6-692	6-980 7-267														30°	
29°	0-996 1-294 1-592	1-899 2-195 2-481	2-775 3-068 3-361	3-654	3-948 4-241 4-533	4-824 5-113 5-401	5-687 5-972 6-255	6-538	6-820 7-100														29°	
28°	0-974 1-268 1-556	1-847 2-137 2-426	2-713 3-000 3-286	3-573	3-860 4-146 4-431	4-716 5-009 5-281	5-861 5-840 6-117	6-394	6-666 6-943														28°	
27°	0-954 1-239 1-524	1-908 2-092 2-374	2-655 2-936 3-216	3-497	3-776 4-055 4-334	4-613 4-890 5-166	5-441 5-714 5-986	6-257	6-526 6-794														27°	
26°	0-935 1-215 1-493	1-771 2-048 2-324	2-600 2-875 3-150	3-425	3-698 3-971 4-244	4-515 4-786 5-057	5-327 5-595 5-862	6-127	6-390 6-653														26°	
25°	0-917 1-191 1-463	1-735 2-006 2-277	2-548 2-818 3-088	3-356	3-624 3-892 4-159	4-425 4-691 4-956	5-220 5-482 5-742	6-001	6-260 6-518														25°	
24°	0-900 1-168 1-435	1-702 1-968 2-234	2-500 2-765 3-020	3-283	3-555 3-821 4-084	4-342 4-598 4-854	5-109 5-361 5-612	5-863	6-112 6-369														24°	
23°	0-884 1-147 1-409	1-671 1-932 2-193	2-454 2-714 2-974	3-233	3-492 3-754 4-047	4-303 4-556 4-807	5-057 5-303 5-547	5-791	6-034 6-277														23°	
22°	0-868 1-126 1-384	1-641 1-898 2-154	2-410 2-667 2-922	3-177	3-431 3-763 4-015	4-266 4-514 4-760	5-005 5-246 5-486	5-726	5-964 6-202														22°	
21°	0-853 1-106 1-359	1-612 1-865 2-118	2-370 2-622		3-373 3-739 3-985	4-230 4-472 4-715	4-956 5-193 5-430	5-665	5-900 6-135														21°	
20°	0-838 1-088 1-337	1-587 1-836 2-085	2-333 2-580		3-318 3-712 3-954	4-197 4-496 4-774	4-909 5-142 5-376	5-610	5-840 6-071														20°	
19°	0-825 1-071 1-317	1-563 1-809																						19°
18°	0-813 1-056 1-299	1-541 1-783																						18°
17°	0-802 1-042 1-281	1-520 1-758																						17°
16°	0-792 1-029 1-265	1-500 1-735																						16°
15°																								15°
14°																								14°
13°																								13°
12°																								12°
11°																								11°
10°																								10°
9°																								9°
8°																								8°

Case II.— $\delta b = 1$ km.Values of u_g

TABLE XXXII.

in seconds.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.	
Lat.																							Lat.	
40°													10.294	10.295	10.297	10.298	10.297							40°
39													9.855	9.856	9.858	9.859	9.858							39
38													9.884	9.885	9.886	9.886	9.885							38
37													8.882	8.883	8.883	8.883	8.882							37
36								8.346	8.347	8.348	8.348	8.349	8.350	8.350	8.351	8.351	8.351	8.350	8.350	8.350	8.351	8.351	8.351	36
35								7.789	7.790	7.790	7.790	7.791	7.791	7.791	7.792	7.792	7.792	7.792	7.792	7.792	7.792	7.791	7.791	35
34								7.204	7.204	7.204	7.204	7.204	7.205	7.205	7.205	7.204	7.204	7.205	7.205	7.205	7.204	7.204	7.204	34
33								6.593	6.593	6.592	6.592	6.592	6.591	6.591	6.590	6.590	6.590	6.590	6.590	6.590	6.590	6.590	6.590	33
32								5.957	5.956	5.955	5.954	5.953	5.952	5.951	5.950	5.950	5.950	5.949	5.949	5.949	5.950	5.950	5.950	32
31								5.292	5.291	5.290	5.288	5.287	5.286	5.285	5.284	5.283	5.282	5.281	5.281	5.281	5.282	5.283	5.283	31
30								4.610	4.610	4.611	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	4.608	30
29								3.908	3.904	3.901	3.897	3.893	3.890	3.887	3.884	3.882	3.880	3.878	3.876	3.874	3.872	3.869	3.870	29
28								3.170	3.168	3.161	3.156	3.152	3.148	3.144	3.139	3.136	3.133	3.129	3.125	3.121	3.117	3.113	3.109	28
27								2.400	2.400	2.398	2.392	2.386	2.381	2.377	2.373	2.369	2.366	2.362	2.358	2.354	2.350	2.346	2.342	27
26								1.623	1.617	1.611	1.604	1.597	1.591	1.585	1.581	1.577	1.574	1.571	1.567	1.563	1.559	1.555	1.551	26
25								0.815	0.808	0.800	0.793	0.786	0.779	0.772	0.767	0.763	0.759	0.755	0.751	0.747	0.743	0.740	0.736	25
24																								24
23																								23
22																								22
21																								21
20																								20
19																								19
18																								18
17																								17
16																								16
15																								15
14																								14
13																								13
12																								12
11																								11
10																								10
9																								9
8																								8
7																								7
6																								6
5																								5
4																								4
3																								3
2																								2
1																								1
0																								0

The sign is - above the horizontal dividing line and + below it.

Values of v_g
in seconds.

TABLE XXXIII.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.
Lat.	N e g a t i v e																						Lat.
40°																							40°
39°																							39°
38°																							38°
37°																							37°
36°																							36°
35°																							35°
34°																							34°
33°																							33°
32°																							32°
31°																							31°
30°																							30°
29°																							29°
28°																							28°
27°																							27°
26°																							26°
25°																							25°
24°																							24°
23°																							23°
22°																							22°
21°																							21°
20°																							20°
19°																							19°
18°																							18°
17°																							17°
16°																							16°
15°																							15°
14°																							14°
13°																							13°
12°																							12°
11°																							11°
10°																							10°
9°																							9°
8°																							8°
Long.	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°	Long.
Lat.	P o s i t i v e																						Lat.
30°																							30°
29°																							29°
28°																							28°
27°																							27°
26°																							26°
25°																							25°
24°																							24°
23°																							23°
22°																							22°
21°																							21°
20°																							20°
19°																							19°
18°																							18°
17°																							17°
16°																							16°
15°																							15°
14°																							14°
13°																							13°
12°																							12°
11°																							11°
10°																							10°
9°																							9°
8°																							8°

Case II.— $\delta b = 1$ km.Values of w_p
in seconds.

TABLE XXXIV.

Long.	60°	61°	62°	63°	64°	65°	66°	67°	68°	69°	70°	71°	72°	73°	74°	75°	76°	77°	78°	79°	80°	81°	Long.
Lat.	P o s i t i v e																			N e g a t i v e			Lat.
40°													0.216	0.177	0.139	0.102	0.066						40°
39													0.194	0.159	0.125	0.091	0.058						39
38													0.171	0.140	0.110	0.080	0.050						38
37													0.147	0.121	0.094	0.068	0.043						37
36								0.214	0.190	0.167	0.144	0.122	0.100	0.078	0.056	0.035	0.014	0.007	0.028	0.049	0.072		36
35								0.171	0.151	0.133	0.115	0.097	0.079	0.062	0.044	0.027	0.010	0.006	0.032	0.039	0.056		35
34								0.127	0.112	0.097	0.084	0.071	0.058	0.046	0.032	0.020	0.008	0.004	0.016	0.028	0.040		34
33								0.081	0.071	0.061	0.052	0.044	0.036	0.028	0.020	0.012	0.005	0.003	0.010	0.018	0.026		33
32								0.036	0.030	0.025	0.021	0.017	0.014	0.011	0.007	0.004	0.001	0.001	0.003	0.006	0.009		32
31	0.091	0.082	0.074	0.067	0.061	0.055	0.049	0.042	0.036	0.030	0.025	0.021	0.017	0.014	0.011	0.007	0.004	0.001	0.001	0.003	0.006	0.009	31
30	0.084	0.084	0.082	0.080	0.076	0.071	0.067	0.064	0.061	0.057	0.052	0.046	0.040	0.032	0.025	0.019	0.013	0.005	0.003	0.010	0.017	0.023	30
29	0.173	0.167	0.160	0.152	0.144	0.135	0.126	0.118	0.110	0.101	0.091	0.080	0.069	0.057	0.044	0.033	0.021	0.009	0.004	0.017	0.029	0.041	29
28	0.263	0.252	0.241	0.229	0.216	0.201	0.187	0.173	0.159	0.145	0.130	0.114	0.098	0.081	0.064	0.047	0.030	0.012	0.006	0.024	0.042	0.059	28
27	0.354	0.338	0.321	0.305	0.287	0.268	0.249	0.230	0.211	0.191	0.170	0.150	0.128	0.106	0.084	0.062	0.039	0.016	0.008	0.032	0.055	0.078	27
26	0.448	0.426	0.404	0.382	0.360	0.337	0.313	0.288	0.263	0.237	0.211	0.185	0.159	0.132	0.104	0.076	0.048	0.019	0.010	0.039	0.068	0.096	26
25	0.546	0.519	0.491	0.463	0.435	0.407	0.378	0.347	0.316	0.285	0.253	0.222	0.190	0.158	0.126	0.092	0.058	0.023	0.012	0.047	0.082	0.115	25
24								0.408	0.371	0.334	0.296	0.259	0.222	0.184	0.146	0.107	0.067	0.027	0.014	0.055	0.094	0.133	24
23								0.469	0.426	0.383	0.341	0.298	0.255	0.212	0.168	0.123	0.077	0.031	0.016	0.063	0.109	0.153	23
22								0.530	0.483	0.433	0.384	0.336	0.288	0.239	0.188	0.137	0.086	0.034	0.018	0.070	0.122	0.173	22
21								0.594	0.540	0.485	0.430	0.376	0.322	0.267	0.210	0.153	0.096	0.038	0.020	0.078	0.136	0.193	21
20													0.356	0.294	0.232	0.169	0.106	0.042	0.022	0.086	0.150	0.212	20
19													0.391	0.322	0.254	0.185	0.116	0.046	0.024	0.094	0.164	0.232	19
18													0.426	0.351	0.276	0.201	0.126	0.050	0.026	0.102	0.178	0.253	18
17													0.462	0.381	0.300	0.219	0.137	0.064	0.029	0.111	0.193	0.275	17
16													0.500	0.412	0.325	0.237	0.148	0.069	0.031	0.121	0.209	0.297	16
15													0.539	0.444	0.350	0.255	0.159	0.063	0.033	0.129	0.225	0.320	15
14													0.578	0.477	0.375	0.273	0.170	0.067	0.035	0.138	0.241	0.343	14
13													0.618	0.510	0.400	0.291	0.182	0.072	0.038	0.148	0.257	0.366	13
12													0.658	0.543	0.426	0.310	0.194	0.077	0.040	0.157	0.274	0.390	12
11													0.698	0.576	0.453	0.320	0.206	0.083	0.043	0.167	0.291	0.414	11
10													0.739	0.609	0.479	0.340	0.218	0.087	0.046	0.177	0.308	0.436	10
9													0.780	0.642	0.505	0.360	0.230	0.092	0.049	0.187	0.325	0.464	9
8													0.821	0.675	0.531	0.389	0.242	0.097	0.052	0.197	0.342	0.489	8

The signs below the horizontal dividing line are opposite to those above it.

Long.	81°	82°	83°	84°	85°	86°	87°	88°	89°	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°	Long.
Lat.	P o s i t i v e																					Lat.	
30°	0.023	0.030	0.037	0.044	0.050	0.055	0.059	0.063	0.066	0.070	0.074	0.078	0.081	0.083	0.084	0.084	0.083	0.081	0.079	0.074	0.069	0.063	30°
29	0.041	0.052	0.065	0.077	0.088	0.098	0.107	0.116	0.124	0.132	0.141	0.150	0.158	0.165	0.171	0.176	0.180	0.184	0.186	0.187	0.186	0.185	29
28	0.059	0.076	0.093	0.110	0.125	0.140	0.155	0.169	0.183	0.197	0.211	0.224	0.237	0.249	0.260	0.270	0.279	0.288	0.295	0.301	0.305	0.309	28
27	0.078	0.100	0.122	0.143	0.164	0.184	0.205	0.225	0.244	0.263	0.281	0.299	0.317	0.334	0.350	0.365	0.380	0.394	0.406	0.417	0.427	0.436	27
26	0.096	0.123	0.150	0.177	0.203	0.229	0.255	0.281	0.305	0.330	0.353	0.376	0.398	0.420	0.442	0.463	0.483	0.502	0.520	0.536	0.551	0.565	26
25	0.115	0.148	0.180	0.212	0.244	0.275	0.307	0.338	0.368	0.398	0.427	0.455	0.482	0.510	0.537	0.564	0.590	0.614	0.636	0.657	0.677	0.697	25
24	0.133	0.172	0.210	0.248	0.285	0.322	0.360	0.396	0.432	0.468	0.502	0.528	0.568	0.605	0.643	0.680	0.717	0.753	0.788	0.823	0.858	0.891	24
23	0.153	0.198	0.242	0.285	0.327	0.370	0.413	0.457	0.499	0.539	0.579	0.527	0.572	0.616	0.659	0.703	0.745	0.786	0.826	0.866	0.907	0.947	23
22	0.173	0.223	0.273	0.321	0.369	0.418	0.467	0.515	0.563	0.611	0.658	0.524	0.575	0.625	0.675	0.724	0.772	0.818	0.864	0.911	0.957	1.003	22
21	0.193	0.249	0.305	0.359	0.413	0.468	0.523	0.577			0.737	0.523	0.580	0.636	0.691	0.747	0.801	0.853	0.905	0.957	1.010	1.062	21
20	0.212	0.274	0.336	0.396	0.459	0.520	0.581	0.642			0.818	0.522	0.585	0.648	0.708	0.769	0.828	0.888	0.947	1.006	1.065	1.124	20
19	0.232	0.300	0.369	0.438	0.505						0.590	0.659	0.725	0.791	0.858	0.925	0.991	1.056	1.121	1.185			19
18	0.253	0.328	0.403	0.478	0.552						0.595	0.670	0.743	0.816	0.890	0.963	1.035	1.107	1.177	1.248			18
17	0.275	0.356	0.437	0.518	0.600						0.601	0.682	0.762	0.842	0.922	1.001	1.080	1.158	1.235	1.311			17
16	0.297	0.385	0.472	0.560	0.648						0.609	0.695	0.782	0.868	0.954	1.040	1.126	1.210	1.293	1.377			16
15															0.989	1.081	1.172	1.264	1.354	1.444			15
14															1.024	1.123	1.221	1.319	1.416	1.513			14
13															1.050	1.165	1.270	1.375	1.479	1.592			13
12															1.095	1.208	1.320	1.431	1.542	1.653			12
11															1.132	1.252	1.371	1.490	1.608	1.726			11
10															1.169	1.297	1.423	1.549	1.674	1.798			10
9															1.207	1.342	1.476	1.610					

Case III.— $u_0=1''$.Values of u, v, w
in seconds.

TABLE XXXV*.

Lat.	Long.	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°
Values of u (positive).														
29°-8°		0.977	0.978	0.969	0.964	0.959	0.954	0.949	0.944	0.938	0.932	0.926	0.919	0.912
Values of v .														
29°	P o s i t i v e	0.119	0.128	0.137	0.147	0.156	0.165	0.175	0.184	0.193	0.202	0.211	0.219	0.228
28		0.114	0.123	0.132	0.141	0.150	0.158	0.167	0.176	0.185	0.194	0.202	0.211	0.219
27		0.109	0.118	0.126	0.135	0.144	0.152	0.160	0.169	0.177	0.185	0.195	0.202	0.210
26		0.104	0.113	0.121	0.129	0.137	0.145	0.153	0.161	0.169	0.177	0.185	0.193	0.201
25		0.100	0.108	0.116	0.124	0.131	0.138	0.146	0.153	0.161	0.169	0.176	0.185	0.192
24		0.095	0.103	0.110	0.118	0.125	0.133	0.140	0.147	0.155	0.162	0.169	0.176	0.183
23		0.091	0.098	0.105	0.112	0.119	0.126	0.134	0.141	0.147	0.154	0.161	0.168	0.175
22		0.087	0.093	0.100	0.107	0.114	0.120	0.127	0.134	0.140	0.147	0.153	0.160	0.166
21		0.082	0.089	0.095	0.102	0.108	0.114	0.120	0.127	0.133	0.140	0.146	0.152	0.158
20		0.078	0.084	0.090	0.096	0.102	0.108	0.114	0.120	0.126	0.132	0.138	0.144	0.150
19		0.074	0.079	0.085	0.091	0.097	0.102	0.108	0.114	0.120	0.125	0.131	0.136	0.142
18		0.069	0.075	0.080	0.086	0.091	0.097	0.102	0.107	0.113	0.118	0.123	0.129	0.134
17		0.065	0.071	0.076	0.081	0.086	0.091	0.096	0.101	0.106	0.111	0.116	0.121	0.126
16		0.061	0.066	0.071	0.076	0.081	0.085	0.090	0.095	0.100	0.104	0.109	0.114	0.118
15		0.057	0.062	0.066	0.071	0.075	0.080	0.084	0.089	0.093	0.098	0.102	0.106	0.110
14		0.053	0.058	0.062	0.066	0.070	0.074	0.079	0.083	0.087	0.091	0.095	0.099	0.103
13		0.049	0.053	0.057	0.061	0.065	0.069	0.073	0.076	0.080	0.084	0.088	0.091	0.095
12		0.045	0.049	0.053	0.056	0.060	0.063	0.067	0.070	0.074	0.077	0.081	0.084	0.088
11		0.041	0.045	0.048	0.051	0.055	0.058	0.061	0.064	0.068	0.071	0.074	0.077	0.080
10		0.038	0.041	0.044	0.047	0.050	0.053	0.055	0.058	0.061	0.064	0.067	0.070	0.073
9		0.034	0.037	0.040	0.043	0.046	0.048	0.050	0.052	0.055	0.057	0.060	0.063	0.065
8		0.030	0.033	0.036	0.039	0.042	0.044	0.045	0.047	0.049	0.051	0.053	0.056	0.058
Values of w .														
29°	P o s i t i v e	0.245	0.264	0.283	0.303	0.322	0.341	0.360	0.379	0.398	0.416	0.435	0.453	0.471
28		0.242	0.261	0.281	0.300	0.319	0.338	0.357	0.375	0.394	0.412	0.431	0.449	0.467
27		0.240	0.259	0.278	0.297	0.316	0.334	0.353	0.372	0.390	0.408	0.427	0.444	0.462
26		0.238	0.257	0.276	0.294	0.313	0.332	0.350	0.369	0.387	0.405	0.423	0.441	0.458
25		0.236	0.255	0.273	0.292	0.311	0.329	0.347	0.365	0.383	0.401	0.419	0.437	0.454
24		0.234	0.253	0.271	0.290	0.308	0.326	0.344	0.363	0.380	0.398	0.416	0.433	0.451
23		0.232	0.251	0.269	0.287	0.306	0.324	0.342	0.360	0.378	0.395	0.413	0.430	0.448
22		0.231	0.249	0.267	0.285	0.304	0.322	0.339	0.357	0.375	0.392	0.410	0.427	0.444
21		0.229	0.247	0.265	0.284	0.302	0.319	0.337	0.355	0.372	0.390	0.407	0.424	0.441
20		0.228	0.246	0.264	0.282	0.300	0.317	0.335	0.352	0.370	0.387	0.404	0.421	0.438
19		0.227	0.244	0.262	0.280	0.298	0.315	0.333	0.351	0.368	0.385	0.402	0.419	0.436
18		0.225	0.242	0.261	0.278	0.296	0.313	0.331	0.348	0.366	0.383	0.400	0.416	0.433
17		0.224	0.241	0.259	0.277	0.295	0.312	0.329	0.347	0.364	0.381	0.398	0.414	0.431
16		0.222	0.240	0.258	0.275	0.293	0.310	0.327	0.345	0.362	0.379	0.395	0.412	0.429
15		0.221	0.239	0.256	0.274	0.291	0.309	0.326	0.343	0.360	0.377	0.393	0.410	0.427
14		0.220	0.238	0.255	0.273	0.290	0.307	0.324	0.341	0.358	0.375	0.392	0.408	0.425
13		0.219	0.237	0.254	0.271	0.289	0.306	0.323	0.340	0.357	0.373	0.390	0.406	0.423
12		0.218	0.236	0.253	0.270	0.288	0.305	0.322	0.339	0.355	0.372	0.388	0.405	0.421
11		0.218	0.235	0.252	0.270	0.287	0.304	0.321	0.338	0.354	0.371	0.387	0.403	0.420
10		0.217	0.234	0.251	0.269	0.286	0.303	0.320	0.337	0.353	0.369	0.386	0.402	0.418
9		0.216	0.233	0.250	0.268	0.285	0.302	0.319	0.336	0.352	0.368	0.385	0.401	0.417
8		0.215	0.232	0.249	0.267	0.284	0.301	0.318	0.335	0.351	0.367	0.384	0.400	0.416

* Extension of tables XVII, XVIII for Burma and Assam.

Case IV.— $w_0=1''$.Values of u, v, w
in seconds,

TABLE XXXVI*.

Lat.	Long.	90°	91°	92°	93°	94°	95°	96°	97°	98°	99°	100°	101°	102°
Values of <i>u</i> (<i>negative</i>).														
29°-8°		0.196	0.212	0.227	0.243	0.258	0.273	0.288	0.303	0.318	0.333	0.348	0.363	0.377
Values of <i>v</i> .														
29°	Positive 													

* Extension of tables XIX, XX for Burma and Assam.

It will be noticed that in tables XXXI, XXXIV discontinuities in the values of w_g in the neighbourhood of lat. 20° , long. 91° , 92° are easily apparent. More careful examination of tables XXIX, XXX, XXXII, XXXIII reveals similar but much less marked discontinuities in the values of u_g and v_g . These are inevitable in view of the method by which the quantities have been found, and the differences are in agreement with those which may be computed by equations (42)—(47) of Chapter I following the two paths (viz. by direct geodesic and by two geodesics through the second origin) to such a point as $L = 91\frac{1}{2}^\circ$, $\lambda = 20^\circ$. The amounts are not however sufficiently large to be of practical importance: and moreover they do not actually occur to the same extent in the actual triangulation of India as in the tables, for the tables have been extended somewhat beyond the triangulation limits for facility of subsequent interpolation. It may be mentioned in passing that the azimuth of a ray of length 40 miles is altered by an amount of order $0''\cdot 1$ when its terminal latitude or longitude is altered by an amount of order $0''\cdot 001$: so that the taking out of azimuth to more than one place of decimals is not really defensible when the coordinates are given to only three places. In the present instance the ordinary procedure of the department has been followed and three places of decimals have been kept, with the idea that at any time the latter two of these may be disregarded.

CHAPTER IV.

Geometrical change from one Spheroid of Reference to another.

1. In selecting a spheroid of reference for the geoid there is no doubt as to the direction of the polar axis; for this is the axis about which heavenly bodies appear to rotate. Hence all possible spheroids of reference are defined by the size of their axes and the position of their centres.

Consider two such spheroids. Let the semi-axes of one be a, b and of the other $a' = a + da, b' = b + db$. Select the origin of coordinates at the centre of the first spheroid and let the coordinates of the centre of the second be $aa, a\beta, b\gamma$ where a, β, γ are small quantities.

In relating a point on a geoid to the spheroid the natural course seems to be to draw the normal through the point to the spheroid and to find out the coordinates of the point where this normal meets the spheroid. So long as the spheroid and geoid are not widely different this normal may, without appreciable error be considered as the vertical to the geoid and also as a straight line. For supposing there is a plumb-line deflection of 1 minute and a separation of the geoid and spheroid by 300 feet, the divergence of the normal from the vertical only amounts to about one inch which only affects coordinates by 0.001 of a second. It is accordingly satisfactory to relate a point on the geoid to one on the spheroid by merely producing the vertical of the geoid until it meets the spheroid. Considering then the relation between the points thus obtained on two reference spheroids corresponding to a point on the geoid, it is clear that all these points may with sufficient accuracy be regarded as being on a straight line, this straight line being normal to one of the three surfaces, whichever is most convenient.

To any triangle formed by three points on the geoid there is a corresponding triangle on any reference spheroid. The angles of these triangles are not identical. Those on the spheroid have different spheroidal excesses. The angles of a triangle observed on the geoid accordingly require correcting before they can be properly applied to a spheroid of reference. If this is properly done then this point relationship given above will hold. It is a fault in reduction of most, if not all, survey observations that geoidal and spheroidal angles have been treated as identical.

2. The coordinates of a point P on the first spheroid may be represented by

$$a \cos \phi \cos L, \quad a \cos \phi \sin L, \quad b \sin \phi$$

while those of a related point P' on the second spheroid may be represented by

$$aa + a' \cos \phi' \cos L', \quad a\beta + a' \cos \phi' \sin L', \quad b\gamma + b' \sin \phi'$$

where $\phi' = \phi + d\phi$, $L' = L + dL$. It is necessary to find expressions for $d\phi$ and dL . It is customary to decide on a point on a spheroid of reference as origin. All spheroids of reference are supposed to pass through this. Suppose that the origin lies in the plane of xz , so that L vanishes at the origin. In notation of previous chapters $dL = v$, $d\phi = u_1$. At the origin these quantities reduce to v_0 and ${}_0u_1$. The value of v_0 is not obtained directly, but is derivable after azimuth has been decided on by the relation

$$w_0 = v_0 \sin \lambda_0 \quad \dots \dots \dots (1)$$

This of course does not show the error of longitude of the origin: it merely shows by how much it will be changed if the azimuth is changed on the supposition of a plumb-line deflection in prime vertical.

Since the origin on either spheroid is identical in position the expression for its coordinates may be equated. Hence putting $\phi = \phi_0$ and $L = 0$

$$\left. \begin{aligned} a \cos \phi_0 &= aa + a' \cos \phi_0' \cos v_0 \\ 0 &= a\beta + a' \cos \phi_0' \sin v_0 \\ b \sin \phi_0 &= b\gamma + b' \sin \phi_0' \end{aligned} \right\} \dots \dots \dots (2)$$

Neglecting second order quantities and substituting from (1) for v_0 (2) may be written

$$\left. \begin{aligned} a + \frac{da}{a} \cos \phi_0 - {}_0u_1 \sin \phi_0 &= 0 \\ \beta + w_0 \frac{\cos \phi_0}{\sin \lambda_0} &= 0 \\ \gamma + \frac{db}{b} \sin \phi_0 + {}_0u_1 \cos \phi_0 &= 0 \end{aligned} \right\} \dots \dots \dots (3)$$

These equations serve to determine a, β, γ in terms of the axes changes and changes at the origin: if the quantities $\frac{da}{a}, \frac{db}{b}$ are multiplied by cosec $1'$ the results are expressed in seconds.

3. Two further conditions are obtained by expressing the fact that the normal to the first spheroid at any point passes through the related point on the other spheroid.

The normal at P is

$$\frac{x - a \cos \phi \cos L}{\cos \phi \cos L} = \frac{y - a \cos \phi \sin L}{\cos \phi \sin L} = \frac{z - b \sin \phi}{\sin \phi}$$

and the conditions that P' should lie on this are

$$\frac{aa + d(a \cos \phi \cos L)}{\cos \phi \cos L} = \frac{a\beta + d(a \cos \phi \sin L)}{\cos \phi \sin L} = \frac{b\gamma + d(b \sin \phi)}{\sin \phi}$$

whence

$$\begin{aligned} \frac{a}{\cos \phi \cos L} + \frac{da}{a} - \tan \phi \cdot u_1 - \tan L \cdot v &= \frac{\beta}{\cos \phi \sin L} + \frac{da}{a} - \tan \phi \cdot u_1 + \cot L \cdot v \\ &= (1 - e^2) \left\{ \frac{\gamma}{\sin \phi} + \frac{db}{b} + \cot \phi \cdot u_1 \right\} \dots (4) \end{aligned}$$

From the first of equations (4)

$$v (\tan L + \cot L) = \sec \phi \left(\frac{a}{\cos L} - \frac{\beta}{\sin L} \right)$$

or

$$v = \sec \phi (a \sin L - \beta \cos L) \quad \dots \quad (5)$$

Eliminating v from (4) it follows that

$$\begin{aligned} \left(\frac{a}{\cos \phi \cos L} + \frac{da}{a} - u_1 \tan \phi \right) \cot L + \left(\frac{\beta}{\cos \phi \sin L} + \frac{db}{a} - u_1 \tan \phi \right) \tan L \\ = (1 - e^2) (\tan L + \cot L) \left(\frac{\gamma}{\sin \phi} + \frac{db}{b} + u_1 \cot \phi \right) \end{aligned}$$

whence

$$(a \cos L + \beta \sin L) \sec \phi + \frac{da}{a} - u_1 \tan \phi = (1 - e^2) \left(\frac{\gamma}{\sin \phi} + \frac{db}{b} + u_1 \cot \phi \right)$$

and expressing the results in seconds this may be written

$$u_1 (1 - e^2 \cos^2 \phi) = (a \cos L + \beta \sin L) \sin \phi - (1 - e^2) \gamma \cos \phi + \sin \phi \cos \phi \left\{ \frac{da}{a} - (1 - e^2) \frac{db}{b} \right\} \operatorname{cosec} 1'' \dots (6)$$

The relation between u_1 and u is given by (16) of Chap. III, and in terms of da and db is

$$u = \left(1 + \frac{e^2}{2} \cos 2\phi \right) u_1 + \frac{1}{2} \left(\frac{da}{a} - \frac{db}{b} \right) \sin 2\lambda \operatorname{cosec} 1'' \dots (7)$$

The quantities a , β , γ , $\frac{da}{a}$, $\frac{db}{b}$ all enter linearly into the equations. Their several effects can accordingly be computed separately and combined afterwards in any desired way. Cases corresponding to each of the four quantities $\frac{da}{a}$, $\frac{db}{b}$, u_1 and w_0 will now be considered.

Case (i) $da = 1 \text{ km.}$ $u_1 = 0$, $(u_0 = 12'' \cdot 063)$

From (3)

$$\begin{aligned} a + A \cos \phi_0 &= 0 & \text{where } A &= \frac{da}{a} \operatorname{cosec} 1'' = 32'' \cdot 3437 \\ \beta &= \gamma = 0 \\ \therefore a &= -29'' \cdot 536 \end{aligned}$$

From (5) and (6)

$$\left. \begin{aligned} v &= a \sec \phi \sin L \\ u_1 (1 - e^2 \cos^2 \phi) &= u_1 \cdot \frac{\sin^2 \phi}{\sin^2 \lambda} = a \sin \phi \cos L + \frac{1}{2} A \sin 2\phi \\ u &= (1 + \frac{e^2}{2} \cos 2\phi) u_1 + \frac{1}{2} A \sin 2\lambda \\ &= \sin 2\lambda \left(\frac{u_1}{\sin 2\phi} + \frac{1}{2} A \right) \end{aligned} \right\} \dots (8)$$

Case (ii) $db = 1 \text{ km.}$ $u_1 = 0$ $(u_0 = -12'' \cdot 1034)$

From (3)

$$\begin{aligned} a &= \beta = 0 & \text{where } B &= \frac{db}{b} \operatorname{cosec} 1'' = 32'' \cdot 4516 \\ \gamma + B \sin \phi_0 &= 0 \\ \therefore \gamma &= -13'' \cdot 2244 \end{aligned}$$

From (5) and (6)

$$\left. \begin{aligned} v &= 0 \\ u_1 (1 - e^2 \cos^2 \phi) &= - (1 - e^2) (\gamma + B \sin \phi) \cos \phi \\ u &= \left(1 + \frac{e^2}{2} \cos 2\phi\right) u_1 - \frac{1}{2} B \sin 2\phi \end{aligned} \right\} \dots \dots \dots (9)$$

Case (iii) ${}_0u_1 = 9''.978$ $(u_0 = 10'')$

From (3) $\alpha - {}_0u_1 \sin \phi_0 = 0$

$\beta = 0$

$\gamma + {}_0u_1 \cos \phi_0 = 0$

$\therefore \alpha = 4''.0659$ $\gamma = -9''.1118$

From (5) and (6)

$$\left. \begin{aligned} v &= \alpha \sec \phi \sin L \\ u_1 (1 - e^2 \cos^2 \phi) &= \alpha \sin \phi \cos L - (1 - e^2) \gamma \cos \phi \\ u &= \left(1 + \frac{e^2}{2} \cos 2\phi\right) u_1 \end{aligned} \right\} \dots \dots \dots (10)$$

Case (iv) $w_0 = 1''$ $v_0 = w_0 \operatorname{cosec} \lambda_0 = 2''.447$

$\alpha = \gamma = 0$

From (3) $\beta + \frac{\cos \phi_0}{\sin \lambda_0} = 0$ $\therefore \beta = -2''.2347$

From (5) and (6)

$$\left. \begin{aligned} v &= -\beta \sec \phi \cos L \\ u_1 (1 - e^2 \cos^2 \phi) &= \beta \sin L \sin \phi \\ u &= \left(1 + \frac{e^2}{2} \cos 2\phi\right) u_1 \end{aligned} \right\} \dots \dots \dots (11)$$

To find the azimuth change w , the following equation holds for all cases

$$v - v_0 = w \operatorname{cosec} \lambda - w_0 \operatorname{cosec} \lambda_0 \dots \dots \dots (12)$$

in which v_0 includes the entire origin change of longitude and is not restricted to that due to plumb-line deflection only. The equation follows from the fact that either side of it gives the difference between spheroidal and geoidal longitude. It is proved otherwise in the following chapter (*vide* equation (4)). With reference to the case IV (or (iv)) it will be noticed that the value of $u \propto \sin L \sin \phi$. The value found by the method of Chapter I was independent of ϕ . The two cases however are not geometrically similar. In the case of Chapter I an azimuth change of origin involves a twist about the normal at the origin. In the present case the fixed axis is the polar axis and any twist introduced to give any desired azimuth change is only a component of a twist round an axis parallel to the polar axis. This makes it clear why the effect on latitude of this azimuth change is zero at the equator, the equator being at right angles to the axis of twist.

The values of u , v , w have been computed for the four cases by means of equations (8) to (12). The values of λ and L are the same as those of tables XXVII, XXVIII, and hence it is easy to make a comparison between the values of u , v , w found by the method of the present chapter which may be denoted by u_r , v_r , w_r (related points on two spheroids) and u_g , v_g , w_g (found by following a geodesic). For this purpose values of $u_r - u_g$ &c. are exhibited in tables XXXVII—XL.

TABLE XXXVII.

Case I.— $\delta\alpha = 1$ km.

L'—L	0°	4°	8°	12°	16°	20°	24°
λ	Values of $(u_r - u_g)$ in seconds.						
38°	0.000	+0.010	+0.049	+0.117	+0.201	+0.315	+0.442
34°	+0.011	+0.021	+0.053	+0.099	+0.165	+0.249	+0.347
30°	+0.005	+0.014	+0.031	+0.061	+0.105	+0.154	+0.216
26°	+0.005	+0.008	+0.013	+0.022	+0.038	+0.053	+0.076
22°	-0.003	-0.010	-0.017	-0.028	-0.043	-0.061	-0.090
18°	-0.004	-0.017	-0.038	-0.070	-0.117	-0.180	-0.255
14°	+0.005	-0.012	-0.048	-0.103	-0.186	-0.290	-0.418
10°	+0.035	+0.013	-0.040	-0.120	-0.239	-0.386	-0.576
λ	Values of $\pm(v_r - v_g)^*$ in seconds.						
38°	0.000	+0.038	+0.080	+0.080	+0.047	-0.005	-0.116
34°	0.000	+0.016	+0.026	+0.020*	-0.019	-0.100	-0.234
30°	0.000	+0.003	+0.003	-0.015	-0.071	-0.162	-0.312
26°	0.000	+0.001	-0.005	-0.026	-0.090	-0.193	-0.351
22°	0.000	-0.002	-0.003	-0.031	-0.091	-0.186	-0.346
18°	0.000	-0.002	-0.004	-0.012	-0.061	-0.153	-0.293
14°	0.000	+0.021	+0.033	+0.034	-0.001	-0.079	-0.204
10°	0.000	+0.047	+0.082	+0.108	+0.100	+0.044	-0.068
λ	Values of $\pm(w_r - w_g)^*$ in seconds.						
38°	0.000	+0.371	+0.735	+1.093	+1.436	+1.756	+2.058
34°	0.000	+0.273	+0.542	+0.806	+1.058	+1.292	+1.509
30°	0.000	+0.168	+0.320	+0.494	+0.641	+0.782	+0.913
26°	0.000	+0.052	+0.107	+0.160	+0.205	+0.238	+0.265
22°	0.000	-0.068	-0.130	-0.197	-0.267	-0.338	-0.411
18°	0.000	-0.191	-0.384	-0.573	-0.760	-0.950	-1.136
14°	0.000	-0.328	-0.648	-0.969	-1.285	-1.598	-1.903
10°	0.000	-0.466	-0.928	-1.380	-1.830	-2.278	-2.707

TABLE XXXVIII.

Case II.— $\delta\delta = 1$ km.

L'—L	0°	4°	8°	12°	16°	20°	24°
λ	Values of $(u_r - u_g)$ in seconds.						
38°	+0.078	+0.073	+0.043	-0.005	-0.082	-0.148	-0.246
34°	+0.017	+0.008	-0.012	-0.044	-0.093	-0.156	-0.234
30°	+0.008	+0.001	-0.011	-0.032	-0.084	-0.105	-0.159
26°	-0.004	-0.006	-0.010	-0.020	-0.032	-0.052	-0.075
22°	+0.003	+0.008	+0.012	+0.017	+0.024	+0.029	+0.035
18°	-0.001	+0.011	+0.025	+0.040	+0.069	+0.098	+0.139
14°	-0.034	-0.022	+0.001	+0.034	+0.084	+0.144	+0.216
10°	-0.114	-0.100	-0.065	-0.017	+0.060	+0.147	+0.259
λ	Values of $\pm(v_r - v_g)^*$ in seconds.						
38°	0.000	-0.068	-0.133	-0.201	-0.266	-0.317	-0.371
34°	0.000	-0.030	-0.064	-0.101	-0.127	-0.152	-0.168
30°	0.000	-0.003	-0.024	-0.037	-0.043	-0.044	-0.043
26°	0.000	-0.003	-0.009	-0.015	-0.005	-0.004	+0.015
22°	0.000	-0.001	-0.011	-0.011	-0.008	-0.007	+0.011
18°	0.000	-0.008	-0.025	-0.043	-0.052	-0.059	-0.056
14°	0.000	-0.033	-0.069	-0.107	-0.147	-0.173	-0.196
10°	0.000	-0.070	-0.137	-0.217	-0.285	-0.348	-0.400
λ	Values of $\pm(w_r - w_g)^*$ in seconds.						
38°	0.000	-0.388	-0.777	-1.170	-1.566	-1.960	-2.363
34°	0.000	-0.280	-0.564	-0.853	-1.143	-1.438	-1.740
30°	0.000	-0.171	-0.341	-0.523	-0.700	-0.889	-1.090
26°	0.000	-0.053	-0.114	-0.177	-0.251	-0.325	-0.416
22°	0.000	+0.067	+0.124	+0.182	+0.228	+0.268	+0.298
18°	0.000	+0.190	+0.379	+0.568	+0.728	+0.885	+1.030
14°	0.000	+0.324	+0.642	+0.953	+1.255	+1.542	+1.815
10°	0.000	+0.463	+0.920	+1.366	+1.805	+2.233	+2.636

TABLE XXXIX.

Case III.— $u_0 = 1''$

TABLE XL.

Case IV.— $w_0 = 1''$

L'—L	0°	4°	8°	12°	16°	20°	24°
λ	Values of $(u_r - u_g)$ in seconds.						
38°	-0.031	-0.029	-0.024	-0.015	-0.002	+0.014	+0.033
34°	-0.016	-0.014	-0.009	+0.001	+0.014	+0.030	+0.050
30°	-0.006	-0.004	+0.002	+0.011	+0.025	+0.042	+0.062
26°	-0.001	+0.002	+0.007	+0.017	+0.031	+0.048	+0.070
22°	0.000	+0.002	+0.008	+0.018	+0.033	+0.051	+0.072
18°	-0.005	-0.002	+0.004	+0.014	+0.029	+0.047	+0.070
14°	-0.014	-0.012	-0.005	+0.005	+0.021	+0.040	+0.063
10°	-0.029	-0.026	-0.020	-0.008	+0.007	+0.027	+0.051
λ	Values of $\pm(v_r - v_g)^*$ in seconds.						
38°	0.000	+0.019	+0.037	+0.065	+0.073	+0.081	+0.108
34°	0.000	+0.013	+0.026	+0.038	+0.051	+0.064	+0.075
30°	0.000	+0.008	+0.016	+0.023	+0.030	+0.037	+0.044
26°	0.000	+0.002	+0.005	+0.007	+0.009	+0.011	+0.014
22°	0.000	-0.003	-0.006	-0.006	-0.009	-0.012	-0.014
18°	0.000	-0.007	-0.014	-0.022	-0.029	-0.035	-0.042
14°	0.000	-0.011	-0.023	-0.035	-0.046	-0.057	-0.068
10°	0.000	-0.017	-0.032	-0.049	-0.065	-0.081	-0.096
λ	Values of $\pm(w_r - w_g)^*$ in seconds.						
38°	0.000	+0.067	+0.133	+0.198	+0.262	+0.326	+0.387
34°	0.000	+0.065	+0.130	+0.194	+0.257	+0.319	+0.380
30°	0.000	+0.065	+0.128	+0.191	+0.253	+0.315	+0.375
26°	0.000	+0.064	+0.127	+0.191	+0.252	+0.313	+0.373
22°	0.000	+0.065	+0.128	+0.191	+0.253	+0.313	+0.371
18°	0.000	+0.065	+0.129	+0.192	+0.254	+0.315	+0.373
14°	0.000	+0.065	+0.129	+0.193	+0.256	+0.317	+0.378
10°	0.000	+0.066	+0.131	+0.196	+0.260	+0.323	+0.383
λ	Values of $\pm(v_r - v_g)$ in seconds.						
38°	0.000	+0.031	+0.064	+0.095	+0.128	+0.158	+0.188
34°	0.000	+0.028	+0.047	+0.069	+0.093	+0.116	+0.138
30°	0.000	+0.018	+0.028	+0.042	+0.056	+0.071	+0.084
26°	0.000	+0.004	+0.009	+0.013	+0.019	+0.024	+0.028
22°	0.000	-0.003	-0.011	-0.016	-0.021	-0.025	-0.030
18°	0.000	-0.018	-0.031	-0.047	-0.061	-0.076	-0.090
14°	0.000	-0.027	-0.053	-0.078	-0.103	-0.127	-0.151
10°	0.000	-0.038	-0.074	-0.110	-0.145	-0.180	-0.214
λ	Values of $\pm(w_r - w_g)$ in seconds.						
38°	+0.083	+0.078	+0.062	+0.037	+0.001	-0.045	-0.100
34°	+0.040	+0.035	+0.020	-0.005	-0.038	-0.065	-0.140
30°	+0.013	+0.009	-0.006	-0.031	-0.066	-0.110	-0.164
26°	+0.001	-0.004	-0.019	-0.044	-0.078	-0.122	-0.176
22°	+0.002	-0.003	-0.017	-0.043	-0.077	-0.121	-0.184
18°	+0.014	+0.008	-0.006	-0.031	-0.066	-0.110	-0.164
14°	+0.038	+0.031	+0.016	-0.010	-0.044	-0.089	-0.143
10°	+0.084	+0.064	+0.047	+0.020	-0.014	-0.059	-0.114
λ	Values of $(u_r - u_g)$ in seconds.						
38°	+0.585	+0.585	+0.580	+0.573	+0.564	+0.550	+0.536
34°	+0.407	+0.406	+0.403	+0.395	+0.383	+0.362	+0.371
30°	+0.235	+0.235	+0.233	+0.231	+0.226	+0.221	+0.214
26°	+0.075	+0.074	+0.073	+0.073	+0.072	+0.070	+0.068
22°	-0.082	-0.082	-0.082	-0.081	-0.079	-0.078	-0.079
18°	-0.234	-0.233	-0.232	-0.229	-0.225	-0.220	-0.214
14°	-0.384	-0.383	-0.380	-0.376	-0.370	-0.362	-0.350
10°	-0.533	-0.532	-0.528	-0.522	-0.512	-0.501	-0.486

* + or - according as point is west or east of origin.

4. The differences $w_r - w_g$ and $v_r - v_g$ are never sufficiently great to have any important effect on geodetic results. In the case of $w_r - w_g$ larger values are met with. The deflections of plumb-line in the prime vertical are affected by the amount $(w_r - w_g) \cot \lambda$, a quantity which may be as much as several seconds. In the practical case however where, according to the most recent determinations $\delta a = 0.924 \text{ km}$ and $\delta b = 0.743 \text{ km}$: the combined effect of Cases I and II in these proportions are not large, the two cases tending to cancel each other; as may be seen from table XLI below, in which the values of $(w_r - w_g) \cot \lambda$ are also given.

TABLE XLI.

λ	$L' - L$	$\cot \lambda$	$0.924 \times \text{Case I}$	$0.743 \times \text{Case II}$	Combined effect	Discrepancy in plumb-line deflection
38°	24°	1.280	1.902	-1.756	0.146	0.187
34	24	1.483	1.394	-1.293	0.101	0.150
30	24	1.732	0.844	-0.810	0.034	0.059
26	24	2.050	0.245	-0.309	-0.064	-0.131
22	24	2.475	-0.380	0.214	-0.166	-0.411
18	8	3.078	-0.355	0.282	-0.073	-0.225
14	4	4.011	-0.303	0.241	-0.062	-0.249
10	4	5.671	-0.431	0.344	-0.087	-0.493

The only case in which the discrepancy in prime vertical deflection would be considerable occurs in low latitudes and this case does not concern Indian Triangulation as in these latitudes there are no great longitude differences: for the case of Burma special treatment, as given in Chapter III, is in any case necessary.

The conclusion is that either this method or that of the preceding chapter could be used with practically satisfactory results. The discrepancies however indicate how far theoretical accuracy has been departed from in failing to project geoidal angles on to the spheroid of reference before introducing them in the computations.

The figures in tables XXXVII—XL may be noticed to be a little irregular. This is doubtless due to the fact that the computations of Chapter III were not made with sufficient accuracy to ensure the last figure always being correct. This was not considered to be sufficiently important to justify the extra labour which would have been necessary. The results are fully accurate enough for all practical purposes to which they can be put.

5. Three methods of finding the change in coordinates due to any proposed changes of the axes of the spheroid and the latitude and azimuth at the origin have now been given. That of Chapter I gives a means of computing these along any path defined by a relation between λ and L . Chapter III gives the results for the special case when the path selected is the geodesic through the origin and the point at which the changes are required: and in the present chapter the geometrical relation between corresponding points on two spheroids is worked out. All these methods give somewhat different results. The latter two have the advantage over the former of being free from any ambiguity due to multiple values and inconsistency: and the differences met with between them are not of amount sufficiently large to be troublesome. The reason for their discrepancy is examined in the following chapter: and the conclusion is arrived at, from theoretical considerations, in view of the methods by which the observations of the triangulation of India have been reduced, that the method of calculation along geodesics as set forth in Chapter III is the correct one to use. The detailed tables XXIX—XXXIV permit this to be done readily.

CHAPTER V.

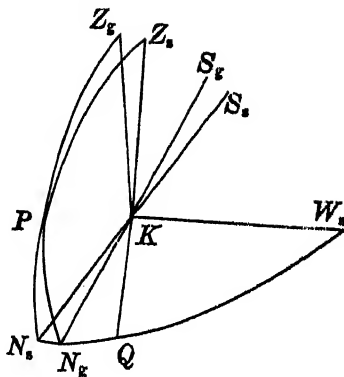
Laplace's Equation and the Choice of a Spheroid of Reference.

1. When a large survey is begun one of the first essentials is the selection of a point as origin. The coordinates of this point have to be decided on. The longitude is of little consequence and the meridian through the point may be taken as that from which all deduced longitudes are measured. The latitude and azimuth can be observed astronomically: but their geodetic values depend on the plumb-line deflection existing at the point. Plumb-line deflection is of course merely the deviation of the vertical from the normal to some assumed figure of reference. In choosing the origin it must be eventually decided whether to consider the deflection there as nil—in other words, choosing a spheroid of reference parallel to the geoid at the origin—or whether considerations of topographical features and irregularities of density justify the adoption of certain values of the deflection in the two components. The question of height above geoid of the point selected as origin also arises. If an error of 10 feet is made in this it is practically equivalent to assuming a spheroid with axes 10 feet different from those actually selected. In the case of the origin of the Indian Survey there is no reason to suppose an error of nearly so much, and so no further consideration will be given to this point here.

Having decided on the origin O , it is next necessary to decide on a figure of reference. This will generally be referred to as the "spheroid" in opposition to the "geoid" or sea level equipotential surface of the earth. It is not implied by this that the figure of reference must be a spheroid, though the almost universal practice is to take a spheroid as a reference figure.

2. Chains of triangulation may now be computed *rigidly* if proper corrections given below in § 16 are applied: and the coordinates, latitude, longitude, height and azimuth at a distant point K may be deduced. Suppose that astronomical azimuth and latitude are observed at K and also that the arc OK is observed as a telegraphic longitude arc. Let λ , L be the geodetic latitude and longitude of K and let A be the geodetic azimuth at K of some reference mark: these are quantities brought up by triangulation. Let η , ξ be the plumb-line deflection at K in meridian and prime vertical (positive for southerly and westerly deflections of the plumb bob) referred to the selected spheroid of reference. These quantities obviously differ for different spheroids of reference.

In the figure suffices s, g refer to spheroidal and geoidal points respectively: or in other words points derived from triangulation and star observations respectively. P is the pole and Z the zenith. The astronomic azimuth of a point Q is clearly $A + \xi \tan \lambda$ being $S_g K S_s$ greater than the geodetic azimuth. Let suffix zero denote quantities appertaining to the origin of the survey: so that the values of $\eta_0 \xi_0$ have been decided in some way or other.



3. First consider the longitude observations. They depend on the interval of time between the meridians of O and K as shown by star transits. If the zenith at a station is displaced in the prime vertical, the meridian is also displaced as a result, the direction of the pole being fixed. Time stars are observed when they transit the plane $Z_g P$ instead of the plane $Z_s P$. With a westerly deflection the zenith is moved towards the east and the result is that stars are observed too soon by the angle $Z_g P Z_s = \xi \sec \lambda$. If ξ is expressed in seconds of arc, the star time as given by the local meridian is early by $\xi \sec \lambda / 15$ seconds of time. Now of the two stations O, K , if O is the more westerly and T is the time interval between the transits of a star at the two meridians, then the time interval between the two spheroidal meridians is

$$T - \frac{\xi \sec \lambda}{15} + \frac{\xi_0 \sec \lambda_0}{15}$$

and this should be the same as $\frac{L - L_0}{15}$, or the difference of longitude in time, on that spheroid on which ξ, ξ_0 represent the plumb-line deflections in prime vertical at K and O .

The quantity T is an observation quantity; suppose its error is δT : also suppose the error in longitude generated in the triangulation is δL . It follows that

$$\xi_0 \sec \lambda_0 - \xi \sec \lambda = L - L_0 - \delta L - 15 (T - \delta T) \quad (1)$$

Now consider the azimuth observations. Let A' be the astronomically observed azimuth which has an error $\delta A'$: A and δA being the geodetic azimuth computed on the spheroid and its error. Then

$$\left. \begin{aligned} A' - \delta A - A + \delta A &= \xi \tan \lambda \\ A'_0 - \delta A_0 - A_0 &= \xi_0 \tan \lambda_0 \end{aligned} \right\} \quad (2)$$

Eliminating ξ, ξ_0 between (1) and (2) it follows that

$$(A'_0 - \delta A'_0 - A_0) \operatorname{cosec} \lambda_0 - (A' - \delta A' - A + \delta A) \operatorname{cosec} \lambda = L - L_0 - \delta L - 15 (T - \delta T) \quad (3)$$

which is an elaborated form of Laplace's equation.

4. Suppose now that the computations had been carried out on a slightly different spheroid. If this had been done rigorously the quantity δL , being itself a small quantity will not be changed appreciably, while $A', \delta A', A'_0, \delta A'_0, T, \delta T$ are all quantities which are not affected by the change in spheroid. The only quantities in (3) which change appreciably are $A, L - L_0$ and λ . The λ terms are multiplied by small coefficients and their variations can be neglected. Hence differentiating (3) for change of spheroid it follows that in the notation of Chapter I where u, v, w represent changes in latitude, longitude and azimuth

$$w \operatorname{cosec} \lambda - w_0 \operatorname{cosec} \lambda_0 = v - v_0 \quad (4)$$

which is equation (12) of Chapter IV.

It might be expected that this equation would be in accordance with those found in Chapter I. As was noticed there, however, the quantities u, v, w are many-valued, a separate set of values appertaining to each route along which the integration is performed. Equation (4) on the other hand is free from any ambiguity and accordingly cannot be in accord with the equations of Chapter I. If numerical quantities are substituted it is at once clear that the relation (4) is not satisfied. Consider the values of

$$v = v_x - f(v_x - v_y) \text{ and } w = w_x - f'(w_x - w_y)$$

where f and f' are fractional quantities. These expressions are the values of v, w computed along routes intermediate to those of v_x and v_y . Taking case where $\delta a = 1 \text{ km.}$, $\lambda = 30^\circ$, $L = 66^\circ$, $v_0 = w_0 = 0$ from the tables VII—X it follows that

$$v \sin 30^\circ = 4.027 - .013 f$$

and

$$w = 3.784 - .504 f$$

which cannot be made equal by any positive fractional values of f and f' . In the same way the tables of Chapter III show that the relation (4) is not satisfied along a geodesic.

5. It has generally been considered that azimuth and longitude observations both give the same information, namely deflection of the plumb-line in prime vertical, and nothing more: and in so far as the results differ by the two methods the reason is that the observations are burdened by errors. Clarke states* that "the observations of the difference of longitude gives "us no information that is not also given by the observation for azimuth". With this principle Colonel Sir Sidney Burrard† has used the longitude observations of India to correct azimuth observations for the accumulation of error due to triangulation, considering the differences of the resulting plumb-line deflection found by the two observations to be entirely accounted for by observation error in the triangulation.

6. The explanation of these apparent inconsistencies was not discovered for some time. Equation (4) is perfectly correct *if the triangulation is properly computed*. The ordinary process of computation is not quite correct. Angles are measured by means of a theodolite and reduced to the horizontal plane of the geoid. This is not quite the same thing in general as the horizontal plane of the spheroid. If the computation is to be effected on the spheroid (on which all the various formulæ are based) the observed angles should be projected on to the selected spheroid of reference, and so will differ according to what spheroid is selected. The actual amount by which the geoidal angle must be altered to get the spheroidal angle depends on two things (vide § 16 below)

(1) the deflection of the plumb-line or inclination of the geoidal (astronomic) vertical to the spheroidal (triangulated) vertical.

(2) the inclination to the horizontal of the rays between which the geoidal angle is measured.

The first of these quantities varies appreciably with change of spheroid and accordingly the correction to the geoidal angle varies according to the spheroid used. The actual case under consideration is represented in symbols by supposing δL in (1) to contain not only the error due to faulty observations but also the error due to failure to correct the geoidal angles to spheroidal angles. This is purely a computation error. The actual "grinding" process has treated these errors as errors of observation.

This perhaps explains why Laplace's equation is in general not satisfied so well as the probable errors of the several observations on which its formation depends would cause to be expected.

* "Geodesy" by Col. A.R. Clarke, p. 291.

† Appendix No. 5 of G.T. Volume XVIII. "On the azimuth observations of the G.T.S. of India".

7. It also explains how it is that the different values of u, v, w arise as noticed in Chapter I. In this case the fact that the closing errors of circuits will differ from one spheroid to another *unless all the geoidal angles are reduced to spheroidal angles* makes the distribution of closing errors have a different effect according as the spheroid is altered.

Equation (4) may be re-written to meet the actual case as follows:

$$[w \operatorname{cosec} \lambda - v] = \Delta L \quad \dots \dots \dots (5)$$

where ΔL is the change in computation error of longitude difference due to the treatment of spheroidal and geoidal angles as identical. The fact that ΔL is not zero would be of more serious importance in the question of change of spheroid had equation (3) been used in all possible cases as a condition for the series of triangulation to satisfy. When the main Indian triangulation was adjusted the longitude arcs either were not available or else were ignored, so that Laplace's condition was not imposed on the triangulation. In correcting the azimuth observations, Colonel Sir Sidney Burrard introduced the condition for the first time in India.

8. Clarke's statement quoted in § 5 was deduced from an equation which arises in his work: but it may be seen to be true without any analysis. The longitude observation fixes the meridian plane at a point, that is the plane through the zenith and the pole, by taking the time of stars transiting this plane. It obviously does no more than fix this plane with relation to another. The azimuth observation practically draws the great circle through the pole and zenith and locates where this cuts the horizon, by means of a horizontal angle measured from a fixed point. The position of the pole and the place of observation being already given the fixing of one other point suffices to fix the meridian plane. Thus longitude and azimuth observations both merely fix the position of the meridian plane and nothing more. The inclination of this plane to the meridian plane deducible from triangulation is the deflection in prime vertical.

9. None the less equation (3) does definitely give some information as to the error of computation generated in the triangulation, and to this extent Clarke's statement needs modification. When the practical case is considered from equation (3) it may be seen that the identity of plumb-line deflection, whether derived from longitude observations or azimuth observations, affords some information concerning the slightly faulty method of computing from geoidal angles instead of from spheroidal angles. For split up the error δL into $\delta_1 L$ due to faulty observation and $\delta_2 L$ due to faulty computation. Suppose next that the spheroid of reference is changed so that it is necessary to substitute $A+w$ for A and $L+v$ for L . The quantity $\delta_1 L$ remains unaltered, but $\delta_2 L$ obviously is a variable according to the spheroid used and from (3) it follows that

$$(A'_0 - A_0 - w_0) \operatorname{cosec} \lambda_0 - (A' - A - w) \operatorname{cosec} \lambda = L - L_0 + v - v_0 - 15T - \delta_2 L + \Delta E \quad \dots \dots (6)$$

in which the only variables are $w, w_0, v - v_0$ and $\delta_2 L$, and ΔE is the combined and fixed effect of observation errors. It is possible to form sixteen equations of the form (6) from the longitude and azimuth observations of India. Expressing $w, w_0, v - v_0$ in terms of $\delta a, \delta b, u_0$ and w_0 it is possible to solve these equations for $\delta a, \delta b, u_0, w_0$ so as to make $\sum (\Delta E - \delta_2 L)^2$ a minimum: i.e. since ΔE is equally likely to be positive or negative $\sum \overline{\Delta E}^2 + \sum \overline{\delta_2 L}^2$ is a minimum. But as the value of ΔE is not being varied, this implies that $\sum \overline{\delta_2 L}^2$ is a minimum. Now when the spheroid differs widely from the geoid it is clear that the computation errors increase: and conversely when the spheroid approximates more closely to the geoid these errors diminish. The fact that $\sum \overline{\delta_2 L}^2$ is made a minimum affords one criterion for the spheroid being in close agreement with the geoid for the area over which the triangulation of India extends. It is of course possible to consider what spheroid suits the actual deflections best: but this is an entirely different point of view from that indicated above, and deals only with the actual localities in which the deflections are measured: and moreover is burdened by the errors of computation involved in treating geoidal and spheroidal angles as identical.

10. The interest in the method is chiefly theoretical. The quantities to be dealt with are very small: and in most cases the effects of observation error may well mask those due to the computation error. Sixteen equations of the form (6) are given below. These can be solved for δa , δb , u_0 , w_0 or, treating δa , δb as known, for u_0 , w_0 only. It was not anticipated that the former course would give reliable values of δa , δb but the solution was none the less made. Afterwards the solution of u_0 , w_0 only taking the latest values of δa , δb was performed. Referring to these two solutions as *A* and *B*, one difficulty of the application of (6) arises in *A*, but to a very much less extent in *B*. This difficulty is the selection of the route along which u , v , w in terms of δa , δb shall be determined. The actual courses of the triangulation series are numerous, and the case seems to be best met by taking the geodesic solution of Chapter III, for this in general leads to a medial path through the triangulation. In solution *B* it so happens that the δa and δb terms very nearly cancel one another. The form of the equations is as follows

$$(v_1 \sin \lambda - w_1) \delta a + (v_2 \sin \lambda - w_2) \delta b + (v_3 \sin \lambda - w_3) u_0 + (v_4 \sin \lambda - w_4) w_0 + A' - A - (A'_0 - A_0) \sin \lambda \operatorname{cosec} \lambda_0 - \sin \lambda (15T - L + L_0) = 0 \quad (7)$$

in which $A'_0 - A_0 = 1'' \cdot 29 - w_0$ (*vide* Chapter I §4).

The sixteen arcs from Kalianpur give rise to sixteen equations which are exhibited in the table.

TABLE XLII.

Azimuth station	Longitude station	Coordinates of Azimuth station		Coefficients of				$A' - A$	$15T - L + L_0$	Absolute term A $- 1 \cdot 29 \operatorname{cosec} \lambda_0 \sin \lambda$ $+ A' - A$ $- (15T - L + L_0) \sin \lambda$	Residual <i>A</i>	Absolute term <i>B</i>	Residual <i>B</i>
		λ	L	δa	δb	u_0	w_0						
Karachi Observatory	Karachi T. O.	24° 49' 50"	67° 1' 35"	+0.080	-0.059	+0.168	+0.042	-1.4	+0.5	-2.9	-1.8	-2.9	-3.0
Dehra Dun Observatory (old)	Dehra Dun Longitude Station	30 19 57	78 3 35	-0.017	+0.018	-0.005	+0.242	-11.9	-25.7	-0.5	-2.5	-0.5	-2.3
Quetta T. O.	Quetta T. O.	30 11 57	67 0 32	+0.457	-0.459	+0.160	+0.251	-4.4	+2.4	-7.2	-3.4	-7.1	-8.8
Calcutta Base-line, S. end	Calcutta	22 30 56	88 22 54	+0.118	-0.118	-0.172	-0.043	-8.9	-11.0	-5.9	-5.4	-5.9	-5.8
Orejhar	Fyzabad T. O.	26 46 56	82 12 8	-0.085	+0.085	-0.071	+0.107	-4.1	-0.5	-5.3	-7.4	-5.8	-6.2
Jalpaiguri	Jalpaiguri	26 31 17	88 44 13	-0.199	+0.196	-0.172	+0.109	-4.7	-20.4	+3.0	-1.0	+3.0	+2.0
Nagarkhana	Chittagong T. O.	22 22 58	91 48 30	+0.226	-0.051	-0.227	-0.043	-8.7	-11.7	-5.4	-0.1	-5.2	-5.1
Bolarum P.W.D. Office	Bolarum	17 30 13	78 31 11	+0.046	-0.045	-0.014	-0.257	-1.1	-3.5	-1.0	+1.4	-1.0	+0.9
Vizagapatam Base-line, N. end	Waltair	18 1 3	83 13 43	+0.268	-0.267	-0.092	-0.235	-1.4	-3.3	-1.4	+2.7	-1.4	+0.2
Karaundi	Jubbulpore T. O.	23 10 40	79 59 43	+0.017	-0.021	-0.088	-0.038	-4.0	-10.2	-1.2	-1.1	-1.2	-1.0
Colaba Observatory	Bombay	18 53 49	72 48 49	-0.198	+0.199	+0.080	-0.199	+1.0	-6.8	+2.2	+2.2	+2.2	+3.7
Deesa T. O.	Deesa T. O.	24 15 30	72 11 6	+0.003	-0.003	+0.086	+0.009	-4.6	+3.6	-7.4	-6.9	-7.4	-7.4
Mangalore	Mangalore	12 52 14	74 50 43	-0.261	+0.261	+0.048	-0.434	-2.8	-2.0	-3.1	-2.2	-3.1	+0.1
Bangalore Base-line, S.W. end	Bangalore	13 0 41	77 35 0	-0.007	+0.007	+0.003	-0.430	-5.3	+2.9	-6.7	-3.4	-6.7	-3.6
St. Thomas' Mount Trestle	Madras	13 0 15	80 11 41	+0.233	-0.233	-0.044	-0.420	-4.0	-7.2	-3.1	+2.4	-3.1	0.0
Kudankulam Observatory	Nagarkoil	8 10 22	77 41 27	+0.005	-0.007	0.000	-0.615	-7.7	+1.8	-8.4	-3.6	-8.4	-3.9
Sum of squares										360.23	203.83	359.88	292.30
Square root of mean square										4.74	3.56	4.74	4.27

The solution *A*, i.e. the most probable value of δa , δb , u_0 , w_0 , is

$$\left. \begin{aligned} \delta a &= 33.08 \text{ km} \\ \delta b &= 22.63 \text{ km} \\ u_0 &= +6''.10 \\ w_0 &= -7''.71 \end{aligned} \right\} \dots \dots \dots A$$

to which correspond the residuals under "Residual *A*" in the table. If 0.92363 km and 0.74273 km are substituted for δa and δb (*vide* Chapter I §8) the following most probable values of u_0 and w_0 are arrived at (solution *B*):—

$$\left. \begin{aligned} u_0 &= +1''.01 \\ w_0 &= -7''.28 \end{aligned} \right\} \dots \dots \dots B$$

and the corresponding residuals are shown in the table under heading "Residual *B*".

12. Solution *A* is obviously of no practical use as the values of δa and δb are much larger than it is possible could be correct. Solution *B* is not unreasonable. A southerly deflection at Kalianpur has previously been inferred, the estimated amount being 4". The value of w_0 indicates an easterly deflection of 16".2. The value computed from the topography, but taking no account of compensation is 10".7 (*vide* Prof. Paper 13, p. 116). The solutions however have been given more as illustrations of a principal than for their numerical values. The residuals show that the solution is not highly successful in satisfying the equations: yet the values of u_0 , w_0 derived from *B* are reasonable and the residuals might fairly be attributed to observation errors.

13. *The choice of a figure of reference for the geoid.* In surveying a surface such as the geoid, in the first place of unknown form, it is necessary at the outset to decide on some figure of reference to which measurements may be referred. This figure of reference may be of any form whatever—a particular case would be any set of three orthogonal planes. In traverse operations a single plane is chosen, on the assumption that for a limited area the geoid is not much different from a plane: and it would be possible to extend the application of the plane of reference by the introduction of a third coordinate, namely that at right angles to the plane. If at each triangulation station the direction of the geoidal vertical is determined by means of latitude and longitude observations the data is sufficient to enable the position of any point to be expressed by means of its three coordinates, quite independently of the shape of the geoid. The statement of the position of numerous points on the geoid in fact determines the shape of the geoid. If however the position of the several points are referred to a figure which approximates in shape and position to the geoid, the actual shape of the geoid is much more readily grasped by the small deviations it exhibits from the well known reference figure. The choice of such a figure is extremely useful and greatly decreases the labour of calculation of the positions of points on the geoid. The closer the approximation between it and the geoid the smaller are the quantities which express the difference of the two surfaces: and, as a result, when these quantities appear in formulæ as square and power terms they may be neglected in many of the computations which arise. It is important none the less to recognise that the two surfaces cannot always be treated as identical, and to examine each case thoroughly. Moreover it is to be borne in mind that the complexity of computation will be much increased if a very complex figure of reference is selected. A balance must be struck between the two considerations, and it has been customary to adopt a spheroid. This is at the same time a comparatively simple figure for computation and also a fairly close approximation to the geoid. This choice need not at all imply that no other geometrical figure can be found which approximates more closely to the geoid. Suppose even that the geoid was in actual fact an ellipsoid (not of revolution) not very different from a spheroid. It would still be strictly accurate to refer it to a spheroid of reference:

and it is probable that this would be the simplest course to follow in dealing with the results of any one survey, for instance the Indian Survey. Or the geoid might be referred with strict accuracy to a sphere: but in this case the residuals in a vertical direction might be inconveniently large.

14. This does not appear to have been quite the point of view usually taken, seeing that much energy has been devoted to finding the spheroid which best fits the whole earth. The origin of this research was doubtless the desire to uphold the Newtonian theory that the earth, being a revolving gravitating mass, should approximate in form to an oblate spheroid: rather than to the prolate spheroid which early measurements led the French school to believe in. This question was finally settled in favour of the Newtonian theory by the measurements of the arcs in Peru and Lapland: and the matters now to be investigated are the relatively minor deviations of the geoid from the oblate spheroid. Given a ready means of converting coordinates from one spheroid to another, each survey may properly select the spheroid most suitable to its own requirements. In any case the several large surveys of the world are expressed in terms of different spheroids, and for purposes of intercomparison it is necessary to develop a method of changing from one spheroid to another. An interesting question is to consider how closely the several spheroids, which best fit the respective surveys, agree *inter se*: to account for any differences: and to see whether a theory of density distribution can be found which will bring all these spheroids into agreement. The same question may be considered by taking the surveys on the spheroids they happen to have been reduced on and afterwards expressing the results in terms of a single spheroid and the local differences of the geoid from this general spheroid. Even if this general spheroid is so selected as to make the differences from the geoid a minimum it still remains only a convenient figure of reference and a more or less close approximation to the geoid.

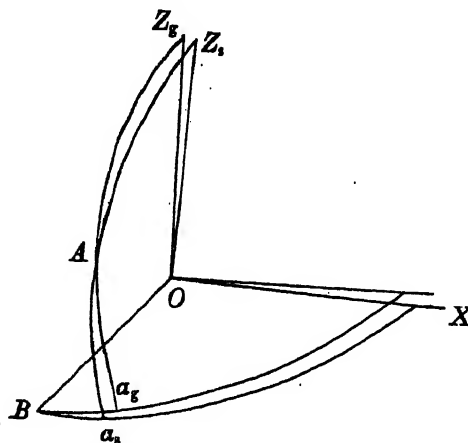
15. In the case of triangulation the usual procedure is as follows: horizontal angles are measured on the geoid, that is to say a theodolite is set up and levelled so that its horizontal circle is tangential to a level surface of the geoid. Spheroidal excess, calculated from the assumed spheroid, is applied to these angles. Further computations of the latitude and longitude of the points of triangulation are then carried out as though the spheroid and geoid were identical.

Now in certain disturbed districts the geoid is of considerably different curvature from the adopted spheroid: and the excess over 180° of the sum of the three angles of a triangle observed on the geoid is not the same as that computed from the spheroid. On account of the relative smallness of this excess in triangles of the size which occur in triangulation, this difference is not of great importance, though it gives rise to the two entirely different methods of Chapters III and IV. But if the rays observed have a considerable elevation, such as 5° , a very appreciable error is introduced, as will shortly be explained. It is necessary to be more precise. The most natural way of relating a point on the geoid to the spheroid is by giving the coordinates (latitude and longitude) of that point of the spheroid the normal—or more strictly for large distances the orthogonal confocal hyperbola—at which passes through the point on the geoid; and by stating the height of the geoid above the spheroid measured along this normal as well as the angle between this normal and the normal to the geoid (deflection of the plumb-line) and the azimuth of the plane containing the two normals.

Defining the position of a geoidal point in this way for the present, the separation of the geoid and spheroid need not be considered. To each point on the geoid there is a corresponding point on the spheroid: and consequently to each geoidal triangle a spheroidal triangle corresponds. It is with such spheroidal triangles that computations of latitude and longitude etc. really deal, the formulæ being deduced from properties of the spheroid. Consider then the relation between the angles of a geoidal triangle and the corresponding spheroidal triangle.

16. Suppose a theodolite is set up at a point O and levelled in the ordinary way: at this point two zeniths may be distinguished, Z_g that of the geoid and Z_s that of the spheroid, the former being indicated by the direction of the theodolite when the altitude is set to 90° .

The plane $Z_g O Z_s$ is the plane of deflection and the line OB at right angles to this plane is parallel to both spheroid and geoid and is chosen as axis of Y : so that OB may be regarded indifferently as belonging to the spheroid or the geoid. Consider another point A and draw the great circles $Z_s A a_s$ and $Z_g A a_g$. Then $O a_s$ and $O a_g$ are the traces of the ray OA on the spheroid and geoid respectively. Suppose that the horizontal angle between OA and OB is required. Observation by the theodolite gives the angle $a_g OB$: but the angle required for computation on the spheroid is $a_s OB$. Denote by a the geoidal angle of elevation of A and by z the azimuthal angle $X_g O a_g$, corresponding quantities for the spheroid being $a + \delta a$, $z + \delta z$. From triangles $Z_g BA$ and $Z_s BA$



$$\cos AB = \sin a \cos Z_g B - \cos a \sin Z_g B \cos(z_g - 90^\circ) = \sin(a + \delta a) \cos Z_s B - \cos(a + \delta a) \sin Z_s B \cos(z_s - 90^\circ)$$

But $Z_g B = Z_s B = 90^\circ$; hence

$$\cos a \sin z = \cos(a + \delta a) \sin(z + \delta z) \quad \dots \dots \dots (1)$$

i.e.

$$\tan a \cdot \delta a = \cot z \cdot \delta z \quad \dots \dots \dots (2)$$

neglecting second order terms.

Now δz , δa are the corrections which should be applied to geoidal quantities to correct them into spheroidal quantities and make them suitable for spheroidal formulæ. δa is $A a_s - A a_g$ and is approximately $\epsilon \cos z$, where ϵ is the total plumb-line deflection which is in the plane $O X_g Z_g$: so that (2) may be written

$$\delta z = \epsilon \tan a \sin z \quad \dots \dots \dots (3)$$

The correction δz accordingly is greatest when $z = 90^\circ$, that is when the observed object is in the plane of no deflection, and its magnitude in this case is $\epsilon \tan a$. Now values of ϵ up to one minute have been observed: and if at the same time a ray of elevation of 4° is observed, the horizontal angle may need a correction of 4 seconds—a very appreciable quantity in geodetic triangulation. The figures here given are roughly applicable to a ray through Jharipani (Dehra Dun district) where the deflection exceeds one minute and considerable angles of depression occur. In the case of a triangle at two of whose corners there is no deflection while a considerable deflection occurs at the third, a large triangular error will be apparent. More usually however the deflection is not so widely different at the three corners and the angular errors partially compensate one another in the sum, thus masking the error, but leaving the triangle distorted.

17. It is of interest to note that deflections have a corresponding effect on the measurements of base lines. Suppose that an element ds is measured along a line inclined at an angle a to the geoidal horizontal and $a + \delta a$ to the spheroidal vertical. Its reduced length is generally taken as $ds \cos a$, whereas reduced to the spheroid it is $ds \cos(a + \delta a)$, so that a correction

of $-ds \sin a \cdot \delta a$ is required. The error on the whole line is $\int ds \cdot \sin a \delta a$. This is equal to $s \sin a_m \delta a_m$ where a_m and δa_m are values which occur at some part of the line. If a is fairly constant $s \sin a$ is approximately the difference of level of the two ends of the base and the error is approximately $(h_2 - h_1) \delta a_m$. Owing to $h_2 - h_1$ being small compared with the length this is only liable to affect the length by a quantity of as much as 1 in 10^6 in extreme cases.

18. Deflections are usually stated in terms of their westerly and southerly components, ξ, η . It is clear that the effect of either component on a ray can be computed independently and then the two results combined. In the case of a ray of azimuth A it follows from (8) that a correction to the geoidal azimuth of amount δA is required where

$$\delta A = (-\xi \cos A + \eta \sin A) \tan a \quad \dots \quad (4)$$

Consider now the case of a traverse. Denote the successive points by 1, 2, 3 . . . n : let a_n be the angle of elevation of $n+1$ from n and let β_n be the elevation of $n-1$ from n . Also let A_n be the azimuth of $n, n+1$ and B_n that of $n, n-1$ and c_n be the arc subtended at centre of earth by $n, n+1$

$$\left. \begin{array}{l} \text{Then} \quad a_n + \beta_{n+1} = -c_n \\ \text{and} \quad A_n = \beta_{n+1} + 180^\circ - K_n \end{array} \right\} \quad \dots \quad (5)$$

where K_n is the convergency.

This traverse may be regarded as the flank of a series of triangulation: and in proceeding along it, the accumulation of azimuth error will be estimated. Now the flank of a triangulation series may, without much loss of generality be considered to proceed along a great circle of the earth (or a geodesic to be more precise). The great circles on the earth which are most conveniently considered are the meridians: but it is clear that by changing the system of coordinates to which points are referred any great circle may be regarded as a meridian of a different system of coordinates. It will accordingly be sufficient to consider the case of a meridian (not necessarily one of the system with the axes of rotation as pole). Along such a meridian the azimuthal angle A is zero or 180° . Suppose then that the traverse 1 2 3 . . . n lies on this meridian and that μ_n is the component of the plumb-line deflection at n in a direction perpendicular to this meridian (but not necessarily east and west as the meridian may be any great circle).

The correction to the angle at n will now be

$$\delta C_n = \mu_n (\tan a_n + \tan \beta_n) \quad \dots \quad (6)$$

which may be written sufficiently accurately for the present purpose

$$\delta C_n = \mu_n (a_n + \beta_n) \quad \dots \quad (7)$$

since a_n and β_n seldom if ever are so large as 5° in triangulation of a geodetic kind.

Let δh_n be the height of $n+1$ above n : then very approximately, if R is the radius of the earth

$$\delta h_n = R c_n (a_n + \frac{1}{2} c_n) = -R c_n (\beta_{n+1} + \frac{1}{2} c_n) \quad \dots \quad (8)$$

and

$$\delta C_n = \frac{\mu_n}{R} \left(\frac{\delta h_n}{c_n} - \frac{\delta h_{n-1}}{c_{n-1}} \right) - \frac{1}{2} \mu_n (c_{n-1} + c_n) \quad \dots \quad (9)$$

The accumulated azimuth error of the side $n, n+1$ is accordingly O_n where

$$\begin{aligned} C_n &= \frac{1}{R} \sum \frac{\delta h_n}{c_n} (\mu_n - \mu_{n+1}) - \frac{1}{2} \sum \mu_n (c_{n-1} + c_n) \quad \dots \quad (10) \\ &= {}_1 C_n + {}_2 C_n. \end{aligned}$$

Some attention to detail of the limits of this summation is necessary to obtain the precise value in a particular case: but the present object is to discuss the accumulation of the error and so this detail need not be considered now. It is clear that the first expression on the right hand side of (10) is not liable to great increase: for δh_n is equally likely to be positive or negative

as also is $\mu_n - \mu_{n+1}$ (considering that μ_n is the deflection at right angles to the line $n, n+1$). The most probable value of the expression ${}_1C_n$ is

$$\frac{\sqrt{n}}{R} \cdot \frac{\delta h}{c} (\mu - \mu')$$

where $\frac{\delta h}{c}$ and $\mu - \mu'$ are values intermediate to the extreme values met with. To get an idea of the magnitude which this might reasonably reach after 25 sides put $\frac{\delta h}{cR} = \frac{1}{50}$ corresponding to an angular elevation of more than 1° and $\mu - \mu' = 10''$. The value is then $5 \times \frac{1}{50} \times 10'' = 1''$. Now angular elevations of 1° are average, but changes of deflection of as much as $10''$ in the distance between two stations are not usual, though occasionally much bigger changes occur. It is felt then that the estimate of $1''$ for 25 rays is fair and that the danger of accumulation of error from this term is not considerable. The second term of (10) remains and its magnitude is liable to be somewhat greater. It may be written

$${}_2C_n = -\mu_m \Sigma c$$

where μ_m is some value intermediate to the extreme values of μ met with: Σc is merely the whole arc subtended by the terminal stations at the earth's centre. Taking $\Sigma c = \frac{1}{10}$ radian which corresponds to a series about 400 miles long we get

$${}_2C_n = -\frac{1}{10} \mu_m$$

In a series along the first range of the Himalayas deflections at right angles to this range of as much as $40''$ are of common occurrence. If $\mu_m = 40''$

$${}_2C_n = -4''$$

Now this is an error of magnitude about what might possibly occur in a single angle at which $\mu = 60''$ and $\tan \alpha = \frac{1}{15}$: so that the conclusion may be drawn that the danger of failing to correct observed angles for deflection of the plumb-line is almost confined to the angles themselves and is not liable to produce a cumulative error of azimuth, if the angles were utilised as in a traverse. The effect however may be felt in a way different from that considered above owing to the distribution of triangular error. Each triangle which contains a station where the deflection differs considerably from those at the other stations is liable to be deformed when the angles are adjusted to equal two right angles plus the spherical excess calculated on the spheroid. The amount of this deformation and the effect on computed coordinates of the stations of the triangulation do not appear to be such as can be estimated for a general case. Its effect in the actual triangulation of India is mixed up with the effect of error of observation and its amount is in general considerably less, as appears from the solution of the modified Laplace equations given above in §§ 10—12.

19. In observations for azimuth the result of using a point at considerable angular elevation about the station of observation as a reference point seems to have always been ignored. Yet the same geometrical fact which causes the horizontal trace of the ray through the pole to be displaced in azimuth also gives rise to an azimuthal deflection of the horizontal trace of the ray through the reference point, unless it happens that the ray is in the azimuth at right angles to that of geoidal deflection. Suppose that at any point the southerly and westerly deflections are η, ξ respectively. The horizontal (spheroidal) trace of the ray through the pole and geoidal zenith will be deflected in azimuth by $+\xi \tan \lambda$, λ being the altitude of the pole. The horizontal trace of a ray in azimuth A and angular elevation α will be deflected by

$$-\tan \alpha (\xi \cos A - \eta \sin A) \quad \text{vide (4)}$$

The difference between astronomic and geodetic azimuth is accordingly

$$\xi \tan \lambda + (\xi \cos A - \eta \sin A) \tan \alpha$$

instead of the simpler expression $\xi \tan \lambda$ usually taken, which is more closely approximated to as α decreases.

20. The advantages and disadvantages of the methods of correction worked out in Chapters I, III, IV may now be considered. The method of Chapter I in which u_y and v_x are taken as the changes of latitude and longitude has the justification and weaknesses referred to on pages 10-12. In a triangulation system where the bulk of the triangulation is along parallels and meridians this solution would be satisfactory were it not for the azimuths. The azimuth computed by the corrections at the ends of a ray of triangulation along a parallel differs by an appreciable amount from those for a ray along a meridian, and at first sight it appears that the difference is the necessary correction to the angle contained by these two rays. This however is not satisfactory as it is clear that the longitudinal and meridional series must be a little bent and that the whole error should not be forced into the junction angles. It would be equivalent to putting all the angular closing errors of a traverse which followed approximately the sides of a rectangle into the four angles at the corners of the rectangle. Moreover the final azimuth will not agree with the longitude as laid down in Laplace's equation. Laplace's equation might be adopted as a mode of determining the azimuth changes: but obviously the result would be inconsistent with the latitude and longitude changes found viz. u_y, v_x . In fact it appears that this method could rightly be applied merely to the junction points of the triangulation series. After changes for these points had been found, the corresponding changes along the series might be adjusted as is done in closing a traverse. This would involve a consideration in detail of all the series and has the disadvantage of being most laborious, and when done it is inconvenient in that the solution for the case of a further change of axes would have to be taken up right from the beginning. It might be supposed to be advantageous in that it takes cognisance of the actual form of the triangulation: but seeing that it is based on a method which is not entirely justifiable there seems to be little advantage in this partial approach to accuracy in the final stages of the reduction.

21. The method of computation along the geodesics at once gets rid of the difficulty of dual values of the changes. It is obvious however that the values obtained for the changes vary according as one origin or another is selected: for the closing errors in a triangle formed by joining any point and two selected origins exist just as much here as in the first method. This closing error however does not occur all at one point as in that method, but is satisfactorily distributed. There is some trouble in computing the geodesics: but this is of minor importance seeing that it has been done once for all and correction tables have been made out from which the coordinates of any point may be deduced by interpolation. These tables permit of the changes due to any desired changes in the elements a, b and latitude and azimuth at the origin being made immediately and admit of further changes being subsequently made when this becomes desirable.

Both this and the first method are based on the idea of the accuracy of the ratios of the sides in the triangulation: this is almost independent of small changes in the spheroid and the consequent minute changes in the spherical excess of any observed triangle. It may be noted here that in the case of an equilateral triangle with observed angles of equal weight the ratios are unaffected by the amount of spherical excess as this would be distributed equally. But the angles of the geoid have been used in place of spheroidal angles and from this some disturbance must have arisen.

While then the ratios of the sides may be regarded as practically perfect so far as corrections due to size of spheroid are concerned, it must at the same time be remembered that the observation errors have a cumulative effect on the ratio of a side to the original base as the side considered is separated more widely from the base: and the treatment of the observed angles as applicable to the spheroid without correction will aggravate this. The magnitude of the errors

so developed is indicated where closure has been made on additional base lines. It is clear that in a network of triangulation these additional bases may be reached from the origin by various routes and the length of these routes must accordingly be duly considered.

The following figures are taken from the circuits and base-lines of the N.W. Quadrilateral* in which there are 5 circuits and 4 measured bases.

TABLE XLIII.

(1) Number of equation	(2) Logarithmic closing error $\times 10^6$	(3) Number of triangles	(2) ² ÷ (3)
1	4.40	51	0.380
2	6.82	96	0.465
3	7.19	36	1.438
4	7.96	95	0.666
5	16.38	123	2.180
6	12.46	185	0.841
7	15.09	88	2.586
8	0.53	138	0.002
			Sum = 8.558
			Mean = 1.070
			Square root of Mean = 1.035

It appears that the mean error per triangle in side ratio is 1.035 in the 6th place of logs which corresponds to an error of one part in 420,000 showing that a high order of accuracy has been attained.

22. In the third method, that of geometrical transformation, the idea of the spheroid as merely a figure of reference is used as a basis for the argument. It is free from the difficulties of multiple values, one for each route traversed, and gives a definite set of values for the changes at any point. Being geometrically correct it naturally satisfies the Laplace condition: but it does not keep the constancy of side ratios, though the departure from constancy is not serious. No attempt is made to correct for the distance between geoid and spheroid which in conjunction with large deflections such as have so far been discovered would make very small changes in the coordinates. It is to be remembered however that the original geoidal triangles have been applied without angular correction to the old spheroid of reference, although considerable corrections must have been necessary in some cases. To put this matter right now, deflections at many stations would need to be observed: and to make use of the information that might be gained by observation, it would be necessary to re-grind the whole triangulation of India. It is the object of the present investigation to avoid this immense piece of work: but as has just been pointed out, it could not be undertaken until many deflection observations had been made. Had the corresponding corrections been made in the first instance, which would have been possible if comparatively rough latitudes and azimuths had been observed at each triangulation station, the method of change of coordinates explained in Chapter IV would have been absolutely correct. The fact that this was not done is the source of the present difficulty and

* Vide G.T. Volume II of the Survey of India, pages 303, 304.

practically disposes of the usefulness of this method. Further the Laplace equations should by right have been applied in the original grinding: they were not. This omission also makes an objection to any method of computing short of regrinding. And so the geometrical accuracy of the method of Chapter IV is vitiated. It is useless to insist on a method which strictly accords with Laplace's equations when the original quantities which are to be corrected fail to satisfy those equations.

23. As remarked above in §7 the reason of the multiple values of u , v , w according to traverse route followed is that in the computations no attempt has been made to correct observed geoidal angles to angles on the particular spheroid which is selected as a reference figure. Had these corrections been applied the method of geometrical change explained in Chapter IV would have given the changes u_r , v_r , w_r which would then have been applicable on changing from one spheroid of reference to another. But seeing that no such corrections were applied, and that the closing errors of circuits were dispersed and treated as errors of observations it is clear that this method is not strictly applicable. The portion of the closing errors due to this lack of correction to the angles is small compared with those due to errors of observation: so in the main no great fault was committed. A greater fault was the neglect of closing on the longitude arcs, or in other words applying Laplace's longitude equation. What is at first sight naturally regarded as a defect of the methods of Chapters I and III is that equation (4) of this chapter is not satisfied. But when it is considered that Laplace's condition, of which equation (4) is an immediate consequence, was not enforced on the computation of the original triangulation, it is clear that there is nothing to be gained by now enforcing equation (4) on to the small changes to be applied an account of change of spheroid. A preferable course is to make these changes and then apply Laplace's condition to the final result as Sir Sidney Burrard has done in his discussion of the Indian azimuth observations. It is concluded then that the two objections to the methods of Chapters I, III cited above have little weight in view of the slight inaccuracies of method by which the Indian triangulation has been reduced. It remains then to decide merely on what route should be followed in deducing the changes of coordinates by the method of Chapter I. This method is applicable to any route if the "closing errors" are applied as explained at the end of that chapter: and in Chapter III although the results are obtained in a special way, yet these results might have been obtained by the method of Chapter I. It is clear that no route can be laid down as rigorously correct and that the best that can be done is to select a route which appears to be the best. Suppose there were four triangulation series all of equal merit forming a square. Then the route which should be followed from one corner to another is the diagonal, and this produces a result intermediate to those which would be found by following either pair of sides. If one pair of sides was distinctly better triangulation than the other, the best route would doubtless be one closer to the good sides than the bad sides. But in the case of a great network of triangulation it is too complicated to go into such detail and so the diagonal would be selected. Now the geodesic corresponds fairly closely in the case of a spheroid to the diagonal of any square or rectangle, and it gives a satisfactory medial path among the triangulation and medial results for the changes deduced. This choice is slightly arbitrary, but seems the best that can be made. Another arbitrary choice is that of the origin from which changes are computed on the point from which all the selected geodesics radiate. It is apparent from the theory of the "closing errors" that different values would be deduced for the change according as this central point is selected; but the differences are not really appreciable in comparison with the errors due to faulty observation: Kalianpur is very centrally situated as regards India, and as it is the origin of the triangulation it appears that it would merely be an unnecessary complication to select a slightly different point for the point from which the geodesics radiate. It would be useless to go into any great refinement as to

the theoretically best centre for this purpose because it would be constantly disturbed as new triangulation is added to the Indian system.

It appears then that, without there being any rigidly accurate reason for adopting the method and results of Chapter III, yet that it meets the present case quite satisfactorily and has no defect of appreciable magnitude: and that any defect of theoretical precision would be present in any alternative method which might be proposed. The conclusion quoted at the end of Chapter IV is accordingly reiterated, namely that the method of calculation along geodesics through Kalianpur as set forth in Chapter III is the correct one to use.

CHAPTER VI.

Strength and Adjustment of Triangulation. Mechanical Analogy.

A criterion of strength of triangulation series.

1. If a mechanical network, which is analogous to a triangulation series in the sense explained in § 10 below, replaces each series in a system of triangulation a mechanical framework is formed. Each mesh in this framework corresponds to a circuit in the triangulation and needs straining to close in a way exactly analogous to the need of adjustment of triangulation circuits. Now in general in Indian triangulation, series follow approximately straight lines and these are generally more or less along meridians or parallels. Consider four series which form a circuit $A B C D$, and their mechanical analogues. To effect closing at A in the mechanical framework strains must be applied: and it seems fairly clear that the strains which will be caused in the side $A B$ will be of the same nature all along this side, but will differ essentially from those caused in $B C$. An example would be that the strains in $A B$ would be such as to increase the length $A B$ while those in $B C$ would be to slightly curve $B C$. On this account it appears desirable to consider strains of a particular type as existing throughout $A B$: but not existing to the same extent in $B C$. The side $A B$ is thought of as having uniform strength which differs from the uniform strength of $B C$. Reverting to the triangulation series it may be remarked that in one series the same strength is aimed at throughout, angles being observed with similar precision and figures of the same type selected as far as possible. When topographical conditions change entirely, as must essentially occur on the passage from plain to hilly country, the series should be considered in sections.

In considering one series of a circuit, it is only necessary to think of a "route" formed by those sides which persist in the general direction of the series, but bearing in mind that the length of each side is expressed in terms of the previous side, and that in any adjustment its relative length to this previous side is the quantity which is to be slightly varied; and similarly its azimuth is relative to the previous side. It will not be very far from the truth in effect if these several sides are for simplicity regarded as of equal length l and practically in the same direction. Suppose the angle between the r^{th} and $(r+1)^{\text{th}}$ side, originally practically 180° , is changed by the small angle η_r : and that the ratio of the length of the $(r+1)^{\text{th}}$ side to the r^{th} , originally unity, is changed to $1+\epsilon_r$. Expressed in terms of the base, the r^{th} side changes in length in the ratio $\prod_1^r (1+\epsilon) - 1 \doteq \sum_1^r \epsilon$ and in direction by $\sum_1^r \eta$. These quantities ϵ and η may be chosen to give any possible small changes in the length and azimuth of the terminal side of the series. Consider the displacements in the terminal

point $r+1$. That in the direction of the series is clearly

$$l\epsilon_1 + l(\epsilon_1 + \epsilon_2) + \dots + l(\epsilon_1 + \dots + \epsilon_r) = \frac{l}{r} (r\epsilon_1 + (r-1)\epsilon_2 + \dots)$$

where s is the length of the series and accordingly $s = rl$.

The most probable value of this is

$$\frac{s\epsilon}{r} \left(r^2 + (r-1)^2 + \dots \right)^{\frac{1}{2}} = \frac{s\epsilon}{r} \sqrt{\frac{r(r+1)(2r+1)}{6}} = s\epsilon \sqrt{\frac{r}{3}} \dots (1)$$

if r is large, ϵ being the most probable value of any of the quantities ϵ_r . Similarly the displacement of the terminal point r in the direction at right angles to the series is

$$l\eta_1 + l(\eta_1 + \eta_2) + \dots = s\eta \sqrt{\frac{r}{3}} \dots (2)$$

Both the quantities ϵ and η depend on the probable error of an angle (adjusted by the triangulation conditions) in the series. General Ferrero introduced the quantity " m " as a criterion of triangulation,

where

$$m = \sqrt{\frac{\sum \Delta^2}{3n}}$$

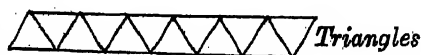
Δ being the triangular error of any triangle and n the number of triangles considered. This quantity " m " is accordingly the error of mean square of one angle of a triangle. The probable error of an observed angle is $\cdot 6745 m$. The probable error of an angle adjusted to satisfy the triangulation condition is

$$a = \sqrt{\frac{2}{3}} \times \cdot 6745 m = \cdot 551 m \dots (3)$$

The formulæ (1) and (2) may accordingly be expressed by saying that the probable displacement of the terminal point of a series of given length in any direction varies as m/\sqrt{l} .

They accordingly show the advantages of figures with long sides. It remains to consider the effect of various types of figures in the series, simple triangles, braced quadrilaterals, central pentagons, hexagons etc. Only regular figures are considered and the sides of each are taken equal to l . To complete a series of each type of given length s suppose there are n_3, n_4, n_5 etc. figures, simple triangles, braced quadrilaterals, pentagons, etc. It is clear that the gains in distance in the required direction are

$n_3 l \cos 60^\circ$	in the case of simple triangles
$n_4 l$	quadrilaterals
$n_5 l (\cos 18^\circ + \frac{1}{2} \cos 54^\circ)$	pentagons †
$2n_6 l \cos 30^\circ$	hexagons



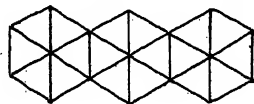
Triangles



Quadrilaterals



Pentagons



Hexagons

and as these must all be equal to s

$$\therefore \frac{1}{2} n_3 = n_4 = \frac{n_5}{1.245} = \frac{n_6}{1.732}$$

Now the probable errors in the determination of a terminal side after a given number of figures of each of the kind mentioned are in the ratio

$$0.82 : 1 : 1.21 : 1.29 *$$

so that at the end of each series of the same length the errors are in the ratio

$$0.82 \sqrt{n_3} : \sqrt{n_4} : 1.21 \sqrt{n_5} : 1.29 \sqrt{n_6}$$

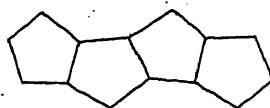
which reduce to

$$1.17 : 1 : 1.08 : 0.98$$

for the cases of triangles, quadrilaterals, pentagons and hexagons respectively.

* Vide Account of the Operations of the G. T. Survey of India, Vol. II, p. 199.

† An alternative arrangement of the pentagon



gives a result practically the same.

Suppose there is a series composed of α simple triangles, β braced quadrilaterals, γ pentagons and δ hexagons, then the ratio of its terminal probable errors to those of a series of the same length composed of quadrilaterals is

$$\sqrt{\frac{1.17^2\alpha + \beta + 1.08^2\gamma + .98^2\delta}{\alpha + \beta + \gamma + \delta}} : 1$$

which may be approximately written $1 + f : 1$

where

$$f = \frac{1}{12} \cdot \frac{2\alpha + \gamma}{\alpha + \beta + \gamma + \delta} \quad \dots \dots \dots (3)$$

Heptagons, nonagons etc. occur rarely and may be treated as pentagons. Octagons, decagons etc. may be treated as hexagons. Combining this result with (1) and (2) the quantity

$$M = (1 + f)^m \sqrt{\frac{18}{l}} \quad \dots \dots \dots (4)$$

is formed in which 18, the average length in miles of sides in the Indian triangulation, is introduced, and " m " is General Ferrero's expression for error of mean square of an angle and l is the average length of side expressed in miles in the series under consideration.

2. This quantity M takes cognizance not only of the probable error of the angles in the triangulation but also of the length of side and type of figure. For a given length of triangulation it gives a relative idea of the errors likely to occur in series of different precision and type: for example if there are several series of the same length, say 300 miles each, for which values $M_1, M_2, M_3 \dots$ have been found by (4), then the probable errors of northing or easting of the terminal point are approximately in ratio $M_1 : M_2 : M_3 \dots$ and the same is true of the probable errors of length or azimuth of the terminal side. " M " gives a criterion of the value of triangulation considering in proper proportion the excellence of observation and the success in choosing well-proportioned figures which has been attained: " m " only gauges the excellence of observation.

The deduction of the quantity M is confessedly based on approximations and simplifications. It would not be expected to be very accurate if applied to badly conditioned figures, and it is not intended that this should be done. In geodetic triangulation such figures are exceptional and figures approximately symmetrical largely predominate: and in these cases M is a practically useful criterion of the excellence or strength of the series.

3. All the triangulation of India has been classified according to values of M (*vide* table XLIV) and the order of merit of the several series deduced. The series are arranged in chronological order and designated by a serial number. Reference to any series can generally be made more conveniently by use of its serial number than by the rather long and frequently artificial names which have been applied. A consideration of the list shows that the principal and secondary triangulation ranges fairly continuously from very high class work in the best of which No. 76 North Baluchistan Series $m = 0.221$ and $M = 0.17$; to the least successful secondary triangulation No. 65 Siam Branch in which $m = 3.711$ and $M = 4.34$. The mean square (*vide* note at foot of table) value of M for the triangulation which was utilised in the grinding of the Indian network is 1.04: that for the whole triangulation 1.51. In some cases so called secondary triangulation proves better than poor principal triangulation: in general there is no marked gap between the two classes. This classification of triangulation into principal and secondary is accordingly dropped after the completion of the series and both are classed as "geodetic" triangulation and placed according to the values of M yielded by them. The further distinction in Indian triangulation is between "geodetic" and "minor" triangulation. The former is always rigorously computed taking account of spherical excess. The latter, which is generally very much rougher, disregards spherical excess.

TABLE XLIV.

Values of "m" and "M" for all Geodetic Series of the Indian Triangulation.

No	Name of Series	Seasons	$\pm m$	l	Number of Independent Figures												f	$\pm M$	Order of Merit
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
1	South Pārasnāth Mer.	1881-39	3.308	24.0	6														
2	Budhon Meridional ...	1883-43	2.242	19.2	25		1	2	4						2	.139	3.26	92	
3	Amūa Meridional ...	1884-38	1.647	18.9	34											.135	2.46	86	
																.167	1.88	77	
4	Rangir Meridional ...	1884-64	1.643	20.6	32*														
5	Calcutta Longitudinal	1884-69	0.369	26.6												.167	1.79	72	
6	Great Arc Meridional, Section 24°-30° ...	1885-66	0.708	22.2	10			4	2	2	2	1	1	1		.051	0.32	8t	
									2	2					1	4	.109	0.71	36b
7	Bombay Longitudinal	1837-63	0.844	27.6	1		1	2	2							2	.077	0.74	38
8	Great Arc Meridional, Section 18°-24° ...	1838-41	0.567	21.3	17		3	4			1								
9	Great Arc Meridional, Section 8°-18° ...	1840-74	0.390	23.7	1		4	2	5	3	1					3	.054	0.36	13b
10	Singī Meridional ...	1842-62	1.187	24.9	18		3									2	.131	1.14	51
11	South Konkan Coast	1842-67	2.176	29.6	16		3										.140	1.93	79
12	Karāra Meridional ...	1843-45	1.507	16.4	21		1, 1								1		.146	1.81	73
13	North Malūncha Mer.	1844-46	1.266	18.4	6				1	2							.130	1.42	60
14	Chendwār Meridional	1844-69	0.841	15.1	17			1	1								.146	1.06	45
15	Gora Meridional ...	1845-47	0.973	15.6	23										1	.161	1.21	52b	
16	Calcutta Meridional...	1845-48	1.173	8.5	45												.167	1.99	82
17	South Malūncha Mer.	1845-53	1.606	15.7	11		2										.141	1.97	81
18	Khānpisura Meridional	1845-62	1.227	27.3			4	1	1	2					1	.071	1.07	46	
19	Gurwāni Meridional...	1846-47	1.165	13.8	32												.167	1.55	65
20	North-East Lon. ...	1846-55	0.446	11.1	96		2	1	3		1						.156	0.65	31b
21	Hurilāong Meridional	1848-52	1.502	14.3	20				1						1	.142	1.92	78	
22	North-West Himalaya	1848-53	0.641	25.3			6	1	2						3	.021	0.55	26	
23	Gurhāgarh Meridional	1848-62	0.914	13.6	70			1	2		1				1	.157	1.21	52b	
24	East Coast ...	1848-63	0.608	16.9	22		6		3		1				2	.116	0.70	34b	
25	Karāchi Longitudinal	1849-53	0.558	15.8			10	2	10	2	2				1	.015	0.60	30	
26	Abu Meridional ...	1851-52	0.617	15.9	1				3							.042	0.68	33	
27	North Pārasnāth Mer.	1851-52	0.895	12.6	20											.167	1.25	54	
28	Kāthiawār Meridional	1852-56	0.990	17.4	7		3	1	1						3	.101	1.11	49	
29	Gujarāt Longitudinal	1852-62	0.859	14.2	31			2								.157	1.12	50	
30	Kāthiawār Minor Lon.	1853	1.481	23.1	1		5									.028	1.34	58	
31	Sābarmati Secondary	1853-54	1.348	5.3	15		2									.147	2.84	88	
32	Great Indus ...	1853-61	0.359	12.7			9	1	12	2	2		3		1	.008	0.43	20b	
33	Rahūn Meridional ...	1853-63	0.327	14.1			2		15	2						.009	0.37	15t	

Mer. = Meridional.

Lon. = Longitudinal.

* One centred A.

† Centred Quadrilateral.

TABLE XLIV.—(Contd.)

Values of “m” and “M” for all Geodetic Series of the Indian Triangulation.

No.	Name of Series	Seasons	±m	Z	Number of Independent Figures												f	±M	Order of Merit
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
34	Assam Longitudinal...	1854-60	0.579	12.5...		2	3	4...	1...							0.025	0.71	36b	
35	Cutch Coast ...	1855-58	0.986	12.5...		5	6...	1...							1	0.074	1.27	55b	
36	Kashmir Principal ...	1855-60	0.884	19.2...		18	1	1...								0.004	0.86	40	
37	Jogī-Tīla Meridional	1855-63	0.481	12.3...				7	1	1...						0.009	0.59	28b	
38	Sambalpur Lon. ...	1856-57	0.806	19.3	7	2		1...								0.117	0.87	41	
39	(Cutch) Coast Line Sec.	1856-60	0.975	10.8	33											0.167	1.47	61	
40	Kāthiāwār Minor																		
	Meridional No. 1 ...	1858-59	0.930	9.1	13	1										0.155	1.51	64	
41	Kāthiāwār Minor																		
	Meridional No. 2 ...	1859-60	1.247	12.1	14	1									1	0.145	1.75	70	
42	Kāthiāwār Minor																		
	Meridional No. 3 ...	1859-60	0.969	10.0	17	3	1									0.139	1.48	62	
43	Bider Longitudinal ...	1859-72	0.311	22.0...		1, 1	2	1...	3...						1	0.064	0.30	5b	
44	Eastern Frontier or ...																		
	Shillong Meridional	1860-64	0.409	13.2...		6	2	1	1...	1...			1			0.028	0.49	23b	
45	Sutlej Meridional ...	1861-63	0.346	10.6	50											0.167	0.53	25	
46	Madras Mer. and Coast	1861-68	0.426	21.6...		3	2	6	3	2...						0.026	0.40	19	
47	Kāthiāwār Minor																		
	Meridional No. 4 ...	1863-64	1.154	10.8	14										1	0.157	1.73	69	
48	East Calcutta Lon. ...	1863-69	0.379	10.7	32			2...								0.157	0.57	27	
49	Mangalore Meridional	1863-73	0.440	20.7	1	1	1	4	2...						2	0.096	0.45	22	
50	Kumaun and Garhwāl	1864-65	1.742	26.7	2	4		1...								0.048	1.50	63	
51	Nāsik Secondary ...	1864-65	2.033	10.5	26											0.167	3.12	91	
52	Burma Coast ...	1864-82	0.380	19.8	14	18	5	5	1	2...					4	0.078	0.39	18	
53	Jabalpur Meridional	1865-67	0.340	22.4...		2		7	1...							0.008	0.31	7	
54	Madras Longitudinal	1865-80	0.384	21.4	1	1	3	6...								0.038	0.37	15t	
55	Assam Valley Triangu-																		
	lation ...	1867-78	1.690	9.5	46	5, 3	1									0.141	2.65	87	
56	Brahmaputra Mer. ...	1868-74	0.564	12.0...			1	6...	1...						1	0.009	0.70	34b	
57	Coimbatore Minor No. 1	1869-71	1.547	13.6	19		1									0.163	2.07	83	
58	Bilāspur Meridional...	1869-73	0.302	15.7...		6	4	6...								0.021	0.33	11	
59	Cuddapah Minor ...	1871-72	0.826	17.6	8	1										0.148	0.96	43b	
60	Hyderabad Minor ...	1871-72	1.405	19.9	9											0.167	1.56	66	
61	Malabar Coast ...	1871, 74, 80	1.532	17.4	12											0.167	1.82	74	
62	Jodhpore Meridional	1873-76	0.291	15.6...		3	1	7	1...	1...						0.019	0.32	8t	
63	South East Coast ...	1875-79	0.522	11.7...		8		11	2...						1	0.007	0.65	31b	

Mer. = Meridional.

Lon. = Longitudinal.

Sec. = Secondary.

|| Centred Quadrilateral.

TABLE XLIV.—(Contd.)

Values of "m" and "M" for all Geodetic Series of the Indian Triangulation.

Series of the Indian Triangulation.

No.	Name of Series	Seasons	$\pm m$	z	Number of Independent Figures												f	$\pm M$	Order of Merit
					3-sided	4-sided	5-sided	6-sided	7-sided	8-sided	9-sided	10-sided	11-sided	12-sided	Compound				
64	Eastern Sindh Mer.	1876-81	0.244	13.0	...	3	2	5	2	0.28	0.30	56	
65	Siam Branch Triangu- lation	1878-81	3.711	16.1	7	4	1.107	4.34	94	
66	Mandalay Meridional	1889-95	0.418	27.0	...	13, 3	1.009	0.35	12	
67	Mong Hsat Secondary	1891-93	3.054	24.0	9	1	1.139	3.01	896	
68	Manipur Longitudinal	1894-99	0.453	28.4	...	5, 2	0.000	0.36	136	
69	Makran Longitudinal	1895-97	0.285	23.9	2	2, 1	1	0.069	0.26	4	
70	Mandalay Lon.	1899-1909	1.696	17.2	8	2	1	1.129	1.96	80	
71	Manipur Minor Mer.	1899-1902 } 1915-16 }	0.750	20.9	31	1	0.161	0.81	39	
72	Great Salween	1900-11	0.404	32.6	...	2, 4	2	1.056	0.32	86	
73	Kidarkanta Secondary	1902-03	1.323	15.2	3	1	1.125	1.62	67
74	Kalat Longitudinal	1904-08	0.365	39.7	...	6, 5	...	1	0.000	0.25	3	
75	Baluchistan Triangu- lation	1908-09	1.348	33.2	...	1	1.083	1.08	476	
76	North Baluchistan	1908-10	0.221	32.7	...	5, 3	1	0.009	0.17	1	
77	Gilgit	1909-11	0.443	31.6	5	3, 1	0.093	0.37	156	
78	Khāsi Hills Secondary	1909-11	2.038	10.7	14	3	1.137	3.01	896	
79	Mawmai Secondary	1909-11	1.575	10.9	41	1	1.163	2.35	85	
80	Upper Irrawaddy	1909-11	0.596	30.6	4	5	0.074	0.49	236	
81	Jaintia Hills Sec.	1910-11	0.986	6.9	23	1.167	1.86	76	
82	Bhīr Secondary	1911-12	0.794	17.4	24	1.167	0.94	42	
83	Ranchi Secondary	1911-12	1.840	15.2	13	1.167	2.34	84	
84	Villupuram Secondary	1911-12	1.184	10.9	18	1.167	1.78	71	
85	Sambalpur Meridional	1911-14	0.250	25.7	...	6, 4	1	1	0.007	0.21	2	
86	Indo-Russian Connec- tion	1912-13	2.790	10.9	11	7, 2	0.092	3.92	93	
87	Khandwa Secondary	1912-13	0.999	15.2	22	1.167	1.27	556	
88	Ashta Secondary	1913-15	1.048	15.3	21	1.167	1.33	57	
89	Buldāna Secondary	1913-14	0.304	12.3	18	1.167	0.43	206	
90	Naldrug Secondary	1913-14	1.465	15.2	27	1	1.161	1.85	75	
91	Nāga Hills Secondary	1913-14	0.913	21.3	7	1	1	1.139	0.96	436	
92	Middle Godaveri Sec.	1914-15	0.913	17.1	14	1	1.156	1.08	476	
93	Kohima Secondary	1914-15	1.094	15.0	13	1.167	1.39	59	
94	Cachar	1914-15	1.077	10.5	10	1.167	1.65	68	

Mer. = Meridional, Lon. = Longitudinal

Mer. = Meridional.

Lon. = Longitudinal.

Sec. = Secondary.

|| Centred quadrilateral.

For 42 Series entering the Simultaneous Grinding (shown in italics above)

$$\Sigma M^2 = 45.3591 \quad \text{Mean Square } M = \pm \sqrt{\frac{45.3591}{42}} = \pm 1.04$$

$$\text{For Series up to No. 94} \quad \Sigma M^2 = 214.3199 \quad \text{Mean Square } M = \pm \sqrt{\frac{214.3199}{94}} = \pm 1.51$$

4. Replace each triangulation series by one of its flanks. The network is then nearly similar to a traverse network, with the addition that closure of the length of the last side is necessary as well as its azimuth and the position of its terminal point. The flank of any series is usually not far from straight: or else consists of two or more portions with approximately straight flanks. Consider each such portion separately and denote it by the name "*triangulation line*."

Each triangulation line is liable to be slightly bent and to have its length slightly altered in the course of adjustment. This is effected by the angle at each station, and by the ratio of successive sides (between stations) of the triangulation line being slightly changed. Triangulation lines may be of different strengths according to the series from which they are derived: but it will be assumed that the strength of any one triangulation line is uniform. In other words, if it is necessary to adjust the azimuth at the end of a triangulation line this would be done correctly by giving the angles at all its stations an equal change: and to adjust the length of terminal side it would be correct to change the ratios of successive sides each by equal percentages.

When several triangulation lines are concerned the angular adjustment at any station of one will in general be different from that at any station of any other on account of both the different strengths of the several triangulation lines as well as their directions. The question of strengths has been considered in some detail above (*vide* § 1) and can be taken into account by means of *M*.

Adjustments of latitude and longitude at the end of a triangulation line may also be effected by a combination of small changes of the angles at the stations and the ratios of sides between stations: but in these cases the most probable adjustment would not be that of changing all the angles and the successive side ratios by equal amounts. The actual difference of this latter course from the most probable one is not very great in triangulation lines of moderate length, and it may be deemed justifiable on the ground of simplicity to make the adjustment by adopting the latter. This would bring the four types of adjustment into one simple scheme: but the more general case will now be explained and the simple case can easily be deduced from this if desired by omission of certain terms.

5. Consider any triangulation line and let the successive stations along the line be denoted by the numbers 0, 1, 2, . . . n , there being altogether n sides in the line. Let A_r be the azimuth at r of $r+1$ and λ_r , L_r the latitude and longitude of r . Denote by c_r the length of the r^{th} side and by $\Delta\lambda_r$, ΔL_r , ΔA_r the increments of latitude, longitude and azimuth along this side ($r-1$, r). Suppose that the angle at the station r is changed by η_r radians and the ratio of the $(r+1)^{\text{th}}$ to the r^{th} side to $\frac{c_{r+1}}{c_r} (1 + \epsilon_r)$ and consider what changes will be caused thereby.

The following expressions hold approximately

$$\left. \begin{aligned} \Delta\lambda_r &= -\frac{c_r}{a} \cos A_{r-1} \\ \Delta L_r &= -\frac{c_r}{a} \sin A_{r-1} \sec \lambda_{r-1} \\ \Delta A_r &= -\frac{c_r}{a} \sin A_{r-1} \tan \lambda_{r-1} \end{aligned} \right\} \dots \dots \dots (5)$$

in which $\Delta\lambda_r$, ΔL_r , ΔA_r are expressed in radians. The differences between ρ , ν the principal radii of curvature and a the mean radius are neglected as only approximate equations are required in what follows.

Differentiate (5) with regard to c_r , A_{r-1} , λ_{r-1} . Denoting the changes in latitude, longitude, back azimuth and forward azimuth at station r by u_r , v_r , w_r , $w_r + \eta_r$ it follows that the change in $\Delta\lambda_r$ is $u_r - u_{r-1}$ etc., so that

$$\left. \begin{aligned} u_r - u_{r-1} &= \Delta \lambda_r \left\{ \frac{\delta c_r}{c_r} - \tan A_{r-1} (w_{r-1} + \eta_{r-1}) \right\} \\ v_r - v_{r-1} &= \Delta L_r \left\{ \frac{\delta c_r}{c_r} + \cot A_{r-1} (w_{r-1} + \eta_{r-1}) + \tan \lambda_{r-1} u_{r-1} \right\} \\ w_r - w_{r-1} - \eta_{r-1} &= \Delta A_r \left\{ \frac{\delta c_r}{c_r} + \cot A_{r-1} (w_{r-1} + \eta_{r-1}) + \sec \lambda_{r-1} \operatorname{cosec} \lambda_{r-1} u_{r-1} \right\} \end{aligned} \right\} \dots (6)$$

In these equations $\cot A_{r-1}$ or $\tan A_{r-1}$ is liable to be inconveniently large: but this is always accompanied by either $\Delta \lambda_r$ or ΔL_r being correspondingly small. It is convenient to eliminate A_{r-1} by means of (5) and to write (6)

$$\left. \begin{aligned} u_r - u_{r-1} &= \Delta \lambda_r \frac{\delta c_r}{c_r} - \Delta L_r \cos \lambda_{r-1} (w_{r-1} + \eta_{r-1}) \\ v_r - v_{r-1} &= \Delta L_r \frac{\delta c_r}{c_r} + \Delta \lambda_r \sec \lambda_{r-1} (w_{r-1} + \eta_{r-1}) + \Delta L_r \tan \lambda_{r-1} u_{r-1} \\ w_r - w_{r-1} &= \Delta L_r \sin \lambda_{r-1} \frac{\delta c_r}{c_r} + \Delta \lambda_r \tan \lambda_{r-1} (w_{r-1} + \eta_{r-1}) + \Delta L_r \sec \lambda_{r-1} u_{r-1} \end{aligned} \right\} \dots (7)$$

In accordance with notation explained above, since $\frac{c_r}{c_0} = \frac{c_r}{c_{r-1}} \cdot \frac{c_{r-1}}{c_{r-2}} \dots \frac{c_1}{c_0}$ and $\frac{c_r}{c_{r-1}}$ is changed into $\frac{c_r}{c_{r-1}} (1 + \epsilon_{r-1})$: then $\frac{c_r}{c_0}$ is changed into $\frac{c_r}{c_0} \prod_0^{r-1} (1 + \epsilon_r)$. Hence $\frac{\delta c_r}{c_r} = E + \sum_0^{r-1} \epsilon$ where c_0 , which may be regarded as the last side previous to the side 01, is supposed to change to $c_0 (1 + E_0)$.

Now suppose that the changes in successive side ratios and angles are in arithmetic progression: i.e.

$$\begin{aligned} \epsilon_0 &= \epsilon_1 - \epsilon' = \epsilon_2 - 2\epsilon' = \dots = \epsilon_r - r\epsilon' = \epsilon \\ \eta_0 &= \eta_1 - \eta' = \eta_2 - 2\eta' = \dots = \eta_r - r\eta' = \eta \end{aligned}$$

and

Then

$$\frac{\delta c_r}{c_r} = E + r\epsilon + \frac{r(r-1)}{2} \epsilon'$$

E and H being the quantities relating to the side from which the triangulation line emanates. Equations (7) may now be written

$$\left. \begin{aligned} u''_r - u''_{r-1} &= \left(E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} l'' \Delta \lambda_r - (w''_{r-1} + \eta + (r-1)\eta') \operatorname{cosec} \lambda_{r-1} \Delta L_r \\ v''_r - v''_{r-1} &= \left(E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} l'' \Delta L_r + (w''_{r-1} + \eta + (r-1)\eta') \sec \lambda_{r-1} \Delta \lambda_r \\ &\quad + u''_{r-1} \tan \lambda_{r-1} \Delta L_r \dots \\ w''_r - w''_{r-1} - \eta - \overline{r-1} \eta' &= \left(E + r\epsilon + \frac{r(r-1)}{2} \epsilon' \right) \operatorname{cosec} l'' \sin \lambda_{r-1} \Delta L_r \\ &\quad + (w''_{r-1} + \eta + (r-1)\eta') \tan \lambda_{r-1} \Delta \lambda_r + u''_{r-1} \sec \lambda_{r-1} \Delta L_r \dots \end{aligned} \right\} \dots (8)$$

in which u, v, w, η are now expressed in seconds and $\Delta \lambda, \Delta L$ are in radians, and $w_0 = H$.

To apply these equations it is necessary to know the values of $E_0, \epsilon, \eta'', u''_0, w''_0$ and v''_0 . The last quantity is simply additive to all values of v . The remaining five quantities give rise to five cases: for they can be considered separately and the results combined afterwards, since second order quantities are being neglected. By successive application of (8) the solution may be obtained in the form

$$\left. \begin{aligned} u_r &= (A_u E + B_u \epsilon + G_u \epsilon') \operatorname{cosec} l'' + C_u \eta + F_u \eta' + D_u u_0 + K_u w_0 \\ v_r - v_0 &= (A_v E + B_v \epsilon + G_v \epsilon') \operatorname{cosec} l'' + C_v \eta + F_v \eta' + D_v u_0 + K_v w_0 \\ w_r &= (A_w E + B_w \epsilon + G_w \epsilon') \operatorname{cosec} l'' + C_w \eta + F_w \eta' + D_w u_0 + K_w w_0 \end{aligned} \right\} \dots (9)$$

The coefficients A, B etc. . . have to be determined for each triangulation line and then the latitude, longitude and azimuth changes are expressed by (9). The r^{th} side is changed in the ratio $1 + E + r\epsilon + \frac{r(r-1)}{2}\epsilon' : 1$; so that the solution is complete.

It is convenient to denote the changes at the end of any, the m^{th} , triangulation line in latitude, longitude, azimuth and side by ${}_mU \quad {}_mV \quad {}_mW \quad {}_mE$: then taking account of different values of ϵ, η which occur in different triangulation lines the change in latitude, longitude and azimuth of the r^{th} station of the m^{th} line may be written

$$\left. \begin{aligned} {}_mU_r &= ({}_mA_{m-1}E + {}_mB_m\epsilon_m + {}_mG_m\epsilon'_m) \operatorname{cosec} 1'' + {}_mC_m\eta_m + {}_mD_mu_0 + {}_mK_mw_0 \\ {}_mV_r - {}_{m-1}V &= ({}_mA_{m-1}E + {}_mB_m\epsilon_m + {}_mG_m\epsilon'_m) \operatorname{cosec} 1'' + {}_mC_m\eta_m + {}_mD_mu_0 + {}_mK_mw_0 \\ {}_mW_r &= ({}_mA_{m-1}E + {}_mB_m\epsilon_m + {}_mG_m\epsilon'_m) \operatorname{cosec} 1'' + {}_mC_m\eta_m + {}_mD_mu_0 + {}_mK_mw_0 \\ {}_mE_r &= {}_{m-1}E + r\epsilon_m + \frac{r(r-1)}{2}\epsilon'_m \end{aligned} \right\} \quad (10)$$

Equations (10) give in the most general form the changes that are effected by introducing the alterations $\eta, \eta', \epsilon, \epsilon'$ different for each triangulation line.

6. We may now consider the probable relative values of ϵ and η (the latter expressed in radians) for this purpose treating ϵ' and η' as zero. Take the case of a series of simple equilateral triangles and let $p-2, p-1, p$ be three successive stations on a flank: also let x_p, y_p, z_p etc. be changes which are applied to the several angles, as indicated in the figure. The ratio of the successive flank sides $Q R/P Q$ being denoted by r , then

$$d \log r = \cot 60^\circ (x_{p-1} + x'_{p-1} + y_p - y'_{p-1} - z'_{p-1} - z_p)$$

Hence

$$\begin{aligned} \Sigma d \log r &= d \log \Pi r \\ &= \cot 60^\circ \left\{ \sum_0^{p-1} (x_r + x'_r - z_{r-1} - z'_{r-1}) + y_p - y'_0 \right\} \\ &\doteq p\epsilon. \end{aligned}$$

Hence the most probable way of getting a particular value for the change in logarithm of the side is by making all the x^s equal to each other and of opposite sign to all the z^s which will also all be equal. The y^s do not come into the case at all except at the two ends, unless an azimuth change is also required.

For the azimuth it is clear that

$$\Sigma \eta_{p-1} = \Sigma (x_p + y'_{p-1} + z_{p-1}) = p\eta$$

is the azimuth change, since the η^s are all to be equal.

Now in the most probable distribution of changes obviously all the x^s are equal as are the y^s and z^s . Hence

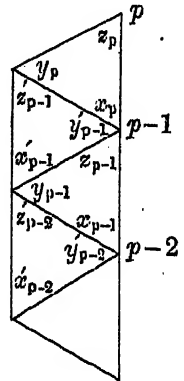
$$p\epsilon = \frac{p}{\sqrt{3}} \left\{ x + x' - z - z' \right\} + \frac{1}{\sqrt{3}} (y - y')$$

$$\text{and} \quad p\eta = p(x + y' + z).$$

$$\text{Now} \quad x + z = -y$$

$$\therefore \quad \eta = y' - y$$

$$\text{and} \quad \epsilon \doteq \frac{1}{\sqrt{3}} \left(x + x' - z - z' \right) \left. \vphantom{\frac{1}{\sqrt{3}}} \right\} \dots \dots \dots (11)$$



The probable values of η and ϵ are accordingly in the ratio of the most probable values of $\frac{y\sqrt{3}}{x-z}$ in which the quantities are subject to the relation

$$x + y + z = 0$$

Following the usual plan of independent multipliers explained below (Chapter VII § 4) we have for the two cases (a) of y and (b) of $x - z$

$$a_1 = a_2 = a_3 = 1$$

$$b_1 = b_2 = b_3 = 0$$

etc.

$$u_1 = u_2 = u_3 = 1$$

and

\therefore

$$3k_1 = 1 \text{ in case (a); } 3k_1 = 0 \text{ in case (b)}$$

$$u_F = 1 - \frac{1}{3} = \frac{2}{3} \text{ in case (a); } u_F = 2 - 0 = 2 \text{ in case (b)}$$

Hence the ratio of probable values of η and ϵ is unity, *i.e.*, η and ϵ have the same weight.

7. Now it is clear that in the adjustment of azimuth and side the most probable solution is obtained by introducing equal values of η , ϵ and by putting $\epsilon' = \eta' = 0$, all round a circuit composed of triangulation lines of equal strength: and if the triangulation lines are of different strengths the value of η , ϵ in any one is inversely proportional to the strength. With latitude and longitude it is clear that the directions of the traverse lines are of essential importance. Thus in the case of a triangulation line along a meridian a change in ϵ will give changes in latitude but no change in longitude. Moreover to get any desired change the most probable values of changes in angle and side ratio would not be equal at all the stations along a triangulation line: though it is probable that they would not alter much on one triangulation line if the circuit was composed of several such lines. Taking the case of a triangulation line along a meridian it is clear that to obtain a given change in longitude the change in the angles at the starting end of the line are more effective than equal changes in the angles near the closing point. If the sides were of equal length the most probable changes would be in arithmetical progression. The complexity of different changes at the various stations of a triangulation line on account of this has been deemed in general to overbalance the slight gain in theoretical accuracy of adjustment and at least in some preliminary adjustments which are about to be made (*e.g.* the incomplete Burma triangulation) the plan will be followed of making η and ϵ uniform along one triangulation line and putting $\eta' = \epsilon' = 0$. When the probable errors of position etc. are considered (*vide* Chapter VII.) it is believed that this plan will be considered fully satisfactory for many cases of geodetic application.

The equations (10) however show how the variation in changes along a triangulation line may be taken into account: and in closing a single triangulation line between two previously fixed sides they give the necessary number of quantities (four) at choice to satisfy the closing condition. It is clear from the theorem stated at the end of § 13 that for the general case of closure the changes at the several stations of a triangulation line should be in arithmetical progression this being a combination of the most probable adjustment firstly for log. side and azimuth and secondly for latitude and longitude. When the number of η 's and ϵ 's at choice is in considerable excess of the number of conditions to be satisfied it is believed that little gain in accuracy is obtained by introducing η 's and ϵ 's: and certainly this doubles the number of unknowns and greatly increases the labour of formation and solution of the normal equations. It also increases the complexity of the solution, and its subsequent application.

8. In the case of a network of circuits including Laplace stations and extra base lines, equations of form of (10) may be formed; and by equating the right hand sides to the several closing

errors which arise, four equations are formed for each circuit together with one extra for each extra base line or Laplace station. These may be used to determine the most probable values of η , ϵ (and η' , ϵ' if it is thought desirable not to put these equal to zero for each triangulation line), due regard being paid to the strength of each triangulation line.

If $m = \sqrt{\frac{\sum \Delta^2}{3n}}$ where Δ is the triangular error of any one of the triangles of a series, then the probable value of η in the case of a series of simple triangles in which the triangular error has been dispersed is, by (11), $\sqrt{\frac{2}{3}} m \sqrt{2} \times .6745 = .779m$. Account may be taken of the series comprising quadrilaterals and other figures by the introduction of a factor $\frac{1+f}{1+\frac{1}{6}}$ (see § 1 above) so that the probable values of η and ϵ (which are equal) are both equal to

$$.779m \frac{1+f}{1+\frac{1}{6}} = 0.668m (1+f) = a \dots \dots \dots (12)$$

in which m is supposed to be expressed in radians.

The equations for η and ϵ accordingly have to be solved subject to the condition that the sum of the squares of the corrections multiplied by their weights is a minimum. In one triangulation line the sum of the squares of the corrections is

$$\begin{aligned} \Sigma(\eta^2 + \epsilon^2) + \Sigma(\eta'^2 + \epsilon'^2) r^2 &= r(\eta^2 + \epsilon^2) + \frac{r(r+1)(2r+1)}{6} (\eta'^2 + \epsilon'^2) \\ &= r(\eta^2 + \epsilon^2) + \frac{r^3}{3} (\eta'^2 + \epsilon'^2) \end{aligned}$$

when r is the number of sides in a triangulation line.

Hence the condition to be satisfied is that

$$\Sigma \left[\left\{ r(\eta^2 + \epsilon^2) + \frac{r^3}{3} (\eta'^2 + \epsilon'^2) \right\} + m^2 (1+f)^2 \right] = \text{a minimum} \dots \dots \dots (13)$$

the summation extending over all the triangulation lines.

Mechanical Analogy

9. Suppose that there is a network of trilateration, in which the lengths of all sides have been determined by direct measurement. Let l be the measured length of any side, $l + \delta l$ the adjusted value, and w the weight of the determination l . In making any adjustment of the network, for example to bring a terminal side into agreement with a line slightly differently placed, the principle of least squares demands that $\Sigma w \delta l^2$ shall be a minimum, all imposed conditions being satisfied.

Now consider a similar framework formed by rods of material obeying Hooke's law of extension and compression, and suppose these rods are freely jointed at their junctions. If this framework is in equilibrium without any strains in action, and, the first side being held fast, the last side is brought into a slightly different position and its length slightly changed, then the several rods will undergo compression or extension and their lengths will be slightly altered. If the unstrained length of any rod is l and its strained length is $l + \delta l$; and the force in action in it causing this strain is F : then the work done on the rod is $\frac{1}{2} F \delta l$. Now by Hooke's law

$$\frac{\delta l}{l} = \frac{F}{aE}$$

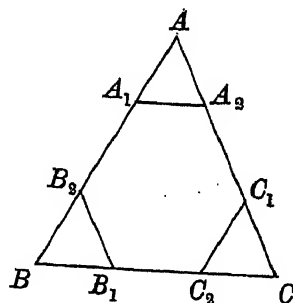
where E is Young's modulus and a is the cross section.

Hence

$$\frac{1}{2} F \delta l = \frac{aE}{2l} \cdot \delta l^2$$

and this represents the work done on the rod. The quantity $\frac{aE}{2l}$ may be varied by suitably choosing a : suppose it is made equal to w . Then the work done on the rod is $w \delta l^2$. The principle of least work immediately shows that for the set of strains applied to bring the last side into the desired position the total work done must be a minimum: whence $\sum w \delta l^2$ is a minimum. The solution accordingly is the same as that of the most probable adjustment of the similar trilateration network.

10. Now trilateration has never been carried out on a large scale for one obvious reason that no large tract of country is suitable for its execution by ordinary methods. Triangulation on the other hand extends over vast tracts and it is its adjustment which is of importance to geodesists. A mechanical analogy can be supposed for triangulation also. It is less simple than that just described for trilateration and to give it practical shape would be a matter of greater difficulty. Imagine a set of rods, the medial parts of which are laterally rigid, but which are longitudinally extensible without the application of (appreciable) force*. Let these rods be freely jointed to form a framework similar to a network of triangulation, the angles of the triangulation being maintained by rigid pieces. For example in the $\triangle ABC$ the rod AB is freely extensible between A_1 and B_2 : a cross piece $A_1 A_2$ maintains the angle A at its proper value and so on. It will be seen then that such a triangle may be enlarged to any size but that it will always remain similar, unless the pieces $A_1 A_2, B_1 B_2, C_1 C_2$ are changed in length. If these pieces $A_1 A_2$ etc. are made of material obeying Hooke's law, by properly choosing their cross section it is possible to make them represent the "strength" of the angles A, B, C . When the system is deformed the work done on $A_1 A_2$ will be proportional to the square of the change in the angle A and accordingly a condition of form $\sum w \delta \theta^2 = \text{minimum}$ must be satisfied so that the total work done shall be a minimum. All geometrical conditions such as triangular conditions, central station conditions and side ratio conditions obviously cannot be avoided in the mechanical analogy, so that the solution of the mechanical problem is precisely the same as that of the triangulation adjustment according to the method of least squares.



11. It is clear that if the change in any observed quantity can be made to correspond to the extension of a rod which obeys Hooke's law; and if a system of such rods are linked up in such a way as to represent the geometrical conditions controlling the observations; then a mechanical analogy for the set of observations is obtainable. The governing fact is that the work done on any rod is proportional to the square of its extension: so that by substituting extension for error of observation the equation of minimum squares is transformed into the principle of least work.

12. Consider a framework representing a network of triangulation and suppose that it is held fast at one or more points. It may be necessary to bring a terminal side into agreement with a predetermined value and position. To do this four conditions have to be satisfied:

- (1) The terminal side must be adjusted to the correct length.
- (2) The terminal side must be adjusted to the correct azimuth.
- (3) One extremity of the terminal side must be moved to the correct latitude.
- (4) This extremity of the terminal side must be moved to the correct longitude.

Now these adjustments may be considered one by one. First strain the terminal side to the correct length and hold it so: then change its azimuth, etc. The adjustments may also be performed

* Approximations to these can readily be conceived, e.g. a rod sliding in a tube.

in any order and the final result is the same. This is immediately obvious mechanically, for it is clear that the final configuration due to small strains has nothing to do with the order in which they are applied.

13. The analogy thus proves an important theorem† in the adjustment of observations, namely that *provided all imposed conditions are maintained, the adjustment conditions may be introduced separately in any order, the previous adjustment conditions in each case being maintained; and the most probable complete adjustment is obtained after the last adjustment condition has been applied.* This enables the circuit adjustments of triangulation to be applied after the figural adjustments, as has been done in the Survey of India. A further theorem is also easily deducible. In the case of the closing of a simple triangulation circuit in which there are four closing conditions to satisfy (*i.e.* the case in which there are no additional base lines or independent azimuth determinations) denote the four closing quantities by X, Y, Z, U . To effect the closing X alone in the most probable manner, changes are concurrently introduced which affect the other quantities by amounts y_x, z_x, u_x : and similarly for the other quantities. From the mechanical analogy it is clear that adjustments as follows should be made:—

$$\begin{aligned} (1) \quad & x \quad y_x \quad z_x \quad u_x \\ (2) \quad & x_y \quad y \quad z_y \quad u_y \\ (3) \quad & x_z \quad y_z \quad z \quad u_z \\ (4) \quad & x_u \quad y_u \quad z_u \quad u \end{aligned}$$

in which

$$\left. \begin{aligned} x + x_y + x_z + x_u &= X \\ y_x + y + y_z + y_u &= Y \\ z_x + x_y + z + z_u &= Z \\ u_x + u_y + u_z + u &= U \end{aligned} \right\} \dots \dots \dots (14)$$

and the quantities with suffixes are geometrically related to the suffix quantity, *i.e.*

$$\left. \begin{aligned} \frac{y_x}{B_x} &= \frac{z_x}{C_x} = \frac{u_x}{D_x} = x \\ \frac{x_y}{A_y} &= \frac{z_y}{C_y} = \frac{u_y}{D_y} = y \\ &\text{etc.} \end{aligned} \right\} \dots \dots \dots (15)$$

where the coefficients $A B C D$ depend only on the form of the triangulation and are independent of the closing errors. It accordingly follows that

$$\left. \begin{aligned} x + A_y y + A_z z + A_u u &= X \\ B_x x + y + B_z z + B_u u &= Y \\ C_x x + C_y y + z + C_u u &= Z \\ D_x x + D_y y + D_z z + u &= U \end{aligned} \right\} \dots \dots \dots (16)$$

and by solving these the following equations may be obtained

$$\left. \begin{aligned} x &= a_x X + a_y Y + a_z Z + a_u U \\ y &= b_x X + b_y Y + b_z Z + b_u U \\ z &= c_x X + c_y Y + c_z Z + c_u U \\ u &= d_x X + d_y Y + d_z Z + d_u U \end{aligned} \right\} \dots \dots \dots (17)$$

If the adjustments x, y, z, u are applied independently and then combined the total effect will be the same as that of the single adjustment X, Y, Z, U taken simultaneously. By use of the quantities x, y, z, u in place of the related quantities X, Y, Z, U it is accordingly possible to treat each closure entirely independently of the remaining three.

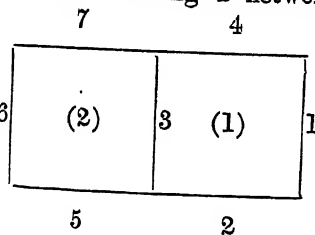
† This theorem was proved analytically in "Account of the Operations of the G.T. Survey of India, Vol. II," Appendix No. 8, pp. 151-158.

If in addition to the 4 ordinary closing quantities additional conditions such as extra bases, or fixings of latitude, longitude or azimuth are introduced, this only makes the relation more involved: it is still possible to express x, y, z, u in terms of the closing errors and proceed as though each of these quantities was independent of the other.

The principle is perfectly general and is applicable to a whole network of triangulation. As this becomes more complex the determination of the coefficients in the equations which correspond to (17) would become more difficult: but this is not necessary at least in some applications of the theorem. It is an important fact that an undetermined portion of each of the closures may be regarded quite independently of the others. The theorem may be stated as follows:—*It is possible to find quantities related to the several closing errors such that each type may be adjusted separately and independently of the others, and such that the combined effects of these several adjustments will be the same as the most probable simultaneous adjustment.*

These related quantities may be called the "independent errors" of each type. If each is adjusted independently of errors of another type, the other type adjustments being allowed to come in just as they will naturally do while the adjustment of the first type is made independently in the most probable manner, then the combined result of the adjustment of the four types will be the most probable adjustment of the closing errors which can be made.

14. Consider now in more detail the case of more than one circuit forming a network which has to be simultaneously adjusted. The case of two circuits which have a common portion is illustrative of this and will be seen to lead to a result generally applicable. Further these circuits may be considered as formed each for four series and no loss of generality occurs in taking these circuits to be of the same strength. The several series may be characterised by the numbers 1—7 and the circuits by (1) and (2).



The adjustments along any side may be made by the introduction of ϵ and η changes: and any one of the four types of closures the ϵ^s and η^s along a triangulation line will be in arithmetical progression (*vide* § 16). For the r^{th} side of the k^{th} line their values may be represented by

$$\epsilon_k + \overline{r-1} \epsilon'_k, \quad \eta_k + \overline{r-1} \eta'_k$$

Consider first the X closure and suppose that ${}_1x, {}_2x$ are the "independent errors" of the two circuits: these quantities have to be determined. Then equations of the following form may be formed:

$$\left. \begin{aligned} \sum_1^4 (A_k \epsilon_k + B_k \epsilon'_k + C_k \eta_k + D_k \eta'_k) &= {}_1x \\ \sum_{3,5,6,7} (A_k \epsilon_k + B_k \epsilon'_k + C_k \eta_k + D_k \eta'_k) &= {}_2x \end{aligned} \right\} \dots \dots \dots (18)$$

in which the coefficients $A B C D$ are independent of the closing errors. The prefixes to x are the circuit numbers.

To obtain the most probable solution the quantities ϵ, η must be chosen so as to make

$$\sum_{k=1}^{k=7} \left\{ \sum_{r=1}^{r=n} (\epsilon_k + \overline{r-1} \epsilon'_k)^2 + \sum_{r=1}^{r=n} (\eta_k + \overline{r-1} \eta'_k)^2 \right\} = \text{minimum.}$$

The solution of this accordingly gives definite value for all the ϵ^s and η^s as linear functions of ${}_1x$ and ${}_2x$. The associated quantities are given by equations of form

$${}_1y_x = {}_1\beta_x {}_1x + {}_1\beta'_x {}_2x$$

$${}_1z_x = {}_1\gamma_x {}_1x + {}_1\gamma'_x {}_2x$$

$${}_1u_x = {}_1\delta_x {}_1x + {}_1\delta'_x {}_2x$$

$${}_2y_x = {}_2\beta_x {}_1x + {}_2\beta'_x {}_2x \quad \text{etc.}$$

Similarly for the other closure

$${}_1x_y = {}_1a_y {}_1y + {}_1a'_y {}_2y \quad \text{etc.}$$

Equations similar to (14) can now be formed for each circuit

$${}_1x + ({}_1a_y {}_1y + {}_1a'_y {}_2y) + ({}_1a_z {}_1z + {}_1a'_z {}_2z) + ({}_1a_u {}_1u + {}_1a'_u {}_2u) = {}_1X \quad \text{etc.} \quad (19)$$

From (19) it is clear that the determination of the "independent errors" ${}_1x$, ${}_2x$ etc. from the known closing conditions can be effected by the solution of simultaneous linear equations of the same number as there are conditions. These equations are soluble without the introduction of the principles of least squares. Their solution however will entail much the same labour as the solution of the normal equations which would arise in the simultaneous adjustment of all the four types of closure. It appears at present to be only of theoretical interest that these closures or rather the adjustment of the related independent errors may be effected, each type independently of the others: for no material reduction in computation would be effected. However it is believed that the mechanical analogy throws some light on the question, and that developments are likely to result from its consideration.

15. The idea of a triangulation line has been introduced in § 4. It is the flank of a nearly straight series of triangulation. For the present, disregard the ϵ changes and fix attention on the η changes. Suppose a set of rigid rods are placed similarly to the several rays of the flank and that these are freely jointed at their ends, which accordingly correspond to the stations on the triangulation line. Introduce constraints at each junction which tend to maintain the angles between successive rods equal to the observed values and such that the force necessary to alter any one of these angles by a stated amount is inversely proportional to the probable error of the angle or directly proportional to the *strength* of the angle. Then it is clear that the work done in varying the angles in any way is proportional to the sum of the weighted squares of the changes of these angles. It is clear then that to bring the system into any given displaced formation the angular changes introduced are such as to make either the total work done, or the sum of the weighted squares of the adjustment in the angles, a minimum. That is to say the mechanical deformation is the same as the most probable adjustment. It is clear from this why when the angular strength is given the strength of the triangulation to resist either angular deflection at the end, or linear deflection, or a combination of the two, is the greater according as the number of stations in the line is less, or, in other words, according as the length of side is greater. The quantity M already introduced takes account of this (see Chapter VII § 1) and for certain considerations of probable error makes it unnecessary to consider the triangulation line in any detail.

The ϵ changes in a triangulation line (due to the extension of its sides) are precisely analogous to the η changes, if for angular deflection, change in ratio of final to initial side is taken, and for linear deflection at right angles to the line, linear deflection in the direction of the line is substituted. In the case of a single triangulation line (which is a straight line) the η changes are

independent of the ϵ changes and the corresponding adjustments can be performed independently: so both could be separately considered by means of this partial mechanical analogy. When however there is a set of triangulation lines in various directions the η and ϵ changes (northing and easting) of the several lines get intermixed and for this a complete analogy is required.

It is clear that for the case of trilateration it would only be necessary to arrange a set of *extensible* rods, jointed as described above, and representing a flank of the trilateration to obtain a complete analogy. But in the case of triangulation it is necessary to arrange that if any element changes length the successive elements change length in the same proportion without any force being involved. A simple mechanical analogy completely representing a triangulation line has not yet been discovered: and for this case it at present appears necessary to consider the analogy of the complete series instead of only one of its flanks (vide § 10). For purposes of probable errors it would be possible to replace the series by a simple series composed of equilateral triangles, in the same way that in the triangulation line it may be a convenient simplification to take all the sides of equal length and persisting in the same direction: and it appears from the mechanical analogy that little accuracy would be lost by so doing.

16. It is clear from the partial analogy given in the preceding section that the best adjustment to make in a triangulation line to obtain a given azimuth change is to change all the angles by amounts proportional to their probable errors: for this corresponds to the mechanical set of rods of which the first is held fast and to the last of which a suitable couple is applied. To obtain a given deflection the most probable adjustment is that the angular changes divided by their probable errors should be in arithmetical progression; as would be the case in the mechanical system, when held at one end and subjected to a simple force at right angles to the line at the other end. The exact analogy between the ϵ^s and η^s shows that similar conditions hold for the ϵ^s . The above statements can be simplified for the case of a triangulation line if the weights of the several angles are considered equal. It appears that a determination of the weights of the angles may best be obtained by a consideration of all the triangular errors, which leads to one value for the probable value of any angle of the series. Any determination of probable error of each angle separately is very much vitiated by the graduation error of the instrument, which is unknown. The simplified statements for a triangulation line with all sides of the same length are:—

- (a) for a given angular deflection or a given change of ratio of final side to initial side, that the η^s or ϵ^s are all equal.
- (b) for a given linear deflection at right angles to or along the line, that the η^s or ϵ^s are in arithmetical progression.

On this account the cases of the η^s and ϵ^s changing in arithmetical progression (which includes (a) as a special case, the constant difference then being zero) have been considered in equations (8) to (10). Some departure from rigid accuracy occurs when the several sides of the triangulation line are of unequal lengths.

CHAPTER VII.

Probable errors of triangulation before and after adjustment.

1. Expressions will now be formed for the probable errors of points and sides generated in one or more series of triangulation, in which only figural conditions have been adjusted.

(1) *Probable errors in logarithm and azimuth of terminal side.*

Equation (12) of chapter VI gives a the probable value of either η or ϵ . The probable error in the logarithm of the terminal side after n sides of a triangulation line is clearly $\sqrt{n} \log_{10}(1+\epsilon) = .4343a \sqrt{n}$; and in azimuth is $a \sqrt{n}$. Considering the triangulation line as practically straight and the distances between stations as equal then, if l is the average length of side and s the length of the line, so that $nl = s$,

$$\begin{aligned} a \sqrt{n} &= .668m(1+f) \sqrt{n} = m(1+f) \sqrt{\frac{18}{l}} \cdot \sqrt{\frac{nl}{18}} \times .668 \\ &= .1575 M \sqrt{s} \end{aligned} \quad (1)$$

Hence if M is expressed (as is always done) in seconds of arc

$$\text{Probable error in azimuth at end of a triangulation line} = 0''.1575 M \sqrt{s}$$

$$\begin{aligned} \text{Probable error in log. side at end of a triangulation line} &= .4343 \sin 1'' \times 0.1575 M \sqrt{s} \\ &= 3.32 \times 10^{-7} M \sqrt{s} \end{aligned}$$

For the case of a number of triangulation lines it is necessary to substitute $\sqrt{\Sigma M^2 s}$ for $M \sqrt{s}$. It is convenient to measure lengths of triangulation lines in units of 100 miles: so replace s by $100 S$ where the unit of S is 100 miles and then

$$\left. \begin{aligned} \text{Probable error in azimuth of the terminal} \\ \text{side of a series of triangulation lines} &= 1''.575 \sqrt{\Sigma M^2 S} \\ \text{Probable error in seventh place of logarithm} \\ \text{of the terminal side of a series of} \\ \text{triangulation lines} \dots &= 33.2 \sqrt{\Sigma M^2 S} \end{aligned} \right\} \dots \dots \dots (2)$$

In the above S is measured along the triangulation, and it is immaterial whether this is straight or not: but if the elements of the summation indicated by Σ are straight, then S may be replaced by L the length of any triangulation line, bringing formulæ (2) into similar terms to those of (5) below.

(2) *Probable errors in easting and northing of terminal points.*

Consider any curve defined by s the distance measured along it and ϕ the angle the tangent makes with OX . Let this curve be divided into elements of length l . Suppose that for purposes of adjustment, or on account of errors, the ratio of the $(m+1)^{\text{th}}$ element to the m^{th} is changed by factor $1 + \epsilon_m$ and the angle at the junction of these two elements by η_m . Then if $P P'$ is the $(m+1)^{\text{th}}$ element the relative shift of P' to P is

$$l \cos \phi_m \sum_1^{m+1} \epsilon - l \sin \phi_m \sum_1^{m+1} \eta \quad \text{in easting}$$

$$\text{and} \quad l \sin \phi_m \sum_1^{m+1} \epsilon + l \cos \phi_m \sum_1^{m+1} \eta \quad \text{in northing}$$

and the total change relative to O of N is given by

$$\Delta x = l \sum_1^n \left(\cos \phi_m \sum_1^{m+1} \epsilon \right) - l \sum_1^n \left(\sin \phi_m \sum_1^{m+1} \eta \right)$$

$$\Delta y = l \sum_1^n \left(\sin \phi_m \sum_1^{m+1} \epsilon \right) + l \sum_1^n \left(\cos \phi_m \sum_1^{m+1} \eta \right)$$

Hence

$$\begin{aligned} \Delta x &= l \epsilon_1 \sum_1^n \cos \phi + l \epsilon_2 \sum_2^n \cos \phi + l \epsilon_3 \sum_3^n \cos \phi + \dots - l \eta_1 \sum_1^n \sin \phi - \dots \\ &= \epsilon_1 (x_n - x_0) + \epsilon_2 (x_n - x_1) + \epsilon_3 (x_n - x_2) \dots - \eta_1 (y_n - y_0) - \dots \end{aligned}$$

since

$$\cos \phi = \frac{x' - x}{l}, \quad \text{and} \quad \sin \phi = \frac{y' - y}{l}$$

The most probable value of Δx is accordingly

$$\left[\epsilon^2 \left\{ (x_n - x_0)^2 + (x_n - x_1)^2 + \dots \right\} + \eta^2 \left\{ (y_n - y_0)^2 + (y_n - y_1)^2 + \dots \right\} \right]^{\frac{1}{2}}$$

and since the probable values of ϵ and η are both equal to a this reduces to $a \sqrt{\sum r^2}$ where r is the radius vector measured from the point N . The probable value of Δy is the same. When the elements are increased in number $\sum r^2$ may be replaced by $\frac{1}{l} \int r^2 ds$: also by (12) of chapter VI and (1)

$$a = .1575 \text{ M sin } 1'' \sqrt{\frac{s}{n}} = .1575 \text{ M sin } 1'' \sqrt{l}$$

and $a \sqrt{\sum r^2}$ becomes equal to $.1575 \text{ M sin } 1'' \sqrt{\int r^2 ds}$

Hence the probable value of Δx or $\Delta y = .1575 \text{ M sin } 1'' \sqrt{\int r^2 ds} \dots \dots \dots (3)$

So far only a single triangulation line has been considered. If there are several it is clear that it is necessary to alter this expression to $.1575 \text{ sin } 1'' \sqrt{\sum M^2 \int r^2 ds}$. This is expressed in units of a mile. To express it in feet multiply by 5280. It is convenient to measure r and s in units of 100 miles: denote their values in units of 100 miles by R and S . Finally if P (feet) is the probable error at a point N in easting or northing, and R is the radius vector measured from N , and S is the distance measured along the triangulation, both R and S being expressed in units of 100 miles, then

$$P = .1575 \times 5280 \times 10^3 \sin 1'' \sqrt{\sum M^2 \int R^2 dS}$$

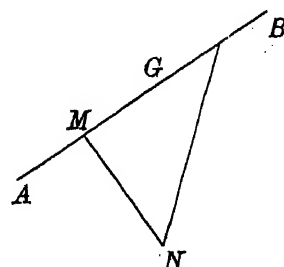
$$= 4.03 \sqrt{\sum M^2 \int R^2 dS} \quad \dots \dots \dots (4)$$

The integral $\int R^2 dS$ may be taken out for each triangulation line (assumed to be a straight line). If AB be one of these and S is measured from M the foot of the perpendicular from N on AB and $NM = p$, then $R^2 = S^2 + p^2$

$$\therefore \int R^2 dS = \left[\frac{1}{3} S^3 + S p^2 \right]_{MA}^{MB} = AB \left(\frac{1}{3} AB^2 + p^2 + MA \cdot MB \right)$$

Now if G is the middle point of AB

$$MA \cdot MB = - \left(\frac{L}{2} + MG \right) \left(\frac{L}{2} - MG \right) = - \frac{L^2}{4} + MG^2$$



$$\text{Hence } \int R^2 dS = L \left(\frac{1}{3} L^2 + p^2 - \frac{1}{4} L^2 + MG^2 \right) = L \left(R_0^2 + \frac{L^2}{12} \right)$$

where R_0 is the radius vector to the middle point of AB and L is the length AB . Therefore for a series of triangulation lines

$$P = 4.03 \sqrt{\sum M^2 L \left(R_0^2 + \frac{L^2}{12} \right)} \quad \dots \dots \dots (5)$$

in which the quantities L, R_0 may be measured off a chart in units of 100 miles. From either (4) or (5) it is clear that the probable closing error in northing or easting is different according as different points of a circuit are selected on which to close and from which to start.

2. Formulæ (2), (5) show that the probable errors in azimuth and logarithm of terminal side increase as the square root of the length of the several triangulation lines involved, while those of easting and northing increase at a much more rapid rate namely as the three halves power of the lengths, the triangulation lines remaining similar and similarly situated. On account of this latter fact it is desirable to have more frequent checks to prevent accumulation of errors than would be necessary if only length and azimuth of side were required. These checks may be obtained by measurement of bases and by forming Laplace stations, that is stations whose longitudes are observed telegraphically and at which astronomical azimuths are also observed. These two checks are of precisely equal importance; and applying only one of them does not serve a very useful purpose. In the Indian triangulation eight base lines have been measured to date (1916), excluding the short Mergui base in Burma, and these have been made use of in the adjustment of the triangulation. The longitude arcs were not available when (previous to 1879) the main adjustment was carried out. They do not all admit of the formation of Laplace equations, as the longitude stations are not coincident with the triangulation stations, nor can they be connected with satisfactory accuracy (as regards azimuth) in all cases. Only latterly* (1906) have they been applied to control the azimuth observations with a view to determining corrected plumb-line deflections in the prime vertical. This application does not improve the probable error in easting and northing of the points concerned or any other points of the triangulation.

3. It is necessary for the full consideration of the problem to find expressions for the probable errors after certain further adjustments, *viz.* closing on extra base lines, closing of circuits or (what has not been done in India) closing on Laplace stations, have been effected. Before treating

* Account of the Operations of the G. T. Survey of India, Vol. XVIII, Appendix 5.

this question quite generally it may be of interest to consider a special case. Suppose a series of triangulation lines closes between two bases. The probable error in easting or northing at any point of it after the adjustment has been performed is required. It is to be noticed in (4) that each portion of the triangulation is fully taken account of by the portion of the integral $\int M^2 R^2 dS$ which applies to it: that is to say, this portion of the integral represents the errors of displacement generated in the corresponding part of the triangulation, as if carried on to the closing point by perfect triangulation.

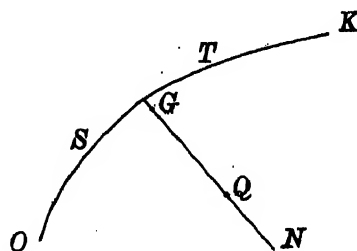
Let O, K be the two bases and suppose ST is the portion whose error as generated at N is required.

As in § 1 the expression for Δx for ST may be written

$$\Delta x = \sum_i^t (\epsilon_r x_r - \eta_r y_r)$$

The adjustment under consideration does not affect the η terms. The condition of closing gives (assuming uniform strength along OK)

$$\sum_1^k \epsilon_r = \text{a known quantity}$$



and in adjusting for this ϵ_r is replaced by $\epsilon_r - \frac{1}{k} \sum_1^k \epsilon_r$.

Denote the adjusted value of Δx by ${}_a\Delta x$. Then

$$\begin{aligned} {}_a\Delta x &= \sum_i^t \left\{ x_r \left(\epsilon_r - \frac{1}{k} \sum_1^k \epsilon_r \right) - \eta_r y_r \right\} \\ &= \sum_i^t \epsilon_r x_r - \frac{t-s}{k} \sum_i^t \sum_1^k \epsilon_r - \sum_i^t \eta_r y_r \end{aligned}$$

where \sum_i^t is the x -coordinate of the centre of gravity of ST . Hence

$${}_a\Delta x = - \frac{t-s}{k} \sum_i^t \left(\sum_1^{s-1} \epsilon + \sum_{t+1}^k \epsilon \right) + \sum_i^t \epsilon_r \left(x_r - \frac{t-s}{k} \sum_i^t X \right) - \sum_i^t \eta_r y_r$$

Hence the probable error of displacement at N , after OK has been adjusted, due to the portion ST is the square root of

$$({}_a\Delta x)^2 + ({}_a\Delta y)^2 = (k-t+s) \left(\frac{t-s}{k} \right)^2 \epsilon^2 NG^2 + \epsilon^2 \sum_i^t R_Q^2 + \eta^2 \sum_i^t R_N^2$$

where G is the centre of gravity of ST and Q lies on NG and $QN/GN = (t-s)/k$: also R_Q and R_N are radius vectors measured from Q and N respectively. Denoting the resultant probable displacement by D , this relation may be written, putting $OK = S_0$, $ST = S$ (measured along the curve)

$$\begin{aligned} D^2 &= k-t+s \left(\frac{S}{S_0} \right)^2 NG^2 \epsilon^2 + \epsilon^2 \sum_i^t R_Q^2 + \eta^2 \sum_i^t R_N^2 \\ &= \epsilon^2 \left\{ n N Q^2 + \sum_i^t R_Q^2 \right\} + \eta^2 \sum_i^t R_N^2 \dots \dots \dots (6) \end{aligned}$$

where $n = k - t + s$ is the number of sides in the line ST . Hence

$$D^2 = a^2 n \left\{ NQ^2 + \frac{1}{S} \int R_Q^2 dS + \frac{1}{S} \int R_N^2 dS \right\}$$

If the portion ST is a single triangulation line this becomes

$$D^2 = a^2 n \left\{ NQ^2 + \rho_0^2 + \frac{L^2}{12} + R_0^2 + \frac{L^2}{12} \right\}$$

where ρ_0 , R_0 are the distances of the centre of gravity of ST from N and Q respectively and S now becomes equal to L . Putting in numerical quantities and expressing the results in feet it follows, as in (4), that

$$D = 4.03 M \sqrt{L \left(R_0^2 + \rho_0^2 + \frac{L^2}{6} + NQ^2 \right)}$$

Finally if ST is composed of a number of triangulation lines, the probable displacement in any direction

$$P = \frac{D}{\sqrt{2}} = 2.85 \sqrt{\Sigma M^2 L \left(R_0^2 + \rho_0^2 + \frac{L^2}{6} + NQ^2 \right)} \quad \dots \quad (7)$$

where the points Q differ for each triangulation line and are found for each as has been done above for ST alone.

If closure had also been effected on Laplace station as well as on base lines at O and K this clearly would become

$$P = 4.03 \sqrt{\Sigma M^2 L \left(\rho_0^2 + \frac{L^2}{12} + NQ^2 \right)} \quad \dots \quad (8)$$

Equations (7) and (8) illustrate the statement made in §2 that closure only on a base and not on a Laplace station does not improve the results nearly so much as closure of both kinds simultaneously will probably do.

Probable errors after adjustment.

4. Consider any function F of quantities $x_1, x_2, x_3 \dots$ which have been found by measurement. If the true values of these quantities are $x_1 + v_1, x_2 + v_2, \dots$ and the true value of the function is $F + dF$ then

$$\begin{aligned} dF &= F(x_1 + v_1, x_2 + v_2, \dots) - F(x_1, x_2, \dots) \\ &\doteq f_1 v_1 + f_2 v_2 + \dots \end{aligned} \quad (9)$$

where

$$f_1 = \frac{dF}{dx_1}, f_2 = \frac{dF}{dx_2}, \text{etc.} \quad \dots$$

Any conditions which may be imposed on x_1, x_2, \dots result in equations of form

$$\left. \begin{aligned} a_1 v_1 + a_2 v_2 + \dots + a_n v_n - l_1 &= 0 \\ b_1 v_1 + b_2 v_2 + \dots + b_n v_n - l_2 &= 0 \\ \text{etc.} \end{aligned} \right\} \quad \dots \quad (10)$$

Then dF may be written

$$dF = f_1 v_1 + f_2 v_2 \quad \dots \quad -k_1 (a_1 v_1 + a_2 v_2 + \dots - l_1) - k_2 (b_1 v_1 + b_2 v_2 + \dots - l_2) - \dots$$

in which any values may be assigned to $k_1, k_2 \dots$. Hence

$$dF = (f_1 - k_1 a_1 - k_2 b_1 - \dots) v_1 + (f_2 - k_1 a_2 - k_2 b_2 - \dots) v_2 + \dots + k_1 l_1 + k_2 l_2 + \dots \quad (11)$$

Let the reciprocal weights of $x_1, x_2 \dots$ be $u_1, u_2 \dots$ and let the reciprocal weight of F be u_F . Then it follows that

$$u_F = (f_1 - k_1 a_1 - k_2 b_1 - \dots)^2 u_1 + (f_2 - k_1 a_2 - k_2 b_2 - \dots)^2 u_2 + \dots \quad (12)$$

If $v_1, v_2 \dots$ are the most probable values of the corrections, then u_F must be a minimum, and accordingly k_1, k_2, \dots are to be determined from the following equations

$$\left. \begin{aligned} [u a a] k_1 + [u a b] k_2 + \dots &= [u a f] \\ [u b a] k_1 + [u b b] k_2 + \dots &= [u b f] \end{aligned} \right\} \dots \quad (13)$$

in which $[u a b]$ is written for $u_1 a_1 b_1 + u_2 a_2 b_2 + \dots$, etc.

Reverting to (12) and developing the squares it is clear that

$$\begin{aligned} u_F &= [u f f] - 2 [u a f] k_1 - 2 [u b f] k_2 - \dots \\ &\quad + k_1 ([u a a] k_1 + [u a b] k_2 + \dots) + k_2 ([u a b] k_1 + [u b b] k_2 + \dots) + \dots \\ &= [u f f] - [u a f] k_1 - [u b f] k_2 - \dots \end{aligned} \quad (14)$$

by using the equations (13)*.

5. In considering the probable errors of triangulation it will be permissible to treat the quantities ϵ and η as errors in observed quantities, and as being independent except in so far as the several closing conditions relate them. Distinguish the several triangulation lines which make up a network of triangulation by the prefixes 1, 2, 3, etc., and the several stations on any triangulation by suffixes 1, 2, 3, etc. It will also be permissible in an investigation into the probable errors after adjustment to replace the triangulation by its projection on a plane, the projection of the general map of India naturally being selected for this purpose. This enables the closing conditions to be written in the following form, either for closure round a circuit or for closure between base lines or Laplace stations:

$$\left. \begin{aligned} \sum \sum r \epsilon_s &= 0 && \text{side closure} \\ \sum \sum r \eta_s &= 0 && \text{azimuth closure} \\ \sum \sum r (x \epsilon - y \eta)_s &= 0 && \text{easting closure} \\ \sum \sum r (y \epsilon + x \eta)_s &= 0 && \text{northing closure} \end{aligned} \right\} \dots \quad (15)$$

the first summation for S corresponding to the number of sides in a triangulation line and the second for r corresponding to the number of triangulation lines: and x, y being the coordinates of any station referred to the closing point of the particular circuit as origin.

* This deduction is taken from p. 229 of "*A Treatise on the Adjustment of Observations*" by T. W. Wright, New York, 1884.

There may be any number of each type of closures, and not necessarily the same number for each type. The triangulation lines over which the summations are taken differ in part for each closure of any type. Thus in the case supposed in Chapter VI §14 the line 3 occurs in both circuits. The quantities whose probable errors are required are of the same form as the left hand sides of equations (15): but the limits are different.

It is clear that the coefficients in (13) may be written, agreeably with the notation already adopted

$$\left. \begin{aligned} [u a a] &= \sum_r [u a a] \\ [u a b] &= \sum_r [u a b] \\ \text{etc.} \end{aligned} \right\} \dots \dots \dots (16)$$

and accordingly the portion relating to each triangulation line may be separately computed. [For the side and azimuth closures the coefficients a , b are all unity: for the other two closures they are x or y . Consider the case of a triangulation line which forms part of a closed circuit, so that there are closures of each type. It may also form part of a closure between base lines and Laplace stations. If so the coefficients of the several ϵ 's and η 's remain the same as in corresponding type of closure in the circuit, and the coefficients reduce to type $[u a a]$. For the case of clearness take the specific case shown in the diagram when there are base lines and Laplace stations at A and B .

Consider the line 2. It enters into the following relations:

	A	1	2	3	B
a	<div style="display: flex; justify-content: space-around; width: 100%;"> 6 4 </div>				
b	<div style="display: flex; justify-content: space-around; width: 100%;"> </div>				
c	<div style="display: flex; justify-content: space-around; width: 100%;"> </div>				
d	<div style="display: flex; justify-content: space-around; width: 100%;"> </div>				
e	<div style="display: flex; justify-content: space-around; width: 100%;"> </div>				
f	<div style="display: flex; justify-content: space-around; width: 100%;"> </div>				

The symbolic coefficients $a b c d e f$ are each written opposite one of these equations.

Along any triangulation line for u write a^2 , a being the probable value of ϵ or η . Suppose that n is the number of sides in the triangulation line, all considered of equal length. Then omitting the prefix 2 for simplicity and considering only the line 2

$$\begin{aligned} [u a a] &= n a^2 \\ [u a b] &= n a^2 \\ [u a c] &= 0 \\ [u a d] &= 0 \\ [u a e] &= a^2 \sum x = n a^2 X \\ [u a f] &= a^2 \sum y = n a^2 Y \end{aligned}$$

where X , Y are the coordinates of the centre of gravity of the line 2. These are typical of all the combinations of $a b c d$ with $a b c d e f$. The remaining typical coefficients are represented by

$$[u e e] = [u f f] = a^2 \sum (x^2 + y^2)$$

If x_1, x_2 are the limits of x and $x_1 = X - x_0$ and $x_2 = X + x_0$

$$\begin{aligned}\Sigma x^2 &= \frac{n}{x_2 - x_1} \int x^2 dx \\ &= \frac{n}{3} (x_2^3 - x_1^3) \\ &= \frac{n}{3} \left((X+x_0)^3 - (X-x_0)^3 \right) \\ &= n \left(X^2 + \frac{1}{3} x_0^2 \right) \dots \dots \dots (17)\end{aligned}$$

Similarly

$$\Sigma y^2 = n \left(Y^2 + \frac{1}{3} y_0^2 \right)$$

Hence

$$\Sigma (x^2 + y^2) = n \left(R^2 + \frac{1}{3} r_0^2 \right) = n \left(R^2 + \frac{L^2}{12} \right) \dots \dots \dots (18)$$

Finally

$$[uee] = na^2 \left(R^2 + \frac{L^2}{12} \right)$$

Also

$$[uef] = a^2 \Sigma (xy - yx) = 0$$

The complete result is given in tabular form:—

Values of ${}_2[uaa] \div na^2$, etc.

	<i>a</i>	<i>b</i>	<i>c</i>	<i>d</i>	<i>e</i>	<i>f</i>		
Side	<i>a</i>	1	1	0	0	<i>X</i>	<i>Y</i>	Base
Side	<i>b</i>	1	1	0	0	<i>X</i>	<i>Y</i>	
Azimuth	<i>c</i>	0	0	1	1	$-Y$	<i>X</i>	Laplace
Azimuth	<i>d</i>	0	0	1	1	$-Y$	<i>X</i>	
Easting	<i>e</i>	<i>X</i>	<i>X</i>	$-Y$	$-Y$	$R^2 + \frac{L^2}{12}$	0	
Northing	<i>f</i>	<i>Y</i>	<i>Y</i>	<i>X</i>	<i>X</i>	0	$R^2 + \frac{L^2}{12}$	

It is clear from this that there are really only four types of quantities, and that a closure on an outer circuit or base gives rise to the same coefficient as the corresponding closure within the circuit. All the quantities of the form ${}_2[uaa]$ are given by the following schemes, in which the four closures, side azimuth, easting, northing are represented by S, A, E, N.

Values of ${}_2[uaa] \div na^2$

	S	A	E	N
S	1	0	<i>X</i>	<i>Y</i>
A	0	1	$-Y$	<i>X</i>
E	<i>X</i>	$-Y$	$R^2 + \frac{L^2}{12}$	0
N	<i>Y</i>	<i>X</i>	0	$R^2 + \frac{L^2}{12}$

..... (19)

The value of na^2 is given by (1) in terms of M and S , or M and L if straight portions are considered separately.

The above scheme deals with the cases in which the coordinates x, y which occur all refer to the same origin. It only remains to take the case where two origins occur and to form the typical expression $[uaf]$. The coefficients a , etc. are either $0, 1, x$ or $\pm y$ in which x, y are the coordinates of any point on the line referred to the point of closure for which the corresponding condition of closure was formed. When it is desired to find the probable error with regard to any point O (for example Kalianpur, the origin of the survey), then the coefficients f , etc. are either $0, 1, x_0$ or $\pm y_0$ where x_0, y_0 are coordinates referred to this point. If the circuit closing point is P let x_p, y_p be the coordinates of P with regard to origin O then

$$x_0 = x_p + x \quad y_0 = y_p + y.$$

The values of $[uaf] + na^2$ are now indicated in tabular form similar to (19). As only the ratios of the coefficients in (13) are required na^2 may be replaced by M^2L by means of (1).

	S_f	A_f	E_f	N_f
S	1	0	$X + x_p = X_0$	$Y + y_p = Y_0$
A	0	1	$-Y - y_p = -Y_0$	$X + x_p = X_0$
E	X	$-Y$	$R^2 + \frac{L^2}{12} + x_p X + y_p Y$	$-x_p Y + y_p X$
N	Y	X	$x_p Y - y_p X$	$R^2 + \frac{L^2}{12} + x_p X + y_p Y$

. (20)

All the quantities X, Y, R are measured from P while x_p, y_p are coordinates of P and X_0, Y_0 are the coordinates of the mid point of the triangulation line relative to O .

6. The method explained in §§ 4,5 will now be applied to the determination of probable errors in the N.W. Quadrilateral of the Indian triangulation, after all adjustments have been carried out. The probable errors most generally required will be those with reference to the origin of the survey at Kalianpur (Sironj Base). In putting down the conditions of closure any closing point may be chosen: but when probable errors with regard to Kalianpur are desired, Kalianpur will naturally be selected as the closing point and origin for X and Y . Chart I shows all the triangulation of India: in charts II . . . V it is represented diagrammatically, each series being replaced by one or more straight lines, which may be regarded as the equivalent triangulation lines. The Indian triangulation was divided for purposes of adjustment into five portions, *viz.* the N.W. Quadrilateral, S.E. Quadrilateral, N.E. Quadrilateral, Southern Trigon, S.W. Quadrilateral. These were adjusted in the order stated, so that the first two were quite independently adjusted while the third was adjusted on the first two: the fourth was adjusted on the second, and the fifth was adjusted on the first, second and fourth. The Burma quadrilateral (chart IV) has just* been adjusted by the methods of Chapter VI and is being adjusted on the eastern series (Shillong Meridional, No. 44) of the N.E. Quadrilateral.

In chart II the series which were taken account of simultaneously in each quadrilateral or trigon are shown in full lines, while some additional series afterwards adjusted on these series are shown in broken lines. The series which are common to adjacent quadrilaterals or trigon are distinguished by heavy lines. The numbers written by the side of each triangulation line in the chart are those which have been applied in Table XLIV to the several series of the triangulation. The eight base lines of the triangulation of India and the Mergui base in Burma are shown: also the points at which it has been possible to form Laplace equations. The circuits are indicated by roman numerals and the points of closure by small arcs at the closing angle, *e. g.* at D . The junction points of the triangulation lines are distinguished by letters $A, B . . . Z, a, b . .$ with suffixes 1, 2, 3, 4, 5, 6 corresponding to the N.W. Quadrilateral, S.E. Quadrilateral, N.E. Quadrilateral, S. Trigon, S.W. Quadrilateral, Burma Quadrilateral.

The first step is to find M^2L for each triangulation line, L being the length in units of 100 miles. Chart II is on the scale of 100 miles = 1 inch, so that L is the length of each line on the chart in inches. As an example take the line between the Sironj base and the Dehra base. This is composed of two triangulation lines representing a portion of the Great Arc Series, No. 6. From table XLIV the value of M is 0.71 and by measurement on the chart the values of L for the two parts are found to be 2.15 and 2.22 inches. Hence the values of M^2L are 1.083 and 1.109 respectively.

The values of M^2L for all the triangulation lines are exhibited in table XLV together with certain related quantities to which reference will be made later. To proceed with the formation of the equations of form (13) which are necessary to determine the several probable errors, all quantities of the types indicated in (20) have to be formed. To any coefficient $[u g h]$ several component terms, each corresponding to a particular triangulation line, may contribute. Some of these are exhibited in table XLV, while the remainder are found in table XLVI. To go more into detail, the N.W. Quadrilateral (*vide* chart II) is divided into five circuits I, II, III, IV, V. In each circuit there are four types of closure—side, azimuth, easting, northing—which may be characterised by suffixes s, a, e, n . These give rise to twenty conditions $I_s, I_a, I_e, I_n, II_s, \dots, V_n$. In addition there are three extra base lines giving conditions VI, VII, VIII. To investigate the additional value of Laplace closures at the same points as the base line closures, giving rise to conditions VI, VII, VIII. The 26 conditions which result make it necessary to determine 26 multipliers $k_1 \dots k_{26}$ by equations of form (13). In table XLVII the coefficients of the left hand sides of these equations are given. The method by which these coefficients are derived will now be given in detail for a few of them. The letters used correspond to those shown marginally in table XLVII.

By (20) $[u a a] + M^2L = 1$; hence $[u a a] = \sum M^2L$ round circuit I = 3.66 from table XLV. Also $[u a b] = 0$, $[u a c] = \sum X M^2L = 2.20$ from table XLV. In the formation of any coefficients involving any pair of a, b, c or d it is clear that the summation extends around the circuit I, for the corresponding conditions relate to this complete circuit. The case is different when a coefficient involving one of the quantities, a, b, c, d and one other quantity, say e , are considered. In this case both circuits I and II are involved and it is only the part common to the two circuits that has to be considered. Moreover in this case two origins are introduced. The portion common to circuits I and II is the line D_1T_1 . It will be seen by (20) that

$$[u a e] = M^2L \text{ for } D_1T_1 = .70 \quad \text{from table XLV.}$$

$$[u a f] = 0$$

$$[u a g] = X_0 M^2L \text{ for } D_1T_1, \quad X_0 \text{ referring to the circuit to which } g \text{ relates, i.e. circuit II}$$

$$= +.41$$

$$[u a h] = Y_0 M^2L \text{ for } D_1T_1 \text{ from circuit II} = -2.30$$

This deals with all the conditions which relate to a portion which occurs also in condition a , except the base line conditions VI, and VII. The most complex case is the coefficient corresponding to northing or easting relations in two circuits which have a portion in common. An example of this is $[u d g]$. The circuits involved I, II have the portion T_1D_1 only in common. Hence from (20) as may be seen opposite the entry I, II in table XLVI, $[u d g] = M^2L (e_p Y - g_p X) = -1.14$. Similarly $[u d h] = M^2L (R^2 + L^2/12 + x_p X + y_p Y) = 7.35$. The quantities in tables XLV, XLVI depend on measurements of L, R, X, Y taken from a chart. These are not quite precise, and so quantities deduced from them are not quite precise. All the quantities which are known by symmetry to be equal in pairs have been separately determined by way of a check. They differ slightly as will be seen in table XLVI, the worst case being the coefficient which occurs in line II, III and also in line III, II of which the values 29.56 and 29.68 are obtained. The quantity made use of is the mean 29.62 shown in block type.

TABLE XLV.
N.W. Quadrilateral.

Base line closures	Circuit	Line	Series	M	L	A= M ² L	Closing point referred to Kalianpur	E ²	L ² 12	E ² +L ² 12	X	Y	A ^x (E ² +L ² 12)	AX	AY	Base line closures	Circuit	Line	A= M ² L	AX	AY
4 1 3	I	D ₁ F ₁	33	2.37	5.10	0.688	D ₁ a _p = -0.91 y _p = +5.18	6.45	2.17	8.62	-0.04	-2.54	6.02	-0.03	-1.77	VI 1	I	A ₁ B ₁	1.083	+0.88	-4.46
		T ₁ A ₁	25	0.60	0.98	0.353		26.50	0.08	26.58	+0.42	-5.13	9.38	+0.15	-1.81			B ₁ C ₁	1.109	+0.98	-2.16
		A ₁ B ₁	6	0.71	2.15	1.083		17.63	0.39	18.02	+0.81	-4.12	19.50	+0.88	-4.46						
		B ₁ C ₁	6	0.71	2.22	1.109		4.87	0.41	4.98	+0.88	-1.95	5.52	+0.98	-2.16						
		C ₁ D ₁	22	0.55	1.38	0.418		0.46	0.16	0.62	+0.53	-0.42	0.26	+0.23	-0.18						
						3.861							40.68	+2.20	-10.88						
	II	E ₁ U ₁	23	1.21	1.36	1.690	E ₁ a _p = -1.54 y _p = +5.94	0.48	0.15	0.63	-0.04	-0.80	1.26	-0.08	-1.37	VII 3	I II IV V	C ₁ D ₁	0.418	+0.22	-0.18
		X ₁ Y ₁	23	1.21	1.85	2.706		5.16	0.29	5.45	-0.11	-2.27	14.75	-0.30	-6.14			D ₁ E ₁	0.306	+0.10	-0.12
		Y ₁ S ₁	23	1.21	1.13	1.654		14.19	0.11	14.40	-0.27	-3.77	23.80	-0.45	-6.23			E ₁ F ₁	0.336	+0.15	-0.10
		T ₁ A ₁	23	1.21	1.44	2.107		25.37	0.17	25.54	-0.27	-5.03	53.80	-0.57	-10.60			F ₁ G ₁	0.163	+0.02	-0.02
		D ₁ E ₁	35	0.60	0.70	0.252		83.53	0.04	83.57	+0.20	-5.79	8.45	+0.05	-1.44						
			22	0.55	1.01	0.306		11.24	2.17	13.41	+0.59	-3.30	9.38	+0.41	-2.30						
5 5 4 4 4	III	J ₁ L ₁	32	0.43	1.99	0.968	J ₁ a _p = -4.26 y _p = +3.36	1.00	0.33	1.33	-0.30	-0.44	111.51	-0.34	-23.22	VIII 5	III V	L ₁ N ₁	0.368	-0.33	-0.16
		L ₁ N ₁	32	0.43	1.87	0.309		6.71	0.23	6.94	-1.88	-1.67	2.14	-0.61	-0.52			N ₁ H ₁	0.309	-0.61	-0.52
		P ₁ Q ₁	25	0.60	2.28	0.824		7.81	0.43	8.24	-1.09	-2.61	6.77	-0.82	-2.15			H ₁ J ₁	0.320	-0.18	-0.24
		Q ₁ R ₁	25	0.60	1.02	0.888		8.53	0.09	8.62	+0.64	-2.65	3.17	+0.24	-1.05						
		S ₁ T ₁	25	0.60	1.46	0.526		12.91	0.18	13.09	+1.85	-3.08	6.88	+0.07	-1.82						
		X ₁ Y ₁	23	1.21	1.44	2.106		12.00	0.17	12.17	+2.45	-2.45	25.61	+5.16	-5.16						
	IV	Y ₁ S ₁	23	1.21	1.13	1.653	F ₁ a _p = -2.50 y _p = +6.55	7.42	0.11	7.53	+2.45	-1.19	12.46	+4.05	-1.97	IX 6	IV V				
		T ₁ A ₁	23	1.21	1.85	2.707		6.90	0.29	7.19	+2.61	+0.30	10.46	+7.06	+0.81						
		D ₁ E ₁	45	0.53	0.98	0.275		5.90	0.08	5.98	+2.21	+1.01	1.65	+0.61	+0.23						
			45	0.53	1.94	0.545		0.92	0.31	1.23	+0.88	+0.39	0.67	+0.43	+0.21						
						9.679							70.30	+16.81	-11.33						
						3.426							8.46	+2.13	-4.28						
5 5 3	V	G ₁ H ₁	32	0.43	1.73	0.320	G ₁ a _p = -2.88 y _p = +6.80	0.87	0.25	1.12	-0.56	-0.75	0.36	-0.18	-0.24	X 7	V				
		J ₁ L ₁	32	0.43	2.03	0.376		7.59	0.34	7.93	-1.20	-2.43	2.93	-0.45	-0.93						
		V ₁ U ₁	45	0.53	1.04	0.545		0.65	0.31	0.96	-0.50	-3.05	5.37	-0.27	-1.66						
		F ₁ G ₁	37	0.59	2.37	0.825		2.25	0.47	2.72	+0.30	-1.45	2.24	+0.32	-1.20						
			22	0.55	0.54	0.163		0.04	0.02	0.06	+0.13	-0.15	0.01	+0.02	-0.02						
						2.229							10.96	-0.50	-4.05						

TABLE XLVI.

N.W. Quadrilateral. Common portions of adjacent circuits.

Circuits		Line	First circuit referred to second		For first circuit			AX_p	AY_p	Sum of last 3 columns		AX_p	AY_p	Sum of last 2 columns	
			x_p	y_p	AX	AY	$A\left(R^2+\frac{L^2}{12}\right)$								
I II	II I	$D_1 T_1$ $T_1 D_1$	+ .63 - .63	- .76 + .76	- .03 + .41	- 1.77 - 2.30	+ 6.02 + 9.38	- 0.02 - 0.26	+ 1.24 - 1.74	+ 7.34 + 7.36	+ 7.35	- 1.11 + 1.45	- 0.02 - 0.31	+ 1.13 + 1.14	\mp 1.14
II	III	$U_1 X_1$ $X_1 Y_1$ $Y_1 S_1$	+ 2.72	+ 2.58	- .30 - .45 - .57	- 6.14 - 6.23 - 10.60	+ 14.75 + 23.80 + 53.80								
III	II	$S_1 Y_1$ $Y_1 X_1$ $X_1 U_1$	- 2.72	- 2.58	- 1.32 + 5.16 + 4.05 + 7.06	- 22.07 - 5.16 - 1.97 + 0.81	+ 92.35 + 25.61 + 12.46 + 19.46	- 3.59 - 50.20	- 50.20	+ 29.56		- 62.45 + 29.62	+ 3.40	- 59.06	\mp 59.10
II IV	IV II	$E_1 U_1$ $U_1 E_1$	+ 0.96 - 0.96	- 0.61 + 0.61	- .08 + 1.35	- 1.37 - 2.59	+ 1.25 + 5.37	- 0.08 - 1.77	+ 0.84 - 1.58	+ 2.01 + 2.02	+ 2.20	- 1.31 + 2.48	- 0.05 - 1.13	- 1.36 + 1.35	\mp 1.36
III IV	IV III	$U_1 V_1$ $V_1 U_1$	- 1.76 + 1.76	- 3.19 + 3.19	+ .61 + .12	+ .28 - .60	+ 1.65 + 1.38	- 1.07 + 0.21	- 0.89 - 1.91	- 0.81 - 0.32	- 0.32	- 0.40 - 1.05	+ 1.94 - 0.38	+ 1.45 - 1.43	\mp 1.44
III V	V III	$V_1 J_1$ $J_1 V_1$	- 1.38 + 1.38	- 3.44 + 3.44	+ .48 - .27	+ .21 - 1.66	+ .67 + 5.37	- 0.66 - 0.37	- 0.72 - 5.71	- 0.71 - 0.71	- 0.71	- 0.29 - 2.29	+ 1.65 + 0.93	+ 1.36 - 1.36	\mp 1.36
IV V	V IV	$F_1 V_1$ $V_1 F_1$	+ 0.38 - 0.38	- 0.25 + 0.25	+ .01 + .32	- .99 - 1.20	+ 1.68 + 2.24	- 0.00 - 0.12	+ 0.25 - 0.30	+ 1.83 + 1.82	+ 1.83	- 0.38 + 0.46	- 0.00 - 0.08	- 0.38 + 0.38	\mp 0.38

The above explanation should make clear the formation of table XLVII, which gives the right hand sides of a set of equations. The left hand sides of these equations are different according as the quantity, whose probable error is sought, is of different form or relative to a different point as origin. To obtain a solution of all the cases which may arise it is desirable to keep the left hand sides of the equation in symbolical form denoting them by $A, B, C \dots Z$. It will be seen that this renders possible solution in a form which afterwards admits of the solution of the same conditions together with any further conditions that may be added.

TABLE XLVII.

N. W. QUADRILATERAL.

	I _a	I _b	I _c	I _d	II _a	II _b	II _c	II _d	III _a	III _b	III _c	III _d	IV _a	IV _b	IV _c	IV _d	V _a	V _b	V _c	V _d	VI _a	VI _b	VII _a	VII _b	VIII _a	VIII _b	IX _a	IX _b	X _a	X _b
I _a	1	3	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
I _b	2	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
I _c	3	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
I _d	4	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
II _a	5	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
II _b	6	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
II _c	7	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
II _d	8	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
III _a	9	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
III _b	10	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
III _c	11	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
III _d	12	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IV _a	13	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IV _b	14	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IV _c	15	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IV _d	16	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
V _a	17	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
V _b	18	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
V _c	19	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
V _d	20	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VI _a	21	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VI _b	22	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VI _c	23	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VI _d	24	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VII _a	25	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VII _b	26	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VII _c	27	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VII _d	28	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VIII _a	29	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VIII _b	30	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VIII _c	31	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
VIII _d	32	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IX _a	33	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IX _b	34	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IX _c	35	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
IX _d	36	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0
X _a	37	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0

7. To find the probable errors at any point of side azimuth easting or northing it is necessary to form the right hand sides of the equations of which the left hand side coefficients are exhibited in table XLVII, *i.e.* to complete the formation of equations (13). The quantities $[uaf]$ etc. are of the same nature as those already formed: but differ in the lines for which they have to be

TABLE XLVIII.

Line	Circuit	M	L	$A = M^2L$	R^2	$\frac{L^2}{12}$	Δ	Y	$R^2 + \frac{L^2}{12}$	$A(R^2 + \frac{L^2}{12})$	ΔX	ΔY	X_0	Y_0
K, L ₁	III	0.43	1.55	0.286	1.03	0.20	-0.91	-0.45	1.23	0.35	-0.26	-0.13	-5.17	+2.91
M, N ₁	III	0.43	0.78	0.144	8.70	0.05	-2.06	-2.11	8.75	1.26	-0.30	-0.30	-6.33	+1.27
O, P ₁	III	0.60	0.57	0.205	7.21	0.03	-0.14	-2.68	7.24	1.48	-0.03	-0.55	-4.40	+0.68
R, S ₁	III	0.60	0.53	0.191	15.27	0.02	+2.31	-3.15	15.29	2.92	+0.44	-0.60	-1.95	+0.21
V ₁ W ₁	III	0.53	0.74	0.208	2.45	0.05	+1.43	+0.64	2.50	0.52	+0.30	+0.13	-3.63	+4.00
V ₁ W ₁	V	0.53	0.74	0.208	7.85	0.05	+0.05	-2.80	7.90	1.64	+0.01	-0.58	-2.68	+4.00
W ₁ J ₁	III	0.53	1.20	0.337	0.37	0.12	+0.55	+0.25	0.49	0.16	+0.19	+0.08	-3.71	+3.61
W ₁ J ₁	V	0.53	1.20	0.337	10.87	0.12	-0.83	-3.19	10.99	3.70	-0.28	-1.08	-3.71	+3.61

TABLE XLIX.

Circuit, Base and Laplace closures.

Line	Circuit	Closing point of circuit referred to Kalianpur		(1) $A = M^2L$	(2) Δx_p	(3) Δy_p	(4) ΔX	(5) ΔY	(6) = (2)+(4) $=\Delta X_0$	(7) = (3)+(5) $=\Delta Y_0$	(8) = (6)+(7) $=\Delta$	(9) ΔX_p	(10) ΔY_p	(11) = (8)+(9) $=\Delta$	(12) ΔX_p	(13) ΔY_p	(14) = (12)+(13) $=\Delta$	(15) ΔX_p	(16) = (15)+ Δ
		x_p	y_p																
A, B ₁	I, VI	-0.91	+5.18	1.083	-0.09	+5.61	+0.88	-4.46	-0.11	+1.15	+19.50	-0.80	-23.10	-4.40	+4.06	-4.58	-0.50	1.53	1.65
B, C ₁	I, VI	-0.91	+5.18	1.109	-1.01	+5.74	+0.98	-2.16	-0.03	+3.58	+5.52	-0.89	-11.10	-6.56	+1.97	-5.08	-3.11	10.64	12.02
C, D ₁	I, VII	-0.91	+5.18	0.418	-0.38	+2.17	+0.22	-0.18	-0.16	+1.99	+0.26	-0.20	-0.03	-0.87	+0.16	-1.14	-0.98	22.96	9.59
D, E ₁	II, VII	-1.54	+5.94	0.306	-0.47	+1.82	+0.10	-0.12	-0.37	+1.70	+0.10	-0.15	-0.71	-0.76	+0.18	-0.59	-0.41	32.38	9.21
E, F ₁	IV, VII	-2.50	+6.55	0.336	-0.84	+2.20	+0.15	-0.10	-0.69	+2.10	+0.13	-0.37	-0.68	-0.90	+0.23	-0.98	-0.73	43.33	14.56
F, G ₁	V, VII	-2.88	+6.80	0.163	-0.47	+1.11	+0.02	-0.02	-0.45	+1.09	+0.01	-0.06	-0.14	-0.19	+0.06	-0.14	-0.06	51.61	8.46
G, H ₁	V, VIII	-2.88	+6.80	0.320	-0.92	+2.18	-0.18	-0.24	-1.10	+1.94	+0.36	+0.52	-1.63	-0.75	+0.69	+1.22	+1.91	48.69	15.58
H, I ₁	V, VIII	-2.88	+6.80	0.376	-1.08	+2.50	-0.45	-0.93	-1.53	+1.63	+2.98	+1.30	-6.32	-2.04	+2.68	+3.06	+5.74	35.65	18.40
K, L ₁	III, VIII	-4.26	+3.36	0.289	-1.22	+0.96	-0.26	-0.13	-1.48	+0.83	+0.85	+1.11	-0.44	+1.02	+0.55	+0.87	+1.42	35.40	10.12
M, N ₁	III, VIII	-4.26	+3.36	0.368	-1.57	+1.24	-0.33	-0.16	-1.90	+1.08	+0.49	+1.41	-0.54	+1.36	+0.68	+1.11	+1.79	35.48	13.06
N, O ₁	III, VIII	-4.26	+3.36	0.144	-0.61	+0.45	-0.30	-0.30	-0.91	+0.18	+1.26	+1.28	-1.01	+1.63	+1.28	+1.01	+2.29	41.56	5.98
L, N ₁	III, VIII	-4.26	+3.36	0.309	-1.32	+1.04	-0.61	-0.52	-1.93	+0.52	+2.14	-2.60	-1.75	+2.99	+2.22	+2.05	+4.27	42.03	12.99
O, P ₁	III	-4.26	+3.36	0.205	-0.87	+0.69	-0.03	-0.55	-0.90	+0.14	+1.48	+0.13	-1.85	-0.24	+2.34	+0.10	+2.44	19.85	4.07
N, P ₁	III	-4.26	+3.36	0.832	-3.50	+2.78	-0.82	-2.15	-4.32	+0.61	+8.77	+3.49	-7.22	+3.04	+9.16	+7.17	+11.92	28.66	23.55
P, Q ₁	III	-4.26	+3.36	0.368	-1.57	+1.24	+0.24	-1.05	-1.33	+0.19	+3.17	-1.02	-3.53	-1.38	+4.47	-0.81	+3.66	18.46	4.95
R, S ₁	III	-4.26	+3.36	0.191	-0.81	+0.64	+0.44	-0.60	-0.37	+0.04	+2.92	-1.87	-2.02	-0.97	+2.56	-1.48	+1.08	3.87	0.74
Q, S ₁	III	-4.26	+3.36	0.526	-2.24	+1.77	+0.97	-1.62	-1.27	+0.15	+6.88	-4.13	-5.44	-2.69	+6.90	-3.26	+3.64	6.06	3.19
S, T ₁	II	-1.54	+5.94	0.252	-0.39	+1.50	+0.05	-1.40	-0.34	+0.04	+8.45	-0.08	-8.67	-0.30	+2.25	-0.30	+1.95	1.81	0.47
T, A ₁	I	-0.91	+5.18	0.353	-0.32	+1.83	+0.15	-1.81	-0.17	+0.62	+9.38	-0.14	-9.37	-0.18	+1.65	-0.78	+0.87	0.32	0.11
T, D ₁	I	-0.91	+5.18	0.698	-0.64	+3.62	-0.03	-1.77	-0.67	+1.85	+6.02	+0.08	-9.17	-3.12	+1.61	+0.16	+1.77	10.04	7.01
T, D ₁	II	-1.54	+5.94	0.698	-1.07	+4.16	+0.41	-2.30	-0.66	+1.85	+9.38	-0.63	-13.66	-4.93	+3.54	-2.44	+1.10		
S, Y ₁	II	-1.54	+5.94	2.107	-3.24	+12.51	-0.57	-10.60	-3.81	+1.91	+53.80	+0.88	-62.96	-8.28	+16.32	+3.39	+19.71	4.28	9.02
S, Y ₁	III	-4.26	+3.36	2.106	-8.97	+7.08	+5.16	-5.16	-3.81	+1.92	+25.61	-21.98	-17.34	-13.71	+21.98	-17.34	+4.64		
Y, X ₁	II	-1.54	+5.94	1.654	-2.55	+9.82	-0.45	-6.24	-3.00	+3.58	+28.80	+0.69	-37.07	-12.58	+9.61	+2.67	+12.28	8.09	13.38
Y, X ₁	III	-4.26	+3.36	1.653	-7.04	+5.55	+4.05	-1.97	-2.99	+3.58	+12.46	-17.25	-6.62	-11.41	+8.39	-13.61	-5.22		
X, U ₁	II	-1.54	+5.94	2.706	-4.17	+16.07	-0.30	-6.14	-4.47	+9.93	+14.75	+0.46	-36.47	-21.26	+9.46	+1.75	+11.24	16.48	44.59
X, U ₁	III	-4.26	+3.36	2.707	-11.53	+9.10	+0.77	-0.81	-4.46	+9.91	+19.46	-30.12	+2.72	-7.94	-3.45	-23.76	-27.21		
U, E ₁	II	-1.54	+5.94	1.990	-3.06	+11.82	-0.08	-1.37	-3.14	+10.45	+1.25	+0.12	-8.14	-6.77	+2.11	+0.45	+2.59	30.21	60.12
U, E ₁	IV	-2.50	+6.55	1.990	-4.88	+13.03	+1.85	-2.59	-3.13	+10.44	+5.37	-4.63	-16.96	-16.22	+6.48	-12.12	-5.64		
U, V ₁	III	-4.26	+3.36	0.275	-1.17	+0.92	+0.61	+0.28	-0.56	+1.20	+1.65	-2.80	+0.94	-0.01	-1.19	-2.05	-3.24	23.38	6.43
U, V ₁	IV	-2.50	+6.55	0.275	-0.69	+1.80	+0.12	-0.60	-0.57	+1.20	+1.38	-0.30	-3.93	-2.85	+1.50	-0.79	+0.71		
V, W ₁	III	-4.26	+3.36	0.208	-0.89	+0.70	+0.30	+0.13	-0.59	+0.83	+0.52	-1.28	+0.44	-0.32	-0.55	-1.01	-1.56	24.05	5.00
V, W ₁	V	-2.88	+6.80	0.208	-0.60	+1.41	+0.01	-0.58	-0.59	+0.83	+1.64	-0.03	-3.94	-2.83	+1.67	-0.07	+1.60		
W, J ₁	III	-4.26	+3.36	0.337	-1.44	+1.13	+0.19	+0.08	-1.25	+1.21	+0.16	-0.81	+0.27	-0.38	-0.34	-0.64	-0.98	26.92	9.07
W, J ₁	V	-2.88	+6.80	0.337	-0.97	+2.29	-0.28	-1.08	-1.25	+1.21	+3.70	+0.81	-7.32	-2.81	+3.11	+1.90	+5.01		
V, J ₁	IV	-2.50	+6.55	0.335	-2.06	+5.40	+0.01	-0.99	-2.05	+4.41	+1.58	-0.03	-6.48	-4.93	+2.48	-0.07	+2.41	35.29	29.11
V, J ₁	V	-2.88	+6.80	0.335	-2.38	+5.61	+0.32	-1.20	-2.06	+4.41	+2.24	-0.92	-8.16	-6.84	+3.46	-2.18	+1.28		

computed. The first step is to compute the necessary quantities for each line; it will only remain then to combine by simple addition the several lines which go to form the route between the point of reference and the point whose relative probable errors are sought. The reference point will be in general A_1 (Kalianpur), though any other point could also be used. The quantities of scheme (20) have to be formed as was done in tables XLV and XLVI. This has to be done for each section. Typical cases are S_1T_1 , T_1D_1 , D_1E_1 . The first S_1T_1 enters only in closing condition of circuit II; T_1D_1 enters in closing conditions of circuits I and II: and D_1E_1 enters into closing condition of circuit II as well as base and Laplace closure between Dehra and Chach, VI. For most of the lines values of A , AX , AY , $A(R^2 + L^2/12)$ are given in table XLV. For the others which are required the values are now exhibited in table XLVIII. In table XLIX all the quantities necessary for forming $[uaf] \dots [uff] \dots$ are given. In the column headed "circuit" the base line closures are indicated, and these require only A , AX , AY , AX_0 , AY_0 , which are the same quantities as for the circuits. To form $[uff]$, $A(R_0^2 + L^2/12)$ is required, where R_0 refers to Kalianpur. This quantity is also given for all lines. It is deduced from values of X_0 , Y_0 .

As an example consider the probable errors at U_1 , selecting the route $A_1 B_1 C_1 D_1 E_1 U_1$. This gives a good example of the method of forming the right hand sides of the equations.

The sections A_1B_1 , B_1C_1 enter into circuits I and VI
 " C_1D_1 I " VII
 " D_1E_1 II " VII
 " E_1U_1 II " IV

TABLE L. (formed by (21))

Equations	Side closure						Azimuth closure	Easting closure						Northing closure.
	$A_1 B_1$	$B_1 C_1$	$C_1 D_1$	$D_1 E_1$	$E_1 U_1$	Total		$A_1 B_1$	$B_1 C_1$	$C_1 D_1$	$D_1 E_1$	$E_1 U_1$	Total	
I 1	+1.08	+1.11	+0.42			+2.61	0	-0.11	-0.03	-0.16			-0.30	+ 6.72
2	0	0	0			0	+2.61	-1.15	-3.58	-1.99			-6.72	- 0.30
3	+0.88	+0.98	+0.22			+2.08	+6.80	-4.40	-6.56	-0.87			-11.83	+ 4.59
4	-4.46	-2.16	-0.18			-6.80	+2.08	-0.50	-3.11	-0.98			-4.59	-11.83
II 5				+0.31	+1.99	+2.30	0				-0.37	-3.14	-3.51	+12.15
6				0	0	0	+2.30				-1.70	-10.45	-12.15	-3.51
7				+0.10	-0.08	+0.02	+1.49				-0.76	-6.77	-7.53	-2.18
8				-0.12	-1.37	-1.49	+0.02				-0.41	+2.59	+2.18	-7.53
III 9						0	0						0	0
10						0	0						0	0
11						0	0						0	0
12						0	0						0	0
IV 13					+1.99	+1.99	0							
14					0	0	+1.99					-3.13	-3.13	+10.44
15					+1.85	+1.85	+2.59					-10.44	-10.44	-3.13
16					-2.59	-2.59	+1.85					-16.22	-16.22	+5.64
V 17						0	0					-5.64	-5.64	-16.22
18						0	0						0	0
19						0	0						0	0
20						0	0						0	0
VI 21	+1.08	+1.11				+2.19	0	-0.11	-0.03				-0.14	+ 4.73
VII 22			+0.42	+0.31		+0.73	0			-0.16	-0.37		-0.53	+ 3.69
VIII 23						0	0						0	0
VI 24						0	+2.19	-1.15	-3.58				-4.73	-0.14
VII 25						0	+0.73			-1.99	-1.70		-3.69	-0.53
VIII 26						0	0						0	0

There are four cases according as probable errors of side, azimuth, easting or northing are sought. The scheme (20) may be rewritten, entering the numbers of the column in table XLIX in place of the actual symbolical quantities.

	S_f	A_f	E_f	N_f
S	(1)	zero	(6)	(7)
A	zero	(1)	-(7)	(6)
E	(4)	-(5)	(11)	-(14)
N	(5)	(4)	(14)	(11)

. (21).

For side error A_1 , B_1 contributes (1), 0, (4) and (5) to equations 1, 2, 3, 4, of circuit I and (1) to equation 21 of VI: to all other equations nothing. The numerical quantities taken from table XLIX are (1) = +1.08, (4) = +0.88, (5) = -4.46. The complete process is shown in table L.

The azimuth closure can at once be written down from the side closure by rearrangement of terms and changing of certain signs in accordance with (21).

8. In a similar way all the quantities occurring on the right hand sides of equations (13) can be formed as required, the necessary data being taken from Table XLIX. The solution of the equations which arise for the N. W. Quadrilateral is effected in the latter portion of the next chapter (VIII). It remains only to refer to the quantities $[u ff]$ occurring in equation (14). From (19) it is clear that the necessary quantities for determining these are $A = M^2 L$, $A X_0$, $A Y_0$, $A (R_0^2 + L^2/12)$, the suffix zero indicating A_1 , or Kalianpur as origin. All these quantities are given in table XLIX in columns (1), (6), (7), (15) for each section of line: and the corresponding quantities for a set of lines are obtained by summation of the sectional quantities.

From (2) and (5) the probable errors before adjustment are proportional to $\sqrt{[u ff]}$ the multiplying factors being 1".575 for azimuth, 33.2 for 7th place of logarithm of side and 4.03 for easting and northing in feet: for it is clear that $[u ff] = \sum M^2 L$ in the cases of azimuth and side and $[u ff] = M^2 L (R_0^2 + L^2/12)$ in the case of easting or northing. By (14) the probable errors after adjustment are proportional to $[u ff] - [u af] k_1 - [u bf] k_2$ the multiplying factors being as just given for the several cases. The ratio of probable error after to adjustment to probable error before adjustment is K where

$$K = \sqrt{1 - \frac{[u af]}{[u ff]} k_1 - \frac{[u bf]}{[u ff]} k_2 - \dots} \quad (22)$$

Numerical values of K will be given in the next chapter.

CHAPTER VIII.

Numerical values of the probable and actual errors in the Indian triangulation. Note on the solution of linear equations.

1. The formulæ (2) and (5) of chapter VII will now be applied to the actual circuits and closures of the Indian triangulation and the numerical results compared with the actual closing errors which have been found in the several circuits. The question is a little complicated by the fact that the triangulation has been adjusted in six portions and so, in the case of some circuits, the closing errors are those due to several series some of which have been adjusted in a neighbouring quadrilateral. In practice this is of little account, as the best quadrilaterals were first adjusted and those adjusted later are of considerably lower quality, so that the probable errors brought in by the adjusted quadrilaterals form only a small part of the total probable error and it is of little account whether the probable errors before or after adjustment are employed. In consideration of this question the fact that the flanking quadrilaterals were previously adjusted will be ignored. What has been said regarding the relative excellence of the several quadrilaterals does not apply to the Burma quadrilateral, which had only been begun when the Indian quadrilaterals were adjusted.

It will be seen from the equations that the quantities required for each line are M^2L and $M^2L\left(R^2 + \frac{L^2}{12}\right)$ whence L is the length of line and R the distance of its mid point from the closing point of the circuit concerned, both expressed in 100 miles. These quantities have already been taken out (table XLV) for the N.W. quadrilateral. It remains to obtain these for the rest of the triangulation. For this purpose charts III, IV, V are given. These as well as chart II are on the scale 100 miles to an inch: so that L and R are the lengths on the charts in inches. See also § 6, chapter VII. The actual measurements and necessary deductions are now shown in table LI. For the Base-line and Laplace closures it is only necessary to compute ΣM^2L along the route. This has already been done for each element of the route which enters into one of the circuits: and only a few remain to be formed for the Laplace closures. The values for the elements are combined to form the necessary values of ΣM^2L for the complete routes. The results are shown in table LII.

Having thus found the probable values of closing errors in all forms of closure, the next step is to compare the results with the closing errors which have actually been found. This is done for the Base-line and Laplace closure in table LII and for the circuit closures in table LIII. In each case the actual error is given and then this is divided by the theoretical error, giving a quantity f , the ratio of the actual error to the probable error. The actual errors of 7th place of logarithm of side, of azimuth in seconds, of easting and northing in feet are denoted by ΔS , ΔA , ΔE and ΔN respectively.

TABLE LI.

S. E. Quadrilateral.												N. E. Quadrilateral.											
Circuit and Closing point	Line	Series	M	L	A = M ² L	R	R ²	L ² /12	R ² + L ² /12 = C	AC	Circuit and Closing point	Line	Series	M	L	A = M ² L	R	R ²	L ² /12	R ² + L ² /12 = C	AC		
I S ₂	S ₂ W ₂ D ₂ A ₁	53	0-31	2-10	0-202	1-05	1-10	0-388	1-47	0-30	I C ₁	C ₁ B ₁ A ₁ C ₁	6	0-71	2-22	1-119	1-11	1-23	0-412	1-04	13-03		
		53	0-31	2-10	0-202	3-22	10-37	0-388	10-74	2-17			6	0-71	2-15	1-084	3-28	10-76	0-388	11-14	12-08		
		5	0-32	1-26	0-129	4-24	17-98	0-132	18-11	2-34			5	0-32	1-02	0-104	4-40	19-36	0-083	19-44	2-02		
	A ₂ T ₂ S ₂ S ₂	8	0-59	4-26	1-483	5-40	11-56	1-510	13-07	10-38		C ₂ C ₂ C ₂ B ₂ C ₂	4	1-79	5-20	10-253	2-92	8-53	0-854	9-38	96-17		
		43	0-30	1-16	0-104	0-58	0-34	0-112	0-45	0-04			20	0-65	0-58	0-376	1-07	1-14	0-066	1-21	0-45		
													20	0-65	0-63	0-266	0-32	0-10	0-039	0-13	0-03		
					2-120					24-23						13-202						123-78	
	R ₂ X ₂ E ₂ D ₂ S ₂	58	0-33	2-85	0-310	1-43	2-04	0-675	2-72	0-84	II C ₂	C ₂ C ₂ C ₂ E ₂ W ₂ D ₂ C ₂	4	1-79	3-20	10-253	1-80	2-56	0-855	3-42	35-07		
		58	0-33	1-33	0-150	3-49	12-18	0-158	13-34	1-85			5	0-32	1-22	0-125	3-26	10-82	0-124	10-94	1-87		
		5	0-32	1-00	0-102	4-25	18-06	0-083	18-14	1-86			5	0-32	0-07	0-423	3-00	9-00	0-078	9-08	81-13		
	D ₂ W ₂ S ₂ S ₂ E ₂	53	0-31	2-10	0-202	3-40	11-56	0-388	11-93	2-41		W ₂ D ₂ D ₂ C ₂	3	1-88	2-13	7-528	1-53	2-34	0-377	2-72	20-48		
		43	0-30	1-10	0-202	1-75	3-06	0-369	3-43	0-70			20	0-65	0-70	0-296	0-35	0-12	0-041	0-16	0-05		
					1-112					7-78						21-630						88-10	
III Q ₂	Q ₂ O ₂ M ₂ L ₂	24	0-70	3-22	1-678	1-61	2-59	0-864	3-45	5-44	III D ₂	D ₂ W ₂ E ₂ E ₂ E ₂	3	1-88	2-13	7-528	1-07	1-14	0-878	1-62	11-44		
		24	0-70	0-96	0-470	3-68	13-18	0-077	13-26	5-23			3	1-88	0-97	8-428	2-54	6-45	0-078	6-53	23-58		
		24	0-70	0-57	0-279	4-31	18-58	0-027	18-61	5-19			12	1-81	2-80	8-518	1-68	2-82	0-564	3-38	28-79		

TABLE LI.

S. Trigon												Burma Quadrilateral.											
Circuit and Closing point		Line	Series	M	L	$A = M^2 L$	R	R^2	$\frac{L^2}{12}$	$R^2 + \frac{L^2}{12} = C$	AC	Circuit and Closing point		Line	Series	M	L	$A = M^2 L$	R	R^2	$\frac{L^2}{12}$	$R^2 + \frac{L^2}{12} = C$	AC
N.E. Quad.												N.E. Quad.											
XI		M.L.	34	0.71	0.87	0.437	0.43	0.19	0.063	0.25	0.11	XII		N.M.	34	0.71	0.87	0.437	0.43	0.19	0.063	0.25	0.11
M.		L.L.	16	1.89	2.56	10.188	1.44	2.07	0.545	2.62	28.57	N.		M.S.	56	0.70	2.04	1.000	1.40	1.96	0.846	2.81	2.81
		L.T.	48	0.87	0.67	0.218	2.32	5.58	0.037	5.42	1.18			S.S.	48	6.57	1.10	0.357	2.07	4.28	0.100	4.38	1.56
		T.S.	48	0.87	0.40	0.180	2.07	4.28	0.018	4.26	0.56			P.Q.	44	0.49	0.47	0.116	1.78	3.17	0.018	3.19	0.37
		S.M.	56	0.70	2.04	1.000	1.02	1.04	0.347	1.89	1.89			Q.P.	44	0.49	0.79	0.137	1.18	1.39	0.052	1.44	0.27
						11.923					29.81			P.S.	44	0.49	0.57	0.209	0.44	0.19	0.063	0.25	0.06
																		2.352				4.72	
S. Trigon												Burma Quadrilateral.											
I		O.	49	0.45	2.04	0.413	1.02	1.04	0.347	1.89	0.57	I		P.	44	0.49	0.80	0.192	0.40	0.16	0.053	0.21	0.04
		O.R.	11	0.74	0.90	0.495	2.20	4.84	0.068	4.91	2.43			Q.R.	44	0.49	0.47	0.113	1.00	1.00	0.018	1.02	0.12
		R.P.	11	0.99	3.37	8.823	1.80	1.69	0.467	1.66	14.63			R.K.	52	0.89	1.45	0.221	1.89	3.57	0.175	8.75	0.88
		P.O.	11	0.99	0.44	0.164	0.22	0.05	0.018	0.07	0.17			K.L.	52	0.89	2.28	0.344	3.57	12.74	0.436	13.17	4.58
						9.900					18.00			L.P.	52	0.89	1.33	0.202	5.02	25.20	0.147	25.35	5.13
														H.G.	68	0.85	1.08	0.132	5.04	25.40	0.117	25.52	3.87
II		B.	9	0.36	3.37	0.436	1.68	2.82	0.045	8.77	1.64			G.F.	68	0.85	0.87	0.107	4.22	17.81	0.063	17.87	1.91
		T.R.	7	0.74	2.10	1.150	3.73	13.91	0.363	14.28	18.42			F.C.	68	0.85	2.78	0.341	2.90	8.41	0.645	9.06	3.09
		R.O.	49	0.45	2.04	0.413	3.46	11.97	0.346	12.32	5.09			C.A.	68	0.85	2.01	0.260	1.50	2.25	0.337	2.59	0.87
		O.N.	49	0.45	1.92	0.388	1.98	3.72	0.307	4.03	1.53			A.P.	44	0.49	0.53	0.120	0.25	0.06	0.021	0.06	0.01
		N.B.	54	0.37	1.46	0.204	0.75	0.56	0.189	0.85	0.17							2.032				19.69	
						2.581					24.88			J.L.	52	0.89	1.23	0.187	0.62	0.38	0.128	0.51	0.10
III		H.	46	0.40	3.23	0.355	1.11	1.23	0.412	1.64	0.58	II		L.G.	52	0.89	1.33	0.202	1.62	8.31	0.147	8.46	0.70
		L.	48	0.46	1.11	0.235	2.78	7.73	0.103	7.83	1.84			H.G.	68	0.85	1.08	0.132	2.22	4.93	0.117	5.05	0.87
		S.T.	43	0.30	1.17	0.105	3.42	11.70	0.114	11.81	1.24			G.F.	68	0.85	0.87	0.107	1.90	3.61	0.063	3.67	0.39
		T.B.	9	0.36	3.37	0.436	2.17	4.71	0.045	5.66	2.47			F.C.	68	0.85	2.78	0.341	2.52	6.35	0.645	7.00	2.39
		B.H.	54	0.37	1.47	0.201	0.73	0.53	0.180	0.69	0.14			C.A.	68	0.85	1.00	0.130	3.50	12.25	0.083	12.83	1.60
						1.332					6.27			B.M.	71	0.81	1.74	1.142	2.50	6.25	0.352	6.50	7.42
IV		D.	63	0.65	0.60	0.379	0.33	0.11	0.096	0.15	0.04			M.L.	71	0.81	1.24	0.816	1.03	1.03	0.128	1.19	0.87
		F.G.	33	0.31	1.28	0.123	1.83	1.74	0.234	2.02	0.37			L.P.	71	0.81	0.53	0.348	0.27	0.07	0.023	0.09	0.03
		G.H.	33	0.31	0.90	0.066	2.62	6.86	0.068	6.93	0.60							3.405				14.27	
		H.B.	54	0.37	1.47	0.201	3.77	7.67	0.160	7.83	1.57			M.E.	70	1.96	1.20	4.610	0.60	0.36	0.120	0.48	2.21
		B.D.	9	0.36	3.64	0.343	1.32	1.74	0.581	2.32	0.80			E.C.	68	0.85	1.00	0.231	1.44	2.07	0.270	2.81	0.52
						1.091					3.38			C.P.	68	0.86	1.00	0.180	1.88	3.53	0.083	3.61	0.47
V		I.	46	0.40	3.43	0.384	1.20	1.44	0.480	1.92	0.74			B.M.	71	0.81	1.74	1.142	0.87	0.70	0.252	0.91	1.04
		Q.S.	43	0.30	3.30	0.207	1.48	2.19	0.441	2.63	0.54							6.103				4.24	
		S.L.	46	0.40	1.11	0.178	0.56	0.31	0.103	0.41	0.07												
						0.769					1.35												

TABLE LII.

Base line closure	Laplace closure	Line	A = M ² L	B = Σ A	33.2 B	1.575 B	Δ S	Δ A	f _s Δ S 33.2 B	f _s Δ A 1.575 B	Reference number
Sironj-Dehra Dun Dehra Dun-Chach Sironj-Karachi	Kalianpur-Dehra Dun	A,B,C, C,D,E,F,G, A,T,S,Q,P,N, A,T,S,Q, Q,P,N, N,L,J,H,G, A,C,K,L, A,D, Juh. L,M,O,Q, A,T, T,S,R-Q, L,L, L,K,J,S,H,E,C,C, A,F, F,Fyz, T,A,B, Q,L,H, H,B, Q,H,B, B,N, B,L,N, A,T, T,S, S,Q, L,R,L,X, L,T,B, R, Chitt, Chitt, K,J, J,L, Prome Prome-Moulmein	2.202 1.221 2.818 1.130 1.188 1.874 1.808 0.158 2.327 1.494 0.813 10.14 3.006 0.780 4.65 0.435 0.739 0.201 0.940 0.204 0.425 1.484 0.854 0.772 0.702 0.050 0.824 0.269 0.375	40.27 36.89 50.66 1.130 1.188 1.874 44.89 0.628 50.63 40.44 18.69 105.71 57.34 5.41 21.91 1.354 0.706 0.711 1.026 21.65 45.85 0.752 0.824 0.897 0.817 0.984	2.338 1.675 1.716 2.115 0.628 2.403 2.403 5.02 3.662 1.354 0.706 0.711 1.026 2.778 1.365 0.897 0.817 0.984	+ 44.0 + 71.9 - 79.6 + 42.5 0.628 - 6.9 - 16.4 - 21.3 + 36.7 - 33.2 + 20.0 + 41.3 + 1.1 + 0.7 - 1.0 - 0.2 - 5.0 + 2.1	+ 0.5 - 4.4 + 1.9 - 5.3 - 1.7 - 2.3 + 1.4 - 3.0 - 2.2 0.017 - 1.0 - 0.2 - 5.0 + 2.1	0.893 1.980 1.574 4.210 0.953 0.137 0.406 1.146 0.347 0.614 0.913 1.283 0.051 0.017 0.733 0.223 6.12 2.179	0.214 2.620 1.106 1.872 1.774 1.447 1.236 3.259 1.970 2.923 0.733 0.223 6.12 2.179	35 36 37 38 39 40 41 42 43 44 45 46 47 48 49 50 51 52 53 54 55 56 57 58 59 60	

The values of ΔA in table LII are taken from the second table of chapter IX in which the accumulated errors of azimuth at Laplace stations are found. These are the errors *after* the adjustment of the triangulation has been performed.

TABLE LIII.

	Circuit	$B^2 = 2\Delta A$	$38 \cdot 2B$	ΔS	$1 \cdot 875 B$	ΔA	$D^2 = 2\Delta C$	$4 \cdot 08 D$	ΔE	ΔN	f_s $\frac{\Delta S}{33 \cdot 2B}$	f_n $\frac{\Delta S}{1 \cdot 875 B}$	f_e $\frac{\Delta E}{4 \cdot 08 D}$	f_n $\frac{\Delta N}{4 \cdot 08 D}$	Reference number
N.W. Quad.	I	3.061	68.512	+ 68.2	3.013	+ 5.906	40.68	25.703	+ 14.613	+ 57.503	1.074	1.960	0.569	2.237	1
	II	9.718	108.484	- 124.6	4.906	+ 1.550	111.61	42.537	+ 18.296	- 39.418	1.205	0.315	0.430	0.826	2
	III	9.679	108.285	- 79.6	4.900	- 3.254	79.30	35.887	+ 26.065	+ 39.067	0.761	0.664	0.726	1.069	3
	IV	3.426	61.453	+ 150.0	2.915	- 4.232	8.46	11.723	- 24.477	+ 3.688	2.456	1.452	2.088	0.310	4
	V	2.229	49.568	- 5.3	2.351	- 3.000	10.96	13.343	- 24.588	- 0.505	0.107	1.276	1.843	0.088	5
S.E. Quad.	I	2.120	48.339	- 54.9	2.293	+ 0.212	24.23	19.836	- 20.676	+ 5.043	1.654	0.092	1.042	0.254	6
	II	1.112	34.993	+ 31.9	1.630	- 4.968	7.73	11.266	+ 19.622	- 21.785	0.961	2.993	1.726	1.838	7
	III	3.306	60.358	- 17.5	2.863	- 3.888	27.04	20.956	+ 23.073	- 14.322	0.627	1.393	1.101	0.683	8
N.E. Quad.	I	13.202	120.640	+ 287.1	5.724	+ 11.598	123.78	44.834	+ 81.933	+ 94.856	2.380	2.026	1.628	2.118	9
	II	21.630	154.413	- 7.8	7.325	- 14.317	88.10	11.961	- 96.904	- 35.551	0.047	1.856	3.102	2.972	10
	III	19.677	147.275	+ 492.8	6.987	- 6.798	62.63	31.893	- 16.360	+ 93.415	3.275	0.973	0.513	2.929	11
	IV	14.681	127.222	- 641.0	6.035	+ 10.177	36.52	24.353	- 5.388	- 122.067	5.038	1.686	0.221	5.013	12
	V	9.601	102.864	+ 169.4	4.879	- 0.791	21.05	18.480	- 21.329	+ 38.675	1.647	0.192	1.164	2.062	13
	VI	12.669	118.192	+ 82.1	5.607	+ 4.843	24.84	20.066	+ 9.377	+ 21.306	0.695	0.864	0.467	1.061	14
	VII	11.530	112.780	- 133.9	5.350	- 5.739	23.13	19.380	- 2.710	- 36.047	1.178	1.073	0.139	1.880	15
	VIII	5.931	80.875	+ 196.7	3.837	+ 4.441	20.44	18.220	- 2.449	+ 30.593	2.432	1.157	0.134	1.679	16
	IX	7.405	90.902	- 284.6	4.312	- 2.279	11.96	13.936	+ 3.269	+ 31.703	3.130	0.528	0.234	2.275	17
	X	15.323	120.978	+ 190.9	6.166	+ 0.648	37.90	24.809	- 10.284	- 21.606	1.468	0.105	0.414	0.871	18
	XI	11.923	114.640	- 193.7	5.438	+ 2.405	29.81	22.004	- 13.044	+ 13.023	0.442	0.592	0.592	0.522	19
	XII	2.392	51.088	+ 102.0	2.421	- 13.140	4.72	8.767	- 57.230	+ 17.465	2.011	5.428	6.542	1.995	20
S. Trigon.	I	9.900	104.447	+ 138.6	4.955	- 0.251	18.00	17.099	+ 7.624	+ 2.218	1.308	0.051	0.446	0.130	21
	II	2.591	53.452	- 23.7	2.536	- 3.680	24.65	20.103	+ 28.479	- 0.101	0.425	1.451	1.416	0.005	22
	III	1.332	38.313	+ 39.9	1.818	+ 4.303	6.27	10.091	- 17.218	0.000	1.041	2.367	1.706	0.000	23
	IV	1.091	34.661	- 0.4	1.644	- 9.040	3.83	7.411	+ 42.500	- 27.515	0.012	5.500	5.734	3.712	24
	V	0.769	28.917	+ 11.7	1.372	- 0.319	1.35	4.663	- 0.973	- 4.235	0.404	0.233	0.208	0.904	25
S.W. Quad.	I	7.677	91.907	+ 189.8	4.364	- 7.806	48.44	28.049	- 54.726	- 64.255	2.063	1.789	1.952	2.291	26
	II	3.989	66.300	- 212.5	3.145	- 7.118	2.38	6.213	+ 38.058	+ 6.862	3.307	2.380	6.121	1.104	27
	III	9.162	100.496	+ 27.4	4.708	- 6.719	35.86	24.132	+ 6.233	- 43.578	0.261	1.408	0.258	1.818	28
	IV	2.934	58.872	- 185.9	2.698	+ 2.957	1.74	5.320	- 1.215	+ 15.238	3.269	1.096	0.226	2.866	29
	V	4.027	66.632	- 62.2	3.181	+ 3.046	3.30	7.323	- 6.712	+ 14.028	0.934	0.984	0.916	1.916	30
	VI	6.536	84.892	+ 257.1	4.027	- 3.912	18.93	17.636	- 23.569	+ 6.057	3.030	0.672	1.344	0.346	31
Burma Quad.	I	2.032	47.343	+ 45.0	2.246	- 2.287	19.69	17.881	- 47.521	+ 41.790	0.951	1.009	2.658	2.337	32
	II	3.405	61.254	+ 47.0	2.906	- 6.002	14.27	13.543	- 17.550	- 17.553	0.767	2.065	1.399	1.399	33
	III	6.103	82.037	- 189.0	3.892	+ 3.606	4.24	6.836	+ 2.914	+ 1.427	2.304	0.627	0.426	0.209	34

2. It is not to be expected that the actual errors will be the same as the probable errors: but in a considerable number of cases, values of the ratio of actual to probable errors, that is f , should be distributed according to the laws of probability. In a given number of cases the probability is that those values of f which are comprised within certain limits will form a certain percentage of the total cases. The probability integral, between the proper ordinates, represents this distribution of errors. It is tabulated in most books on minimum squares*. By means of this it is seen that 10 per cent of the errors will most probably fall in each of the regions $A, B, \dots J$ of table LIV defined by limiting values of f .

The actual values of f found in 167 cases of closures are classified in these columns. Each value of f is followed by a number in brackets which refers to the corresponding closing condition in tables LII or LIII.

* *Vide* Wright's "Adjustment of Observations", § 213.

TABLE LIV.

Values of f from to	A 0 -188	B -185 -976	C -375 -572	D -572 -777	E -777 1-000	F 1-000 1-249	G 1-240 1-539	H 1-539 1-900	I 1-900 2-438	J 2-438 ∞
Side (i) Circuit ...	-012 (24) -047 (10) -107 (6)	-261 (28)	-425 (23) -404 (26) -527 (8)	-005 (14) -781 (3) -787 (33)	-834 (30) -961 (82) -991 (7)	1-011 (23) 1-074 (1) 1-178 (16) 1-205 (2)	1-908 (21) 1-408 (18)	1-647 (13) 1-054 (6) 1-690 (19)	2-011 (20) 2-068 (26) 2-304 (34) 2-380 (9) 2-438 (16)	2-456 (4) 3-080 (31) 3-130 (17) 3-207 (27) 3-269 (29) 3-275 (11)
(ii) Base ...	3 -017 (56) -051 (54) -137 (43) 3	1 -347 (46) -406 (44) 1	3 -614 (47) 1	3 -803 (35) -913 (40) -953 (41) 3	3 1-146 (5) 1	4 1-283 (52) 1	2 1-574 (37) 1	3 1-980 (36) 1	5 5-088 (12) 7 4-210 (40) 1	
Asimuth (i) Circuit ...	-051 (21) -092 (8) -104 (18) -162 (13)	-233 (25) -315 (2)	-412 (10) -528 (17)	-664 (3)	-861 (14) -927 (34) -964 (30) -972 (31) -973 (11)	1-009 (33) 1-073 (15) 1-303 (8) 1-090 (20) 1-167 (16)	1-276 (5) 1-303 (8) 1-408 (28) 1-451 (32) 1-452 (4)	1-696 (12) 1-769 (26)	1-956 (10) 1-990 (1) 2-026 (9) 2-085 (33) 2-260 (27) 2-367 (23)	2-963 (7) 5-428 (20) 5-500 (24)
(ii) Laplace ...	4 0	2 -214 (35) -223 (58) 2	2 0	1 -733 (57) 1	5 -702 (55) 1	4 1-106 (30) 1	5 1-256 (50) 1-447 (48) 2	2 1-774 (16) 1-872 (43) 2	6 1-917 (42) 1-970 (53) 2-179 (60) 3	3 2-629 (38) 2-790 (41) 2-923 (54) 3-259 (51) 6-12 (58) 5
Easting ...	-134 (16) -139 (15)	-208 (25) -221 (12) -228 (39) -234 (17) -258 (28)	-414 (18) -420 (34) -430 (2) -446 (31) -467 (14) -513 (11) -560 (1)	-502 (10) -720 (3)	-916 (30)	1-042 (6) 1-101 (8) 1-154 (13)	1-314 (31) 1-399 (33) 1-416 (32)	1-706 (23) 1-726 (7) 1-828 (9) 1-843 (5)	1-952 (26) 2-088 (4)	2-653 (32) 5-734 (24) 6-121 (27) 6-542 (20) 8-102 (10)
Northing ...	2 -000 (23) -006 (22) -038 (5) -130 (21) 4	5 -209 (34) -254 (6) -310 (4) -346 (31) 4	7 -502 (10) -683 (8)	2 -871 (18) -904 (25) -925 (2)	1 1-061 (14) 1-080 (3) 1-104 (37)	3 1-399 (33) 1-679 (16) 1-818 (28) 1-860 (15)	3 1-916 (30) 1-936 (7) 1-995 (20) 2-092 (13) 2-118 (9) 2-237 (1) 2-275 (17) 2-291 (26) 2-337 (32) 9	4 2-866 (29) 2-929 (11) 2-972 (10) 3-712 (24) 5-013 (12)	2 5	
Side and Azi- muth exclu- ding Laplace	10	4	6	5	11	9	8	6	12	11
Northing and Easting ...	6	9	7	4	4	6	4	7	11	10
All except Laplace ...	16	13	13	9	15	15	12	13	23	21

3. On examination of table LIV it is immediately noticeable that the cases in which the actual error is greater than the computed probable error are more numerous than the cases when it is less. Considering all the 167 cases, f is less than unity for 70 cases and greater than unity for the remaining 97 cases. This inequality is largely attributable to the Laplace closures,

although in this case values of azimuth, adjusted for all circuit conditions, have been used. In the Laplace closures there are 13 cases of actual error greater than computed probable error and only 4 cases of actual error less than computed error. It is believed that the explanation of this is that given in § 6, Chapter V. The error due to acceptance of geoidal angles uncorrected, instead of spheroidal angles, to some extent magnifies the triangular error and so increases the value of M , and as a result the closures of azimuth in circuits and of the deduced quantities side, northing and easting are in better agreement with the formulæ than the closures on Laplace points; since the former do not depend on absolute errors while the latter do. If the Laplace closures are ignored the number of cases, less than and greater than the formulæ give, are 66 and 84 respectively. If the formulæ values were increased in the ratio 1.1 to 1.0* the figures would become 74 and 76 respectively. It appears then that the formulæ give values of the probable error which are some ten per cent below what the 150 cases would lead to. This is not a serious deviation from the facts and, apart from mere chance, may be attributed partly to

- (1) the use of geoidal instead of spheroidal angles.
- (2) the fact that M is based on certain simplifying assumptions regarding the regularity of the triangles and polygons in the series of triangulation*.

The total number of cases falling in each class A, ... J is shown at the bottom of the table and from this it is seen that the errors are fairly distributed in the various classes, except that in classes I, J a considerable excess of cases occur. The excess of large errors over the number which is given by the formulæ, viz. $45 - 30 = 15$ or 50%, in classes I, J is to be attributed to the neglect of certain sources of error. One such source of error is that, already mentioned, of treating geoidal and spheroidal angles as identical: and it may be that other undetected sources also exist. However the formulæ give practically a satisfactory indication of the probable accuracy of side, azimuth, easting or northing. They should be a useful guide to the care which ought to be expended on observing and selecting a series in order that a result of any stated precision may be arrived at. As work on such a series progresses, the value of M may be taken out and observations increased in number, or rays increased in length, until the value of M is reduced to a quantity sufficiently small to give the proper precision.

4. It has just (June 1917) been noticed that the question of probable errors of side, azimuth, easting and northing generated in a chain of triangles were considered by General Walker and Mr. W. H. Cole in 1882†. The deduction is based on the equations by which the simultaneous reduction of the triangulation of India had been effected, and the equations obtained—*vide* xxviii, xxix, xxx *ibid*—are somewhat complex. These equations are comparable with (2) and (5) of Chapter VII of this work. Dealing with the case of a simple chain of equilateral triangles (on p. 104) with sides of 15 miles and chain of length 8° of arc, it is found in the Appendix that the

$$e. m. s. \left(i. e. \frac{\text{probable error}}{.6745} \right) \text{ in azimuth} = 6''.93\epsilon, \text{ average value}$$

$$\text{latitude } 0''.55\epsilon$$

$$\text{longitude } 0''.59\epsilon$$

the first quantity being somewhat dependent on the direction of the chain. It appears that ϵ is the quantity now denoted by m . To obtain results by the method of present work put $M =$

$$\frac{7}{8} m \sqrt{\frac{18}{15}} = 1.278m. \text{ Then taking } 8^\circ \text{ as equivalent to 550 miles,}$$

$$\text{Probable error in azimuth} = 1.575 \times 1.278\epsilon \sqrt{5.5} = 4''.72\epsilon$$

$$\text{Mean error in azimuth} = \frac{4.72}{.6745} \epsilon = 6''.99\epsilon$$

$$\text{Probable error in easting or northing} = 4.03 \times 1.278\epsilon \sqrt{5.5} \times \frac{5.5}{\sqrt{3}} \text{ feet} = 38.3\epsilon \text{ feet.}$$

* See also § 4 below.

† *Vide* G.T.S. Vol. VII, Appendix No. 3.

if it is multiplied by $-\frac{r_1}{a_1}$ and added to the r th equation x_1 is eliminated. The following equations are formed

$$\left. \begin{aligned} \left(b_2 - \frac{b_1}{a_1} a_2\right) x_2 + \left(b_3 - \frac{b_1}{a_1} a_3\right) x_3 + \dots &= B - \frac{b_1}{a_1} A \\ \left(r_2 - \frac{r_1}{a_1} a_2\right) x_2 + \left(r_3 - \frac{r_1}{a_1} a_3\right) x_3 + \dots &= R - \frac{r_1}{a_1} A \end{aligned} \right\} \quad (2)$$

To eliminate x_2 the same process is applied, the multipliers in this case all having the denominator $b_2 - \frac{b_1}{a_1} a_2$. It is to be observed that the denominator of the multiplying factors is always the first coefficient of the first equation of the set being operated on. The successive denominators are a_1 , $b_2 - \frac{b_1}{a_1} a_2$, $c_3 - \frac{c_1}{a_1} a_3 - \frac{c_2 - \frac{c_1}{a_1} a_2}{b_2 - \frac{b_1}{a_1} a_2} \left(b_3 - \frac{b_1}{a_1} a_3\right)$ etc.

In the solution of normal equations the diagonal coefficients are generally larger than the others. Being of the forms $\sum u a^2$ and $\sum u a b$ respectively the component parts of the first form are all positive while those of the second form are equally likely to be positive and negative, and accordingly tend to cancel. Accordingly in more cases than not $a_1 > a_r$ ($r \neq 1$), $b_2 > b_r$ ($r \neq 2$) . . . so that $b_2 - \frac{b_1}{a_1} a_2$ is not likely to be small compared with b_2 . But as the denominators become more complex there is more possibility of their becoming small. To avoid this, as well as may be foreseen before the actual computations are carried out, it is accordingly convenient to rearrange the equations in such order that the diagonal coefficients are of increasing magnitude. Before doing this however it is desirable that all these diagonal coefficients should be brought up to as near as may be the same order of magnitude. It is inconvenient to have quantities entering the computation, some with many figures preceding the decimal point while in others there are no figures before the decimal point and a number of zeros following the decimal point.

Any coefficient say f_r can be changed to $10 f_r$ or $\frac{1}{10} f_r$ if at the same time for x_r is written $\frac{1}{10} x_r$ or $10 x_r$, this being done in all the equations; and solution subsequently being performed for $\frac{1}{10} x_r$ or $10 x_r$ as the case may be. It is convenient to use as multiplier a power of ten, as this involves no loss of precision in the coefficient and no labour in transforming it. If the process thus suggested is carried out in the case of symmetrical equations—such as normal equations—the symmetry is destroyed. As the symmetry is advantageous it is desirable to avoid destroying it, and this may be arranged for as follows. When any column is multiplied by a power of ten the corresponding row should be multiplied by the same quantity. Then the equations (1) may be written

$$\left. \begin{aligned} a_1 x_1 + a_2 x_2 + \dots + 10^a a_r X_r + \dots &= A \\ 10^a r_1 x_1 + 10^a r_2 x_2 + \dots + 10^{2a} r_r X_r &= R \cdot 10 \end{aligned} \right\} \quad (3)$$

in which $X_r = \frac{x_r}{10^a}$ which are quite symmetrical.

In the solution of (1) the quantities which should be brought up to about the same order are the diagonal coefficients, as these enter more than the others into the computations. It may be seen from (3) that the diagonal coefficients can only be conveniently changed by powers (positive or

negative) of 10^3 . The first step then is to apply this process of dealing with the diagonal coefficients: the second is to rearrange the equations in such order that the modified diagonal coefficients occur in increasing order of magnitude. It can always be arranged that the largest diagonal coefficient is not so much as one hundred times the smallest.

6. Suppose now that (1) represents a set of equations dealt with and arranged as explained above. They are now in an order as favourable for solution as can be arranged for by mere inspection.

In proceeding with the elimination as explained in §5, notice that the R. H. S. quantities A, B, C do not in any way affect the elimination. The work on the left hand side of the equation is entirely independent of the values A, B, C Suppose then that a solution has been taken out, which is essentially only approximate: the degree of approximation depends on the number of figures retained in the arithmetical processes. Instead of the correct quantities x_1, x_2 . . . this solution will determine slightly different quantities $x_1 - \delta x_1, x_2 - \delta x_2$. On substituting these in the L.H.S of (1) the values found will be $A - \delta A, B - \delta B$ Hence

$$\left. \begin{aligned} a_1 \delta x_1 + a_2 \delta x_2 + \dots &= \delta A \\ b_1 \delta x_1 + \dots &= \delta B \end{aligned} \right\} \dots \dots \dots (4)$$

These equations have the same coefficients on the L.H.S. as those of the original equations (1). Hence the process of elimination in order to determine $\delta x_1, \delta x_2$. . . is the same as that already performed for x_1, x_2 . . . : and it is only necessary to change the portion of the computation involving A, B, C . In this way a second approximation is easily arrived at. Clearly this result may again be treated in the same way, and so successive sets of higher approximation may be obtained, without any necessity of increasing the accuracy of the elimination process. The gain in accuracy is arrived at by means of accurate substitution, and this substitution may be made absolutely perfect by keeping all the figures resulting from the substitution. As an example, if the original solution consists of numbers of 4 significant figures, and the coefficients are given to 4 figures the products will consist of 7 or 8 significant figures. It may be pointed out here that if the coefficients are given to much higher accuracy it is not necessary to use the full amount of figures for the elimination process: but this must be done in the substitution.

The upshot of this is that the various multiplications and divisions which arise in the process of elimination may all be performed by slide rule, which greatly facilitates the work. The substitutions must be carried out with higher, or even absolute, accuracy, as can most conveniently be done by arithmometer.

7. As remarked in §5 it is sometimes convenient to have the solution in terms of symbolic values of the R.H.S. This gives rise to another general method of procedure as regards higher approximation to any desired degree, which will now be described. Other attendant advantages will be seen to arise in this method.

Suppose the solution of (1) is expressed in the form

$$\left. \begin{aligned} x_1 &= a_1 A + b_1 B + c_1 C + \dots \\ &\vdots \\ x_r &= a_r A + b_r B + c_r C + \dots \end{aligned} \right\} \dots \dots \dots (5)$$

Then a_1, a_2, \dots are the solutions of

$$\left. \begin{aligned} a_1 x_1 + a_2 x_2 + \dots &= 1 \\ b_1 x_1 + b_2 x_2 + \dots &= 0 \\ r_1 x_1 + r_2 x_2 + \dots &= 0 \end{aligned} \right\}$$

and x_1, x_2, \dots are the solutions of

$$\left. \begin{aligned} a_1 x_1 + a_2 x_2 &= 0 \\ r_1 x_1 + r_2 x_2 &= 1 \\ &= 0 \end{aligned} \right\} \dots \dots \dots (6)$$

In (6) all the quantities on the R.H.S. are zero, except the r th, which is unity. The elimination processes are identical for all values of r , but the work on the R.H.S. differs for each value of r . The accurate solution of all the sets of form (6) gives values of all the quantities x_i : but any actual solution will give quantities slightly different, viz. $x_i - \delta x_i$. On substituting these, instead of getting the quantities unity and zero as values of the R.H.S. slightly different quantities are obtained, as indicated.

Values of R.H.S. of equations (6) resulting from an approximate solution.

Values of R.H.S.				Approximate solutions.			
	1	2	3	1	2	3	
1	$1 + a_1$	a_2	a_3	$a u_1$	$a u_2$	$a u_3$	
2	β_1	$1 + \beta_2$	β_3	$b u_1$	$b u_2$	$b u_3$	
3	γ_1	γ_2	$1 + \gamma_3$	$c u_1$	$c u_2$	$c u_3$	

If the original equations are symmetrical, so also are the quantities of the approximate solution.

It is clear that if the approximate calculation has been properly carried out all the quantities a, β, γ are small compared with unity.

8. Suppose the solutions 1, 2, 3 . . . are combined in any way, taking for x_1 the value $A a u_1 + B b u_2 + \dots$ and similarly for the other x . Then it is clear that the corresponding values of the R. H. S. will be

$$\begin{aligned} A(1 + a_1) + B a_2 + C a_3 & \dots \dots \dots , \\ A \beta_1 + B(1 + \beta_2) + C \beta_3 & \dots \dots \dots , \\ A \gamma_1 + B \gamma_2 + C(1 + \gamma_3) + & \dots \dots \dots , \end{aligned}$$

Putting $A = 1 - a_1$. . $B = -\beta_1$. . $C = -\gamma_1$ these become

$$\begin{aligned} 1 - a_1^2 - \beta_1 a_2 - \gamma_1 a_3 & \dots \dots \dots , \\ - a_1 \beta_1 - \beta_1 \beta_2 - \gamma_1 \beta_3 & \dots \dots \dots , \end{aligned}$$

Only products of the small quantities a, β . . . now occur except in the quantity unity in the first line: so that a higher approximation is readily found in this way. The process can obviously be repeated as often as is desirable.

9. The question can also be considered otherwise. Suppose the true value of any quantity x is $u-v$, u being a value obtained by solution and v a small correction. Then the solutions of equations

$$\begin{array}{lcl} a_1 x_1 + a_2 x_2 + \dots & = & \begin{vmatrix} A & B & C \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{vmatrix} \quad \text{are} \quad \begin{vmatrix} A & B & C \\ a_1 x_1 & a_2 x_2 & a_3 x_3 \\ a_2 x_2 & a_3 x_3 & a_4 x_4 \end{vmatrix} \dots \dots (7) \\ b_1 x_1 + b_2 x_2 + \dots & = & \end{array}$$

and as above

$$\begin{array}{lcl} a_1 v_1 + a_2 v_2 + \dots & = & \begin{vmatrix} A & B & C \\ a_1 & a_2 & a_3 \\ \beta_1 & \beta_2 & \beta_3 \end{vmatrix} \\ b_1 v_1 + b_2 v_2 + \dots & = & \begin{vmatrix} A & B & C \\ \beta_1 & \beta_2 & \beta_3 \\ \gamma_1 & \gamma_2 & \gamma_3 \end{vmatrix} \\ \dots & & \dots \end{array}$$

Whence by means of (7) the solutions are represented by

$$\left. \begin{array}{l} s v_1 = a_s a x_1 + \beta_s b x_2 + \gamma_s c x_3 + \dots \\ s v_2 = a_s a x_2 + \beta_s b x_3 + \gamma_s c x_4 + \dots \\ s v_3 = a_s a x_3 + \beta_s b x_4 + \gamma_s c x_5 + \dots \end{array} \right\} \dots \dots (8)$$

Hence

$$s v_r = a_s a x_r + \beta_s b x_{r+1} + \gamma_s c x_{r+2} + \dots \dots \dots (9)$$

The quantities $a x_r$ etc. may as a first approximation be replaced by the determined quantities $a x_r$ etc. The solution then takes the form

$$s u'_r = s u_r - a_s a x_r - \beta_s b x_{r+1} \dots \dots \dots (10)$$

and if further approximation is required this value may be substituted in (8) for $a x_r$ giving the next approximation to $s v_r$. Any number of successive approximations may be made in this manner. The $R. H. S.$ corresponding to (9) may be written down. They are

$$\begin{array}{lcl} 1 + a_1 - a_1(1 + a_1) - a_2 \beta_1 - a_3 \gamma_1 & , & a_2 - a_2(1 + a_1) - a_2 \beta_2 - a_3 \gamma_2 \\ \beta_1 - \beta_1(1 + a_1) - \beta_2 \beta_1 - \beta_3 \gamma_1 & , & 1 + \beta_2 - a_2 \beta_1 - \beta_2(1 + \beta_2) - \gamma_2 \beta_3 \end{array} \dots \dots$$

$$\text{or} \quad \begin{array}{lcl} 1 - a_1^2 - a_2 \beta_1 - a_3 \gamma_1 & , & -a_1 a_2 - a_2 \beta_2 - a_3 \gamma_2 \\ -\beta_1 a_1 - \beta_2 \beta_1 - \beta_3 \gamma_1 & , & 1 - \beta_1 a_2 - \beta_2^2 - \beta_3 \gamma_2 \end{array} \dots \dots$$

which may be written with abbreviated form

$$\begin{array}{ccc} 1 + a'_1 & a'_2 & a'_3 \\ \beta'_1 & 1 + \beta'_2 & \beta'_3 \end{array} \dots \dots$$

where

$$a'_1 = -a_1^2 - a_2 \beta_1 - a_3 \gamma_1, \dots \dots \text{etc.}$$

These residuals α', β' etc., are composed of binary products of α, β etc. and accordingly are of a smaller order than α, β Starting with the second approximate values α', β' and these residuals α', β' another approximation may be made: and so on as far as is necessary to attain the accuracy of solution desired.

10. It may at times be convenient to split up a set of equations and apply the above process to any portion consisting of an equal number of rows and columns. This may be done with advantage when a considerable number of the coefficients are zero. It is to be remembered that the actual numerical labour of solving n equations varies as the cube of n , so that a group of say 30 equations presents a formidable piece of computation. By the method now proposed perhaps this labour may be considerably reduced: but one certain advantage is a substitution check at a comparatively early stage of the computation. This is a check against actual computation blunders, as well as against accumulation of error due to lack of absolute exactness in the calculation on account of the necessity of limiting the number of figures employed. In this way, as has been shown above, much of the work can be performed readily with a slide rule, greatly accelerating the work.

In a large class of equations, *e.g.* the normal equations which occur in the method of least squares, there is complete symmetry about a diagonal. This reduces the work of elimination to about one-half. It is important then that in dealing with equations of this class that the symmetry should be preserved. Gauss's arrangement secures this for his method of solution, and it will now be shown that symmetry is maintained when the equations are split up as just suggested.

Denote the equations by

$$\left. \begin{aligned} (1,1) k_1 + (1,2) k_2 + \dots + (1,n) k_n &= (i) \\ (2,1) k_1 + (2,2) k_2 + \dots + (2,n) k_n &= (ii) \\ (n,1) k_1 + (n,2) k_2 + \dots + (n,n) k_n &= (n) \end{aligned} \right\} \dots \dots \dots (11)$$

Suppose the solution of

$$\left. \begin{aligned} (1,1) k_1 + (1,2) k_2 \dots + (1,r) k_r &= [i] \\ (2,1) k_1 + (2,2) k_2 \dots + (2,r) k_r &= [ii] \\ (r,1) k_1 + (r,2) k_2 \dots + (r,r) k_r &= [r] \end{aligned} \right\} \dots \dots \dots (12)$$

where r is less than n , is

$$\left. \begin{aligned} k_1 &= {}_1k_1 [i] + {}_2k_1 [ii] + \dots + {}_rk_1 [r] \\ k_2 &= {}_1k_2 [i] + \dots + {}_rk_2 [r] \\ k_r &= {}_1k_r [i] + \dots + {}_rk_r [r] \end{aligned} \right\}^* \dots \dots \dots (13)$$

Then from the first r equations of (11) it is seen that

$$\left. \begin{aligned} [i] &= (i) - (1,r+1) k_{r+1} - (1,r+2) k_{r+2} - \dots - (1,n) k_n \\ [ii] &= (ii) - (2,r+1) k_{r+1} - (2,r+2) k_{r+2} - \dots - (2,n) k_n \\ [r] &= (r) - (r,r+1) k_{r+1} - (r,r+2) k_{r+2} - \dots - (r,n) k_n \end{aligned} \right\} \dots (14)$$

Substituting from (14) in (13) it follows that

* This may be a first or higher approximation as appears most suitable.

$$\left. \begin{aligned}
 k_1 &= {}_1k_1(i) + {}_2k_1(ii) + \dots + {}_rk_1(r) \\
 &\quad - {}_{r+1}k_1 \sum {}_sk_1(s, r+1) - {}_{r+2}k_1 \sum {}_sk_1(s, r+2) - \dots - {}_nk_1 \sum {}_sk_1(s, n) \\
 k_i &= {}_1k_i(i) + {}_2k_i(ii) + \dots + {}_rk_i(r) \\
 &\quad - {}_{r+1}k_i \sum {}_sk_i(s, r+1) - {}_{r+2}k_i \sum {}_sk_i(s, r+2) - \dots - {}_nk_i \sum {}_sk_i(s, n) \\
 &\dots \dots \dots
 \end{aligned} \right\} \dots (15)$$

the summation indicated by Σ referring to s to which all values from 1 to r are to be given.

The values k_1, k_2, \dots, k_r are now to be substituted in the latter $n-r$ equations of (11). It is clear that thereby the coefficients of $k_{r+1}, k_{r+2}, \dots, k_n$ are altered. The original coefficients are clearly symmetrical, and it is only necessary to show that the change in the coefficient of k_u in the $r+1^{\text{st}}$ equation—the first of the equations dealt with—is the same as the change in the coefficient of k_{r+1} in the u^{th} equation. By (15) and (11) it is seen that the change in coefficient of k_u in the $r+1^{\text{st}}$ equation is

$$-(r+1, 1) \sum {}_sk_1(s, u) - (r+1, 2) \sum {}_sk_2(s, u) - \dots - (r+1, r) \sum {}_sk_r(s, u) \dots (16)$$

while the change in the coefficient of k_{r+1} in the u^{th} equation is

$$-(u, 1) \sum {}_sk_1(s, r+1) - (u, 2) \sum {}_sk_2(s, r+1) - \dots - (u, r) \sum {}_sk_r(s, r+1) \dots (17)$$

Remembering that ${}_sk_t = {}_tk_s$, notice that the sum of the coefficients of ${}_sk_t$ and ${}_tk_s$ in (16) is $-(r+1, t)(s, u) - (r+1, s)(t, u)$ and the corresponding quantity in (17) is $-(u, t)(s, r+1) - (u, s)(t, r+1)$. These quantities are the same, since $(u, t) = (t, u)$ etc.

The equations resulting from this method of solution are accordingly symmetrical.

11. In some cases after the solution of a set of normal equations has been effected, additional conditions may have to be introduced. The form of the equations, *vide* (13), is only modified thereby by the addition of a number of terms at the end of the original equations and an addition of the same number of equations at the end. If then the original equations have been solved in the manner explained above, it is possible to proceed immediately to derive the solution of the larger number of equations, making use of the solution already obtained. If the ordinary method of solution of the original equations had been followed this would have been of little help in proceeding to the solution of the larger number of equations.

These methods will now be given effect to in the solution of the 26 equations of p. 116. In table XLVII the coefficients are marked off in the stages for which solution will be performed.

In cases where a highly accurate solution is not desired the values of the residuals are not required. It is however of importance to verify that the solution does not contain any blunders, as may easily occur in the numerical work. A check on this is obtained by substituting the values obtained for solution A, B, \dots in the last equation. This equation only enters into the final eliminant from which the value of the last unknown is determined, and not into any of the previous equations used for the actual solution—that is the first of each group of equations formed by successive elimination of the first, second, \dots unknowns. If then this last equation is satisfied with satisfactory precision it is an indication that no blunder has been committed. It is not certain from this that the residuals of the other equations are equally small, and nothing short of substitution in each of these will make this point quite clear: but it is a sure check against any serious blunder.

12. Application of the method suggested above will now be made to the equation whose solution is necessary for the determination of the several probable errors of the N.W. Quadrilateral after adjustment. The L.H.S. of the equations are indicated in table XLVII. Conformably with §10 only a portion of the complete set is dealt with at first. It is at once clear that the first 8 equations and the next 12 equations form convenient groups. The first step is to solve the first 8 equations for the quantities $k_1 \dots k_8$ ignoring for the present quantities $k_9 \dots k_{16}$ which occur in the 5th—8th equations. The R.H.S. are taken as zero or unity, *vide* (6). As an example of (3) multiplications by powers of 10 are introduced and the order of equations arranged to make the diagonal quantities in increasing order of magnitude. In this particular case there is little gained by the former procedure, which is introduced merely to illustrate the method. The arrangement of the work is shown in tabular form in table LV, of which detailed explanation is now given.

TABLE LV.

Equation Number	Right Hand Side	Left Hand Side							
		1 k_3	2 k_4	3 k_7	4 k_8	5 $\frac{1}{10}k_1$	6 $\frac{1}{10}k_2$	7 $\frac{1}{10}k_5$	8 $\frac{1}{10}k_6$
1	+1.000	+40.68	0	+7.35	+1.14	+22.0	+103.8	-0.8	+17.7
1(1)	+4.1335		0	-0.096	-0.001	-1.354	-2.993	-0.001	+0.247
2	0		+40.68	-1.14	+7.35	-103.8	+22.0	-17.7	-0.8
3	0		0	0	0	0	0	0	0
3	0			+111.51	0	+4.1	+23.0	-8.4	+282.2
4	-0.1807			-1.33	-0.206	-3.975	-18.76	+0.54	-3.2
4	0				+111.51	-23.0	+4.1	-282.2	-8.4
5	-0.0280				-0.3	-0.62	+2.906	+0.084	-0.496
6	0					+306.0	0	+70.0	0
6	-0.541					-11.9	-56.16	+0.162	-9.576
6	0						+306.0	0	+70.0
7	-2.552						-264.9	+0.7656	-45.17
7	0							+971.0	0
8	+0.0074							-0.002	+0.131
8	0								+971.0
8	-0.435								-7.7
9	0	+1.000	+40.68	-1.14	+7.35	-103.8	+22.0	-17.7	-0.8
9(1)	-0.0006			+0.0046	-0.0008	-0.6389	-0.6344	-0.0073	-0.0004
9(2)	+4.1333			-0.0001	-0.0296	-2.993	-1.352	+0.247	-0.001
10	-0.1807	0		+110.18	-0.206	+0.125	+4.24	-8.346	+279.0
11	0	+0.0280		-0.3	+0.206	-2.906	+0.62	-0.496	-0.01
11	-0.0280	0			+111.48	-23.62	+1.194	-282.2	-8.896
12	0	-0.1807			-1.33	+18.76	-3.975	+3.2	+0.54
12	-0.541	0				+354.1	-56.16	+70.16	-9.576
13	0	+2.552				-264.9	+56.144	-45.17	-7.656
13	-2.552	0					+101.1	+0.7656	+24.83
14	0	-0.541					-11.9	+9.576	+0.16
14	+0.0074	0						+971.0	+0.131
15	0	+0.435						-7.7	-0.131
15	-0.435	0							+963.3
15	0	+0.0074							-0.002
16	-0.1807	+0.0280	+1.000	+110.15	0	-2.781	+4.86	-8.842	+279.0
16(1)	-0.4433					+0.0171	-0.1401	-0.0036	+0.3892
16(2)		+0.0116				-0.0802	-0.0299	+0.0123	+0.1142
16(3)			+3.7709			+0.0005	+0.0043	-0.0027	-2.773
17	-0.0280	-0.1807	0		+110.15	-4.86	-2.781	-279.0	-8.842
18	0	0	0		0	0	0	0	0
18	-0.541	+2.552	0			+89.2	-0.016	+24.99	-10.342
19	-0.0046	+0.0007	+0.02525			-0.07	+0.1227	-0.223	+7.045
19	-2.552	-0.541	0				+89.2	+10.342	+24.99
20	+0.008	-0.0012	-0.04412				-0.21	+0.39	-12.31
20	+0.0074	+0.435	0					+963.3	0
21	-0.0145	+0.0022	+0.0803					-7.1	+22.40
21	-0.435	+0.0074	0						+963.3
21	+0.4577	-0.0709	-2.533						-706.7

TABLE LV (Continued).

Equation Number	Right Hand Side				Left Hand Side				
					$\frac{4}{k_8}$	$\frac{5}{10 k_1}$	$\frac{6}{10 k_2}$	$\frac{7}{10 k_3}$	$\frac{8}{10 k_4}$
22	- .028	- .1807	0	+ 1.000	+110.15	- 4.86	- 2.781	-270 0	- 8.849
22(1)	- .0114					+ .0299	+ .0802	- .1144	- .0132
22(2)		- .4436				- .1401	+ .0171	+ .3895	- .0036
22(3)			+ .0002			+ .0008	- .0025	- .0864	+ .0879
22(4)				+ 3.7709		+ .0043	+ .0005	- 2.773	- .0027
23	- .5456	+ 2.5527	+ .02525	0		+ 89.13	+ .1067	+ 24.77	- 3.897
	- .0012	- .0080	0	+ .04412		- .21	- .1227	- 12.31	- .390
24	- 2.544	- .5422	- .0441	0			+ 88.99	+ 10.732	+ 12.68
	- .001	- .0046	+	+ .02525			- .07	- 7.045	- .21
25	- .0071	+ .4372	+ .0808	0				+ 962.6	+ 22.40
	- .0709	- .4577	0	+ 2.533				- 706.7	- 22.40
26	+ .0227	- .0635	- 2.533	0					+ 255.6
	- .0022	- .0145	0	+ .0803					- .71
27	- .5468	+ 2.5447	+ .02525	+ .04412	+ 1.000	+ 88.92	- .016	+ 12.46	- 3.687
27(1)	- .5473						+ .0005	+ .0051	- .0051
27(2)		+ 2.5635					+ .0001	- .0174	- .0018
27(3)			- .01525				0	+ .0039	+ .0366
27(4)				- .0787			0	+ .1239	- .0011
27(5)					+ 1.0075		0	- .0069	- .0006
28	- 2.545	- .5468	- .04412	+ .02525	0		+ 88.92	+ 3.687	+ 12.48
	- .0001	+ .0005	0	+ .00001	+ .00018		0	+ .002	- .001
29	- .078	- .0205	+ .0803	0				+ 255.9	0
	+ .0766	- .3566	- .0035	- .006	- .1401			- 1.75	+ .5165
30	+ .0205	- .078	- 2.533	+ .0803	0				+ 255.9
	- .0227	+ .1055	+ .001	+ .0018	+ .04146				- .15
31	- 2.5451	- .5463	- .04412	+ .02525	+ .00018	+ 1.000	+ 88.92	+ 3.687	+ 12.459
31(1)	- 2.564							+ .0015	+ .0174
31(2)		- .5463						- .0051	+ .0051
31(3)			+ .0786					+ .0011	- .1238
31(4)				- .0153				+ .0367	+ .0039
31(5)					+ .00018			- .0020	+ .0020
31(6)						+ 1.0075		- .0006	- .0069
32	- .0014	- .3771	+ .0768	+ 2.527	- .1401	0		+ 254.15	+ .5165
	- .1026	- .0227	+ .0018	- .001	0	- .04149		- .153	+ .5169
33	- .0022	+ .0275	- 2.532	+ .0821	+ .04146	0			+ 255.75
	+ .3566	+ .0766	+ .006	- .0035	0	- .1401			- 1.75
34	+ .1042	- .3544	+ .0786	+ 2.526	- .1401	- .04149	+ 1.000	+ 254.0	- .0004
34(1)	+ .1042								0
34(2)		- .3544							0
34(3)			+ .0786						0
34(4)				+ 2.526					0
34(5)					- .1401				0
34(6)						- .04149			0
34(7)									0
35	+ .3544	+ .1040	- 2.526	+ .0786	+ .04146	- .1401	+ 0.000		+ 254.0
	0	0	0	0	0	0	+ .000002		0
36	- .3544	+ .1040	- 2.526	+ .0786	+ .04146	- .1401	+ 0.000	+ 1.000	+ 254.0
Case	I	II	III	IV	V	VI	VII	VIII	
Resolution	Case I	k_1	k_4	k_7	k_8	$\frac{1}{10} k_1$	$\frac{1}{10} k_2$	$\frac{1}{10} k_3$	$\frac{1}{10} k_4$
1st approximation.	II	+ .1018	- .00001474	- .004025	- .0001085	- .006155	- .028885	+ .00041	+ .001395
	III		+ .1016	+ .0001053	- .004027	+ .028835	- .001396	- .001396	+ .004094
	IV			+ .03423	+ .00000182	- .00017	+ .000884	+ .0008095	- .00994
	V				+ .03423	- .000875	- .000172	+ .00994	+ .0008095
	VI					+ .01133	+ .00000202	- .000552	+ .000163
	VII						+ .01133	- .0001635	- .000552
	VIII							+ .009937	0
									+ .003937

The first, second, fifth and sixth lines and columns of table XLVII are multiplied by 10 and their rows and columns are arranged in order 3, 4, 7, 8, 1, 2, 5, 6: giving rise to equations whose coefficients are written down in column headed "Left Hand Side 1 . . . 8," and in lines whose equation numbers are 1 to 8. The Right Hand Sides of these equations are 1 | 0 | 0
0 | 1 | 0
0 | 0 | 1
It is however only necessary to write down the first column under heading R.H.S. To eliminate the first quantity k_3 from these 8 equations, the first equation is multiplied successively by $-\frac{0}{40.68}=0$, $-\frac{7.35}{40.68}=-.1807$, $-\frac{1.14}{40.68}=-.280$

and the results written below the 2nd, 3rd, 4th equations in old face type. These multipliers are also applied to the right hand side of the first equation which is unity, and therefore appear in this column. The work is performed by sliderule, making one setting representing division by 40.68, and reading off opposite the quantities 0, 7.35, 1.14 etc. these quantities being at once entered in R.H.S. column (shown in old face type). In completing the multiplication of the first equation by these several factors, it is noticed that the quantity to be entered is the product of the factor of the particular line by the coefficient of the particular column in equation No. 1, *e. g.* old face figures -3.975 (line 3, column 5) $= -.1807 \times 22.0$. The old face figures thus formed for all the equations Nos. 2—8, are added to the corresponding coefficients and give rise to 7 equations (numbered 9—15) from which k_3 has been eliminated. The process as regards the right hand side has only actually been applied for case I in which the right hand side of the first equation is unity and the rest are zero (*vide* § 7): but it will be easily seen that for the other cases all the old face quantities would be zero having zero as a factor. Case II is accordingly brought in conveniently after the elimination of k_3 has been completed.

The necessary multipliers for the next elimination, that of k_4 , are now duly entered of the R.H.S. for case II. They are the old face figures $+0.280$, $-.1807$, $+2.552$. . . : and the process of elimination is proceeded with. In this way the groups of equations 16—21, 22—26, 27—30, 31—33, 34—35, 36 are formed successively in the last of which only $\frac{1}{10} k_4$ occurs. From this values of $\frac{1}{10} k_4$ are written down at the bottom of the table for each of the eight cases. These values would be substituted in 34, except that the coefficient of k_4 is so small that they are negligible. Thus equations 34 (1) . . . 34 (7) are formed giving values of $\frac{1}{10} k_5$ for seven cases. Then the values of $\frac{1}{10} k_5$ and $\frac{1}{10} k_4$ are substituted in 31, giving 31 (1) . . . 31 (6) from which six values of $\frac{1}{10} k_5$ are formed. These several substitutions are shown in old face. The results of the terms on the left hand sides are combined and with changed sign applied to the right hand side for the corresponding case, *e. g.* 31 (3) $+0.786 = -.04412 - .0011 + .1238$. All the values obtained are exhibited at the foot of the table: they are the values of x_s in the notation of § 7. Since the equations are symmetrical the solution is also symmetrical and $x_s = x_r$: so that it is only necessary to take out half of the quantities. Had the equations not been symmetrical all work below the diagonal in each group of eliminants would have had to be completed and the full number of cases, eight, would have been necessary in substituting in each of equations 1, 9, 16, 22, 27, 31, 34, 36.

18. An approximate solution of the first eight equations has now been found. As it will be necessary to substitute from the solution in the remaining equations, it is desirable in any case at this stage to check the solution; but this substitution at the same time enables a higher approximation to be reached by (8). This is desirable; although a high order of accuracy of final solution is not desired, yet it is proper to avoid the introduction of computation inaccuracy at an early stage in the work. The first step is to substitute in the equations and so to find the quantities α, β . . . of § 7. In performing this substitution a sliderule cannot be used, as greater precession is desired. If an arithmometer is used, then there is little extra labour in taking out the work to the full number of figures which occur. When this has been done (10) gives a means of an infinite number of successive approximations. As exemplifying this, full accuracy is kept in this substitution, which is reproduced for Case I only in table LVI. It is to be remarked that there is no advantage in keeping the solution of

table LV to as many figures as has been done, as the latter figures cannot be accurate: and their presence adds to the labour of multiplication. However in the present case the substitution had already been carried out by the computer before this simplification could be given effect to.

TABLE LVI.

Equation	1	2	3	4	5	6	7	8
Case I	+4.188,088,00 0	0	+746,780,00	+115,824,00	+2.235,200	+10.546,080,00	-030,480	+1.798,320,00
	-0.029,588,75	-000,589,823,2	+000,016,808,6	-000,108,889	+0.001,580,012	-0.000,324,28	+000,280,898	+0.000,004,422
	-0.000,117,99	+004,688,85	-448,827,75	0	-0.016,802,5	-0.002,575	+038,810	-1.185,856
	-0.186,410	-000,780,725	0	-011,541,285	+0.002,880,5	-0.000,424,38	+028,207,7	+0.000,869,4
	-2.988,078	+688,889	-025,235,5	+141,585	-2.252,730	0	-430,850	0
	-0.000,133	-634,370	-688,205	+118,288,5	0	-10.558,610	0	-2.018,450
	+0.024,691,5	-007,287	-008,444	-115,702	+0.028,700	0	+398,110	0
Sum	+0.999,471,78	+000,072,001,8	-000,268,446,4	+000,085,876	-0.001,421,988	+0.007,650	+000,058,598	-0.000,566,178

Table LVII gives the results of the substitution for all cases, that is the quantities $\alpha \beta \gamma$. . . in notation of § 7. To avoid constant repetition of zeros, they have all been multiplied by 10^8 . It is now perfectly straightforward to substitute in (10) and obtain a second approximation. Since the solution is known to be symmetrical, it is not necessary to perform the substitution on both sides of the diagonal: but as a check it may sometimes be useful to do so. When two determinations of what is known by symmetry to be one quantity differ slightly, the mean can be taken.

TABLE LVII.

r	1	2	3	4	5	6	7	8
$10^8 \alpha_r$	-528,24	+072,001,8	-268,448,4	+095,876	-1.421,988	-3.208,68	+058,598	-566,178
$10^8 \beta_r$	-128,468,2	-235,112	-019,758	-180,788,6	+2.538,45	-464,812	+153,802	+248,962
$10^8 \gamma_r$	-158,075,2	-127,389	-749,292	+077,903,2	+268,00	-436,688	-077,414	-1.443,378
$10^8 \delta_r$	+217,747	-1.219,784,8	+334,408,2	-977,14	+3.968,062	-230,44	+2.077,762	+842,254
$10^8 \epsilon_r$	-112,024	+482,29	+067,01	+1.172,782	-916	+132,82	-3.072,0	+196,4
$10^8 \zeta_r$	-829,39	-100,006	-160,988	-028,98	+251,92	-474,2	-105,0	-618,7
$10^8 \eta_r$	-012,175	-087,41	+022,785	-359,65	+131,75	-122,5	+1.105,4	+075,7
$10^8 \theta_r$	+015,73	+016,817	+370,834	+028,735	-020,22	+014,75	-099,78	+1.087,88

The result of the second approximation, found by means of (8), is given in table LVIII which has been rearranged in the order of the original quantities.

TABLE LVIII.

	k_1	k_2	k_3	k_4	k_5	k_6	k_7	k_8
1	+1.1825	+000,006	-081,44	+288,135	-055,14	+016,27	-001,702	-008,885
2		+1.132,5	-288,135	-081,44	-016,27	-055,14	+008,835	-001,702
3			+101,552	0	+004,079	+018,938	-004,021	-000,108,5
4				+101,552	-018,938	+004,079	+000,108,5	-004,021
5					+398,68	0	+008,092	+099,41
6						+398,68	-099,41	+008,092
7							+034,241	0
8								+034,241

14. This solution corresponds to that indicated in (13). The next step is to form (15) with a view to substitution for k_1 . . . k_8 in the equations Nos. 9—20. For this only values of k_5 . . . k_8 are required since the coefficients of k_1 . . . k_4 are zero in these equations. However values of k_1 . . . k_4 are also required at a later stage, so the complete series of quantities k_1 to k_8 are expressed

according to (15). Denote the quantities given in table LVIII by ${}_sK_r$ where r and s have all values from 1 to 8. It is necessary to compute all the quantities $\Sigma {}_sK_r(s, t)$ where s has all values from 1 to 8 and is the quantity to which Σ refers, for each value of r from 1 to 8 and each value of t from 9 to 20: (s, t) in the notation of (11) indicate the coefficients shown in table XLVII, while ${}_sK_r$ are the quantities of table LVIII. Owing to zero coefficients it has only to be taken for values 9 to 16 and so the number of quantities $\Sigma {}_sK_r(s, t)$ actually to be computed is $8 \times 8 = 64$; and among them there is a sort of skew symmetry due to equality in pairs of coefficients in table XLVII, making altogether only 32 independent quantities. Further the summation Σ although relating to 8 values of s , actually only gives rise to four terms owing to zero coefficients. The details of the computation are given in full in table LIX. The computation is taken out to full accuracy to exhibit the symmetry which afterwards occurs when substituting in equations Nos. 9—20.

TABLE LIX.

Values of ${}_sK_r(s, t)$ and of $\Sigma {}_sK_r(s, t)$, the latter in old face type.

r	s	$t = 9$	10	11	12	13	14	15	16
1	5	-356,755,80	0	-567,127,80	+348,484,8	-109,728,80	0	-102,000,00	+142,812,80
	6	0	+105,266,80	+102,826,40	+284,712,9	0	+032,377,30	+042,189,80	+030,099,50
	7	+002,246,64	-089,064,84	-050,418,24	-100,588,2	+000,186,16	-002,381,74	-003,488,04	-002,314,72
	8	+202,939,96	+011,662,20	+622,145,50	-281,692,7	+012,103,05	+000,706,80	+012,015,80	-017,546,70
	Sum	-151,569,21	+077,834,16	-322,566,14	+250,916,8	-097,488,49	+030,752,36	-051,292,14	+152,750,68
2	5	-105,266,80	0	-284,712,9	+102,826,40	-032,377,30	0	-030,099,50	+042,189,80
	6	0	-356,755,80	-348,484,8	-827,127,80	0	-109,728,80	-142,812,80	-102,000,00
	7	-011,662,20	+202,939,96	+281,692,7	+522,145,50	-000,706,80	+012,103,05	+017,546,70	+012,015,80
	8	+089,064,84	+002,246,64	+100,588,2	-050,418,24	+002,381,74	+000,186,16	+002,314,72	-003,488,04
	Sum	-077,834,16	-151,569,21	-250,916,8	-322,566,14	-030,752,36	-097,488,49	-152,750,68	-051,292,14
3	5	+026,391,130	0	+086,365,33	-025,779,28	+008,117,210	0	+007,546,15	-010,584,61
	6	0	+000,165,02	+088,075,52	+228,738,72	0	+027,732,04	+088,064,24	+025,781,80
	7	+005,307,720	-092,362,37	+119,102,02	-237,641,10	+000,321,680	-005,503,77	-005,122,42	-005,468,56
	8	+002,492,245	+000,143,22	+006,412,35	-003,213,77	+000,148,645	+000,008,68	+000,147,56	-000,219,17
	Sum	+034,191,095	-002,053,23	+041,751,18	-039,895,43	+008,587,535	+022,232,55	+035,665,53	+009,529,26
4	5	-090,105,92	0	-228,738,72	+088,075,52	-027,732,04	0	-025,781,80	+036,094,24
	6	0	+026,391,130	+025,779,28	+086,365,33	0	+008,117,210	+010,584,61	+007,546,15
	7	-000,143,22	+002,492,245	+003,213,77	+006,412,35	-000,008,68	+000,148,645	+000,219,17	+000,147,56
	8	+092,362,37	+005,307,720	+237,641,10	-119,102,02	+005,503,77	+000,321,680	+005,468,56	-008,122,42
	Sum	+002,053,23	+034,191,095	+039,895,43	+041,751,18	-022,232,55	+008,587,535	-009,529,26	+035,665,53
5	5	+2646,788,10	0	+6404,360,10	-2487,741,6	+783,323,70	0	+728,215,50	-1010,501,70
	6	0	0	0	0	0	0	0	0
	7	-004,081,44	+071,033,24	+001,585,04	+182,737,2	-000,247,36	+004,236,04	+006,245,84	+004,205,12
	8	-2383,447,70	-131,221,20	-5875,131,00	+2044,524,2	-138,191,70	+007,952,80	-185,197,60	+200,808,20
	Sum	+259,256,96	-060,197,96	+620,814,14	+639,519,8	+646,884,64	-003,716,76	+599,263,74	-814,488,38
6	5	0	0	0	0	0	0	0	0
	6	0	+2546,788,10	+2487,741,6	+6404,360,10	0	+783,323,70	+1010,501,70	+728,215,50
	7	+131,221,20	-2383,447,70	-2044,524,2	-5875,131,00	+007,952,80	-138,191,70	-200,808,20	-185,197,60
	8	-071,033,24	-004,081,44	-182,737,2	+001,585,04	-004,236,04	-000,247,36	-004,205,12	+006,245,84
	Sum	+060,197,96	+259,256,96	-639,519,8	+620,814,14	+003,716,76	+646,884,64	-814,488,38	+599,263,74
7	5	+030,005,24	0	+050,306,84	-019,541,44	+006,153,08	0	+008,720,20	-008,008,28
	6	0	-643,182,70	-828,271,20	-1417,400,70	0	-187,825,80	-287,471,90	-188,908,50
	7	-045,108,12	+786,516,77	+1014,218,42	+2023,643,10	-002,739,28	+016,910,17	+060,166,82	+046,567,76
	8	0	0	0	0	0	0	0	0
	Sum	-025,192,88	+143,333,07	+436,254,06	+386,700,96	+003,413,80	-150,915,73	-182,584,88	-145,349,02
8	5	+643,182,70	0	+1617,400,70	-828,271,20	+187,825,80	0	+188,908,50	-287,471,90
	6	0	+020,005,24	+019,541,44	+050,306,84	0	+006,153,08	+008,720,20	+008,008,28
	7	0	0	0	0	0	0	0	0
	8	-786,516,77	-045,108,12	-2023,643,10	+1014,218,42	-046,910,17	-002,739,28	-046,567,76	+089,168,82
	Sum	-143,333,07	-025,192,88	-386,700,96	+436,254,06	+150,915,73	+003,413,80	+145,349,02	-182,584,88

15. Values of k_5, k_6, k_7, k_8 given in (15) contain terms in $k_4 \dots k_{16}$ of which the coefficients have just been found in table LIX. These are to be substituted in equations 9—16, they do not occur in equations 17 to 20. The formation of the products and the collecting of coefficients is carried out in Table LX. In this the values of $\Sigma {}_sK_r(s, t)$ are rewritten with sign changed at the top, kept to six places, while the multiplying coefficients are all shown in the first column. The previously existing

coefficients of k_9 . . . k_{20} are also included. The complete coefficients of k_9 to k_{20} after including the portions due to substitution of k_5 to k_8 , are shown in Table LX in old face. Thus twelve symmetrical equations relating k_9 to k_{20} have been formed, k_1 to k_8 having been eliminated. The solution of these twelve equations is performed in a manner similar to that employed for the solution of the first eight. The equations are first rearranged in increasing order of the diagonal coefficients, the whole process being given in table LXI. It does not seem likely that a second approximation is necessary in this case, so a verification, as described in § 11, is carried out in table LXII showing the degree of precision with which the last equation of the group of 12 is satisfied for each of the 12 cases.

TABLE LX.

Substitution of k_5 — k_8 in equations 9 to 20.

r	t	9	10	11	12	13	14	15	16	17	18	19	20	
5 6 7 8	$\Sigma^* K_r(s,t)$	-259,257 -060,198 +025,183 +143,333	+080,198 -259,257 -143,333 +025,193	-620,814 +639,520 -436,254 +386,701	-639,520 -620,814 -386,701 -436,254	-646,885 -003,717 -646,885 -003,414	+003,717 -646,885 +150,916 -003,414	-599,264 -814,488 +182,585 -145,349	+814,488 -599,264 +145,349 +182,585					
Equation 9	Multiplier + 6.47 0 - 1.32 - 22.97 (8,t) + 9.68 Sum	- 1.6774 0 - .0833 - 3.2924 + 9.68 + 4.6770	+ .3895 0 + .1892 - .5787 0 0	- 4.0167 0 + .5759 + 8.8825 + 16.81 + 4.4867	- 4.1377 0 + .5104 + 10.0208 - 11.33 - 4.9365	- 4.1853 0 + .0045 + 3.4685 + 0.28 - 0.4343	+ .0240 0 - .1992 + .0784 0 - 0.0968	- 3.8772 0 + .2410 + 3.3387 + 0.12 - 0.2838	+ 5.2697 0 - .1919 - 4.1940 0.60 + 0.2838	+ 0.55	0	- 0.27	- 1.66	
10	0 + 6.47 + 22.97 - 1.32 (10,t) Sum	0 - .3895 + .5787 - .1892 0 + 4.4867	0 - 1.6774 - 3.2924 + .0833 + 9.68 + 4.6770	0 + 4.1377 - 10.0208 + .0333 + 11.33 + 4.9365	0 - 4.0167 - 8.8825 + .5759 + 16.81 + 4.4867	0 + .0240 - .0784 + .1992 + .0045 - 0.28 - 0.0968	0 - 4.1853 + 3.4685 - .1992 + .0784 + .28 - 0.4343	0 - 5.2697 + 4.1940 - .1919 + .2410 + .60 - 0.2838	0 - 3.8772 + 3.3387 - 4.1940 0.60 + 0.2838	0	0	0	0	
11	+ 16.27 + 6.32 + 29.83 - 59.10 (11,t) Sum	- 4.2181 - .3804 + .7462 - 8.4710 + 16.81 + 4.4867	+ .9794 - 1.6385 - 4.2455 - 1.4889 + 11.33 + 4.9365	- 10.1008 - 4.0418 - 12.0218 - 22.8510 + 79.30 + 37.4654	- 10.4050 - 3.9235 - 11.4541 - 25.7828 0 - 1.2303	+ 3.0418 - 10.1008 - 22.8510 - 12.0218 + 79.30 + 37.4654	+ 4.0883 - .0235 + .1011 + 8.6191 + .61 - 1.2303	- .0235 + 4.0883 + 4.4701 + 8.6191 + .28 + 0.3640	+ 3.8778 - 5.1478 + 5.4082 - 10.7908 - .32 + 1.2193	+ 5.1478 - 3.8778 + 5.4082 - 10.7908 + 1.44 + 1.5388	+ 0.48	- 0.21	- 0.71	- 1.36
12	- 6.32 + 16.27 + 59.10 + 29.83 (12,t) Sum	+ 1.6385 - .9794 + 1.4889 + 4.2455 - 11.33 - 4.9365	- .3804 - 1.6385 - 8.4710 - 7.462 + 16.81 + 4.4867	+ 3.9235 + 10.4050 - 25.7828 + 11.4541 0 + 37.4654	+ 3.0418 - 10.1008 - 22.8510 - 12.0218 + 79.30 + 37.4654	+ 4.0883 - .0235 - 20.18 - 4.4701 + .28 - 0.3640	- .0235 + 4.0883 + 8.6191 - .1011 + .61 + 1.2303	+ 3.8778 - 5.1478 + 5.4082 - 10.7908 + 1.44 + 1.5388	+ 5.1478 - 3.8778 + 5.4082 - 10.7908 + 1.44 + 1.5388	+ 0.21	+ 0.48	+ 1.36	- 0.71	
13	+ 1.09 0 - 0.08 - 1.37 (13,t) Sum	- .5159 0 - .0020 - .1984 + .28 - 0.4343	+ .1198 0 + .0115 - .0945 0 + 0.0968	- 1.2354 0 + .0349 - .5298 + .61 - 1.2303	- 1.2726 0 + .0309 - .5977 + .28 - 0.3640	- 1.2873 0 + .0003 + .2068 + 3.43 + 2.3497	+ .0074 0 - .0121 - .0047 0 + 1.2320	- 1.1925 0 + .0146 - .1991 + 2.13 + 1.2200	+ 1.6208 0 - .0116 - .2501 + 4.28 + 2.9209	+ 0.83	0	+ 0.32	- 1.20	
14	0 + 1.09 + 1.37 - 0.08 (14,t) Sum	0 + .1198 + .0845 + .0116 0 - 0.0968	- .5159 + .1114 - .0845 - .0020 + .28 - 0.4343	+ 1.2726 - 1.2354 - .5977 - .0309 + .28 + 0.3640	- 1.2354 0 + .5298 - .0349 + .61 - 1.2303	- 1.2873 0 - .0047 + 0.2068 + .0121 + 3.43 + 2.3497	0 - 1.2873 + 0.2068 - .0003 + 3.43 + 2.3497	- 1.8208 + 1.6208 + .2501 + 0.116 + 4.28 + 2.9209	- 1.1925 + 1.6208 - .0116 - .2501 + 4.28 + 2.9209	0	0	0	0	
15	+ 1.85 + 2.59 + 2.02 - 1.36 (15,t) Sum	- .4796 + .1559 + .0509 + .1049 + .12 - 0.6595	+ .1114 - .8715 + .2895 + .0343 + .60 - 0.2839	- 1.1485 + 1.6564 - .8812 + .5933 + 1.44 + 1.5388	- 1.1831 - 1.6079 - .7811 + .2052 + 2.13 + 1.2200	- 1.1967 + .0089 - .0089 + .2052 + 2.13 + 1.2200	+ .0089 - 1.6754 - .3048 + .0046 + 4.28 + 2.9209	- 1.1086 - 2.1095 + .3698 + .1977 + 8.46 + 5.8084	- 1.6068 + 1.5521 + .2636 - .2483 0 0	+ 0.01	+ 0.99	+ 1.83	+ 0.38	
16	- 2.59 + 1.85 + 1.36 + 2.02 (16,t) Sum	+ .4796 - .1114 - .0343 + .2895 + .80 + 0.2839	- .1559 + .8715 - .2895 + .0343 + .12 - 0.6595	+ 1.6079 + 1.1831 - .5933 + .7811 + 1.44 + 1.5388	+ 1.6564 - 1.1485 + .5933 - .8812 + .32 - 1.2192	+ 1.6754 - .0089 - .2052 + .3048 + 2.13 + 2.9210	- .0089 + 1.6754 + .0046 - .3048 + 4.28 + 2.9210	+ 1.5521 - 1.5068 + .2483 - .2636 + 8.46 + 5.8084	- 2.1095 + 1.5068 + .1977 - .2483 + 8.46 + 5.8084	- 0.99	+ 0.01	- 0.38	+ 1.83	
17	(17,t)	+ 0.55	0	+ 0.48	+ 0.21	+ 0.83	0	+ 0.01	- 0.99	+ 2.23	0	- 0.56	- 4.05	
18	(18,t)	0	+ 0.55	- 0.21	+ 0.48	0	+ 0.83	+ 0.99	+ 0.01	0	+ 2.23	+ 4.05	- 0.56	
19	(19,t)	- 0.27	+ 1.66	- 0.71	+ 1.36	+ 0.32	+ 1.20	+ 1.83	- 0.38	- 0.56	+ 4.05	+ 10.96	0	
20	(20,t)	- 1.66	- 0.27	- 1.36	- 0.71	- 1.20	+ 0.32	+ 0.38	+ 1.83	- 4.05	- 0.56	0	+ 10.96	

TABLE LXI.—(Continued).

Equation Number	Right Hand Side.				Left Hand Side.								
					k ₁₄	k ₈	k ₁₀	k ₁₂	k ₁₆	k ₁₈	k ₂₀	k ₂₂	k ₂₄
35	- .3823	0	+ .311	+ .049		+4.3486	+ .0311	- .3154	- .2693	+ .0324	- .5655	+ 3.0077	- 5.1283
	0	- .01824	- .2468	0		- .0049	- .0311	+ .1250	+ .0547	+ .01506	+ .0250	+ .0215	- .0036
36	+ .01824	- .2468	- .049	0			+4.3391	- .5789	- .5375	- .6352	- .14006	+ 5.0586	+ 4.8831
	0	- .1157	0	+ .311			- .1974	+ .7935	+ .3471	- .0956	+ .1643	+ .1303	- .4030
37	+ .199116	- .4439	- .547	0				+4.76	+ .13555	- .2545	+ .4787	- .41885	- 1.51434
	0	+ .4653	0	- .125				-3.19	- .13955	+ .3843	+ .6606	- .5478	+ .10235
38	- .0214	- .004484	+ .125	0					+2.181	+ .0138	+ .4188	+ .13054	- 1.07075
	0	+ .2036	0	- .547					- .6105	+ .1681	+ .280	+ .2397	+ .7103
39	+ .3475	- .1816	- .2589	0						+ 3.3276	- .0708	+ .1281	+ .0544
	0	- .0561	0	+ .1506						- .0463	+ .0706	+ .0660	+ .1056
40	+ .18721	+ .2511	- .1506	0							+ 3.4181	- .3454	- .1421
	0	+ .0964	0	- .2589							- .1308	- .1134	+ .3302
41	- .452	+ .09417	+ .6363	0								+30.5205	- .2748
	0	+ .0799	0	- .2147								- .0041	+ .2748
42	- .17407	- .2152	+ .2147	0									+37.2522
	0	- .2368	0	+ .6363									- .8204
43	- .3823	- .01824	+ .311	+ .049	+1.00	+4.3417	0	- .1904	- .2146	+ .01734	- .5300	- 3.0002	- 5.1010
	- .12764					(1)	0	- .03578	+ .01007	+ .00187	- .31364	- .01870	- .07503
		+ .07423	+ .71022	- .1041		(2)	0	- .01692	- .04033	+ .01008	- .05834	- .05705	- .02445
						(3)	0	+ .06487	+ .1876	- .00217	- .04183	- .04215	- .2730
						(4)	0	+ .16644	+ .07311	- .001344	- .01707	- .21015	- .05408
						(5)	0	- .00408	- .01512	- .00020	- .02420	- .1340	- .25308
44	+ .01824	- .3823	- .049	+ .311	0	+4.3417	+ .2146	- .1904	+ .5306	+ .01734	+ 5.1010	+ 3.0002	+ 5.1010
	0	- .01824	- .547	- .125	0	0	0	0	0	0	0	0	0
45	+ .199116	+ .0214	- .547	- .00215	+ .04385		+ .157	0	+ .1298	- .1810	- .00045	+ .10818	+ .10818
	- .01589	- .0008	+ .01364	0	0		- .00835	- .0004	+ .00076	- .0337	+ .1740	+ .2277	+ .2277
46	- .0214	+ .199116	+ .125	+ .0024	+ .04943			+ .1570	+ .1208	+ .1208	- .10010	- .10010	- .10010
	- .0179	- .00001	+ .0154	+ .0024	+ .04943			- .0106	+ .0009	- .0267	+ .1072	+ .2800	+ .2800
47	+ .3475	- .18721	- .2589	+ .1506	0				+ 3.2813	0	+ .1011	+ .1548	+ .1548
	+ .00145	+ .0001	- .00124	- .0002	+ .003995				- .0001	+ .00215	- .01503	+ .0707	+ .0707
48	+ .18721	+ .3475	- .1506	- .2589	0					+ 3.2813	- .4684	+ .1041	+ .1041
	- .0450	- .0023	+ .03865	+ .0061	+ .1242					- .0170	+ .4050	+ .6450	+ .6450
49	- .452	+ .17407	+ .6363	- .2147	0						- 3.064	+ 4.700	+ 4.700
	+ .333	+ .01675	- .2589	- .045	- .9188							+30.4264	+ 6.2040
50	- .17407	- .452	+ .2147	+ .6363	0								
	- .4323	- .0218	+ .3713	+ .0585	+ .196								
51	+ .01824	- .3823	- .049	+ .311	0	+1.00	+4.3417	+ .2146	- .1904	+ .5306	+ .01734	+ 5.1010	+ 3.0002
	- .07423	- .12764			0		(1)	+ .04033	+ .01602	+ .05834	+ .01008	- .02445	- .05705
			+ .1041	+ .71022	0		(2)	+ .01007	- .03578	- .31364	+ .00187	- .01870	- .07503
					0		(3)	- .07311	+ .16644	- .06707	- .001344	- .05408	- .21045
					0		(4)	+ .1876	+ .06487	+ .04183	- .00217	- .2730	- .04215
					0		(5)	+ .00408	- .01341	- .00009	- .00078	- .1808	+ .1040
					0		(6)	- .01512	- .00408	- .02420	- .00020	- .25308	- .1340
52	+ .13323	+ .0206	- .5334	- .124785	+ .04385	0	+1.5616	- .0084	+ .1336	- .2056	- .7015	- .1185	- .1185
	- .00090	+ .0179	+ .0024	- .0154	0	- .04943	- .0106	+ .0004	+ .0267	- .0009	- .2800	- .0707	- .0707
53	- .0893	+ .1982	+ .12654	- .5446	+ .0484	0		+1.5504	+ .1298	- .1031	- .1084	- 1.22705	- 1.22705
	+ .0008	- .0159	- .00215	+ .01364	+ .04385	0		- .00835	+ .0237	+ .00076	+ .2277	+ .1740	+ .1740
54	+ .840	- .1872	- .2601	+ .1504	- .00399	0			+ 3.2812	+ .00215	+ .1740	+ .1740	+ .1740
	- .0023	+ .045	+ .0061	- .0386	0	+ .1242			- .0670	- .00215	- .6450	- .4050	- .4050
55	+ .1827	+ .3452	- .112	- .2528	+ .1242	0				+ 3.2143	+ .1088	- .4540	- .4540
	- .0001	+ .00145	+ .0002	- .00124	- .00399	0				- .0001	- .0207	- .0150	- .0150
56	- .119	+ .1908	+ .3505	- .2597	- .9188	0					+32.7024	+ 4.700	+ 4.700
	- .0218	+ .4323	+ .0585	- .3713	0	- .196					- 6.2040	+ 30.2178	+ 30.2178
57	- .60687	- .4738	+ .588	+ .6948	+ .196	0							
	- .01675	+ .333	+ .045	- .2589	0	- .9188							
58	+ .1823	+ .0885	- .531	- .126325	+ .04385	- .0484	+1.00	+1.551	0	+ .1030	- .2056	- 1.0441	- .3157
	+ .29144	+ .13785	- .52842	- .135583	+ .03326	- .10925		(1)	+ .01123	+ .01200	- .004930	+ .004582	+ .004582
								(2)	0	- .00030	- .03232	- .01515	- .001487
								(3)	0	+ .01303	+ .01001	+ .0111	- .01001
								(4)	0	+ .00085	+ .0250	+ .05820	+ .003343
								(5)	0	- .00117	- .00013	+ .03040	- .01543
								(6)	0	- .00068	- .00233	+ .05121	- .01009
								(7)	0	- .00186	- .00014	- .0274	- .002772
59	- .0835	+ .1823	+ .126325	- .531	+ .0484	+ .04385	0	+ 1.551	+ .2005	+ .1030	+ .3157	- 1.0441	- .3157
	0	- .1827	- .254	+ .1118	- .00399	- .1242	0	0	+ 3.2142	0	- .4684	+ .0161	+ .0161
60	+ .8467	- .0122	+ .0026	+ .0356	+ .0847	+ .00294	- .06703				+ .0702	+ .00210	+ .00210
	+ .1827	+ .3467	- .1118	- .254	+ .1242	- .00399	0			+ 3.2142	+ .0161	- .4684	- .4684
61	+ .0243	+ .0051	+ .0716	- .1681	+ .00584	+ .00657	+ .1331			- .0275	- .1305	- .0420	- .0420
	- .1408	+ .6281	+ .409	- .681	- .9188	- .196	0				+26.5538	0	0
62	+ .1231	+ .0260	+ .3588	- .853	+ .0206	- .0334	+ .6754				- .7078	- .2133	- .2133
	- .6281	- .1408	+ .681	+ .409	+ .196	- .9188	0					+26.5538	+ 26.5538
63	+ .0371	+ .0078	- .108	- .257	+ .0089	- .0101	+ .2034					- .0442	- .0442

TABLE LXI.—(Concluded).

Equation number	Right Hand Side.									Left Hand Side.				
										k_{18}	k_{19}	k_{20}	k_{11}	k_{12}
64	- .0885 + .13785	+ .1823 + .29144	+ 1.26325 + 1.35583	- .531 - .52842	+ .0494 + .10925	+ .04385 + .03326	0 0	+1.00 +1.04117	+1.551 (1) (2) (3) (4) (5) (6) (7) (8)	+ .2065 + .02232 + .120 + .0259 + .01601 + .00233 + .0093 + .00309 + .00914	+ .1089 + .00639 + .01123 + .00805 + .01303 + .00468 + .00117 + .00460 + .00186	+ .3157 + .001487 + .004562 + .003343 + .01666 + .01099 + .01543 + .00825 + .002772	- 1.0181 + .01515 + .004936 + .05529 + .01110 + .05121 + .03649 + .00920 + .0274	
65	+ .3345 + .0051	- 1.8296 + .0243	- .2184 + .1681	+ .1985 + .0716	- .00693 + .00657	- .1209 + .00584	- .06703 0	0 - .1331	+3.2072 - .0275	+ .01884 + .01324 +3.1867 + .0070	- .3986 + .0420 + .1234 + .02116	+ .00506 + .1395 + .5088 + .0702		
66	+ 1.8518 + .0026	+ .3518 + .0122	- .1834 + .0847	+ .4221 + .0356	+ .13. 01 + .0033	- .01056 + .00294	+ .1381 0	- .06703 0						
67	+ .0177 + .0078	+ .0491 + .0371	+ .0502 + .257	- 1.484 + .108	- .8892 + .0101	- 1.2294 + .0089	+ .6754 + .2034	0 0						
68	- .586 + .026	- .133 + .1231	+ .523 + .853	+ .152 - .3588	+ 1.2049 + .0334	- .9289 + .0296	+ .2034 0	+ .6754						
69	+ .3396 + .34375	- 1.8539 - 1.8482	- .3865 - .39877	+ .2681 + .24649	- .0135 - .03583	- .12674 - .14314	- .06703 - .05683	- .1331 - .14073	+1.00 +1.00261	+3.1797 (1) (2) (3) (4) (5) (6) (7) (8) (9)	0 0 0 0 0 0 0 0 0	- .4386 + .002065 + .006338 + .004045 + .02314 + .01527 + .02143 + .01147 + .003851 + .00235	+ .14456 + .002089 + .000681 + .007626 + .001531 + .007063 + .005034 + .001269 + .003779 + .000250	
70	+1.8539 0	+ .3396 0	- .2681 0	- .3865 0	+ .12674 0	- .0135 0	+ .1331 0	- .06703 0	0 0 0	+3.1797 0	- .14456 +25.7818 + .0065	- .4386 0 + .01995		
71	- .0099 + .0468	+ .612 + .2558	- .2068 + .0533	- 1.376 + .037	- .8993 + .00186	- 1.2383 + .0175	+ .6754 + .00924	- .2034 + .01836	+ .13795 + .0754					
72	- .0154 + .0154	+ .0842 + .0176	+ .1.376 + .0176	- .0122 + .0006	+ .00576	+ .00305	+ .00605	- .04545						
73	+ 1.8539 + 1.8482	+ .3396 + .34375	- .2681 - .24649	- .3865 - .39877	+ .12674 + .14314	- .0135 - .03583	+ .1331 + .14073	- .06703 - .05683	0 +1.00261	+1.00 +3.1797 (1) (2) (3) (4) (5) (6) (7) (8) (9) (10)	- .14456 + .006338 + .002089 + .001531 + .007626 + .005034 + .007063 + .003851 + .001269 + .003779 + .00078 + .000256	- .4386 + .006338 + .002089 + .001531 + .007626 + .005034 + .007063 + .003851 + .001269 + .003779 + .00078 + .000256		
74	+ .0360 + .0842	+ .3562 + .0154	- .2601 + .0122	- 1.339 + .0176	- .90116 + .00576	- 1.2558 + .0006	+ .66616 + .00605	- .22276 + .00305	+ .13795 0	0 + .04545	+25.7213 + .0066	+ .01985 + .01995		
75	- .6274 + .2558	+ .0743 + .0468	+ 1.8936 + .0370	- .219 + .0533	+ 1.2389 + .0175	- .89354 + .00186	+ .20645 + .01836	+ .68145 + .00924	- .04545 0	+ .13795		- .0605		
76	+ .1211 0	+ .3716 0	- .2723 0	- 1.3566 0	- .8954 0	- 1.2564 0	+ .6722 0	- .2258 0	+ .13795 0	+ .04545 0	+1.00 0	+25.7147 0		
77	- .3716 0	+ .1211 0	+ 1.3566 0	- .2723 0	+ 1.2564 0	- .8954 0	+ .2259 0	+ .6722 0	- .04545 0	+ .13795 0	0 0	+25.7147 0		
78	- .3716 0	+ .1211 0	+ 1.3566 0	- .2723 0	+ 1.2564 0	- .8954 0	+ .2258 0	+ .6722 0	- .04545 0	+ .13795 0	0 0	+25.7147 0		
		k_{17}	k_{18}	k_{19}	k_{14}	k_{15}	k_{10}	k_{13}	k_{11}	k_{19}	k_{20}	k_{11}	k_{12}	
Case I	+1.6926	0	- .5202	- .33746	- .0294	- .0171	+ .1879	- .08888	+ .1081	+ .58125	+ .004709	- .01445		
II	+1.6926	+1.6926	+ .83746	- .5202	+ .0171	- .0294	+ .1879	+ .08888	+ .1081	+ .58125	+ .004709	+ .01445	+ .004709	
III			+1.8675		+ .1636	+ .02398	- .3407	+ .87417	- .1254	- .07752	+ .01059	+ .052756		
IV				+1.8675	+ .02398	+ .1636	- .87417	- .3407	+ .07752	- .1254	- .052756	+ .01059		
V					+ .3308	0	+ .02144	+ .07044	- .01127	+ .04502	- .03482	+ .04886		
VI							+ .07044	+ .02144	- .04502	- .01127	- .04886	+ .03482		
VII							+ .6713	0	- .04426	+ .02614	+ .008781	+ .02614		
VIII								+ .6713	- .04426	- .01787	+ .008781	+ .02614		
IX									+ .3153	+ .0053646	+ .0017675	+ .0017675		
X										+ .3153	+ .0017675	+ .0053646		
XI											+ .038888	0		
XII												+ .038888		

* As the coefficient of k_{12} vanishes no elimination is required, and this equation gives k_{11} direct.

TABLE LXII.

Verification of solution of 12 equations.

+ .8554400	0	- .1092420	- .0708888	- .0061740	- .0095910	+ .0394590	- .0186648	+ .0227010	+ .1220825	+ .00098889	- .00303450
0	+ .8124480	+ .1619808	- .2496960	+ .0082080	- .0141120	+ .0420624	+ .0901920	- .2790000	+ .0518880	+ .00898900	+ .00226032
+ .1872720	- .1214856	- .6728000	0	- .0588660	- .0089328	+ .1226520	- .8147012	+ .0451440	+ .0279072	+ .00381240	- .01899216
+ .3779552	+ .5826240	0	- 2 .0916000	+ .0298576	- .1832320	+ .9790704	+ .8915840	- .0868224	+ .1404480	+ .05908672	+ .01186080
+ .1452360	- .0844740	- .8081840	+ .1184612	- .6341520	0	- .1058138	- .8479738	+ .0556738	- .2323988	+ .17201080	- .24186340
- .0767790	- .1320060	+ .1076702	+ .7345640	0	+ 1 .4852920	- .3162756	+ .0962656	- .2021398	- .0508023	- .21988140	- .15634180
- .2989600	- .1368752	+ .5246780	+ 1 .3463218	- .0830176	+ .1084776	- 1 .0388020	0	+ .0275198	- .0681604	- .04025560	- .01352274
+ .1094396	- .2262397	- 1 .0864874	+ .4156540	- .0859368	- .0261568	0	- .8189860	+ .0584972	+ .0218014	+ .01071382	- .03189090
+ .1470160	- .7905000	- .1705440	+ .1054272	- .0153272	- .0612272	- .0243032	- .0801936	+ .4288080	0	+ .00729586	- .00240380
- .4126875	- .0767510	+ .0550892	+ .0890940	- .0810642	+ .0080017	- .0814246	+ .0126877	0	- .2238630	- .00125493	- .00380887
- .5414415	0	0	0	0	0	0	0	0	0	0	0
- .5414415	+ .17844628	+ 1 .67676732	- .8980878	+ 1 .8307842	- 1 .8047064	+ .82602407	+ .9794658	- .06622823	+ .20101156	0	+ 1 .45713396
+ .0010848	+ .00018843	- .00062188	+ .0004903	+ .0003820	+ .0001141	- .00114887	- .0003241	- .00034663	+ .00009416	+ .00015156	+ .99989141

16. The residuals in table LXII are sufficiently small. Accordingly the values of $k_9 \dots k_{20}$ have been found satisfactorily for the 12 latter cases. It remains to find their values for the first 8 cases, and also values of k_1 to k_8 for all cases. In this the work is much simplified by the known symmetry of the solution. The introduction of cases 1 to 8 —i.e. giving the R.H.S. of the first 8 equations values 1, 0 . . . , (case 1) 0, 1, . . . (case 2) etc. causes the R.H.S. of the latter 12 equations to take the values $-\Sigma^s (s, t) {}_r K_s$ for case r and equation t, s being taken from 1 to 8: and these quantities accordingly have to be found for each of the cases 1 to 8. Values of these quantities with sign reversed have already been given in table LIX. It is necessary then to combine these cases 9 to 20 in such a way as to give the solution for k_9 to k_{20} for these related cases. For case r the value of k_u is ${}_r k_u = -\Sigma^t {}_t k_u \left\{ \Sigma^s (s, t) {}_r K_s \right\}$, t being given all values from 9 to 20: but in fact $\Sigma^s (s, t) {}_r K_s$ vanishes for values of t above 16. The process is carried out in table LXIII.

The next step is to find ${}_u k_u$ for values of t and u from 1 to 8. Having found values of ${}_u k_u$ for all values of u and values of t from 9 to 20, it is possible to write down values of ${}_u k_t$ by symmetry. Equation (15) then enables the remaining quantities to be found as is done in table LXIV. The symmetry occurring largely simplifies the process while still affording a check.

This leads up to the solution of the combined 20 equations for all the 20 fundamental cases. The results of the solution, compiled from tables LXI, LXIII, LXIV, are given in table LXV. The solution of these 20 equations enables the probable errors of the N.W. Quadrilateral after adjustment of circuit conditions only, to be written down. By the incorporation of the next three equations, corresponding results can be given for the case when circuits and base line closures have been made (the actual adjustment carried out). Finally by incorporation of the last three equations the corresponding results obtainable if Laplace closures were introduced at each extra base can be given. Accordingly before giving the application of table LXV, the further solution of 23 and 26 equations will be carried out.

17. In table LXV are also shown certain multipliers. They are the coefficients of k_{21}, k_{22}, k_{23} in table XLVII. It is necessary to find k_u in terms of k_{21}, k_{22}, k_{23} , for all values of u between 1 and 20 by means of (15), with a view to substituting in equations 21, 22, 23. For this values of $\Sigma^s {}_s k_r (s, t)$ are required, where ${}_s k_r$ are the values given in table LXV and t has values 21, 22, 23, so that (s, t) are the multipliers just alluded to. Each of these products ${}_s k_r (s, t)$ are given in table LXVI, and their sums $\Sigma^s {}_s k_r (s, t)$ for values of s from 1 to 20, the latter in old face type. These are the coefficients of $-k_u$ in the expressions for $k_1 \dots k_{20}$ as found from the first 20 equations. These quantities have to be substituted in equations 21, 22, 23, and accordingly multiplied by the respective coefficients. The process is carried out in table LXVII, where the coefficients of the three equations giving k_{21}, k_{22}, k_{23} , are formed. The process has been carried out with full accuracy to illustrate the complete symmetry of the resulting equations. The solution of these equations is very simple and is given in table LXVIII, with verification at the foot of the table.

TABLE LXIII.

x	t	9	10	11	12	13	14	15	16	17 to 20	u
1 2 3 4 5 6 7 8	1	+ .1516	— .0778	+ .3226	— .2500	+ .0976	— .0808	+ .0613	— .1628	Columns 17 to 20 are not required	14
	2	+ .0778	— .1516	+ .2609	— .8226	+ .0908	— .0976	+ .0613	— .0613	only zeros	15
	3	+ .0842	— .0921	+ .0418	— .0899	— .0098	— .0232	— .0857	— .0985		16
	4	— .0021	— .0842	— .0899	— .0418	— .0232	— .0098	— .0857	— .0985		17
	5	— .2500	— .0778	— .3226	— .8226	— .0908	— .0976	— .0613	— .0613		18
	6	— .0808	— .0976	— .0613	— .0613	— .0613	— .0613	— .0613	— .0613		19
	7	— .0098	— .0232	— .0857	— .0985	— .0985	— .0985	— .0985	— .0985		20
	8	— .0613	— .0613	— .0613	— .0613	— .0613	— .0613	— .0613	— .0613		
9 10 11 12 13 14 15 16 17 18 19 20	9	+ .3908	0	— .0348	+ .0480	— .1886	— .0240	+ .0214	+ .0704	K_n to K_{20} are not required	15
	10	— .0348	0	— .0398	— .0480	— .0240	— .1886	— .0214	+ .0704		16
	11	— .0498	— .0348	0	— .0588	0	— .0106	— .0828	— .0281		17
	12	— .1698	— .0498	0	— .0888	+ .0106	— .0698	— .0106	+ .0408		18
	13	— .0240	— .1698	— .0498	— .0888	— .0106	— .0698	— .0106	+ .0408		19
	14	— .0214	— .0240	— .0528	— .0281	— .0106	— .0698	— .0106	+ .0408		20
	15	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214		
	16	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214	— .0214		
1 2 3 4	1	+ .05015	0	— .01128	— .01227	+ .01896	— .00744	+ .01110	— .01076	— .15834	17
	2	— .01128	0	— .00878	— .01227	— .00878	— .00744	+ .01110	— .01076	— .15834	18
	3	— .00878	0	— .00445	— .01095	— .00445	— .00744	+ .01110	— .01076	— .15834	19
	4	— .00445	0	— .00130	— .00920	— .00130	— .00744	+ .01110	— .01076	— .15834	20
	5	— .00130	0	— .00000	— .00828	— .00000	— .00744	+ .01110	— .01076	— .15834	
	6	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	7	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	8	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
9 10 11 12 13 14 15 16 17 18 19 20	9	+ .05015	0	— .01128	— .01227	+ .01896	— .00744	+ .01110	— .01076	— .15834	17
	10	— .01128	0	— .00878	— .01227	— .00878	— .00744	+ .01110	— .01076	— .15834	18
	11	— .00878	0	— .00445	— .01095	— .00445	— .00744	+ .01110	— .01076	— .15834	19
	12	— .00445	0	— .00130	— .00920	— .00130	— .00744	+ .01110	— .01076	— .15834	20
	13	— .00130	0	— .00000	— .00828	— .00000	— .00744	+ .01110	— .01076	— .15834	
	14	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	15	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	16	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
1 2 3 4	1	+ .05015	0	— .01128	— .01227	+ .01896	— .00744	+ .01110	— .01076	— .15834	17
	2	— .01128	0	— .00878	— .01227	— .00878	— .00744	+ .01110	— .01076	— .15834	18
	3	— .00878	0	— .00445	— .01095	— .00445	— .00744	+ .01110	— .01076	— .15834	19
	4	— .00445	0	— .00130	— .00920	— .00130	— .00744	+ .01110	— .01076	— .15834	20
	5	— .00130	0	— .00000	— .00828	— .00000	— .00744	+ .01110	— .01076	— .15834	
	6	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	7	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	8	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
9 10 11 12 13 14 15 16 17 18 19 20	9	+ .05015	0	— .01128	— .01227	+ .01896	— .00744	+ .01110	— .01076	— .15834	17
	10	— .01128	0	— .00878	— .01227	— .00878	— .00744	+ .01110	— .01076	— .15834	18
	11	— .00878	0	— .00445	— .01095	— .00445	— .00744	+ .01110	— .01076	— .15834	19
	12	— .00445	0	— .00130	— .00920	— .00130	— .00744	+ .01110	— .01076	— .15834	20
	13	— .00130	0	— .00000	— .00828	— .00000	— .00744	+ .01110	— .01076	— .15834	
	14	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	15	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	16	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
1 2 3 4	1	+ .05015	0	— .01128	— .01227	+ .01896	— .00744	+ .01110	— .01076	— .15834	17
	2	— .01128	0	— .00878	— .01227	— .00878	— .00744	+ .01110	— .01076	— .15834	18
	3	— .00878	0	— .00445	— .01095	— .00445	— .00744	+ .01110	— .01076	— .15834	19
	4	— .00445	0	— .00130	— .00920	— .00130	— .00744	+ .01110	— .01076	— .15834	20
	5	— .00130	0	— .00000	— .00828	— .00000	— .00744	+ .01110	— .01076	— .15834	
	6	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	7	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	
	8	— .00000	0	— .00000	— .00744	— .00000	— .00744	+ .01110	— .01076	— .15834	

TABLE LXIV.

t	9	10	11	12	13	14	15	16			
r=1	+ .1516	-.0778	+ .8226	-.2509	+ .0975	-.0908	+ .0513	-.1528	$\Sigma^* K_s(t)$ from table LIX.		
2	+ .0778	+ .1516	+ .2509	+ .8226	+ .0308	+ .0975	+ .1528	+ .0513			
3	-.0942	+ .0021	-.0418	+ .0899	-.0036	-.0222	-.0357	-.0085			
4	-.0021	-.0342	-.0399	-.0418	+ .0222	-.0036	+ .0095	-.0357			
5	-.2593	+ .0602	-.6208	-.6395	-.6469	+ .0037	-.5993	+ .8145			
6	-.0602	-.2593	+ .6395	-.6208	-.0037	-.6469	-.8145	-.5993			
7	+ .0252	-.1433	-.4383	-.3887	-.0034	+ .1509	+ .1826	+ .1453			
8	+ .1433	+ .0252	+ .3887	-.4383	-.1509	-.0034	-.1826	+ .1453			
u=1	+ .0337	-.0424	+ .0144	+ .0023	+ .0373	-.0811	+ .0431	-.0072	tK_u or uK_t		
2	+ .0424	+ .0337	-.0023	+ .0144	+ .0811	+ .0373	+ .0072	+ .0431			
3	-.0102	-.0002	-.0001	-.0010	-.0132	-.0041	-.0032	-.0073			
4	+ .0002	-.0102	+ .0010	-.0001	+ .0041	-.0132	+ .0073	-.0032			
5	-.1568	+ .1172	-.0342	-.0579	-.3600	+ .3090	-.2168	-.0482			
6	-.1172	-.1568	+ .0579	-.0342	-.3090	-.3600	+ .0482	-.2168			
7	+ .0146	+ .0023	-.0153	-.0052	+ .0433	+ .0758	-.0123	+ .0366			
8	-.0023	+ .0146	+ .0052	-.0153	-.0758	+ .0433	-.0366	-.0123			
u	r	$tK_u \Sigma^* K_s(t)$							ΣK_u from LVIII.	rK_u	
1	1	+ .00511	+ .00330	+ .00464	-.00058	+ .00364	+ .00250	+ .00221	+ .00110	+ 1.13850	= + 1.15442
	2	+ .00362	-.00643	+ .00341	+ .00074	+ .00115	-.00791	+ .00658	-.00037	+ .00001	= 0
	3	-.00115	-.00009	-.00060	+ .00009	-.00032	+ .00180	-.00154	+ .00007	-.06144	= - .06318
	4	-.00007	+ .00145	-.00057	-.00010	+ .00083	+ .00070	+ .00041	+ .00026	+ .28814	= + .29104
	5	-.00674	-.00255	-.00894	-.00147	-.02413	-.00030	-.02582	-.00586	-.05514	= - .13294
	6	-.00203	+ .01100	+ .00921	-.00143	-.00014	+ .05247	-.03510	+ .00431	+ .01627	= + .05454
	7	+ .00085	+ .00608	-.00628	-.00089	-.00013	-.01224	+ .0787	-.00105	-.00170	= + .00748
	8	+ .00483	-.00107	+ .00587	-.00100	-.00583	+ .00028	-.00620	-.00132	-.00884	= - .01346
2	1	+ .00642	-.00262	-.00058	-.00360	+ .00790	-.00115	+ .00037	-.00658	+ .00001	= 0
	2	+ .00330	+ .00511	-.00074	+ .00463	+ .00249	+ .00363	+ .00110	+ .00221	+ 1.13850	= + 1.15442
	3	-.0145	+ .00007	+ .00009	+ .00037	-.00070	-.00083	-.00020	-.00041	-.28814	= - .29104
	4	-.00009	-.00115	+ .00010	-.00060	+ .00180	-.00032	+ .00007	-.00154	-.06144	= - .06318
	5	-.0109	+ .00203	+ .00147	-.00918	-.05245	+ .00014	-.00433	+ .03509	-.01627	= - .05454
	6	-.00255	-.00873	-.00143	-.00981	-.00030	-.02410	-.00589	-.02582	-.05514	= - .13294
	7	+ .00107	-.00483	+ .00089	-.00555	-.00028	+ .00562	+ .00132	+ .00026	+ .00884	= + .01346
	8	+ .00607	-.00008	-.00100	-.00626	-.01224	-.00013	-.00105	+ .00787	-.00170	= + .00748
3	1	-.00158	+ .00002	-.00004	+ .00024	-.00148	+ .00012	-.00017	+ .00112	-.00144	= - .06318
	2	-.00080	-.00008	-.00003	-.00031	-.00047	-.00040	-.00049	-.00037	-.28814	= - .29104
	3	+ .00035	-.00090	+ .00001	-.00004	+ .00013	+ .00009	+ .00012	+ .00007	+ .10155	= + .10227
	4	+ .00009	+ .00001	+ .00000	+ .00004	-.00034	+ .00004	-.00003	+ .00026	0	= 0
	5	+ .00235	-.00001	+ .00008	+ .00061	+ .00983	-.00002	+ .01194	-.00595	+ .00406	= + .01322
	6	+ .00033	+ .00005	-.00006	+ .00060	+ .00006	+ .00263	+ .00264	+ .00437	+ .01394	= + .02484
	7	-.00026	+ .00003	+ .00006	+ .00037	+ .00005	-.00061	-.00059	-.00106	-.00402	= - .00604
	8	-.00146	-.00001	-.00005	+ .00042	+ .00239	-.00001	-.00047	-.00133	-.00011	= + .00024
4	1	+ .00003	+ .00079	+ .00032	+ .00003	+ .00040	+ .00047	+ .00037	+ .00049	+ .28814	= + .29104
											= - .06318
											= 0
											= + .10227
The same numbers occur as for u = 3, differently arranged.											= - .02481
											= + .01322
											= - .00024
											= - .00604
5	1	-.02377	-.00912	-.01104	+ .01452	-.03510	-.00952	-.01112	+ .00736	-.05514	= - .13294
	2	-.01220	+ .01777	-.00858	-.01867	-.01106	+ .03012	-.03312	-.00247	-.01627	= - .05454
	3	+ .00536	+ .00000	+ .00143	-.00231	+ .00310	-.00686	+ .00774	+ .00046	+ .0048	= + .01322
	4	+ .00033	-.00401	+ .00136	+ .00242	-.00800	-.02266	-.00206	+ .00172	-.01394	= - .02481
	5	+ .04065	+ .00706	+ .02123	+ .03703	+ .29290	+ .00114	+ .12993	-.03924	+ .39363	= + .82433
	6	+ .03844	-.03040	-.02188	+ .03594	+ .00133	-.19990	+ .17658	+ .02888	0	= 0
	7	-.00395	-.01680	+ .01492	+ .02239	+ .00122	+ .04603	-.03959	-.00700	+ .00309	= + .02091
	8	-.02247	+ .00295	-.1322	+ .02527	+ .05433	-.00105	+ .03150	-.00680	+ .09441	= + .16791
6											= + .05454
											= - .13294
											= + .02484
											= + .01325
The same numbers occur as for u = 5, differently arranged.											= 0
											= + .82433
											= - .16791
											= + .02091
7	1	+ .00221	-.00018	-.00493	+ .00130	+ .00422	-.00233	-.00063	-.0544	-.00170	= - .00748
	2	+ .00114	+ .00035	-.00324	-.00168	+ .00137	+ .00739	-.00188	+ .0183	+ .00824	= - .01346
	3	-.00050	+ .00000	+ .00084	-.00021	-.00033	-.00168	-.00044	-.00034	-.00402	= - .00604
	4	-.00003	-.00008	+ .00061	+ .00022	+ .00096	-.00065	-.00012	-.00127	+ .00011	= - .00024
	5	-.00379	+ .00014	+ .00950	+ .00333	-.02800	+ .00028	+ .00737	+ .02900	+ .00309	= + .02091
	6	-.00038	-.00080	-.00878	+ .00323	-.00016	-.04904	+ .01002	-.02133	-.09441	= + .16791
	7	+ .00037	-.00033	+ .00667	+ .00201	-.00015	+ .01144	-.00225	+ .00517	+ .03424	= + .05717
	8	+ .00209	+ .00006	-.00592	+ .00227	-.00653	-.00026	+ .00179	+ .00650	0	= 0
8											= - .01348
											= - .00748
											= + .00025
											= - .00604
The same numbers occur as for u = 7, differently arranged.											= + .16795
											= + .02092
											= 0
											= + .05717

TABLE LXV.

Values of k_r for 20 conditions.

[illegible]

VALUES OF $k_r(s, t)$ AND $\Sigma^s k_r(s, t)$, s FROM 1 TO 20, THE LATTER IN OLD FACE TYPE.

t	s	r=1	2	3	4	7	8	9	10	11	12	14	15	16	17	18	19	20
21	1	+2.3231798	-1.1883648	+6373776	-2011896	+1104436	-0.168311	-0.294774	+0.737811	-0.028122	+0.815775	-0.177553	+0.345452	-0.015837	-0.038083	+0.726204	-0.010401	+0.180456
2	2	-0.1715148	-541.9344	0	+0.245382	+0.462924	-0.112834	+0.004464	-0.190092	-0.003960	-0.017835	-0.0283534	-0.075516	-0.035789	-0.018636	+0.048918	-0.007812	+0.011718
3	3	-1.4986848	0	-6770274	+1.644408	-0.675164	+0.015888	+0.898248	-0.0018902	+0.076564	-0.008903	-0.268772	+1.005578	-0.485260	-0.0214438	+0.0415796	-0.0047106	-0.027804
4	4	+4.89892	-1.193688	+0.51836	-0.936498	-0.01085	+0.019538	-0.019538	+0.033817	+0.048514	+0.041120	+0.026460	-0.845500	-0.309028	-0.0070520	+0.039119	+0.035936	+0.064370
Sum		+4.845664	-0.223968	+1.232968	-0.658348	+0.292008	-0.031416	-0.048393	+0.141498	-0.177966	+0.009860	+0.156450	-0.045536	-0.009498	-0.008294	-0.189273	-0.000118	+0.018430
22	1	+0.196666	-0.640288	0	+0.029163	-0.013288	-0.029164	-0.000528	-0.000528	-0.004692	-0.002388	-0.000112	-0.008938	-0.007128	-0.0116060	-0.0019816	+0.0004824	+0.0011880
2	2	-0.0523973	+0.113724	0	-0.079004	+0.010272	-0.004878	-0.001580	-0.004878	-0.001728	-0.000234	-0.002708	-0.027749	-0.011904	+0.0005620	-0.004734	+0.011904	+0.0007158
3	3	+0.143114	-0.168974	+0.046682	-0.077003	+0.345543	-0.0064521	-0.030321	-0.036324	-0.036343	-0.016682	-0.116092	-0.857776	-0.071856	-0.149651	-0.006389	-0.048970	-0.0167648
4	4	-0.007430	-0.000640	-0.000240	-0.020810	-0.167810	-0.057170	+0.014530	+0.003270	-0.001530	-0.000300	+0.043240	+0.073770	-0.019340	+0.0056310	-0.022500	-0.009380	+0.018900
5	5	-0.0016152	+0.000288	-0.007243	-0.001492	-0.025092	-0.068894	-0.002724	-0.007724	-0.003247	-0.018336	-0.006234	-0.003906	+0.042723	+0.011918	-0.074550	+0.007100	-0.000816
6	6	-0.139650	-0.267572	-0.051648	-0.013804	-0.105044	+0.147356	-0.237013	-0.053624	-0.078368	+0.078368	-0.645650	0	-1.118350	-0.076890	-0.118784	-0.026800	-0.028358
7	7	-0.0094620	-0.0010845	-0.004880	-0.0010850	-0.027215	-0.015810	+0.053415	+0.053160	-0.016660	+0.0089210	-0.011370	-0.611650	-0.131255	-0.006850	-0.0281850	+0.015320	-0.020605
8	8	-0.007290	-0.0043080	-0.0079300	-0.0003240	-0.0216700	-0.0035410	-0.0012510	-0.007440	-0.021440	-0.008140	-0.674170	-0.840700	0	-0.0671300	-0.088880	-0.017960	+0.017870
9	9	-0.024112	-0.0053050	-0.010048	-0.026114	-0.221920	-0.006900	-0.0063405	-0.007049	-0.023120	-0.023930	-0.639938	-0.859938	-0.800640	-0.142918	-0.270180	+0.044230	+0.069400
10	10	-0.000658	-0.0001648	-0.0000654	-0.000128	-0.001676	-0.001388	-0.0002960	-0.000254	-0.0009004	-0.001072	-0.015504	+0.015504	-0.002580	-0.008852	-0.012629	-0.0118250	0
11	11	-0.001648	-0.000658	-0.0000684	-0.0001676	-0.0010816	-0.0002360	-0.0001889	-0.0000004	-0.002254	-0.0003554	-0.0001072	-0.000580	-0.000852	-0.0003574	-0.0116950	-0.0091690	-0.060360
12	12	+3.95456	-0.484511	-0.014908	+0.991862	-0.6626897	-0.81661	+0.013900	+0.017271	-0.0193	-0.016450	+34.5310	-0.810495	-0.344112	+20.13765	+1.606666	+0.666668	+0.033337
Sum		+0.294009	-0.268164	-0.009408	+0.001428	-0.1066512	-0.071764	+0.009408	-0.015438	-0.032948	+0.035248	+0.112480	-0.163004	+0.145782	+0.073982	-0.019620	+0.076536	+0.800136
23	9	+0.134684	-0.021690	-0.001192	+0.000924	+0.321688	-0.014893	-0.053700	+0.0459284	-0.085168	0	+0.085044	-0.449304	-0.024574	-0.018932	-0.044574	-0.015380	+0.016698
10	10	-0.015648	-0.007648	-0.000528	-0.000884	-0.003880	-0.023696	+0.13804	-0.685248	0	-0.284452	-0.355768	+0.072012	-0.006704	+0.177752	+0.068260	-0.0032028	+0.012036
17	17	+0.05490	-0.232120	+0.043900	-0.01810	+1.142930	-0.070900	-0.175500	+0.277410	-0.095800	-0.011150	-0.841440	-2.862220	+1.811500	-0.023160	+1.184200	+0.755700	+0.408750
18	18	-0.060177	+0.051912	-0.002946	-0.003969	-0.540704	-0.015558	-0.053780	+0.071011	-0.253692	-0.038708	-0.070020	-0.488378	+0.112381	-0.273838	-0.681160	+0.8631875	-1.988380
19	19	-0.006408	+0.050043	-0.007371	-0.001914	+0.0095046	-0.652736	-0.0155610	-0.651081	-0.626734	-0.1311569	-0.006854	+1.467180	-0.517949	+0.204079	-0.800025	-1.294770	-0.3889010
20	20	-0.003253	+0.087991	-0.0217693	+0.026661	+0.458268	-0.822888	+0.032441	+0.012127	+1.838967	+0.991845	-0.051840	-0.024915	-0.078524	+0.754111	+0.249550	+4.20611	+0.245537
Sum		-0.003253	+0.087991	-0.0217693	+0.026661	+0.458268	-0.822888	+0.032441	+0.012127	+1.838967	+0.991845	-0.051840	-0.024915	-0.078524	+0.754111	+0.249550	+4.20611	+0.245537
																	+1.844074	+0.632190

TABLE LXVIII.

Multipliers for equation				From LXVI: coefficients of				Equation 21				Equation 22				Equation 23			
r	21	22	23	k ₂₁	k ₂₂	k ₂₃		21	22	23		21	22	23		21	22	23	
1	+2.19	+42		-488602	-3854036	+0049253		-1.0549106388				-203271684	-1.66048512	+003916626					
2	0	0		+1290528	+0485111	-00277691		0		0		0							
3	+1.86	+22		-0518830	+0044916	-0077803		-084453880	+006392858	+005171558		-011405760	+006387976	+000011666					
4	-6.62	-18		+0306498	-0691682	+0026067		-2824516776	+656612644	-017258554		-007186964	+017858516	-000469206					
5	+81			+1021068	-0638587	-0546208						+031653666	-029454947	-016694448					
6	0	0		-0751286	+0451691	-0323288						+006602880	+001890010	-062824510					
7	+10			+0640263	-0159001	-0232451						+001314456	+002072588	+001895594					
8	-12			-0106538	-0179774	-0153127													
9	+68			-0539317	-0106878	-1582967													
10				+0254464	-0108471	-0991845													
11	-84			-0248514	+0049066	-0623849													
12	-68			-0041120	+0016549	+0084915													
13	+84			-0264460	-9258310	+1091198													
14	+15			-0845500	-0810495	-0725524													
15				-0800228	-0941112	-0750411													
16	-10			+0079629	-2169765	-0246529													
17	+16			+0069119	-1809889	-4230911													
18	+68			-0359886	-06460068	-2844527													
19	+02			+0055583	-00893927	+1394674													
20	-02	-1.17		-0164570	-0675276	-0632790													
				Coefficients from table XLVII															
				-1.418854194	+2.19			-1.418854194	0			0	+1.22	0		0	0	+1.87	

Left Hand Side.			
No.	k ₂₁	k ₂₂	k ₂₃
1	+1+1.00	+ .7711	- .2010
2	+1+0.9885		- .05885
3	+1+0.9452		- .04000
4	+1+0.9032		- .0324
5	+1+0.8618		- .0254
6	+1+0.8210		- .01901
7	+1+0.7807		- .01365
8	+1+0.7407		- .00885
9	+1+0.7010		- .00455
10	+1+0.6617		- .00081
11	+1+0.6227		- .00037
12	+1+0.5842		- .00017
13	+1+0.5461		- .00008
14	+1+0.5085		- .00004
15	+1+0.4714		- .00002
16	+1+0.4348		- .00001
17	+1+0.3987		- .00000
18	+1+0.3631		- .00000
19	+1+0.3279		- .00000
20	+1+0.2932		- .00000
Verification			
+1	+1.000001177	+1.058851987	- .0689467701
+2	+1.000028246	- .276007170	- .276140684
+3	+1.000068828	- .011867811	- .014267596
+4	+1.00007463	+ .226754947	- .026934962
+5	+1.000079824	- .06894677	+ .11.61.694784
+6	+1.000089043	- .054817894	- .06395869
+7	+1.00009204	- .009429691	- .054768176
+8	+1.00009436	- .00946713	- .0117087
+9	+1.00009636	- .009493	- .054768176
+10	+1.000098134	- .00002656	- .00986969
+11	+1.00009960	- .00002656	- .00986969
+12	+1.00009960	- .00002656	- .00986969
+13	+1.00009960	- .00002656	- .00986969
+14	+1.00009960	- .00002656	- .00986969
+15	+1.00009960	- .00002656	- .00986969
+16	+1.00009960	- .00002656	- .00986969
+17	+1.00009960	- .00002656	- .00986969
+18	+1.00009960	- .00002656	- .00986969
+19	+1.00009960	- .00002656	- .00986969
+20	+1.00009960	- .00002656	- .00986969

TABLE LXIX.

18. To pass to the complete solution of the 23 equations the process is precisely similar to that already followed after the solution of 12 equations in §15: the notation given in the corresponding tables concerned—*viz* tables LXIX, LXX—explains itself and the solution, keeping only 4 decimal places, is exhibited in table LXXI.

In this table are given also the value of Σ^*, k_r for all values of s from 1 to 23. These obviously should correspond to the case where all the R. H. S. of the 23 equations are unity. This solution, which depends on all the fundamental cases, can be verified in the original equations, and affords a complete check of the work. The substitution is carried out in table LXII. It will be seen that the values of the R.H.S. obtained by substitution differ slightly from unity, the greatest difference being .035, showing that no gross error has been committed. It is of interest to consider to what these differences may be attributed. Each value of Σ^*, k_r is given to four decimal places: and accordingly may be in error by .00005. The sum of 23 such quantities *may* be wrong by .00115: and in substituting in the original equation such an error is multiplied by coefficients, the largest of which is 111 51, which would admit of the corresponding term containing an error of .1. That this extreme value should be obtained is most unlikely: but it is clear that the actual discrepancies obtained may easily be attributable to this cause. This line of argument shows how many figures it will be necessary to keep to be absolutely sure of all errors being less than a stated amount. For the present object, the solution as found may be considered sufficiently precise.

19. Table LXXI also contains the necessary multipliers to proceed to the solution of the complete 26 equations. The process is exactly similar to that already described in passing from 20 to 23 equations. All results are given in tables LXXIII—LXXIX, and the verification, similar to that of LXXII, in table LXXX. This completes the solution of the equations of table XLVII.

r	t	21	22	23	
1		-.48398	-.39540	+.00939	
2		+.12308	+.04835	+.00580	
3		+.05186	+.00449	+.00275	
4		+.03965	+.00919	+.00261	
5		+.10211	+.00639	+.05482	
6		+.07813	+.04817	+.08229	
7		+.02803	+.01890	+.02325	
8		+.01095	+.01723	+.01521	
9		+.05338	+.01094	+.15830	
10		+.02555	+.01085	+.09918	
11		+.02485	+.00491	+.06238	
12		+.04111	+.00165	+.00849	
13		+.02645	+.32563	+.10911	
14		+.08456	+.08105	+.09785	
15		+.08999	+.08441	+.07504	
16		+.00796	+.20138	+.04495	
17		+.00391	+.16067	+.42206	
18		+.03594	+.06461	+.23455	
19		+.00554	+.00323	+.13447	
20		+.01844	+.08753	+.06323	
	Σ^*, k_r	+.137317	+.29277	+.00927	
	Σ^*, k_r	+.29277	+.12562	+.05870	
	Σ^*, k_r	+.00320	+.05870	+.12448	
	Σ^*, k_r	+.00320	+.05870	+.12448	Values of Σ^*, k_r
21	1	-.664587	-.115761	+.000030	-.780318
	2	+.169010	+.014155	+.000028	+.183137
	3	-.071213	+.001815	+.000009	-.069889
	4	+.054446	+.029040	+.000008	+.025414
	5	+.140214	+.019406	+.000175	+.120631
	6	-.107286	+.014103	+.000263	-.092920
	7	+.035744	+.004070	+.000074	+.031600
	8	-.015036	+.005069	+.000040	-.020144
	9	+.073800	+.003203	+.000506	+.077009
	10	+.035064	+.003177	+.000317	+.031590
	11	-.034123	+.001438	+.000260	-.032435
	12	+.005644	+.000483	+.000027	+.005134
	13	-.036320	+.003335	+.000340	-.131306
	14	+.116115	+.023729	+.000313	+.140157
	15	-.054913	+.010074	+.000240	-.045079
	16	+.010930	+.008958	+.000080	+.048106
	17	+.003369	+.047039	+.001350	+.043120
	18	+.043352	+.018916	+.000750	+.069013
	19	+.007607	+.000946	+.000430	+.008909
	20	-.022575	+.019771	+.000203	-.042546
22	1	-.141695	+.445070	+.000548	-.586217
	2	+.038034	+.054424	+.000517	+.089041
	3	-.015133	+.005054	+.000163	-.009866
	4	+.011606	+.111650	+.000153	+.009588
	5	+.029895	+.074617	+.003218	+.047940
	6	-.022874	+.054221	+.004330	+.036177
	7	+.007621	+.015646	+.001365	+.006390
	8	-.003206	+.019451	+.000893	-.023550
	9	+.015638	+.012314	+.009292	+.037234
	10	+.007480	+.012213	+.005822	+.015655
	11	-.007275	+.005527	+.003632	+.001014
	12	+.001203	+.001557	+.000498	+.001152
	13	-.007744	+.886536	+.008405	-.367875
	14	+.024757	+.001232	+.005744	+.121733
	15	-.011708	+.038733	+.004408	+.022621
	16	+.003330	+.228677	+.001495	+.225812
	17	+.001145	+.180853	+.021775	+.204453
	18	+.010523	+.072726	+.013768	+.067016
	19	+.001622	+.003636	+.007893	+.013161
	20	-.004813	+.076013	+.008716	+.084541
23	1	+.001549	+.023210	+.010491	+.014266
	2	+.000394	+.002338	+.008905	+.006863
	3	+.000163	+.000264	+.003126	+.003224
	4	+.000127	+.005822	+.002835	+.002760
	5	+.000327	+.003891	+.001644	+.005208
	6	+.000250	+.002328	+.002533	+.005111
	7	+.000083	+.000816	+.002144	+.002697
	8	+.000035	+.000124	+.017103	+.018152
	9	+.000171	+.000432	+.179005	+.178513
	10	+.000082	+.000637	+.111526	+.112031
	11	+.000080	+.000289	+.070145	+.070353
	12	+.000013	+.000097	+.005547	+.006931
	13	+.000085	+.019114	+.123622	+.103493
	14	+.000271	+.004753	+.110030	+.115059
	15	+.000123	+.002020	+.084381	+.082489
	16	+.000025	+.011821	+.022056	+.039852
	17	+.000013	+.009431	+.474598	+.434016
	18	+.000115	+.003733	+.263747	+.267655
	19	+.000018	+.000190	+.151209	+.151417
	20	+.000053	+.003964	+.071157	+.075174

TABLE LXX.

t	21	22	23		t	21	22	23	r_k from LXV	r_k	t	21	22	23	r_k from LXV	r_k		
r=1	21	22	23		r	u	$r_k \Sigma r_k (s, t)$				r	u	$r_k \Sigma r_k (s, t)$					
1	-0.4840	-0.8954	+0.0093		2	16	+0.00147	-0.01811	+0.00017	+0.04368	+0.02890	6	11	+0.02831	+0.00018	+0.00589	+0.08785	+0.06633
2	+0.1231	-0.0484	-0.0088		17	+0.00471	-0.01445	+0.00283	-0.03310	-0.04411	12	+0.00083	+0.00006	+0.00081	-0.03422	-0.03304		
3	-0.0519	+0.0045	+0.0028		18	-0.00354	-0.00551	+0.00157	-0.01507	-0.02859	13	+0.02345	-0.01179	+0.01089	-0.00896	-0.3079		
4	+0.0396	-0.0992	+0.0026		19	+0.0101	+0.00329	-0.00090	-0.00824	-0.0078	14	-0.00786	+0.02933	+0.00931	-0.90042	-0.3553		
5	+0.1021	-0.0663	-0.0548		20	-0.00300	-0.00307	+0.00342	-0.00479	-0.0134	15	+0.00872	+0.00125	-0.00713	+0.04831	+0.0460		
6	-0.0781	+0.0482	+0.0823		3	1	+0.03383	+0.03935	+0.00003	-0.08318	-0.0254	16	-0.00074	-0.00729	-0.00288	-0.21676	-0.2272	
7	+0.0280	-0.0139	-0.2322		2	+0.00800	-0.00048	-0.00093	-0.29104	-0.3002	17	-0.00036	-0.00582	-0.04014	+0.18870	+0.0924		
8	-0.0110	-0.0173	-0.0152		3	+0.00303	-0.00004	+0.00001	+0.10227	+0.1059	18	+0.00334	-0.00234	-0.02351	+0.16319	+0.1419		
9	-0.0534	-0.0109	-0.1583		4	-0.0277	+0.00309	+0.00001	0	-0.0018	19	-0.00051	+0.00012	+0.01279	-0.00838	+0.0205		
10	+0.0255	-0.0108	-0.0992	$\Sigma r_k (s, t)$ from Table LXVI	5	-0.0714	+0.00096	-0.00018	+0.01322	+0.0066	20	+0.00152	-0.00244	-0.00602	+0.05408	+0.0471		
11	-0.0249	+0.0049	+0.0624		6	+0.00540	-0.00048	+0.00026	+0.02484	+0.0301	7	1	-0.01528	+0.00372	-0.00025	-0.00748	-0.0183	
12	-0.0041	+0.0017	+0.0085		7	+0.00182	+0.00014	-0.00007	-0.00604	-0.0078	2	+0.00858	-0.00046	+0.00024	+0.01349	+0.0171		
13	-0.0264	-0.0256	+0.1091		8	+0.00077	+0.00017	-0.00005	+0.00024	+0.0011	3	-0.0104	-0.00004	-0.00008	-0.00604	-0.0078		
14	+0.0946	+0.0810	+0.0979		9	+0.00373	+0.00011	-0.00051	-0.01022	-0.0069	4	+0.00125	+0.00003	-0.00007	-0.00024	+0.0019		
15	-0.0400	+0.0344	-0.0750		10	+0.00178	+0.00011	-0.00032	-0.00021	-0.0022	5	+0.00323	+0.00062	+0.00147	+0.02091	+0.0262		
16	+0.0080	-0.0104	-0.0250		11	+0.00174	-0.00005	+0.00020	-0.00018	+0.0018	6	-0.00247	-0.00045	-0.00222	-0.16791	-0.1780		
17	+0.0039	-0.0078	-0.0221		12	+0.00029	-0.00002	+0.00003	-0.00006	-0.0007	7	+0.00082	+0.00013	+0.00062	+0.05717	+0.0587		
18	-0.0359	-0.0646	-0.2346		13	+0.00185	+0.00326	+0.00035	-0.01519	-0.0007	8	-0.00035	+0.00018	-0.00041	0	+0.0002		
19	+0.0055	+0.0032	+0.1345		14	-0.00501	-0.00081	+0.00031	-0.00406	-0.0105	9	-0.00169	+0.00010	-0.00046	-0.01456	+0.0172		
20	-0.0164	-0.0075	-0.0633		15	+0.00280	-0.00034	-0.00024	-0.00324	-0.0010	10	+0.00081	+0.00010	-0.00267	+0.02227	+0.0058		
u=1	-0.7803	-0.5932	-0.0143		16	-0.00056	+0.00201	-0.00006	-0.00730	-0.0059	11	-0.00079	-0.00005	-0.00168	-0.01538	-0.0178		
2	+0.1831	-0.0899	-0.0077		17	-0.00227	+0.00181	-0.00135	-0.00628	+0.0032	12	-0.00013	-0.00002	-0.00293	-0.00530	-0.0053		
3	-0.0029	-0.0100	+0.0032		18	-0.00251	-0.00003	-0.00075	-0.00363	+0.0050	13	-0.00083	+0.00008	-0.00234	+0.04334	+0.0428		
4	+0.0254	-0.0099	-0.0028		19	-0.00038	-0.00003	+0.00043	-0.00442	+0.0004	14	+0.00267	-0.00078	-0.00263	+0.07577	+0.0751		
5	+0.1206	-0.0479	-0.0652		20	+0.00115	+0.00008	-0.00020	+0.00003	+0.0023	15	-0.00129	-0.00032	-0.00202	-0.01234	-0.0119		
6	-0.0029	+0.0362	+0.0951		4	1	-0.01229	+0.03050	-0.00003	+0.0104	+0.3182	16	+0.00025	+0.00139	+0.00067	+0.03561	+0.0384	
7	+0.0316	-0.0094	-0.0809		2	+0.00313	-0.00484	+0.00026	-0.00318	-0.0049	17	+0.00012	+0.00131	+0.01136	-0.00250	-0.0005		
8	-0.0201	-0.0236	-0.0182		3	-0.00132	-0.00045	-0.00001	0	-0.0018	18	-0.00113	+0.00081	-0.00031	-0.00903	-0.0039		
9	-0.0770	-0.0372	-0.1788		4	+0.00101	+0.00081	-0.00001	+0.00227	+0.0132	19	+0.00017	-0.00003	-0.00032	+0.00683	+0.0035		
10	+0.0316	-0.0106	-0.1121	r_k from LXX	5	+0.00259	+0.00082	+0.00015	-0.02484	-0.0155	20	-0.00052	+0.00063	+0.00170	-0.01530	-0.0115		
11	-0.0325	+0.0019	+0.0704		6	-0.00198	-0.00481	-0.00023	+0.01322	+0.0062	8	1	+0.00773	+0.00033	-0.00017	-0.01346	+0.0055	
12	-0.0051	+0.0012	+0.0096		7	+0.00066	+0.00139	+0.00006	-0.00024	+0.0019	2	-0.02347	-0.00114	+0.00017	-0.00748	-0.0109		
13	-0.0131	-0.0279	+0.1035		8	-0.00028	+0.00173	+0.00004	-0.00004	-0.0046	3	+0.00104	-0.00011	-0.00005	+0.00234	+0.0011		
14	+0.1402	+0.1217	+0.1151		9	-0.00136	+0.00109	+0.00044	+0.00021	+0.0004	4	-0.00060	+0.00234	-0.00005	-0.00604	-0.0046		
15	-0.0451	+0.0230	-0.0825		10	+0.00065	+0.00108	+0.00028	-0.01022	-0.0082	5	-0.00205	+0.00157	+0.00100	+0.16791	+0.1684		
16	-0.0491	-0.2258	-0.0309		11	-0.00063	-0.00040	-0.00017	+0.00066	-0.0003	6	+0.00157	-0.00114	-0.00150	+0.02091	+0.0198		
17	-0.0430	-0.2045	-0.4840		12	-0.00010	-0.00017	-0.00002	-0.00013	-0.0004	7	-0.00052	+0.00033	-0.00042	0	+0.0002		
18	-0.0090	-0.0970	-0.2077		13	-0.00067	+0.00353	-0.00031	+0.00406	+0.0036	8	+0.00023	+0.00041	+0.00028	+0.05717	+0.0581		
19	+0.0090	+0.0132	+0.1514		14	+0.00215	-0.00009	-0.00027	-0.01519	-0.0014	9	+0.00107	+0.00026	+0.00289	-0.00227	+0.0019		
20	-0.0425	-0.0845	-0.0752		15	-0.00102	-0.00344	+0.00021	+0.00730	+0.0031	10	-0.00051	+0.00026	+0.00181	+0.01456	+0.0161		
r	u	$r_k \Sigma r_k (s, t)$	r_k from LXV	r_k	16	+0.00200	+0.02012	+0.00007	-0.00324	+0.0171	11	+0.00050	-0.00012	-0.00114	+0.00520	+0.0045		
1	1	+0.37767	+0.23178	-0.00133	+0.15442	=+0.17037					12	+0.00008	-0.00004	-0.00015	+0.01538	-0.0154		
2	+0.06005	-0.08337	+0.00013	0	-0.1242						13	+0.00063	+0.00788	-0.00199	-0.07577	-0.0895		
3	+0.04050	-0.0264	-0.00004	-0.06318	-0.0254						14	-0.00170	-0.00191	-0.00179	-0.04334	+0.0379		
4	-0.03000	+0.05815	-0.00004	+0.0104	+0.3182						15	+0.00080	-0.00081	+0.00137	-0.03561	-0.0343		
5	-0.07967	+0.03887	+0.00078	-0.13294	-0.1730						16	-0.00016	+0.00475	+0.00045	-0.01234	-0.0073		
6	+0.06094	-0.02825	-0.00118	+0.05454	+0.0801						17	-0.00008	+0.00379	+0.00788	+0.03903	+0.0510		
7	-0.02039	+0.0815	+0.00033	-0.00748	-0.0193						18	+0.00073	+0.00152	+0.00427	-0.02250	-0.0160		
8	+0.00858	+0.01014	+0.00022	-0.01340	+0.0055						19	-0.00011	-0.00008	-0.00215	+0.01330	+0.0107		
9	+0.04167	+0.06339	+0.00236	+0.03069	+0.0840						20	+0.00033	+0.00159	+0.00115	+0.00663	+0.0100		
10	-0.01900	+0.00633	+0.00142	-0.04233	-0.0545						9	1	+0.03727	+0.01471	-0.00166	+0.03300	+0.0840	
11	+0.01043	-0.00237	-0.00039	+0.01430	+0.0300						2	-0.00438	-0.00180	+0.00157	+0.04233	+0.0327		
12	+0.00820	-0.00100	-0.00012	+0.00230	+0.0041						3	+0.00400	-0.00017	-0.00050	-0.01022	-0.0069		
13	+0.02060	+0.19037	+0.00156	+0.03725	+0.2472						4	-0.00305	+0.00309	-0.00046	+0.00021	+0.0004		
14	-0.00601	-0.07448	-0.00140	-0.08108	-0.1960						5	-0.00786	+0.00247	+0.00980	-0.15694	-0.1524		
15	+0.03121	-0.02017	+0.00107	+0.04308	+0.0552						6	+0.00601	-0.00179	-0.01472	-0.11723	-0.1277		
16	-0.00624	+0.11800	+0.00036	-0.00723	+0.1050						7	-0.00200	+0.00053	+0.00415	+0.01456	+0.0172		
17	-0.00304	+0.09420	+0.00003	-0.01507	+0.0821						8	+0.00085	+0.00064	+0.00272	-0.00227	+0.0010		
18	+0.02801	+0.03757	+0.00035	+0.03310	+0.1024						9	+0.00411	+0.00041	+0.00230	+0.03080	+0.0363		
19	-0.04230	-0.00188	-0.00102	-0.00479	-0.0120						10	-0.00196	+0.00040	+0.01774	0	+0.0162		
20	+0.01280	+0.03057	+0.00091	+0.00824	+0.0615						11	+0.00192	-0.00018	-0.01116	-0.03482	-0.0442		
2	1	-0.08902	-0.03555	-0.00006	0	-0.1242					12	+0.00032	-0.00006	-0.00153	+0.04886	+0.0476		
3	+0.02254	+0.04435	+0.00001	+0.15442	+0.1814						13	+0.00196	+0.01311	-0.01351	+0.16360	+0.1582		
4	-0.00950	+0.00040	-0.00002	-0.20104	-0.3002						14	-0.00651	-0.00301	-0.01750	-0.02398	-0.0510		
5	+0.00735	-0.00632	-0.00002	-0.03318	-0.0640						15	+0.00308	-0.00138	+0.01341	+0.02144	+0.0367		
6	+0.01899	-0.00506	+0.00037	-0.05454	-0.0414						16	-0.00062	+0.00749	+0.00447	+0.07044	+0.0818		
7	-0.01430	+0.00433	-0.00055	-0.13294	-0.1435						17	-0.00030	+0.00508	+0.00747	-0.02940	+0.0518		
8	-0.00476	-0.00125	-0.00016	+0.01340	+0.0171						18	+0.00276	+0.00240	+0.01495	+0.01710	+0.0642		
9	-0.00201	-0.00156	-0.00010	-0.00748	-0.0109						19	-0.00042	-0.00012	-0.00405	-0.01127	-0.0350		
10	-0.00678	-0.00095	+0.00106	+0.04233	+0.0327						20	+0.00126	+0.00251	+0.01132	+0.04502	+0.00		

TABLE LXX.—(Continued).

t						21						22						23						r _{ku} from LXV						r _{ku}						t						21						22						23						r _{ku} from LXV						r _{ku}																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																							
r						u						t						u						r _{ku} Σ r _{ku} (s,t)						r _{ku}						r						u						t						u						r _{ku} Σ r _{ku} (s,t)						r _{ku} from LXV						r _{ku}																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																	
10						6						7						8						9						10						11						12						13						14						15						16						17						18						19						20																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																															
-0.00247						-0.00051						-0.00023						-1.5684						-1.6800						-0.00082						-0.00015						-0.00280						+0.00227						+0.00065						-0.00035						+0.00018						+0.00170						+0.01456						-0.00169						+0.00012						+0.01776						0						+0.0162						+0.00081						+0.00011						+0.01112						+0.33080						-0.00079						-0.00005						-0.00700						-0.04886						-0.0567						-0.00018						-0.00002						-0.00065						-0.08182						-0.0366						-0.00089						+0.00045						-0.01233						+0.02898						+0.0145						+0.00267						-0.00086						-0.01097						+0.16860						+0.1544						-0.00126						-0.00036						+0.00841						-0.07044						-0.0637						+0.00025						+0.00214						+0.00280						+0.03144						+0.0266						+0.00012						+0.00170						+0.04732						+0.01710						+0.0320						-0.00113						+0.00068						+0.02830						+0.02840						+0.0036						+0.00017						-0.00003						-0.01509						-0.04502						-0.0600						-0.00052						+0.00072						+0.00710						-0.01127						-0.004																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																									
11						1						2						3						4						5						6						7						8						9						10						11						12						13						14						15						16						17						18						19						20																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																	
+0.01574						-0.00075						+0.00085						+0.01436						+0.0300						-0.00400						+0.00009						-0.00062						-0.00230						-0.00065						+0.00169						+0.00001						+0.00020						-0.00013						+0.0018						-0.00129						-0.00019						+0.00018						+0.00006						-0.00003						-0.00386						-0.03422						-0.03422						+0.000254						+0.00009						+0.00879						+0.05785						+0.0663						-0.00085						-0.00003						-0.00163						-0.01528						-0.0179						+0.00036						-0.00003						-0.00107						+0.00520						+0.0046						+0.00174						-0.00002						-0.01114						-0.03422						-0.0442						-0.00085						-0.00002						-0.00698						-0.04886						-0.0567						+0.00081						+0.00001						+0.00439						+0.03889						+0.0441						+0.00013						+0.00000						+0.00060						0						+0.0007						+0.00086						-0.00082						+0.00798						-0.01059						-0.002						-0.00275						+0.00015						+0.00830						-0.05276						-0.0485						+0.00130						+0.00007						+0.00538						+0.02814						+0.0222						-0.00026						-0.00038						-0.00176						-0.00878						-0.0112						-0.00013						-0.00081						-0.03972						+0.00471						+0.0254						-0.00117						-0.01851						+0.01445						+0.0010						-0.00018						+0.00001						+0.00947						+0.00588						+0.0147						+0.00053						-0.00013						+0.00446						+0.00177						+0.0023																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																									
12						1						2						3						4						5						6						7						8						9						10						11						12						13						14						15						16						17						18						19						20																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																																	
+0.00247						-0.00047						+0.00009						+0.00230						+0.0044						-0.00063						+0.00006						-0.00008						+0.01438						+0.0137						+0.00026						+0.00001						+0.00003						-0.00096						-0.0007						-0.00020						-0.00012						+0.00002						-0.00013						-0.00004						-0.00003						-0.00053						-0.05785						-0.0504						+0.00040						+0.00006						+0.00079						-0.03422						-0.0300						-0.0066						-0.00013						-0.00002						-0.00022						-0.00520						-0.0056						+0.00009						-0.00002						-0.00015						-0.01528						-0.0154						+0.00027						-0.00001						-0.0152						+0.04886						+0.0476						-0.00013						-0.00001						-0.00096						-0.03422						-0.0362						+0.00013						+0.00000						0						+0.0007						+0.00002						+0.00009						+0.03889						+0.0380						+0.0007						+0.00002						+0.00009						+0.05276						+0.0538						-0.00043						+0.00010						+0.00094						+0.01059						-0.0100						+0.00020						+0.00004						+0.00072						+0.00878						+0.0083						-0.00004						-0.00024						+0.00024						+0.02614						+0.0256						-0.00002						-0.00019						-0.00405						-0.01445						+0.0187						+0.0028						-0.00008						+0.00000						+0.00129						+0.00177						+0.0005						+0.00008						-0.00068						-0.00061						+0.00536						+0.0048						+0.000355						+.14547						+0.00096						+0.03725						+0.2472						-0.01816						-0.01781						-0.00091						+0.08108						+0.0402						+0.00681						-0.00166						+0.00020						-0.01519						-0.0097						-0.00820						+0.03850						+0.00027						+0.00408						+0.0356						-0.01341						+0.02439						+0.00687						-0.38002						-0.3547						+0.01025						-0.01773						+0.00852						-0.30896						-0.3079						+0.00341						+0.00511						-0.00240						+0.04334						+0.0426						+0.00144						+0.00836						-0.00157						-0.07577						-0.0694						+0.00701						+0.00401						+0.01638						+0.16860						+0.1582						-0.00336						+0.00397						-0.01027						+0.02398						+0.0143						+0.00327						-0.00180						+0.00646						-0.01059						-0.0027						+0.00054						-0.00063						+0.00083						+0.05276						+0.0520						+0.00347						+0.11979						+0.01129						+1.88750						+0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013						0						-0.0908						-0.01111						-0.02890						+0.01013											

TABLE LXXI. Values of k_r for 23 conditions.

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23		
1	+3.5107	0	+2.1102	-9.9564	+6714	0	+3833	-2.2062													+2.1006	+	-4029	0	
2	+3.3010	+	+9.3617	+1.9843	0	+6313	+3.0744	+5898													-	0	0	0	
3	-6367	-3.0040	-11.7728	0	+0087	-5132	-3.1271	-3239													-	-5393	-	0	
4	-3.4690	+7552	0	+13.5393	+	-5915	-9810	+2.4564													-	-2.2124	-	0	
5	+	+3.4276	0	-0.149	+4.8820	0	-4171	-14.0141													0	+	+1539	0	
6	0	-0824	-2068	+0095	0	-1.1429	-3.3315	+0889													0	0	0	0	
7	-0084	-0121	-0610	+0095	+0070	-2512	-9255	0	+0110	-1807	-2858	-4965	+0007	-0113	-0168	-0113					-	-0008	0	0	
8	-38202	+0678	+1884	+1.2150	-4.6645	-1.1889	0	+13.4326	-3.7969	-2183	-9.7692	+4.8602	-2355	-0132	-3248	+3839					0	-	-0198	0	
9					+1.9753	0	-4080	-7.0127	+3.9353	+	+5.1321	-3.4590	+0855	0	+0806	-1393	+1679	0	-0894	-5068	0	0	+2076	0	
10					+1.4538	+5.1614	-2966	0	+2.1751	-2.6459	+3.7772	0	+0039	+	-0348	+0270	+1236	0	-1236	-0270	0	0	0	0	
11					-5531	-2183	-1.0160	+2.0271	-5768	-3861	-2.7200	0	-0239	+0096	+0110	+0484	+0168	+0072	+0244	+0465	0	0	+0832	0	
12					-1394	+4380	+1.6339	-7388	-3036	+	4505	0	+3.1262	+0075	-0165	-0686	+0066	+0129	+0864	+0190	0	0	-0183	0	
13					+3.3498	0	-1347	-2.3061	+4713	0	+1.0268	+4713	+5.7737	0	+8.5864	-7.2045	+1.8071	0	-5397	-2.0200	0	+	-5723	0	
14					-1510	-1040	+0081	0	-0213	+0313	-0463	0	-2603	-3249	-4.0202	-1617	0	-0630	-0011	-0943	0	0	0	0	
15					-6791	-1.2306	-6599	+6463	-0570	-2851	+1521	-1.0122	-2.0336	-4.0202	0	+9.6410	-1.1353	+0043	-4704	-5866	-1306	-	-0713	0	
16					-2.8516	+2.1063	+1.5499	+2.3020	-6838	+1368	-1.6410	-3847	-4.8775	+2.4273	0	+8.5864	-7.2045	+1.8071	+0113	-4350	+2.0356	0	-	-1140	0
17									+8845	+	+8156	+3568	+1.4108	0	+0170	-1.6821	+3.7890	0	-9515	-6.8814	0	+	-2719	+1.1894	0
18									+6604	-2521	-5763	0	+8668	+1.1837	+0190	0	+2.6776	+	+4.5383	-6724	0	0	0	0	
19									+0535	-3392	-1.408	-2897	-0635	-2380	-3939	+0754	+1110	-8631	-3.1754	0	0	-	-0040	+1246	0
20									-1.4615	-2377	-1.1073	-6251	-1.0365	+2317	-3346	-1.6111	-3.5656	-4630	0	+9.6492	0	-	-0176	+1.0301	0
21									0	0	0	0	0	0	0	0	0	0	0	0	+1.6510	0	0	0	
22									0	0	0	0	0	0	0	0	0	0	0	0	0	-0.497	0	0	
23									0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	0	
Sum	+1.0020	+9985	+9966	+9886	+1.0071	+9996	+9690	+9723	+1.0043	+9925	+1.0062	+9648	+9950	+1.0037	+1.0049	+9986	+1.0013	+1.0032	+1.0064	+9954	+1.0009	+9999	+9985	9985	

TABLE LXXII. Verification of solution in Table LXXI.

TABLE LXXVI.

r	t=24	25	26	u	t=24	25	26	Σk_u	t=24	25	26	Σk_u	t=24	25	26	Σk_u
1	-1517	-0102	+0030	1	-208006	-002993	+000011	-211889	-044509	-011520	+000176	-056853	-000545	-000500	+003374	+002230
2	-4792	-0404	+0104	2	-659906	-011922	+000037	-778491	-140597	-456816	+000610	-596803	-001725	-023732	+011897	-018760
3	-0403	+1026	-0029	3	-055497	+030103	-000010	-025404	-011824	+115876	-000170	-103822	-000145	+000203	-003262	+002616
4	-0563	+0033	+0024	4	-077306	+000968	+000309	-078529	-018577	+003727	+000141	-012709	-000203	+000194	+002699	+002690
5	-0758	-0538	-0831	5	+104384	-015785	-000209	-088300	+022240	-080762	-004878	-000273	-003188	-003468	-009483	-009348
6	+1085	-0622	-0552	6	+142530	-012340	-000199	+124082	+030367	-070249	-003240	-043123	+000373	-003151	-062083	-064861
7	+0107	+0159	-0153	7	+014666	+004650	+000055	+019371	+003125	+017901	+000898	+021924	+000038	-000090	+017208	+002690
8	+0251	-0128	-0235	8	+034593	-003758	-000085	+030723	+007894	-014456	-001879	-008471	+000090	-000761	-026430	-027091
9	-0267	+0154	-0987	9	-036766	+004518	+000355	-031893	-007834	+017303	+005794	+015353	-000096	+000974	+111008	+118166
10	-0536	-0120	-1581	10	-073807	-003521	-000589	-077897	+015726	-018553	-006280	-038558	-000198	-000704	-177815	-178712
11	+0035	-0004	-0057	11	+004819	-000117	-000031	+004671	+001027	-000103	-000511	+000054	+000013	-000023	-009785	-009785
12	-0246	+0051	-0624	12	-038374	+001496	+000225	-032153	-007218	+005780	+003863	+002205	-000089	+000289	+070161	+070391
13	-1032	-0751	-1000	13	-142106	-022034	-000360	-164500	-030279	-084818	-005870	-120987	-000872	-004408	-112470	-117250
14	-0208	-3530	+1102	14	-028642	-067702	+000397	-125947	-006103	-376090	+006469	-375724	-000075	-018547	-123942	-104320
15	-0068	+2039	+0249	15	-008364	+059824	+000090	+050550	-001995	+230285	+001482	+228752	-000024	+011969	+023006	+039950
16	-0510	+0370	-0763	16	-070332	+010656	-000275	-059651	-014863	+041788	-004479	+022346	-000184	+003172	-088815	-088815
17	+0282	+0689	+2335	17	+001790	-000840	-000840	+005889	+003874	-017816	-013708	+099790	+000102	+004044	+263917	+263917
18	+0004	-1561	-4223	18	+000551	-045800	-001522	-046771	+000117	-176290	-024818	-201000	+000001	-009163	-475523	-484686
19	+0107	+0684	+0638	19	+022998	+019488	+000228	+042708	+004900	+074092	+003718	+038908	+000060	+008898	+071194	+075152
20	+0018	+056	-1340	20	+001790	-000163	-000182	+003915	+000381	-000326	+007888	+014672	+000005	+000329	+150710	+151044
21	+0145	-0671	+0068	21	+019988	-019687	+000024	+000267	+004254	-077833	-000399	-071180	+000052	-003839	+007648	+003871
22	+0551	-0145	+0068	22	+075878	-004254	+000023	+071647	+016166	-016378	+000370	+000160	+000198	-000851	+007086	+006433
23	+0013	-0052	+0003	23	-001790	-001526	+000001	-003815	-000381	-005879	+000018	-006236	-000005	-000305	+000337	+000027
Σk_u	+1.3771	+2.9934	+0.0366													
24	+1.3771	+2.9934	+0.0366													
25	+2.9934	+1.1294	+0.0567													
26	+0.0366	+0.0567	+1.1247													

TABLE LXXVII.

t	21	22	23	Σk_u from LXXI	Σk_u	t	21	22	23	Σk_u from LXXI	Σk_u
r	u	$\Sigma k_u \Sigma k_u (s,t)$				r	u	$\Sigma k_u \Sigma k_u (s,t)$			
1	1	-03215	+00057	+00001	+1.7637	3	3	+0010	+0107	+0000	+1.059
2	2	+10154	+02280	+00002	-1.242	4	4	+0014	+0003	+0000	-0.018
3	3	+00854	-00574	-00001	-0.254	5	5	-0019	-0056	-0002	+0.066
4	4	+01197	-00018	+00001	+1.182	6	6	-0028	-0065	-0001	+0.301
5	5	-01606	+00301	-00018	-1.730	7	7	-0003	+0018	+0000	-0.078
6	6	-02193	+00348	-00012	+0.661	8	8	-0006	-0013	-0001	+0.011
7	7	-00227	-00089	+00003	-0.189	9	9	+0007	+0016	+0003	+0.069
8	8	-00532	+00073	-00005	+0.655	10	10	+0014	-0012	-0004	-0.022
9	9	+00566	-00086	+00022	-0.840	11	11	-0001	+0000	+0000	+0.018
10	10	+01138	+00067	-00035	-0.545	12	12	+0006	+0005	+0002	-0.007
11	11	-00074	+00002	-00002	+0.800	13	13	+0028	-0078	-0008	-0.097
12	12	+00521	-00029	+00014	+0.044	14	14	+0005	-0046	-0003	-0.105
13	13	+02187	+00420	-00022	+2.729	15	15	+0002	+0212	+0001	-0.010
14	14	+00441	+01861	+00024	-1.960	16	16	+0019	+0038	-0002	-0.059
15	15	+00144	-01140	+00005	+0.552	17	17	-0007	+0072	+0006	+0.032
16	16	+01081	-00207	-00017	+1.050	18	18	-0000	-0162	-0011	+0.050
17	17	-00598	-00385	+00051	-0.821	19	19	-0004	+0080	+0002	-0.004
18	18	-00008	+00573	-00003	+1.024	20	20	-0000	+0006	+0003	+0.023
19	19	-00354	-00371	+00014	-0.129						
20	20	-00028	-00031	+00029	+0.615						
21	1	+11810	+00008	-00004	-1.242	3	3	+0031	-0018	+0000	-0.018
22	2	+37308	+24121	-00014	+1.1814	4	4	+0043	+0000	+0000	+1.182
23	3	+03137	-06121	+00004	-3.902	5	5	-0058	+0007	-0002	-0.155
24	4	+04339	-00197	-00003	-0.649	6	6	-0079	+0008	-0001	+0.062
25	5	-05001	+03210	+00115	-0.414	7	7	-0008	-0002	+0000	+0.019
26	6	-08057	+03711	+00076	-1.435	8	8	-0010	+0002	-0001	-0.046
27	7	-00833	-00948	-00021	-0.171	9	9	+0021	-0002	+0003	+0.004
28	8	-01954	+00764	+00032	-0.109	10	10	+0041	+0002	-0004	-0.082
29	9	-02079	-00919	-00136	+0.927	11	11	-0003	+0000	+0000	-0.008
30	10	+04173	+00710	+00218	+0.381	12	12	+0019	-0001	+0002	-0.004
31	11	-00272	+00024	-00012	-0.068	13	13	+0079	+0010	-0008	+0.356
32	12	+01015	-00304	-00086	+0.137	14	14	+0016	+0042	+0003	-0.021
33	13	+08034	+04480	+00138	+0.462	15	15	+0005	-0028	+0001	+0.031
34	14	+01610	-01897	-00132	-0.594	16	16	+0039	-0005	-0002	-0.171
35	15	+00529	-12185	-00034	-0.035	17	17	-0022	-0009	+0006	+0.147
36	16	+03970	-02207	+00105	+0.266	18	18	-0000	+0020	-0012	+0.125
37	17	-02195	-04111	-00322	-0.441	19	19	-0013	-0008	+0002	-0.012
38	18	-00031	+03913	-00584	-0.259	20	20	-0001	-0001	+0004	+0.066
39	19	-01900	-03031	-00057	-0.078						
40	20	-00101	-00334	-00185	-0.134						

TABLE LXXVII.—(Continued).

t	21	22	23	rk _a from LXXI	rk _a	t	21	22	23	rk _a from LXXI	rk _a	t	21	22	23	rk _a from LXXI	rk _a
r	u	rk _a Σ rk _a (s, t)				r	u	rk _a Σ rk _a (s, t)				r	u	rk _a Σ rk _a (s, t)			
5	5	+0067	+0023	+0080	+8484	8	12	-0008	-0000	-0017	-0154	13	13	+0170	+0091	+0117	+2-0020
	6	+0091	+0027	+0083	-0171		13	-0032	+0006	+0027	-0695		14	+0034	+0403	-0129	-0308
	7	+0009	-0007	-0015	-0382		14	-0006	+0028	-0030	-0379		15	+0011	-0217	-0029	-3559
	8	+0023	+0006	+0023	-1684		15	-0002	-0017	-0007	-0343		16	+0084	-0045	+0089	+9448
	9	-0024	-0007	-0006	-1524		16	-0016	-0003	+0021	-0073		17	-0046	-0089	-0274	-5053
	10	-0047	+0005	+0152	+1273		17	+0009	-0006	-0063	+0510		18	-0001	+0180	+0496	+3417
	11	+0008	-0000	+0006	-0415		18	-0000	+0013	+0115	-0160		19	-0027	+0080	-0074	-1134
	12	-0023	-0003	-0080	-0590		19	+0005	-0006	-0017	+0107		20	-0002	-0007	-0157	-0871
	13	-0091	+0083	+0096	-3547		20	-0000	-0000	-0036	+0100						
	14	-0018	+0144	-0106	+3089												
6	15	-0006	-0088	-0024	-2183	9	9	+0009	+0002	+0110	+3636	14	13	+0180	+0282	-0104	-0308
	16	-0045	-0018	+0073	-0380		10	+0017	-0002	-0177	+0162		14	+0028	+1251	+0115	+1-0005
	17	+0026	-0080	-0235	+1990		11	-0001	-0000	-0010	-0442		15	+0009	-0766	+0026	-8842
	18	-0000	+0068	+0407	+1246		12	+0008	+0001	+0070	+0176		16	+0084	-0139	-0089	-8670
	19	+0015	-0029	-0061	+0458		13	+0033	-0012	-0112	+1582		17	-0086	-0259	+0244	+4050
	20	+0001	-0002	-0129	-0050		14	+0007	-0051	+0123	-0510		18	-0001	+0586	-0441	-5601
							15	+0002	+0031	+0028	+0367		19	-0021	-0249	+0066	+0942
							16	+0016	-0006	-0086	+0818		20	-0002	-0021	+0140	-1432
							17	-0009	+0011	+0261	+0518						
							18	-0000	-0024	-0473	+0642						
7	10	-0004	-0000	-0006	-0415	10	10	-0005	+0010	+0071	-0359	15	15	-0008	+0469	+0010	+6801
	11	+0008	-0000	+0006	-0415		20	-0000	+0001	+0150	+0601		16	-0026	+0085	-0631	-0029
	12	-0023	-0003	-0080	-0590		9	+0021	-0006	-0176	+0162		17	+0014	+0158	+0093	+2189
	13	-0091	+0083	+0096	-3547		10	+0042	+0005	+0283	+3428		18	-0000	-0859	-0169	-1084
	14	-0018	+0144	-0106	+3089		11	-0003	-0000	+0136	-0567		19	+0008	+0153	+0035	-0291
	15	-0006	-0088	-0024	-2183		12	+0019	-0002	-0111	-0359		20	+0001	+0013	+0054	+0487
	16	-0045	-0018	+0073	-0380		13	+0080	+0029	+0179	+0143						
	17	+0026	-0080	-0235	+1990		14	+0016	+0129	-0197	+1544		16	+0004	+0045	-0021	-0029
	18	-0000	+0068	+0407	+1246		15	+0005	-0079	-0045	-0637		17	+0030	+0008	+0064	+7174
	19	+0015	-0029	-0061	+0458		16	+0040	-0014	+0136	-0286		18	-0017	+0015	-0196	-0559
8	5	+0067	+0023	+0080	+8484	11	11	-0000	0	+0001	+0441	17	17	+0017	+0069	+0033	+1-8296
	6	+0091	+0027	+0083	-0171		12	-0001	0	-0006	-0007		18	0	-0156	-1123	+1283
	7	+0009	-0007	-0015	-0382		13	-0005	0	+0010	-0027		19	+0010	+0086	+0169	+0421
	8	+0023	+0006	+0023	-1684		14	-0001	0	-0011	-0485		20	+0001	+0006	+0358	+6264
	9	-0024	-0007	-0006	-1524		15	0	0	-0002	-0232						
	10	-0047	+0005	+0152	+1273		16	-0002	0	+0007	-0112		18	-0018	-0138	-1132	+1283
	11	+0008	-0000	+0006	-0415		17	+0001	0	-0023	-0254		19	0	+0314	+2042	+1-7641
	12	-0023	-0003	-0080	-0590		18	0	0	+0011	-0010		20	-0008	-0133	-0307	-6179
	13	-0091	+0083	+0096	-3547		19	+0001	0	-0006	+0147						
	14	-0018	+0144	-0106	+3089		20	0	0	-0013	-0028						

TABLE LXXVIII. Values of k_r for 26 conditions.

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26
1	+1.7860	0	-0.2223	+3500	-1.8922	+0.0773	-0.2224	+0.0090	+0.0301	+0.0291	+0.0035	+0.2739	-0.2739	-0.1727	+0.4533	+1.1524	+0.7228	-1.1022	-0.1082	+0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022
2	0	+1.7860	-0.2223	-3500	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
3	-0.2223	-3500	-0.0773	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
4	+3500	-0.2223	0	+1.7860	-0.0773	-0.0773	+0.0090	+0.0301	+0.0291	+0.0035	+0.2739	+0.2739	+0.1727	+0.4533	+1.1524	+0.7228	+1.1022	+0.1082	+0.0113	+0.7735	+0.0596	+0.1038	+0.2119	+0.5530	-0.0022	
5	-1.8922	-0.0773	-0.0773	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
6	+0.0773	-1.8922	-0.0773	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
7	-0.0773	-1.8922	-0.0773	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
8	+0.0090	-0.0291	-0.0035	0	-0.0773	-0.0773	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022
9	+0.0301	-0.0291	-0.0035	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
10	+0.0291	-0.0301	-0.0035	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
11	+0.0035	-0.0291	-0.0035	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
12	+0.0035	-0.0291	-0.0035	0	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
13	+0.2739	-0.1727	-0.1727	-0.0773	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
14	-0.2739	-0.1727	-0.1727	-0.0773	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
15	+0.4533	-0.2739	-0.2739	-0.0773	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
16	+1.1524	-0.4533	-0.4533	-0.0773	-0.0773	-0.0773	-0.0090	-0.0301	-0.0291	-0.0035	-0.2739	-0.2739	-0.1727	-0.4533	-1.1524	-0.7228	-1.1022	-0.1082	-0.0113	-0.7735	-0.0596	-0.1038	-0.2119	-0.5530	+0.0022	
17	+0.7228	-1.1022	-1.1022	-0.1082	-0.1082	-0.1082	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
18	+1.1022	-0.7228	-0.7228	-0.1082	-0.1082	-0.1082	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
19	-0.1082	-1.1022	-1.1022	-0.1082	-0.1082	-0.1082	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
20	+0.0113	-0.1082	-0.1082	-0.1082	-0.1082	-0.1082	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
21	-0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	
22	+0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
23	-0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	
24	+0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	
25	-0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	+0.0113	
26	+0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	-0.0113	

TABLE LXXIX. Verification of solution in Table LXXVIII.

[illegible]

TABLE LXXI.

Values of k_r from Table LXXI.

For 23 conditions

Side closure

r	1	3	4	5	7	8	13	15	16	21	22	
s	[urf]	Values of [urf] _s k _r										
1	+ 2.61	+ 4.603	- .066	+ .831	- .451	- .050	+ .014	+ .645	+ .144	+ .274	- 2.037	- 1.530
3	+ 2.08	- .053	+ .220	- .004	+ .014	- .016	+ .002	- .020	- .002	- .012	- .145	- .021
4	- 6.80	- 2.164	+ .012	- .770	+ .105	- .013	+ .031	- .242	- .021	- .116	- .173	+ .679
5	+ 2.30	- .398	+ .015	- .036	+ 1.940	+ .060	+ .387	- .816	- .502	- .083	+ .277	- .110
7	+ 0.02	.000	.000	.000	+ .001	+ .001	.000	+ .001	.000	+ .001	+ .001	.000
8	- 1.49	- .008	- .002	+ .007	- .251	- .000	- .087	+ .104	+ .051	+ .011	+ .030	+ .035
13	+ 1.99	+ .492	- .019	+ .071	- .706	+ .085	- .138	+ 3.984	- .708	+ 1.880	- .261	- .782
15	+ 1.85	+ .102	- .002	+ .008	- .404	- .022	- .063	- .658	+ 1.258	- .005	- .083	+ .042
16	- 2.59	- .272	+ .015	- .044	+ .093	- .099	+ .019	- 2.447	+ .008	- 1.858	+ .125	+ .585
21	+ 2.19	- 1.709	- .153	+ .056	+ .264	+ .069	- .044	- .288	- .099	- .105	+ 3.007	+ .641
22	+ 0.73	- .428	- .007	- .073	- .035	- .007	- .017	- .269	+ .016	- .165	+ .214	+ .822
Sum = k _r		+ .165	+ .013	+ .044	+ .570	+ .008	+ .104	- .006	+ .145	- .178	+ .955	+ .411
Multiplier = [urf]		+ 2.61	+ 2.08	- 6.80	+ 2.30	+ 0.02	- 1.49	+ 1.99	+ 1.85	- 2.59	+ 2.19	+ 0.73
[urf]k _r		+ .431	+ .027	- .299	+ 1.311	.000	- .155	- .012	+ .268	+ .461	+ 2.091	+ .800

$$u_f = 4.91$$

$$u_F = u_f - \sum [urf] k_r = 4.91 - 4.42 = .49$$

$$K = \sqrt{\frac{u_F}{u_f}} = .32$$

Azimuth closure at U₁

$$[urf] = 2.61 + 2.30 = 4.91$$

$$k_r = 2.61 {}_2k_r + 6.80 {}_3k_r + 2.08 {}_4k_r + 2.30 {}_6k_r + 1.49 {}_7k_r + 0.02 {}_8k_r + 1.99 {}_{13}k_r + 2.59 {}_{16}k_r + 1.85 {}_{15}k_r.$$

This holds for either 20 or 23 conditions. But the values of k_r are different in the two cases. Values of r required are 2, 3, 4, 6 16.

For 20 conditions.

Comparing terms in k_1 side closure and k_3 azimuth closure

$$\begin{aligned} 2.61 {}_1k_1 &= 2.61 {}_3k_3 \text{ since } {}_1k_1 = {}_3k_3 \\ 2.08 {}_3k_1 &= 2.08 {}_4k_2 \text{ ,, } {}_3k_1 = {}_4k_2 \\ -6.80 {}_4k_1 &= 6.80 {}_3k_2 \text{ ,, } {}_4k_1 = -{}_3k_2 \\ 2.30 {}_6k_1 &= 2.30 {}_6k_2 \text{ ,, } {}_6k_1 = {}_6k_2 \end{aligned}$$

etc.

so that k_3 (azimuth) is same as k_1 (side).

As regards k (side) and k_4 (azimuth)

$$2.61 {}_1k_3 = 2.61 {}_2k_4$$

etc.

and k_4 (azimuth) = k_3 (side).

For k_4 (side) and k_3 (azimuth)

$$2.61 {}_1k_4 = -2.61 {}_2k_3$$

so that k_3 (azimuth) = k_4 (side)

etc.

TABLE LXXIV.

Easting closure

r	1	2	3	4	5	6	7	8	13	14	15	16													
s	[urf]	Values of [urf] _r k _r																							
1	- .30	- .346	0	+	-.019	- .087	+	-.040	-	-.016	+	-.002	+	-.004	-	-.011	+	-.024	-	-.018	+	-.002			
2	- 6.72	0	- 7.768	+	1.956	+	+.425	+	+.387	+	+.803	+	+.090	+	+.050	+	+.545	+	+.250	+	+.049	+	+.289		
3	-11.83	+	+.747	+	3.443	-	1.210	0	-	-.157	-	-.204	+	+.071	-	-.003	+	+.180	+	+.048	+	+.038	+	+.036	
4	- 4.59	- 1.336	+	+.291	0	-	-.469	+	+.114	-	-.061	+	+.001	+	+.028	-	-.019	+	+.070	-	-.034	+	+.015		
5	- 3.51	+	+.437	+	1.191	-	-.046	+	+.087	-	2.883	0	-	-.078	-	-.559	+	1.264	-	1.084	+	+.781	+	1.169	
6	-12.15	-	-.683	+	1.615	-	-.302	-	-.161	0	-	10.016	+	2.040	-	-.254	+	3.754	+	4.374	-	-.586	+	2.634	
7	- 7.53	+	+.058	-	1.101	+	+.045	+	+.002	-	1.157	+	1.264	-	-.430	0	-	-.326	-	-.571	+	+.093	-	+.268	
8	+	2.18	-	-.029	-	-.016	+	-.001	-	-.013	+	+.386	+	+.046	0	+	1.125	-	1.165	+	-.004	-	-.072	-	-.027
13	- 3.13	-	-.117	-	-.254	+	+.046	-	-.013	+	1.127	+	+.967	-	1.131	+	2.37	-	5.845	0	+	1.066	-	2.736	
14	-10.44	+	+.846	+	+.386	+	+.042	+	+.159	-	3.226	+	3.759	-	1.791	-	452	0	-	10.497	+	9.120	+	3.557	
15	-16.22	-	-.699	-	-.117	+	+.053	-	-.118	+	3.616	-	+.782	+	+.200	+	+.578	+	5.586	+	14.179	-	10.883	0	
16	- 5.64	+	+.041	-	-.243	+	+.041	+	+.018	+	+.272	+	1.223	-	+.201	+	+.070	-	4.030	+	1.922	0	-	3.786	
Sum = k _r		-1.038	-3	-.339	+	+.647	-	170	-	630	-	3.017	+	563	-	206	-	1.117	-	691	-	564	-	643	
Multiplier = [urf]		- .30	- 6.72	- 11.83	-	4.59	- 3.51	- 12.15	-	7.53	+	2.18	-	3.13	-	10.44	-	16.22	-	5.64					
[urf] k _r		+	310	+	23.438	-	7.654	+	780	+	2.211	+	36.657	-	4.463	-	449	+	3.496	+	7.214	+	9.148	+	3.627
u _f = 93.29													u _R = u _f - Σ [urf] k _r = 93.29 - 73.31 = 19.98										K = $\sqrt{\frac{u_R}{u_f}}$ = .46		

TABLE LXXXV.

Easting closure

r	1	2	3	4	5	6	7	8	13	14	15	16	21	22			
s	[urf]	Values of [urf] _r k _r															
1	- .30	- .628	+ .037	+ .008	- .005	+ .052	- .023	+ .008	- .002	- .074	+	- .059	- .017	- .032	+ .234	+ .176	
2	- 6.72	+ .835	- 7.939	+ 2.017	+ .496	+ .278	+ .964	- .115	+ .078	- .310	- .399	-	- .024	- .179	- 1.230	- .804	
3	- 11.93	+ .300	+ 3.551	- 1.263	+ .021	- .078	- .358	+ .029	- .013	+ .115	+ .124	+	- .012	+ .070	+ .827	+ .118	
4	- 4.69	- 1.461	+ .298	+ .008	- .530	+ .071	- .023	- .009	+ .021	- .163	+ .098	-	- .014	- .078	- .117	+ .459	
5	- 3.51	+ .607	+ .145	- .023	+ .054	- 2.960	+ .060	- .092	- .591	+ 1.245	- 1.084	+	+ .766	+ .126	- .423	+ .168	
6	- 12.15	- 1.046	+ 1.744	- .366	- .075	+ .208	- 10.221	+ 2.102	- .241	+ 3.741	+ 4.321	+	- .559	+ 2.760	+ 1.139	- .440	
7	- 7.63	+ .145	- .129	+ .969	- .014	- .197	+ 1.303	- .442	- .002	- .321	- .566	+	- .090	- .280	- .238	+ .071	
8	+ 2.18	+ .012	- .024	+ .002	- .010	+ .367	+ .043	0	+	- .127	- .152	+	- .033	- .075	- .016	- .044	- .051
13	- 3.13	- .774	- .145	+ .030	- .111	+ 1.110	+ .964	- .133	+ .218	- 6.266	+ .096	+ 1.114	- 2.857	+ .411	+ 1.152		
14	- 10.44	+ 2.048	- .620	+ .110	+ .223	- 3.225	+ 3.712	- .784	- .396	+ .322	- 19.841	+ 9.231	+ 3.831	- 1.464	- 1.271		
15	- 16.22	- .895	- .057	+ .016	- .050	+ 3.541	- .740	+ .193	+ .556	+ 5.779	+ 14.342	- 11.031	+ .047	+ .732	- .867		
16	- 5.64	- .592	- .160	+ .033	- .096	+ .203	+ 1.281	- .217	+ .041	- 5.328	+ 2.070	+ .016	- 4.046	+ .271	+ 1.274		
21	- .14	+ .109	- .026	+ .011	- .004	- .017	+ .013	- .004	+ .008	+ .018	- .020	+ .006	+ .007	- .192	- .041		
22	- .53	+ .311	- .049	+ .005	+ .053	+ .025	- .019	+ .005	+ .013	+ .195	- .065	- .012	+ .120	- .155	- .597		
Sum = k _r	- .932	- 3.363	+ .657	- .188	- .522	- 3.056	+ .602	- .193	- 1.205	- .782	- .497	- .636	- .259	+ .047			
Multiplier=[urf]	- .30	- 6.72	- 11.93	- 4.59	- 3.51	- 12.15	- 7.63	+ 2.18	- 3.13	- 10.44	- 16.22	- 5.64	- .14	- .53			
[urf] k _r	+ .280	+ 22.599	- 7.772	+ .863	+ 1.832	+ 37.130	- 4.533	- .421	+ 3.772	+ 8.164	+ 8.061	+ 3.687	+ .036	- .026			

$u_f=93.29$ $u_F=u_f-\sum [urf]k_r=93.29-73.57=19.72$

$K=\sqrt{\frac{u_F}{u_f}}=4.6$

TABLE LXXXVI.

For 23 conditions

At U_1

Nothing closure

	r	1	2	3	4	5	6	7	8	18	14	15	16	21	22	
s	[ur]	Values of [ur] k _r														
1	+	6.72	+11.882	-.885	-.171	+2.138	-1.163	+ .579	-.130	+ .037	+ 1.661	-.132	+ .371	+ .071	- 5.244	-3.938
2	+	.80	-.087	-.854	+.090	+.019	+.012	+.043	-.005	+.003	-.014	-.018	-.001	-.008	-.055	-.027
3	+	4.59	-.117	-1.378	+.488	-.008	+.030	+.138	-.036	+.005	-.045	-.048	-.005	-.027	-.321	-.046
4	+	11.88	-3.764	+.768	+.021	-1.839	+.183	-.073	-.032	+.054	-.421	+.253	-.037	-.202	-.300	+1.182
5	+	12.15	-2.102	-.508	+.090	-.188	+10.247	-.208	+.318	+2.048	-4.310	+3.753	-2.652	-.437	+1.465	-.582
6	+	8.51	-.802	+.504	-.106	-.022	+.080	-2.863	+.607	-.069	+1.081	+1.248	-.181	+.787	+.826	-.127
7	+	2.18	+.042	-.087	+.017	-.004	-.057	+.377	-.128	-.000	-.093	-.164	+.026	-.084	+.069	+.020
8	+	7.53	-.041	+.082	-.008	+.035	-1.268	-.149	-.002	-.437	+.523	-.285	+.258	+.055	+.151	+.178
13	+	10.44	-2.581	+.482	-.101	+.372	-3.703	-3.214	+.445	-.726	+20.901	-.322	-3.716	+9.882	-1.371	-3.841
14	+	3.13	+.613	-.146	+.033	-.877	-.987	+1.113	-.235	-.119	+.096	-5.049	+2.788	+1.149	+.439	-.381
15	+	5.64	-.811	+.020	-.006	-.017	+1.231	+ .259	-.067	-.193	-2.007	-4.937	+3.836	-.016	+.254	+.127
16	+	16.22	-1.703	-.431	+.036	-.277	+.584	+3.635	-.623	+.118	-15.321	+5.953	+.047	-11.634	+.750	+3.622
21	+	4.73	-3.681	+.886	-.831	+.120	+.570	-.439	+.149	-.095	-.621	+.063	-.213	-.228	+6.495	+1.325
22	+	3.69	-2.169	+.332	-.037	-.369	-.177	+.134	-.035	-.087	-1.358	+.449	+.083	-.833	+1.060	+4.183
Sum=k _r		+1.553	-.670	+.063	+.561	+3.120	-.708	+.236	+.537	+.072	+.414	+.604	-1.537	+2.244	+1.764	
Multiplier=[ur]		+6.72	-.30	+4.59	-11.83	+12.15	-3.51	-2.18	-7.53	+10.44	-3.13	+5.64	-16.22	+4.73	+3.69	
[ur] k _r		+10.436	+.201	+.289	-6.637	+37.906	+2.485	-.514	-4.044	+.752	-1.296	+3.407	+24.930	+10.614	+6.509	

$u_f=93.29$ $u_F=u_f-\sum[ur]k_r=93.29-85.04=8.25$

$K=\sqrt{\frac{u_F}{u_f}}=.30$

TABLE LXXXVII.

For 26 conditions

At U_1

Easting closure

Easting closure

	r	1	2	3	4	5	6	7	8	13	14	15	16	21	22	24	25	
s	[u s f]	Values of [u s f] k_r																
1	- .30	- .538	0	- .007	- .099	+ .056	- .020	+ .007	- .000	- .082	+ .052	- .014	- .034	+ .234	+ .179	+ .064	+ .017	
2	- 6.73	0	- 13.069	+ 2.218	+ .153	+ .452	+ 1.251	+ .008	+ .151	- 1.161	- 1.334	+ .762	- .304	- 1.424	- .876	+ 5.232	+ 4.009	
3	- 11.83	+ .270	+ 3.904	- 1.890	0	+ .012	- .246	+ .077	+ .011	+ .180	+ .523	- .241	+ .012	+ .909	+ .150	+ .800	- 1.239	
4	- 4.59	- 1.515	- .105	0	- .539	+ .095	+ .005	- .004	+ .030	- .803	+ .070	- .005	- .094	- .117	+ .477	- .853	+ .058	
5	- 3.51	+ .654	+ .236	+ .004	+ .073	- 3.020	0	- .098	- .809	+ 1.232	- 1.092	+ .806	+ .122	- .436	+ .151	- .310	+ .152	
6	- 12.15	- .818	+ 2.762	- .283	+ .012	0	- 10.454	+ 2.107	- .304	+ 3.779	+ 4.265	- .423	+ 2.796	+ 1.073	- .527	- 1.508	+ .524	
7	- 7.53	+ .169	+ .007	+ .046	- .007	+ .188	+ 1.306	- .448	0	- .279	- .523	+ .053	- .277	- .231	+ .064	- .146	- .165	
8	+ 2.18	+ .002	- .049	- .042	- .014	+ .578	+ .055	0	+ .130	- .151	+ .081	- .080	- .015	- .042	- .048	+ .067	- .019	
13	- 3.13	- .854	- .541	+ .048	- .138	+ 1.099	+ .973	- .116	+ .217	- 6.384	0	+ 1.197	- 2.996	+ .894	+ 1.176	+ .515	+ .379	
14	- 10.44	+ 1.808	- 2.849	+ .421	+ .159	- 3.247	+ 3.864	- .725	- .387	0	- 21.294	+ 9.994	+ 5.992	- 1.717	- 1.263	+ 1.814	+ 3.922	
15	- 16.22	- .735	+ 1.839	- .331	- .016	+ 3.732	- .564	+ .114	+ .597	+ 6.203	+ 15.527	- 11.802	0	+ .985	- .362	- .819	- 3.727	
16	- 5.64	- .640	- .255	+ .006	- .116	+ .196	+ 1.298	- .208	+ .039	- 5.899	+ 2.157	- 11.802	0	- 4.104	+ .266	+ 1.296	+ .337	
21	- .14	+ .109	- .030	+ .018	- .064	- .017	+ .012	- .004	+ .003	+ .018	- .023	+ .008	+ .007	- .198	- .041	- .000	+ .010	
22	- .53	+ .316	- .090	+ .007	+ .055	+ .023	- .023	+ .005	+ .012	- .189	- .064	- .012	- .122	- .156	- .599	- .058	- .000	
24	- 4.73	+ 1.062	+ 3.682	+ .120	+ .363	- .418	- .587	- .092	- .015	+ .778	+ .596	- .239	+ .882	- .001	- .839	- 6.514	- 1.388	
25	- 3.69	+ .296	+ 2.201	- .383	+ .047	+ .160	+ .169	- .081	+ .031	+ .446	+ 1.886	- .849	- .082	+ .262	- .001	- 1.088	- 4.167	
Sum = k_r		- .570	- 1.587	+ .579	- .070	- .687	- 3.171	+ .550	- .094	- .824	- .178	- .842	- .578	- .192	- .063	- 2.236	- 1.780	
Multiplier = [u s f]		- .30	- 6.73	- 11.83	- 4.59	- 3.51	- 12.15	- 7.53	+ 2.18	- 3.13	- 10.44	- 16.22	- 5.64	- .14	- .53	- 4.73	- 3.69	
[u s f] k_r		+ .171	+ 10.663	- 6.860	+ .321	+ 2.411	+ 38.528	- 4.142	- .205	+ 2.579	+ 1.806	+ 13.657	+ 3.232	+ .027	+ .093	+ 10.576	+ 6.468	

$u_f = 93.29$
 $u_R = u_f - \sum [u s f] \cdot k_r = 93.29 - 79.27 = 14.02$

$K = \sqrt{\frac{u_R}{u_f}} = .39$

To consider the probable errors at some other points the values of the R.H.S. of the corresponding equations have to be found as was done in table *L* for U_1 . The details are given below in table *LXXXVIII*.

TABLE LXXXVIII.

Circuits	Equations	At N ₁ computed along route A, T, S, Q, P, N ₁														At J ₁ along A, to N ₁ as before and L ₁ J ₁									
		A ₁ T ₁	T ₁ S ₁	S ₁ Q ₁	Q ₁ P ₁	P ₁ N ₁	Side	Az.	A ₁ T ₁	T ₁ S ₁	S ₁ Q ₁	Q ₁ P ₁	P ₁ N ₁	East- ing	North- ing	N ₁ L ₁	L ₁ J ₁	Side	Az.	N ₁ L ₁	L ₁ J ₁	East- ing	North- ing		
I	1	+ .35					0	- .17						- .17	+ .02			+ .35	0			- .17	+ .02		
	2	0					+ .35	- .02						- .02	- .17			0	+ .35			- .02	- .17		
	3	+ .15					+1.81	- .13						- .13	- .87			+ .15	+1.81			- .13	- .87		
	4	-1.81					-1.81	+ .15	+ .87					+ .87	- .13			-1.81	+ .15			+ .87	- .13		
II	5		+ .25				0							- .34	+ .04				0				- .34	+ .04	
	6		0				+ .25							- .04	- .34			+ .25	0			- .04	- .34		
	7		+ .05				+ .05	+1.46						- .30	- 1.95				0	+ .25			- .30	- 1.95	
	8		-1.46				-1.46	+ .05	+ .46					+ 1.95	- .30			+ .05	+1.46			+ 1.95	- .30		
III	9			+ .53	+ .37	+ .82	+1.72	0					-1.27	-1.83	+ 4.32	+ 6.92	+ .95	+ .31	+ .37	+3.40	0	-1.93	-1.90	-10.76	+ 3.55
	10			0	0	0	0	+1.72					- .15	- .19	- .61	- .95	- 6.92	0	0	+ 2.40	- .52	-1.08	- 2.55	-10.76	
	11			+ .07	+ .24	- .82	+ .39	+4.82					-2.69	-1.38	+ 3.04	- 1.03	-19.22	- .61	- .33	- .55	+5.50	+3.99	+1.36	+ 3.32	-25.28
	12			-1.62	-1.05	-2.15	-4.82	+ .39					+3.64	+3.66	+11.92	+19.22	- 1.03	- .52	- .16	-5.50	- .55	+4.27	+1.79	+25.28	+ 3.32
VIII	23																	+ .31	+ .37	+ .68	0	-1.93	-1.90	- 3.63	+ 1.60
	26																	0	+ .68	- .62	-1.08	- 1.60	- 3.68		

Circuits	Equations	At G, along A, to J, as before and H, G.								Circuits	Equations	At C, along A, B, C.							
		J ₁ H ₁	H ₁ G ₁	Side	Az.	J ₁ H ₁	H ₁ G ₁	East- ing	North- ing			A, B, 1	B, C, 1	Side	Az.	A, B, 1	B, C, 1	East- ing	North- ing
I	1			+ .35	0			- .17	+ .02	I	1	+1.08	+1.11	+2.19	0	- .11	- .03	- .14	+ 4.78
	2			0	+ .35			- .02	- .17		2	0	0	0	+2.19	-1.15	-3.58	- 4.78	- .14
	3			+ .15	+1.61			- .13	- .87		3	+ .88	+ .98	+1.88	+6.62	-4.40	-6.56	-10.96	+ 3.61
	4			-1.81	+ .15			+ .87	- .13		4	-4.46	-2.16	-6.62	+1.86	- .50	-3.11	- 3.61	-10.96
II	5			+ .25	0			- .34	+ .04	VI	21	+1.08	+1.11	+2.19	0	- .11	- .03	- .14	+ 4.78
	6			0	+ .25			- .04	- .34		24	0	0	0	+2.19	-1.15	-3.58	- 4.78	- .14
	7			+ .05	+1.46			- .30	- 1.95										
	8			-1.46	+ .05			+ 1.95	- .30										
III	9			+2.40	0			-10.75	+ 2.55										
	10			0	+2.40			- 2.55	-10.75										
	11			- .55	+5.50			+ 3.32	-25.25										
	12			-5.50	- .55			+25.25	+ 3.32										
V	17	+0.38	+ .32	- .70	0	-1.53	-1.10	- 2.63	+ 3.57										
	18	0	0	0	+ .70	-1.63	-1.94	- 3.57	- 2.63										
	19	- .45	- .18	- .63	+1.17	-2.04	- .75	- 2.79	- 7.65										
	20	- .98	- .24	-1.17	- .63	+5.74	+1.91	+ 7.65	- 2.79										
VIII	23	+ .38	+ .32	+1.38	0	-1.53	-1.10	- 6.46	+ 5.17										
	26			0	+1.38	-1.63	-1.94	- 5.17	- 6.46										

Note.—The values, in columns of Side, Azimuth, Easting and Northing at J, and G, have been brought from the corresponding columns at N, by adding to them the quantities N, L₁, L₁ J, to find those at J; and similar quantities J, H₁, H₁ G, have been added to the quantities at J, to find those at G.

Note.—The values, in columns of Side, Azimuth, Easting and Northing at J, and G, have been brought from the corresponding columns at N, by adding to them the quantities N, L, L, J, to find those at J; and similarly the quantities J, H, H, G, have been added to the quantities at J, to find those at G.

The computations of k_r for the following are omitted, and only the values of u_f and k , for Side, Azimuth, Easting and Northing closures are given in the following Table.

TABLE LXXXIX

C ₁	G ₁ —(Continued)	G ₁ —(Continued)	J ₁ —(Continued)	N ₁ —(Continued)
<p>20 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 + .483 + 2.19 + 1.058</p> <p>2 + .032 + 1.86 + .097</p> <p>3 + .040 + 6.62 + .285</p> <p>$\Sigma [ur] k_r = + 1.440$</p> <p>$u_f = 2.19$ $K = .69$</p> <p>23 conditions—A. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>2 - .125 + .35 - .044</p> <p>3 + .092 + 1.81 + .187</p> <p>4 + .018 + .15 - .003</p> <p>5 + .033 + .25 + .008</p> <p>6 - .059 + 1.46 - .088</p> <p>7 + .078 + .05 + .004</p> <p>8 + .435 + 2.40 + 1.044</p> <p>9 + .118 + 5.50 + .638</p> <p>10 - .118 - .55 + .065</p> <p>11 + .865 + .70 + .256</p> <p>12 - .089 + 1.17 - .104</p> <p>13 - .184 - .63 + .105</p> <p>$\Sigma [ur] k_r = + 2.048$</p> <p>$u_f = 3.70$ $K = .67$</p> <p>20 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .531 - .14 + .073</p> <p>2 - 2.042 - 4.73 + 9.859</p> <p>3 + .285 - 10.96 - 2.904</p> <p>4 - .111 - 3.61 + .401</p> <p>$\Sigma [ur] k_r = + 7.220$</p> <p>$u_f = 13.67$ $K = .69$</p> <p>23 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .421 - .14 + .059</p> <p>2 - 2.073 - 4.73 + 9.805</p> <p>3 + .279 - 10.96 - 3.065</p> <p>4 - .131 - 3.61 + .473</p> <p>21 - .273 - .14 + .088</p> <p>$\Sigma [ur] k_r = + 7.317$</p> <p>$u_f = 13.67$ $K = .68$</p> <p>23 conditions—N. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 + 1.089 + 4.73 + 5.151</p> <p>2 - .259 - .14 + .038</p> <p>3 - .007 + 3.61 - .025</p> <p>4 + .387 - 10.96 - 4.242</p> <p>21 + 2.048 + 4.73 + 9.687</p> <p>$\Sigma [ur] k_r = + 10.607$</p> <p>$u_f = 13.67$ $K = .47$</p> <p>26 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .081 - .14 + .011</p> <p>2 - 1.144 - 4.73 + 5.411</p> <p>3 + .407 - 10.96 - 4.481</p> <p>4 - .003 - 3.61 + .011</p> <p>21 - .229 - .14 + .032</p> <p>24 - 2.247 - 4.73 + 10.628</p> <p>$\Sigma [ur] k_r = + 11.632$</p> <p>$u_f = 13.67$ $K = .39$</p> <p>G₁</p> <p>20 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .103 + .35 - .036</p> <p>2 - .021 + .15 - .003</p> <p>3 + .081 - 1.81 + .147</p> <p>4 + .013 + .25 + .003</p> <p>5 + .072 + .05 + .004</p> <p>6 + .053 - 1.46 - .077</p> <p>7 + .452 + 2.40 + 1.085</p> <p>8 - .122 - .55 + .087</p> <p>9 - .103 - 5.50 + .587</p> <p>10 + .425 + .70 + .288</p> <p>11 - .149 - .63 + .064</p> <p>12 + .105 - 1.17 - .133</p> <p>$\Sigma [ur] k_r = + 2.026$</p> <p>$u_f = 3.70$ $K = .67$</p>	<p>23 conditions—A. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>2 - .125 + .35 - .044</p> <p>3 + .092 + 1.81 + .187</p> <p>4 + .018 + .15 - .003</p> <p>5 + .033 + .25 + .008</p> <p>6 - .059 + 1.46 - .088</p> <p>7 + .078 + .05 + .004</p> <p>8 + .435 + 2.40 + 1.044</p> <p>9 + .118 + 5.50 + .638</p> <p>10 - .118 - .55 + .065</p> <p>11 + .865 + .70 + .256</p> <p>12 - .089 + 1.17 - .104</p> <p>13 - .184 - .63 + .105</p> <p>$\Sigma [ur] k_r = + 2.048$</p> <p>$u_f = 3.70$ $K = .67$</p> <p>20 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .098 - .17 + .011</p> <p>2 - .095 - .02 + .002</p> <p>3 + .079 - .13 - .010</p> <p>4 + .050 + .87 + .044</p> <p>5 - .297 - .34 + .100</p> <p>6 + .481 - .04 - .020</p> <p>7 - .231 - .30 + .084</p> <p>8 - .340 + 1.95 - .683</p> <p>9 - 2.002 - 10.75 + 21.522</p> <p>10 - 1.706 - 2.55 + 4.350</p> <p>11 + .557 + 3.32 + 1.949</p> <p>12 + .907 + 25.28 + 15.345</p> <p>13 - .272 - 2.63 + .715</p> <p>14 - 3.528 - 3.57 + 12.599</p> <p>15 + 1.232 - 2.79 - 3.421</p> <p>16 + 2.202 + 7.65 + 1.515</p> <p>$\Sigma [ur] k_r = + 54.152$</p> <p>$u_f = 87.31$ $K = .62$</p> <p>23 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .482 - .17 + .082</p> <p>2 + .018 - .02 - .000</p> <p>3 + .025 - .18 - .003</p> <p>4 + .039 + .87 - .031</p> <p>5 - .075 - .34 + .026</p> <p>6 + .180 - .04 - .007</p> <p>7 - .193 - .30 + .053</p> <p>8 - .297 + 1.95 - .579</p> <p>9 - 1.502 - 10.75 + 16.147</p> <p>10 - 1.353 - 2.65 + 3.463</p> <p>11 + .381 + 3.32 + 1.199</p> <p>12 + .578 + 25.28 + 14.561</p> <p>13 + 1.078 - 2.63 - 2.830</p> <p>14 - 2.868 - 3.57 + 10.232</p> <p>15 + .673 - 2.79 - 1.892</p> <p>16 + .388 + 7.65 - 2.988</p> <p>17 - 3.306 - 6.46 + 21.357</p> <p>$\Sigma [ur] k_r = + 64.813$</p> <p>$u_f = 87.31$ $K = .61$</p> <p>23 conditions—N. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .020 + .02 - .001</p> <p>2 - .054 - .17 + .008</p> <p>3 - .062 - .67 + .054</p> <p>4 + .069 - .19 - .012</p> <p>5 - .607 + .94 - .024</p> <p>6 - 1.093 - .94 + .372</p> <p>7 + 2.84 - 1.95 - .554</p> <p>8 - .328 - .30 + .098</p> <p>9 + 1.281 + 3.65 + 3.189</p> <p>10 - 2.626 - 10.75 + 24.381</p> <p>11 - .437 - 25.28 + 11.047</p> <p>12 + .609 + 3.32 + 2.022</p> <p>13 + 2.331 + 3.57 + 3.322</p> <p>14 - .954 - 2.63 + 2.509</p> <p>15 + 1.854 - 7.65 - 1.408</p> <p>16 + .841 - 2.79 - 2.625</p> <p>17 + 2.831 + 5.17 + 14.802</p> <p>$\Sigma [ur] k_r = + 63.130$</p> <p>$u_f = 87.31$ $K = .54$</p>	<p>26 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .448 - .17 + .076</p> <p>2 + .125 - .02 - .003</p> <p>3 + .050 - .13 - .007</p> <p>4 + .046 + .87 + .043</p> <p>5 + .139 - .34 - .047</p> <p>6 + .296 - .04 - .012</p> <p>7 - .238 - .30 + .071</p> <p>8 - .243 + 1.45 - .474</p> <p>9 - 1.778 - 10.75 + 19.081</p> <p>10 - .889 - 2.55 + 2.287</p> <p>11 + .832 + 3.32 + 1.268</p> <p>12 + .402 + 25.28 + 10.163</p> <p>13 + .430 - 2.63 - 1.131</p> <p>14 - 1.675 - 3.57 + 5.623</p> <p>15 + .493 - 2.70 - 1.875</p> <p>16 + .002 + 7.65 + .015</p> <p>17 - 4.618 - 6.46 + 20.818</p> <p>18 - 2.861 - 5.17 + 14.791</p> <p>$\Sigma [ur] k_r = + 80.168$</p> <p>$u_f = 87.31$ $K = .29$</p> <p>J₁</p> <p>20 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .088 + .35 - .080</p> <p>2 - .024 + .15 - .004</p> <p>3 - .078 - 1.81 + .143</p> <p>4 - .077 + .25 - .019</p> <p>5 + .076 + .05 + .004</p> <p>6 + .041 - 1.46 - .080</p> <p>7 + .512 - 2.40 + 1.243</p> <p>8 - .120 - .55 + .066</p> <p>9 - .088 - 5.50 + .424</p> <p>$\Sigma [ur] k_r = + 1.830$</p> <p>$u_f = 3.00$ $K = .62$</p> <p>23 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 + .139 + .35 + .048</p> <p>2 - .002 + .15 - .000</p> <p>3 - .128 + .15 - .163</p> <p>4 - .040 - .85 - .032</p> <p>5 + .089 + .05 + .003</p> <p>6 + .048 - 1.46 - .063</p> <p>7 + .500 + 2.40 + 1.200</p> <p>8 - .029 - 5.50 - .051</p> <p>9 - .034 - 5.50 - .462</p> <p>10 + .254 + .08 + .173</p> <p>$\Sigma [ur] k_r = + 2.005$</p> <p>$u_f = 3.00$ $K = .58$</p> <p>23 conditions—A. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>2 - .103 + .35 - .037</p> <p>3 + .038 + 1.81 + .161</p> <p>4 - .028 + .15 - .004</p> <p>5 - .080 + .35 - .015</p> <p>6 - .046 + 1.46 - .067</p> <p>7 + .077 + .05 + .004</p> <p>8 - .508 - 2.40 + 1.214</p> <p>9 + .099 + 5.50 + .545</p> <p>10 + .116 - .55 + .064</p> <p>$\Sigma [ur] k_r = + 1.865$</p> <p>$u_f = 3.00$ $K = .61$</p> <p>26 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 + .118 + .35 + .041</p> <p>2 - .001 + .15 - .000</p> <p>3 - .088 - 1.81 + .177</p> <p>4 - .112 + .25 - .028</p> <p>5 + .061 + .05 + .003</p> <p>6 + .048 - 1.46 - .070</p> <p>7 + .457 + 2.40 + 1.169</p> <p>8 - .092 - .55 + .051</p> <p>9 - .094 - 5.50 + .517</p> <p>10 + .253 + .08 + .172</p> <p>$\Sigma [ur] k_r = + 2.032$</p> <p>$u_f = 3.00$ $K = .67$</p>	<p>20 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .064 - .17 + .011</p> <p>2 - .222 - .02 + .004</p> <p>3 + .088 - .13 - .011</p> <p>4 + .060 + .87 + .052</p> <p>5 - .148 - .34 + .050</p> <p>6 + 1.048 - .04 - .042</p> <p>7 - .840 - .30 + .108</p> <p>8 - .832 + 1.95 - .647</p> <p>9 - 2.393 - 10.75 + 25.725</p> <p>10 - 1.896 - 2.55 + 4.535</p> <p>11 + .852 + 3.32 + 2.165</p> <p>12 + .540 + 25.28 + 13.651</p> <p>$\Sigma [ur] k_r = + 45.901$</p> <p>$u_f = 58.33$ $K = .46$</p> <p>23 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .443 - .17 + .075</p> <p>2 - .126 - .02 + .003</p> <p>3 + .051 - .13 - .007</p> <p>4 + .057 + .87 + .050</p> <p>5 - .017 - .34 + .006</p> <p>6 + .873 - .04 - .034</p> <p>7 - .312 - .30 + .064</p> <p>8 - .315 + 1.95 - .614</p> <p>9 - 2.167 - 10.75 + 23.305</p> <p>10 - 1.721 - 2.55 + 4.389</p> <p>11 + .534 + 3.32 + 1.773</p> <p>12 + .523 + 25.28 + 13.221</p> <p>13 - 1.631 - 3.53 + 6.247</p> <p>$\Sigma [ur] k_r = + 43.497$</p> <p>$u_f = 58.33$ $K = .41$</p> <p>23 conditions—N. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 + .071 + .02 + .001</p> <p>2 - .082 - .17 + .006</p> <p>3 - .074 - .87 + .064</p> <p>4 + .090 - .13 - .012</p> <p>5 - 1.070 + .04 - .043</p> <p>6 - .100 - .34 + .034</p> <p>7 + .320 - 1.95 - .624</p> <p>8 - .877 - .30 + .113</p> <p>9 + 1.746 + 2.55 + 4.452</p> <p>10 - 2.467 - 10.75 + 26.520</p> <p>11 - .494 - 25.28 + 12.438</p> <p>12 + .657 + 3.32 + 3.181</p> <p>13 + .819 + 1.60 + 1.310</p> <p>$\Sigma [ur] k_r = + 46.490$</p> <p>$u_f = 58.33$ $K = .45$</p> <p>26 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .414 - .17 + .070</p> <p>2 + .020 - .02 - .000</p> <p>3 + .053 - .13 - .007</p> <p>4 + .088 + .87 + .059</p> <p>5 + .041 - .34 - .014</p> <p>6 + .901 - .04 - .036</p> <p>7 - .325 - .30 + .098</p> <p>8 - .805 + 1.95 - .595</p> <p>9 - 2.243 - 10.75 + 24.112</p> <p>10 - 1.675 - 2.55 + 4.016</p> <p>11 + .539 + 3.32 + 1.789</p> <p>12 + .473 + 25.28 + 11.957</p> <p>13 - 1.629 - 3.53 + 6.239</p> <p>14 - .819 - 1.60 + 1.310</p> <p>$\Sigma [ur] k_r = + 48.998$</p> <p>$u_f = 58.33$ $K = .40$</p> <p>N₁</p> <p>20 conditions—S. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .092 + .35 - .032</p> <p>2 - .017 + .15 - .003</p> <p>3 - .079 - 1.81 + .143</p> <p>4 - .048 + .25 - .010</p> <p>5 + .042 + .05 + .002</p> <p>6 + .037 - 1.46 - .054</p> <p>7 + .285 + 1.72 + .507</p> <p>8 - .080 + .39 - .023</p> <p>9 - .094 - 5.50 + .453</p> <p>$\Sigma [ur] k_r = + .983$</p> <p>$u_f = 2.32$ $K = .76$</p>	<p>23 conditions—A. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>2 - .114 + .35 - .040</p> <p>3 + .059 + 1.81 + .161</p> <p>4 - .019 + .15 - .003</p> <p>5 - .021 + .25 - .005</p> <p>6 - .043 + 1.46 - .063</p> <p>7 + .049 + .05 + .002</p> <p>8 + .278 + 1.72 + .478</p> <p>9 - .107 + 4.62 + .516</p> <p>10 - .057 + .39 - .022</p> <p>$\Sigma [ur] k_r = + 1.024$</p> <p>$u_f = 2.32$ $K = .75$</p> <p>20 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .080 - .17 + .014</p> <p>2 - .082 - .02 + .002</p> <p>3 + .053 - .13 - .007</p> <p>4 + .042 + .87 + .037</p> <p>5 - .061 - .34 - .021</p> <p>6 + .303 - .04 - .018</p> <p>7 - .202 - .30 + .081</p> <p>8 - .247 + 1.95 - .482</p> <p>9 - 1.270 - 6.92 + 8.788</p> <p>10 - .943 - .95 + .896</p> <p>11 + .271 - 1.03 - .279</p> <p>12 + .435 + 19.22 + 8.361</p> <p>$\Sigma [ur] k_r = + 17.358$</p> <p>$u_f = 32.23$ $K = .63$</p> <p>23 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .420 - .17 + .071</p> <p>2 - .017 - .02 - .000</p> <p>3 + .080 - .13 - .004</p> <p>4 + .038 + .87 + .033</p> <p>5 - 1.07 - .34 + .036</p> <p>6 + .395 - .04 - .016</p> <p>7 - .229 - .30 + .069</p> <p>8 - .279 + 1.95 - .544</p> <p>9 - 1.528 - 6.92 + 10.574</p> <p>10 - 1.076 - .95 + 1.022</p> <p>11 + .348 - 1.03 - .358</p> <p>12 + .446 + 19.22 + 8.572</p> <p>$\Sigma [ur] k_r = + 19.455$</p> <p>$u_f = 32.23$ $K = .68$</p> <p>23 conditions—N. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .088 + .02 - .002</p> <p>2 - .083 - .17 + .006</p> <p>3 - .063 - .87 + .055</p> <p>4 + .061 - .13 - .008</p> <p>5 - .229 + .04 - .009</p> <p>6 - .147 - .34 - .050</p> <p>7 + .274 - 1.95 - .534</p> <p>8 - .194 - .30 + .053</p> <p>9 + 1.039 + .95 + .986</p> <p>10 - 1.188 - 6.92 + 8.221</p> <p>11 - .490 - 19.22 + 9.418</p> <p>12 + .264 - 1.03 - .272</p> <p>$\Sigma [ur] k_r = + 17.969$</p> <p>$u_f = 32.23$ $K = .67$</p> <p>26 conditions—E. closure</p> <p>r k_r $[ur]$ $[ur]k_r$</p> <p>1 - .373 - .17 + .063</p> <p>2 + .160 - .02 - .003</p> <p>3 + .097 - .13 - .005</p> <p>4 + .059 + .87 + .051</p> <p>5 - .139 - .34 - .064</p> <p>6 + .345 - .04 - .014</p> <p>7 - .230 - .30 + .063</p> <p>8 - .306 + 1.95 - .597</p> <p>9 - 1.448 - 6.92 + 10.020</p> <p>10 - 1.178 - .95 + 1.114</p> <p>11 + .340 - 1.03 - .350</p> <p>12 + .497 + 19.22 + 9.553</p> <p>$\Sigma [ur] k_r = + 19.958$</p> <p>$u_f = 32.23$ $K = .62$</p>

The results obtained in tables *LXXX* to *LXXXIX* are now collected in the following table showing values of u_f and u_F for Side, Azimuth, Easting and Northing closures of several points of N.W.Q., the latter for 20, 23 and 26 conditions.

TABLE XC.

	C ₁ or Dehra base				U ₁				G ₁ or Chach base				J ₁				N ₁ or Karachi base			
	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F
Side	2.19	.77	0	0	4.91	1.19	.49	.46	3.70	1.67	0	0	3.00	1.17	1.00	.97	2.32	1.34	0	0
Azimuth	2.19	.77	.76	0	4.91	1.19	2.42	.46	3.70	1.67	1.67	0	3.00	1.17	1.14	.97	2.32	1.34	1.30	0
Easting	13.67	6.44	6.35	2.04	93.29	19.98	19.72	14.02	87.31	33.16	22.50	7.14	58.33	12.43	9.83	9.33	32.28	14.92	12.83	12.32
Northing	13.67	6.44	3.06	2.04	93.29	19.98	6.25	14.02	87.31	33.16	34.18	7.14	58.33	12.43	11.84	9.33	32.28	14.92	14.31	12.32

These values all seem reasonable. Rather unexpected results are 2.42 and 8.25 for 23 conditions for U_1 .

It appears that the greatest probable error of the adjustment of Easting or Northing is $4\sqrt{33} \doteq 23$ feet: in terms of deflection this is negligible and of the order of probable error of latitude (astronomic) result.

As regards azimuths for 20 or 23 conditions the worst case is $1.6\sqrt{2.4}$ i.e. probable error of $2''.4$. Error of $7''$ is in this case possible and liable to occur.

In N. E. Quadrilateral where triangulation is not so good there will be greater errors. Closed on Laplace stations, however, errors are probably reduced to $1.6\sqrt{.5}$ and $1.6\sqrt{1.0}$ i.e. probable error to 1.6 and possible to $5''$.

In the above the "possible error" is regarded as three times the probable error.

As further discussion of these results is at present impossible, for reasons explained in the preface the chapter is concluded with a tabular statement of the probable errors of log. side to 7th place of decimals, azimuth in seconds and easting and northing in feet; these are obtained as explained in § 20.

TABLE XCI

	C ₁ or Dehra base				U ₁				G ₁ or Chach base				J ₁				N ₁ or Karachi base			
	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F	u_f	20 u_F	23 u_F	26 u_F
Side	49	29	0	0	74	36	23	23	64	43	0	0	57	36	34	33	50	39	0	0
Azimuth	2.33	1.89	1.37	0	3.50	1.72	2.46	1.07	3.02	2.03	2.03	0	2.72	1.70	1.69	1.54	2.39	1.83	1.60	0
Easting	feet 14.81	feet 10.24	feet 10.16	feet 5.76	feet 38.93	feet 18.01	feet 17.89	feet 15.07	feet 37.64	feet 23.21	feet 19.10	feet 10.76	feet 30.79	feet 14.23	feet 12.63	feet 12.29	feet 22.89	feet 15.66	feet 14.43	feet 14.15
Northing	14.81	10.24	7.05	5.76	38.93	18.01	11.67	15.07	37.64	23.21	19.83	10.76	30.79	14.23	13.86	12.29	22.89	15.66	15.23	14.15

CHAPTER IX.

Deflections of the Plumb-line and values of "g" derived from observations of the Survey of India.

1. The first use of the tables derived in the earlier chapters will now be made use of to display in convenient form all the data of plumb-line deflections available up to the time of writing (May 1917). As regards the deflections in meridian no comment is necessary. The results of observation are immediately available. With the deflections in prime vertical the case is different. It has been stated already that the triangulation of India was not adjusted on the longitude arcs, which have been subsequently observed or reduced. The triangulation accordingly is burdened with an accumulation of error in azimuth which may be largely reduced by adjustment on the longitude arcs. For this purpose it is not essential for the present purpose to reopen the adjustment of the whole triangulation except as regards the azimuth, and the process followed will be substantially that followed by Colonel Sir Sidney Burrard in Appendix 5, G.T.S. Volume XVIII. The numerical results will however be slightly different owing to the improved methods of computing the effect of a change in azimuth at the origin which have been developed in Chapters I—III; the differences will depend mainly on the taking into account of the effect of a change on azimuth at the origin on longitudes of points considerably removed from the origin.

When the triangulation of India was adjusted General Walker decided to adopt a value of the fundamental azimuth (of Surantal from Kalianpur) which differed from the observed value by a small amount (*vide* Chapter I, § 4) and this of course implied a deflection in prime vertical at Kalianpur. This has given rise to a little confusion as regards the longitudes of India. The astronomic longitude of Kalianpur has been determined with reference to Greenwich, but no account of the implied deflection in prime vertical has been hitherto considered.* Colonel Sir Sidney Burrard in adjusting the azimuth observations eliminated the effect of this oversight by returning to an observed value of the fundamental azimuth. The deflection in meridian remained in terms of the Everest spheroid with Walker's initial azimuth. In dealing with the deflections as a whole it will accordingly be better to keep the azimuths in the same terms as the latitudes, and to recognise that a deflection in prime vertical at Kalianpur is thereby implied. After the adjustments on the longitude arcs have been performed, the results of both azimuth and latitude deflections will be in common terms of Everest spheroid and Walker's origin.

Quantities for correcting all the deflections to refer to any other spheroid and origin are given in table XCV in which all the results are exhibited, as well as the deflections corrected to the special case of Helmert's spheroid and the latest observed value of latitude and azimuth at Kalianpur, as derived from observations at a group of stations surrounding Kalianpur.†

* *Vide*, p. xv G.T.S. Vol. XVII, Survey of India.

† *Vide*, pp. 7,9 Professional Paper No. 5, Survey of India.

2. As the correction for azimuths has already been treated by Colonel Sir Sidney Burrard *loc. cit.* it will not be necessary to state afresh the various practical difficulties which arose owing to longitude stations not being in general identical with azimuth stations. The observation results exhibited by him will be taken unaltered, and immediately applied to Laplace's equation. This equation has been given in somewhat amplified form in (3) of Chapter V. There is now no occasion to consider observation errors of astronomic azimuths or their determination. The accumulated error of geodetic longitude determination is certainly small compared with that of geodetic azimuth and so will be neglected. The equation may accordingly be written

$$(A - G - \delta G) \operatorname{cosec} \lambda - (A_0 - G_0) \operatorname{cosec} \lambda_0 = A - G \quad \dots \dots \dots (1)$$

where the notation has been changed in conformity with the usual practice and A, A and G, G signify astronomic and geodetic determinations respectively, roman letters referring to azimuth and italic letters to longitude determinations: δG is the correction necessary to the geodetic value of azimuth: as this is the quantity required it will be convenient to rewrite (1). From Chapter I § 4 it is seen that $A_0 - G_0 = +1.29$, $\lambda_0 = 24^\circ 7' 12''$, $(A_0 - G_0) \operatorname{cosec} \lambda_0 = 3.16$. Hence

$$\delta G = A - G - (3.16 + A - G) \sin \lambda \quad \dots \dots \dots (2)$$

which serves to determine δG . So long as Walker's value of azimuth is adhered to all *geodetic* longitudes of India require a correction of $-3''.16$.

For the Helmert spheroid an additional correction $\delta_2 G$ is necessary. The corresponding equation to (1) is

$$(A - G - \delta G - \delta_1 G - w) \operatorname{cosec} \lambda = A - G - v \quad \dots \dots \dots (3)$$

since G, G are changed by v and w respectively and the astronomic and geodetic azimuths at the origin have been made identical. Subtracting (3) from (2) it follows that

$$\delta_2 G = 3.1571 \sin \lambda + v \sin \lambda - w \quad \dots \dots \dots (4)$$

The solutions of (2) and (4) and the deduction of δG and $\delta_2 G$ are now shown in table XCIII.

3 Having obtained the values of δG and $\delta_2 G$ at all Laplace stations it is next necessary to find values at the intervening azimuth stations. This is done in table XCIV, interpolating according to the number of removes from terminal stations. The Laplace stations are shown in block type. Azimuth stations between which adjustments have been performed are shown in italics.

4. The precision of deflections in prime vertical so far as is due to the astronomical observation, obtained from azimuth observation, is much less than those in meridian.* As may be seen from results obtained in Chapter VII the probable error of azimuth generated in triangulation is much greater than that in latitude or longitude, all being expressed in seconds. The deflection in prime vertical is derived from the azimuth anomaly by multiplication by $\cot \lambda$ —a quantity which ranges from 7 to 1.4 in Indian latitudes—and this further increases the lack of precision. Considerable improvement on the other hand should result by the use of Laplace stations, which has been made. A further source of weakness, of varying amount, is the actual azimuth observation itself. The observation is not nearly so satisfactory as that for latitude and involves graduation error of the instrument which, especially in the older observations, introduces a serious uncertainty. It is desirable then to consider the relative degree of reliability of the azimuth observations. On account of the other sources of error, mainly that of accumulation of error in triangulation it is not useful to do this in very great detail, and it is considered sufficient to work out the probable error from the

* For probable errors in astronomic latitude vide *G.T.S. Vol. XI, pages 882—982* and *G.T.S. Vol. XVIII App. 7 Table III*. These are seldom so great as $0''.2$. The worst case (Gogipatri) is $\pm 0''.68$.

results obtained on the several zeros. This takes no account of errors in star places, a defect which was more serious in the early days of the survey than it is at present. Probably graduation error in the instruments has improved at much the same rate as the error of star place, and a fair estimate of probable error will be obtained by consideration of the probable error due to graduation only. The formula used is

$$p.e. \text{ of mean of observations on } n \text{ zeros} = .6745 \sqrt{\frac{\sum \delta^2}{n(n-1)}} \quad \dots \quad (5)$$

where δ is the discrepancy of the value derived from any zero from the mean result. The results of the application of this formula are given in table XCIV. Some idea of the probable errors in the geodetic values of latitude, longitude and azimuth is given in Chapter VIII, where the N.W. Quadrilateral is considered in detail.

5. In the table XCV the deflections derived from observations of the Survey of India are given. These have been arranged by degree sheets. On the left hand page of the table the data are expressed in terms of the Everest Spheroid, using the observed value of latitude and the deduced value of azimuth at Kalianpur which General Walker adopted: on the right hand page of the table the quantities are given which must be applied properly to the triangulated values to express in terms of any other spheroid. There are four variables to be considered, giving rise to four cases: these are (1) change of semi-major axis, δa , (2) change of semi-minor axis, δb , (3) change of latitude of origin, u_0 and (4) change of azimuth at origin, w_0 . The cases given correspond to $\delta a = 1 \text{ km.}$, $\delta b = 1 \text{ km.}$, $u_0 = 1''$, $w_0 = 1''$, and to obtain the general case these must be combined as follows:—

$$\delta a \times \text{case I} + \delta b \times \text{case II} + u_0 \times \text{case III} + w_0 \times \text{case IV}$$

in which δa , δb are expressed in kilometres and u_0 , w_0 are expressed in seconds. Thus in the case of the Helmert Spheroid in which $a = 6378.2 \text{ km.}$ and $1/e = 298.3$ and with revised values of latitude and azimuth at Kalianpur as given on p. 2, $\delta a = .924$, $\delta b = .743$, $u_0 = .31$, $w_0 = 1.29$. Deflections in terms of this spheroid are given on the right of the right hand page of the table: but those in terms of any other spheroid may be easily found by making use of different values of δa , etc. It is clear in the notation of this work that the correction to latitude deflection is $-u$, to prime vertical deflection $-v \cos \lambda$ or $-w \cot \lambda$ according as the deflection is derived from longitude or azimuth observations. Values of these quantities have been taken from tables XVII—XX, XXIX—XXXVI.

It frequently happens that latitude observations have been made on a site not quite identical with the triangulation station, but at some (small) distance from it on the prime vertical through it. Similarly the longitude observations done by wire-telegraphy were made at the telegraph offices and not exactly on the station sites. The coordinates of the triangulation stations are generally the quantities given. But at latitude stations the value of latitude given is that of the latitude station, and in longitude stations the longitude of the longitude station. When a change is made to a different spheroid, since the corrections do not satisfy Laplace equation, slight discrepancies occur between deflection derived from longitude and azimuth observations. This point has been explained in Chapter V and has been taken into account in the case of the Helmert spheroid in table XCIII.

The elevation of the referring mark affects the result of azimuth observations (*vide* §6 Chapter V) and accordingly this has been given in the table except for a few cases where the data could not be found.

Longitude arcs by means of wireless telegraphy were observed in collaboration with the expedition of Cav. de Filippi in 1914 between Dehra Dun and eight stations. Their names and the values of A are appended. The Dehra Dun observations were made in the transit room adjoining the

Dome Observatory (new) and the longitude (geodetic) of the transit instrument is $7'' \cdot 18$ (equivalent linear measurement being 628·8 feet*) less than that of the Dehra Dun Haig Observatory where all previous longitude observations were made. The geodetic values of the eight stations and the astronomic values of the latter four are not yet available (1917) and deflections cannot be given.

TABLE XCII.

	Skardu	Kargil	Lamayaru	Leh	Depsang	Suget Karaul	Yarkand	Kashgar
Δ	$75^{\circ} 38' 22'' \cdot 80$	$76^{\circ} 7' 40'' \cdot 65$	$76^{\circ} 46' 32'' \cdot 01$	$77^{\circ} 34' 53'' \cdot 78$	$77^{\circ} 56' 49'' \cdot 4$	$78^{\circ} 12' '' \cdot 4$	$77^{\circ} 15' 55'' \cdot 4$	$76^{\circ} 6' 47'' \cdot 4$
G
Lat.	$35^{\circ} 18' 40'' \cdot 4$	$34^{\circ} 33' 40'' \cdot 4$	$34^{\circ} 17' 1'' \cdot 4$	$34^{\circ} 10' 4'' \cdot 4$	$35^{\circ} 17' 20'' \cdot 4$	$36^{\circ} 20' 56'' \cdot 4$	$38^{\circ} 25' 1'' \cdot 4$	$39^{\circ} 24' 26'' \cdot 4$

TABLE XCIII.

Azimuth station Latitude= λ	Azimuth (Everest)	(1) $A-G$	Longitude station $\sin \lambda$	Longitude from Kalianpur (Everest)	$A-G$	(2) $(A-G) + 9'' \cdot 16 \sin \lambda$	(1)-(2) $= \delta G$	(3) $3 \cdot 1571 \times \sin \lambda$	(4) $\nu \sin \lambda$	(5) ν	(Helmert) $(3)+(4)-(5) = \epsilon_2 G$
Kalianpur H.S. $24^{\circ} 7' 11''$	A 190 27 6·39 G 190 27 5·10	+ 1·29	Kalianpur ... ·4086	A 0 0 0 G 0 0 0	0	+ 1·29	0·0	+ 1·29	0·00	+ 1·29	0·00
Karachi Observatory S. $24^{\circ} 49' 50''$	A 221 39 9·5 G 221 39 10·9	- 1·4	Karachi T.O. ... ·4200	A - 10 38 24·8 G - 10 38 24·3	- 0·5	+ 1·1	- 2·5	+ 1·33	+ 2·41	+ 3·62	+ 0·12
Dehra Dun Obs. (old) S. $30^{\circ} 19' 57''$	A 165 10 58·8 G 165 11 10·7	- 11·9	Dehra Dun ·5050	A + 0 23 39·9 G + 0 24 4·6	- 25·7	- 11·4	- 0·5	+ 1·59	- 0·03	+ 1·25	+ 0·81
Quetta T.O. S. $30^{\circ} 11' 57''$	A 166 31 12·1 G 166 31 17·0	- 4·9	Quetta T.O. ... ·5030	A - 10 38 48·3 G - 10 38 45·8	- 2·5	+ 0·8	- 5·2	+ 1·59	+ 3·06	+ 4·23	+ 0·42
Calcutta Base S., T.S. $22^{\circ} 36' 56''$	A 177 10 27·3 G 177 10 36·2	- 8·9	Calcutta ... ·3846	A + 10 42 0·3 G + 10 42 11·3	- 11·0	- 3·0	- 5·9	+ 1·21	- 2·20	- 0·90	- 0·09
Orejhar S. $26^{\circ} 48' 56''$	A 308 36 18·9 G 308 36 23·0	- 4·1	Fyzabad T.O. ... ·4506	A + 4 28 50·1 G + 4 28 50·6	- 0·5	+ 1·2	- 5·3	+ 1·42	- 1·06	+ 0·22	+ 0·12
Jalpaiguri s. $26^{\circ} 51' 17''$	A 321 33 25·3 G 321 33 30·0	- 4·7	Jalpaiguri ... ·4465	A + 11 4 34·8 G + 11 4 55·2	- 20·4	- 7·7	+ 3·0	+ 1·41	- 2·68	- 1·32	+ 0·06
Nagarkhana H.S. $22^{\circ} 22' 53''$	A 155 47 13·3 G 155 47 23·5	- 10·2	Chittagong T.O. ... ·3808	A + 14 10 47·4 G + 14 10 59·1	- 11·7	- 3·3	- 6·9	+ 1·30	- 2·87	- 1·71	+ 0·04
Dattaung H.S. $20^{\circ} 13' 14''$	A 171 27 28·8 G 171 27 38·1	- 9·8	Akyah T.O. ... ·3455	A + 15 14 21·0 G + 15 14 32·1	- 11·1	- 2·7	- 7·1	+ 1·09	- 2·77	- 1·03	+ 0·25
Kyaunggyi S. $16^{\circ} 49' 13''$	A 109 26 42·1 G 109 26 58·1	- 16·0	Prome ... ·3226	A + 17 33 24·6 G + 17 33 39·9	- 15·3	- 3·9	- 12·1	+ 1·02	- 2·87	- 2·27	+ 0·32
Taungun H.S. $16^{\circ} 25' 49''$	A 31 16 18·9 G 31 16 32·7	- 13·8	Moulmein ... ·2829	A + 19 58 5·9 G + 19 58 22·5	- 16·6	- 3·8	- 10·0	+ 0·89	- 2·82	- 2·54	+ 0·51
Bolarum P.W.D. S. $17^{\circ} 30' 13''$	A 25 57 35·8 G 25 57 36·9	- 1·1	Bolarum ... ·3008	A + 0 51 50·3 G + 0 51 53·6	- 3·3	0·0	- 1·1	+ 0·95	- 0·18	+ 1·10	- 0·33
Vizagapatam Base N., S. $18^{\circ} 1' 3''$	A 203 44 24·5 G 203 44 25·9	- 1·4	Waltair ... ·3093	A + 5 39 42·6 G + 5 39 45·8	- 3·2	0·0	- 1·4	+ 0·98	- 0·05	+ 0·83	- 0·30
Karandi H.S. $23^{\circ} 10' 40''$	A 206 23 35·6 G 206 22 39·6	- 4·0	Jabalpur T.O. ... ·3936	A + 2 17 34·8 G + 2 17 45·0	- 10·2	- 2·8	- 1·2	+ 1·24	- 0·49	+ 0·79	- 0·04
Colaba Observatory S. $18^{\circ} 53' 47''$	A 288 5 27·7 G 288 5 26·7	+ 1·0	Bombay ... ·3238	A - 4 50 21·8 G - 4 50 28·6	+ 6·8	+ 3·2	- 2·2	+ 1·02	+ 0·77	+ 2·05	- 0·26
Deesa T. O. s. $24^{\circ} 16' 29''$	A 241 16 15·3 G 241 16 19·9	- 4·6	Deesa T.O. ... ·4108	A - 5 28 16·4 G - 5 28 12·7	- 3·7	- 0·2	- 4·4	+ 1·30	+ 1·20	+ 2·46	+ 0·04
Mangalore S. $12^{\circ} 52' 15''$	A 205 52 50·8 G 205 52 53·6	- 2·8	Mangalore ... ·2227	A - 2 48 32·9 G - 2 48 35·1	+ 2·2	+ 1·2	- 4·0	+ 0·70	+ 0·26	+ 1·55	- 0·59
Bangalore Base S.W., S. $13^{\circ} 0' 41''$	A 224 31 21·7 G 224 31 27·0	- 5·3	Bangalore ... ·2252	A - 0 4 20·3 G - 0 4 17·6	- 2·7	+ 0·1	- 5·4	+ 0·71	- 0·05	+ 1·22	- 0·56
St. Thomas's Mount S. $13^{\circ} 0' 15''$	A 12 30 5·3 G 12 30 9·3	- 4·0	Madras ... ·2350	A + 2 35 29·6 G + 2 35 36·6	- 7·0	- 0·9	- 3·1	+ 0·71	- 0·35	+ 0·89	- 0·53
Kudankulam Obs. S. $8^{\circ} 10' 22''$	A 185 55 18·8 G 185 55 26·5	- 7·7	Nagarkoll ... ·1431	A - 0 13 15·8 G - 0 13 14·2	- 1·6	+ 0·2	- 7·9	+ 0·45	- 0·03	+ 1·19	- 0·77

Obs.=observatory, T.O.=Telegraph office. * Vide G.T.S. Vol. XV, p. (5). † Approximate values. ‡ A, Δ =Astronomic values; G, G=Geodetic values. § This is the additional correction for Helmert's spheroid. || Derived from unadjusted values of Quetta Secondary Series. This does not enter into the azimuth correction.

TABLE XCIV. (See Index pp. 170-172)

Serial Number	Station	Corrections		Prob- able Error ±	Date of Observation	Serial Number	Station	Corrections		Prob- able Error ±	Date of Observation
		Everest's Spheroid	Helmert's Spheroid*					Everest's Spheroid	Helmert's Spheroid*		
1	Kalianpur H.S.	0.0	0.0	0.31	{ 1836	25	Karachi Obs. S.	-2.5	+0.1	0.19	1855
2	Losalli S.	-0.1	0.0	0.44	{ 1898	30	Karachi Base S., S.	-2.4	+0.1	0.30	1853
3	Salot H.S.	-0.2	0.0	0.27	1849	40	Yusuf S.	-2.0	+0.2	0.53	1858
4	Māta-ka-hūra H.S.	-0.3	0.0	0.34	1849	41	Bhanar T.S.	-1.9	+0.2	0.27	1859
5	Gurāria H.S.	-0.4	0.0	0.44	1849	42	Miāni T.S.	-1.8	+0.2	0.32	1859
6	Rāmpūra H.S.	-0.5	0.0	0.49	1849	43	Dājil S.	-1.6	+0.2	0.22	1860
7	Aramlia S.	-0.6	0.0	0.32	1850	44	Dera Dīn Panāh S.	-1.4	+0.2	0.25	1859
8	Nānd H.S.	-0.7	0.0	0.90	1850	45	Jharkil T.S.	-1.2	+0.2	0.27	1859
9	Tiki H.S.	-0.8	0.0	0.38	1851	46	Umarkhel H.S.	-1.0	+0.3	0.20	1909
10	Kānnagar H.S.	-0.9	0.0	0.50	1850	47	Jāoli H.S.	-0.9	+0.3	0.80	1851
11	Gurā Sikkar H.S.	-1.0	0.0	0.45	1850	48	Medwāni H.S.	-0.7	+0.3	0.50	1853
12	Birana S.	-1.1	0.0	0.46	1851	33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853
13	Khanakharā S.	-1.2	0.0	0.33	1851	25	Karachi Obs. S.	-2.5	+0.1	0.19	1855
14	Saria S.	-1.2	+0.1	0.25	1851	49	Andar H.S.	-2.4	+0.1	0.27	1895
15	Didāwa H.S.	-1.3	+0.1	0.21	1851	50	Piara H.S.	-2.4	+0.1	0.27	1896
16	Virāria H.S.	-1.4	+0.1	0.19	1851	33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853
17	Lūnki H.S.	-1.5	+0.1	0.19	1851	49	Andar H.S.	-2.4	+0.1	0.27	1895
18	Kujhā H.S.	-1.6	+0.1	0.29	1851	51	Gandpāhar H.S.	-2.1	+0.1	0.13	1906
19	Chinga H.S.	-1.7	+0.1	0.28	1852	52	Zawa H.S.	-1.9	+0.1	0.20	1905
20	Mairāb-ka-Shahar T.S.	-1.8	+0.1	0.56	1852	53	Mashelak H.S.	-1.9	+0.2	0.34	1908
21	Khorī T.S.	-1.9	+0.1	0.53	1852	54	Gundak H.S.	-1.7	+0.2	0.22	1910
22	Alamkhān T.S.	-2.0	+0.1	0.22	1852	55	Salighar H.S.	-1.6	+0.2	0.19	1910
23	Chāth T.S.	-2.1	+0.1	...	1853	56	Tounsa T.S.	-1.5	+0.2	0.31	1910
24	Kārothol H.S.	-2.2	+0.1	0.31	1853	44	Dera Dīn Panāh S.	-1.4	+0.2	0.25	1859
25	Karachi Obs. S.	-2.5	+0.1	0.19	1855	52	Zawa H.S.	-1.9	+0.1	0.20	1905
1	Kalianpur H.S.	0.0	0.0	0.31	{ 1836	57	Kisanen Chappār H.S.	-1.9	+0.1	0.18	1907
26	Pahārgarh H.S.	-0.1	0.0	0.22	{ 1898	58	Tuzgi H.S.	-1.9	+0.1	0.36	1907
27	Keeri H.S.	-0.1	+0.1	0.39	1836	59	Koh-i-Malik Siah H.S.	-1.9	+0.1	0.30	1907
28	Usra H.S.	-0.2	+0.1	0.18	1836	60	† Quetta T.O. S.	-5.2	+0.4	0.28	1904
29	Noh T.S.	-0.3	+0.2	0.53	1837	5	Gurāria H.S.	-0.4	0.0	0.44	1849
30	Datāiri T.S.	-0.3	+0.2	0.40	1836	61	Kānkra H.S.	-0.4	0.0	0.34	1862
31	Kālāna S.	-0.4	+0.3	0.33	1836	62	Bānskho H.S.	-0.5	+0.1	0.14	1862
32	Banog H.S.	-0.5	+0.3	0.28	{ 1836	63	Tāsing H.S.	-0.5	+0.1	0.70	1861
33	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	{ 1907	64	Rākhi T.S.	-0.6	+0.2	0.19	1860
34	Dehra Dun Obs. (old) S.	-0.5	+0.3	0.34	1853	65	Kheri T.S.	-0.7	+0.2	0.37	1856
35	Rājpur h.s.	-0.5	+0.3	0.26	1853	66	Bowra T.S.	-0.7	+0.2	0.52	1853
36	Sour point h.s.	-0.5	+0.3	0.29	1914	48	Medwāni H.S.	-0.7	+0.3	0.50	1853
37	Museo-ree Dome Obs. H.S.	-0.5	+0.3	0.25	1914	7	Aramlia S.	-0.6	0.0	0.32	1850
38	Nag Tibā H.S.	-0.5	+0.3	0.45	1912	67	Rājgarh H.S.	-0.6	0.0	0.56	1863
39	Banog H.S.	-0.5	+0.3	0.33	1903	68	Garinda S.	-0.7	+0.1	0.42	1863
		-0.5	+0.3	0.23	{ 1836	69	Sirsa S.	-0.7	+0.1	0.34	1861
				{ 1907	70	Sangatpur T.S.	-0.8	+0.2	0.54	1860	
					71	Jāoli H.S.	-0.9	+0.3	0.80	1851	
					72	Murree h.s.	-0.9	+0.3	0.77	1860	
					73	Ganga Choti H.S.	-0.9	+0.3	0.16	1910	
					74	Poshkar H.S.	-0.9	+0.3	0.65	1862	
					75	Gogipatri H.S.	-0.9	+0.3	0.84	1862	
					76	Rustamgarhi h.s.	-0.9	+0.3	0.67	1862	
					77	Sangatpur T.S.	-0.8	+0.2	0.54	1860	

Note.—In the azimuthal observations, level corrections were introduced from 1850.

Appendix 9, p. 73 and Diurnal Aberration corrections from 1850.

1902, p. 74.

Note. — In the azimuthal observations, Level corrections were introduced from 1868, vide G.T.S. Vol. II Appendix 9, p. 73 and Diurnal Aberration corrections from 1902 vide Handbook of the Trigonometrical Branch 1902, p. 74.

(Obs. = observatory, T.O. = Telegraph Office.
 S = Station, H.S. = hill station, T.S. = tower station of principal triangulation. The same small letters refer to minor triangulation.
 * Additional correction for Helmert's spheroid.
 † Computed from the unadjusted values of longitude and azimuth of Quetta T.O. station: not used for adjusting any azimuth observations.

TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station		Corrections		Probable Error \pm	Date of Observation	Serial Number	Station		Corrections		Probable Error \pm	Date of Observation
			Everest's Spheroid	Helmert's Spheroid*						Everest's Spheroid	Helmert's Spheroid*		
11	Guru Sikhar	H.S.	-1.0	0.0	0.45	1850	90	Amūa	H.S.	-1.4	0.0	0.94	1834
76	Thob	H.S.	-1.0	0.0	0.38	1873	111	Nimkār	T.S.	-2.3	+0.1	0.33	1838
77	Jambo	H.S.	-1.0	+0.1	0.21	1874	103	Ramuapur (old)	T.S.	-2.7	+0.2	0.52	1838
78	Mugrāla	H.S.	-1.0	+0.1	0.40	1875	91	Karāra	H.S.	-1.8	0.0	0.69	1842
79	Lādimsir	T.S.	-1.0	+0.1	0.45	1862	112	Pabhosa	H.S.	-2.4	+0.1	1.26	1845
80	Mandressa	T.S.	-0.9	+0.2	0.17	1862	113	Sora	T.S.	-2.9	+0.1	0.44	1845
81	Jhambhera	T.S.	-0.9	+0.2	0.63	1862	104	Māsi	T.S.	-3.8	+0.2	0.37	1850
82	Akbar	S.	-0.9	+0.2	0.79	1857							
47	Jauli	H.S.	-0.9	+0.3	0.80	1851	92	Gurwāni	H.S.	-2.2	0.0	0.64	1845
18	Rojhra	H.S.	-1.6	+0.1	0.29	1851	114	Marār	T.S.	-3.7	+0.1	0.48	1846
83	Malar	H.S.	-1.5	+0.1	0.32	1877	115	Bisaul	T.S.	-5.1	+0.1	0.60	1847
84	Asu	H.S.	-1.4	+0.1	0.49	1880	106	Orejhar	S.	-5.3	+0.1	0.26	1904
85	Vijnot	T.S.	-1.3	+0.1	0.22	1881	93	Gora	H.S.	-2.7	0.0	0.52	1845
86	Dāowāla	T.S.	-1.3	+0.1	0.20	1881	116	Hirlepur	T.S.	-3.0	0.0	0.58	1846
87	Paphra	T.S.	-1.1	+0.1	0.25	1861	117	Samenda	T.S.	-3.4	0.0	1.13	1846
79	Lādimsir	T.S.	-1.0	+0.1	0.45	1862	118	Rājabārī	T.S.	-3.9	+0.1	1.52	1847
1	Kalianpur	H.S.	0.0	0.0	0.31	1836	105	Bāsadela	T.S.	-4.6	+0.1	0.32	1849
88	Budhon	H.S.	-0.4	0.0	0.36	1893	106	Orejhar	S.	-5.3	+0.1	0.26	1904
89	Rangir (old)	H.S.	-0.8	0.0	0.64	1864	119	Naunangarhi	T.S.	-2.3	+0.1	0.34	1852
90	Amūa	H.S.	-1.4	0.0	0.94	1834	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
91	Karāra	H.S.	-1.8	0.0	0.69	1842	121	Rānganj	T.S.	+2.4	+0.1	0.64	1853
92	Gurwāni	H.S.	-2.2	0.0	0.64	1845	122	Jalpaiguri	S.	+3.0	+0.1	0.33	1904
93	Gora	H.S.	-2.7	0.0	0.52	1845	94	Hurilāong	H.S.	-3.1	-0.1	0.44	1849
94	Hurilāong	H.S.	-3.1	-0.1	0.44	1849	123	Mednipur	T.S.	-2.9	0.0	0.34	1850
95	Chendwār (old)	H.S.	-3.5	-0.1	0.67	1843	124	Jalālpur	T.S.	-2.6	0.0	0.62	1852
96	Pārasnāth	H.S.	-3.7	-0.1	0.35	1850	119	Naunangarhi	T.S.	-2.3	+0.1	0.34	1852
97	Tilabani	H.S.	-3.9	-0.1	0.76	1845	95	Chendwār (old)	H.S.	-3.5	-0.1	0.67	1843
98	Malūncha	H.S.	-4.3	-0.1	0.57	1844	125	Pota	T.S.	-2.8	0.0	0.46	1846
99	Madhpur	T.S.	-4.9	-0.1	0.49	1868	119	Naunangarhi	T.S.	-2.3	+0.1	0.34	1852
100	Aknāpur	T.S.	-5.2	-0.1	0.83	1869	96	Pārasnāth	H.S.	-3.7	-0.1	0.35	1850
101	Calcutta Base-line S. End	T.S.	-5.9	-0.1	0.37	1845	126	Bichwi	H.S.	-3.2	-0.1	0.30	1851
33	Dehra Dun Obs. (old)	S.	-0.5	+0.3	0.34	1853	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
102	Kaliānpur	T.S.	-1.8	+0.2	0.64	1850	98	Malūncha	H.S.	-4.3	-0.1	0.57	1844
103	Ramuapur (old)	T.S.	-2.7	+0.2	0.52	1838	127	Sirkanda	T.S.	-1.7	0.0	0.43	1846
104	Māsi	T.S.	-3.8	+0.2	0.37	1850	120	Chūni	T.S.	+0.9	+0.1	0.69	1846
105	Bāsadela	T.S.	-4.6	+0.1	0.32	1849	101	Calcutta Base-line S. End	T.S.	-5.9	-0.1	0.37	1845
106	Orejhar	S.	-5.3	+0.1	0.26	1904	128	Anandbās	T.S.	-4.2	-0.1	0.45	1846
88	Budhon	H.S.	-0.4	0.0	0.36	1864	129	Madhpur	T.S.	-2.9	0.0	0.63	1846
107	Gūrmī	T.S.	-0.4	+0.1	0.50	1842	122	Jalpaiguri	S.	+3.0	+0.1	0.33	1904
108	Sankrāo	T.S.	-0.4	+0.2	0.39	1843	101	Calcutta Baseline S. End	T.S.	-5.9	-0.1	0.37	1845
109	Sirsa	T.S.	-0.5	+0.2	0.70	1843	130	Daulatpur	T.S.	-6.3	-0.1	0.15	1868
33	Dehra Dun Obs. (old)	S.	-0.5	+0.3	0.34	1853	131	Gangapur	T.S.	-6.5	0.0	0.34	1866
89	Rangir (old)	S.	-0.8	0.0	0.64	1834	132	Lakhinagar	T.S.	-6.5	0.0	0.14	1866
110	Mohammadabad	T.S.	-1.4	+0.1	0.44	1840	133	Semu Tār	H.S.	-6.8	0.0	0.39	1865
102	Kaliānpur	T.S.	-1.8	+0.2	0.64	1850	134	Nagarkhana	H.S.	-6.9	0.0	1.29	1905

* Additional correction for Helmert's Spheroid.

TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station	Corrections		Probable Error \pm	Date of Observation	Serial Number	Station	Corrections		Probable Error \pm	Date of Observation
		Everest's Spheroid	Helmert's Spheroid*					Everest's Spheroid	Helmert's Spheroid*		
130	Daulatpur T.S.	- 6.3	- 0.1	0.15	1868	160	Ubyetaung H.S.	- 6.8	+ 0.1	0.31	1894
135	Tepri T.S.	- 4.2	- 0.1	0.30	1869	167	Sinpitaung H.S.	- 6.8	+ 0.1	0.36	1901
136	Aloakāndi T.S.	- 2.2	0.0	0.51	1873	168	Loi Hpa Lang H.S.	- 6.8	+ 0.1	0.16	1903
137	Halkāchar T.S.	- 1.1	0.0	0.21	1873	169	Loi Hpatwān H.S.	- 6.8	+ 0.1	0.20	1907
138	Alangjāni T.S.	+ 0.6	0.0	0.79	1874	170	Loi Kiipma H.S.	- 6.8	+ 0.1	0.17	1908
139	Ataro Bānki T.S.	+ 1.3	+ 0.1	0.34	1856	171	Loi Hsām Hsum H.S.	- 6.8	+ 0.1	0.21	1911
142	Jalpaiguri s.	+ 3.0	+ 0.1	0.33	1904	172	Kumtum Bum H.S.	- 6.8	+ 0.1	0.22	1910
143	Senn Tūn H.S.	- 6.8	0.0	0.39	1865	173	Kumon Bum H.S.	- 6.8	+ 0.1	0.17	1911
140	Dawa H.S.	- 5.3	0.0	0.26	1864	1	Kallianpur H.S.	0.0	0.0	0.31	1836
141	Rangsano H.S.	- 2.7	0.0	0.32	1861	174	Ahmadpur H.S.	- 0.1	0.0	0.37	1838
142	Baikusni H.S.	- 0.4	0.0	0.29	1858	175	Bhimbhat H.S.	- 0.2	- 0.1	0.39	1838
143	Alangjāni T.S.	+ 0.6	0.0	0.79	1874	176	Nilgarh H.S.	- 0.3	- 0.1	0.23	1839
144	Nagarkhana H.S.	- 6.9	0.0	1.29	1905	177	Badgaon H.S.	- 0.5	- 0.1	0.22	1839
143	Fi Tān H.S.	- 6.9	+ 0.1	0.52	1865	178	Sakri H.S.	- 0.6	- 0.2	0.30	1838
144	Dattaung H.S.	- 7.1	+ 0.3	0.35	1866	179	Somtana H.S.	- 0.7	- 0.2	0.27	1838
144	Dattaung H.S.	- 7.1	+ 0.3	0.35	1866	180	Dāmargida Obs. H.S.	- 0.9	- 0.3	0.40	1838
145	Retkanauk H.S.	- 8.0	+ 0.3	0.41	1916	181	Bolarum P. W. D. Office s.	- 1.1	- 0.3	0.31	1904
146	Kyaunggyi s.	- 12.1	+ 0.3	0.34	1904	181	Bolarum P. W. D. Office s.	- 1.1	- 0.3	0.31	1904
147	Taungun H.S.	- 10.0	+ 0.5	0.76	1884	182	Pirmulo H.S.	- 1.2	- 0.3	0.23	1869
148	Southern Moscos H.S.	- 10.0	+ 0.5	0.62	1877	183	Vānākonda H.S.	- 1.2	- 0.3	0.20	1869
149	Mergui Base-line E. End T.S.	- 10.0	+ 0.5	0.18	1882	184	Singāwāram H.S.	- 1.3	- 0.3	0.19	1871
150	Mergui Base-line W. End T.S.	- 10.0	+ 0.5	0.32	1882	185	Kālingkonda H.S.	- 1.3	- 0.3	0.24	1872
151	Natkalintaung H.S.	- 10.0	+ 0.5	0.20	1881	186	Sānjib H.S.	- 1.4	- 0.3	0.18	1860
152	Minthangtaung H.S.	- 10.0	+ 0.5	0.24	1881	187	Vizagapatam Base-line N. End S.	- 1.4	- 0.3	0.24	1863
146	Kyaunggyi s.	- 12.1	+ 0.3	0.34	1904	101	Calcutta Base-line S. End T.S.	- 5.9	- 0.1	0.37	1845
153	Myayabelungkyo H.S.	- 11.1	+ 0.3	0.45	1889	188	Patna T.S.	- 4.6	- 0.2	0.19	1852
154	Toungoo s.	- 10.3	+ 0.2	0.23	1890	189	Chandipur T.S.	- 4.3	- 0.2	0.30	1854
155	Letpataung H.S.	- 9.6	+ 0.2	0.27	1891	190	Cuttaok H.S.	- 3.4	- 0.2	0.20	1854
156	Taungpila H.S.	- 8.8	+ 0.2	0.21	1891	191	Khundābolo H.S.	- 3.0	- 0.2	0.35	1857
157	Mingun H.S.	- 8.0	+ 0.2	0.19	1892	192	Rawal H.S.	- 1.8	- 0.3	0.30	1860
158	Shienmaga H.S.	- 7.8	+ 0.2	0.36	1892	187	Vizagapatam Base-line N. End S.	- 1.4	- 0.3	0.24	1863
159	Male H.S.	- 7.3	+ 0.1	0.19	1892	192	Rawal H.S.	- 1.8	- 0.3	0.30	1860
160	Ubyetaung H.S.	- 6.8	+ 0.1	0.31	1894	198	Deodonger H.S.	- 2.0	- 0.3	0.11	1914
161	Thonbinzin H.S.	- 6.5	+ 0.1	0.18	1894	194	Sindur H.S.	- 2.3	- 0.2	0.16	1913
162	Seikpa H.S.	- 6.3	+ 0.1	0.24	1895	195	Andhari H.S.	- 2.6	- 0.2	0.20	1913
163	Tamunja H.S.	- 5.7	+ 0.1	0.30	1896	196	Bhursu H.S.	- 2.9	- 0.1	0.17	1912
164	Thyliching H.S.	- 5.5	+ 0.1	0.19	1898	94	Hurilāong H.S.	- 3.1	- 0.1	0.44	1849
165	Loijing H.S.	- 5.0	+ 0.1	0.17	1899	197	Karaundi H.S.	- 1.2	0.0	0.33	1865
141	Rangranobe H.S.	- 2.7	0.0	0.32	1861	198	Sarandi Pat H.S.	- 1.2	- 0.1	0.33	1865
144	Dattaung H.S.	- 7.1	+ 0.3	0.35	1866	199	Bhimsain H.S.	- 1.2	- 0.1	0.18	1866
166	Yepuetaung H.S.	- 6.9	+ 0.3	0.60	1916	200	Diwai H.S.	- 1.2	- 0.2	0.16	1867
163	Tamunja H.S.	- 5.7	+ 0.1	0.30	1896	201	Burgpalli H.S.	- 1.1	- 0.2	0.18	1867
						181	Bolarum P.W.D. Office s.	- 1.1	- 0.3	0.31	1904

Additional correction for Helmert's Spheroid.

TABLE XCIV.—(Continued). (See Index pp. 170-172)

Serial Number	Station	Corrections		Probable Error \pm	Date of Observation	Serial Number	Station	Corrections		Probable Error \pm	Date of Observation
		Everest's Spheroid	Helmert's Spheroid*					Everest's Spheroid	Helmert's Spheroid*		
91	Karāra	H.S.	- 1.8	0.0	0.69	1842	230 Mangalore S.	- 4.0	-0.6	0.58	1873
202	Pathāidi	T.S.	- 1.7	-0.1	0.39	1871	231 Nughallibēta H.S.	- 4.7	-0.6	0.23	1871
203	Ramai	H.S.	- 1.6	-0.1	0.50	1872	229 Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
204	Karā	H.S.	- 1.5	-0.2	0.20	1873					
186	Sānjib	H.S.	- 1.4	-0.3	0.18	1860	229 Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
97	Tilabani	H.S.	- 3.9	-0.1	0.76	1845	232 Anandalamalai H.S.	- 3.9	-0.5	0.10	1866
205	Kalsibhanga	T.S.	- 4.4	-0.2	0.32	1849	233 Injambākam H.S.	- 3.3	-0.5	0.17	1880
188	Patna	H.S.	- 4.6	-0.2	0.19	1852	234 St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
181	Bolarum P.W.D. Office	S.	- 1.1	-0.3	0.31	1904	234 St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
206	Achola	H.S.	- 1.4	-0.3	0.25	1840	235 Kistama H.S.	- 2.7	-0.5	0.37	1864
207	Nitali	H.S.	- 1.6	-0.3	0.27	1840					
208	Kanheri	H.S.	- 1.7	-0.3	0.32	1837	236 Dānapa H.S.	- 2.4	-0.4	0.21	1863
209	Alsunda	H.S.	- 1.7	-0.3	0.32	1863	237 Dhūlipalla S.	- 2.3	-0.4	0.17	1868
210	Khānpisura	H.S.	- 1.8	-0.3	0.42	1846	238 Parampūdi H.S.	- 1.9	-0.4	0.13	1861
211	Dhauleshvar	H.S.	- 1.9	-0.3	0.18	1838	187 Vizagapatam Base-line N. End S.	- 1.4	-0.3	0.24	1863
212	Māndvi	H.S.	- 2.1	-0.3	0.32	1841					
213	Karanja	H.S.	- 2.1	-0.3	...	1839	214 Colaba Obs. S.	- 2.2	-0.3	...	1839
214	Colaba Obs.	S.	- 2.2	-0.3	...	1839	239 Pāchvad H.S.	- 2.7	-0.4	0.16	1865
215	Deesa T.O.	S.	- 4.4	0.0	0.26	1904					
216	Sonāda	T.S.	- 3.8	-0.1	0.39	1851	240 Karabgati H.S.	- 3.0	-0.4	0.22	1865
217	Patangdi	H.S.	- 3.3	-0.2	0.30	1861	241 Koramūr H.S.	- 3.6	-0.5	0.21	1873
218	Sāler	H.S.	- 2.7	-0.2	0.60	1845	230 Mangalore S.	- 4.0	-0.6	0.58	1873
219	Pārnera	H.S.	- 2.6	-0.3	0.47	1843	214 Colaba Obs. S.	- 2.2	-0.3	...	1839
220	Kalsubai	H.S.	- 2.4	-0.3	0.64	1842	242 Mirya H.S.	- 2.5	-0.3	0.89	1844
214	Colaba Obs.	S.	- 2.2	-0.3	...	1839	243 Chauloka H.S.	- 2.6	-0.3	0.83	1843
25	Karachi Obs.	S.	- 2.5	+0.1	0.19	1855	244 Kumbhari H.S.	- 2.7	-0.4	0.61	1844
221	Hāthria	H.S.	- 3.1	0.0	0.54	1856	240 Karabgati H.S.	- 3.0	-0.4	0.22	1865
222	Dungarpur	H.S.	- 3.4	-0.1	0.31	1852	229 Bangalore Base-line S.W. End S.	- 5.4	-0.6	0.15	1870
223	Ingridi	T.S.	- 3.6	-0.1	0.32	1852					
216	Sonāda	T.S.	- 3.8	-0.1	0.39	1851	245 Bangalore Base-line N.E. End S.	- 5.5	-0.6	0.23	1870
222	Dungarpur	H.S.	- 3.4	-0.1	0.31	1852	246 Kanjamalai H.S.	- 6.3	-0.7	0.21	1869
224	Kunkāvāv	T.S.	- 3.4	-0.1	0.27	1853	247 Pachupālaiyam S.	- 6.7	-0.7	0.26	1870
7	Aramlia	S.	- 0.6	0.0	0.32	1850	248 Kutipārai S.	- 7.2	-0.7	0.33	1873
225	Indrāwan	T.S.	- 1.0	-0.1	0.29	1847	249 Rādhāpuram S.	- 7.8	-0.8	0.14	1869
226	Valvādi	H.S.	- 1.4	-0.2	0.56	1846	250 Kudankulam Obs. S.	- 7.9	-0.8	0.21	1869
210	Khānpisura	H.S.	- 1.8	-0.3	0.42	1846					
181	Bolarum P.W.D. office	S.	- 1.1	-0.3	0.31	1904	234 St. Thomas's Mount Trestle S.	- 3.1	-0.5	0.27	1880
227	Kodangal	S.	- 2.2	-0.4	0.18	1872	251 Kallapat Trestle S.	- 3.9	-0.5	0.17	1879
228	Darur	H.S.	- 2.7	-0.4	0.91	1871	252 Nayinipiriyān „ S.	- 4.5	-0.6	0.24	1870
229	Bangalore Base-line S.W. End S.		- 5.4	-0.6	0.15	1870	253 Pātharankota S.	- 5.1	-0.6	0.14	1877
							254 Manēgandi S.	- 6.0	-0.7	0.25	1876
							255 Rāmnad S.	- 6.4	-0.7	0.39	1875
							250 Kudankulam Obs. S.	- 7.9	-0.8	0.21	1869

* Additional correction for Helmert's Spheroid.

INDEX TO DEFLECTION STATIONS.

Name of Station	Reference Number		No. of Series	Name of Station		Reference Number		No. of Series	Name of Station	Reference Number		No. of Series
	Table XCV	Table XCIV				Table XCV	Table XCIV			Table XCV	Table XCIV	
Achola H.S.	231	206	7	Birond H.S.		179		20	Dehra Dun Base, E., S.	146		6
Agra-group E. Point	201		6	Bisaul T.S.		318	115	19	Dehra Dun Haig Obs. S.	170		6
Agra-group N. Point	199		6	Bithnok H.S.		71		62	Dehra Dun Obs. (Old) S.	169	83	6
Agra-group S. Point	203		6	Black s.		288		9	" No. VI } between	165		20
Agra-group W. Point	187		6	Bolarum P.W.D. Long. S.		250	181	60	" No. V } Dehra and	166		20
Agra Longitude S.	200		6	Bolkonda H.S.		255		43	" No. IV } Rājpur	167		20
Agra Parade Point	202		6	Bombay Colaba Long. S.		111		7	" No. III }	168		20
Ahmadpur H.S.	214	174	8	Bombay Colaba Obs. S.		112	214	7	Deodonger H.S.	382	193	85
Akampalle h.s.	239		9	Bommasandra s.		267		9	Deo Dongri H.S.	106		18
Akbar S.	63	82	37	Bosān T.S.		155		6	Dera Dīn Panāh S.	29	44	32
Aknāpur T.S.	402	100	5	Bowra T.S.		140	66	33	Devanūr s.	238		9
Akyab Longitude S.	419		44	Budhon H.S.		209	68	5	Devaragat h.s.	248		9
Alamkhān T.S.	35	22	25	Bulāwāla h.s.		150		6	Dewarsān T.S.	300		3
Alamvādi H.S.	103		10	Bulbul H.S.		367		5	Dhaigaon S.	127		18
Alangjāni T.S.	395	138	34	Burgpāli H.S.		253	201	53	Dhamanva T.S.	94		26
Algi H.S.	206		2	Calcutta Base, S., T.S.		403	101	5	Dhānura s.	221		8
Alibagh Observatory S.	115		9	Calcutta Longitude s.		404		5	Dhaulesvar H.S.	122	211	7
Aloakāndi T.S.	397	136	56	Chamardi H.S.		53		30	Dhūlipalla S.	337	237	46
Alsunda H.S.	129	209	7	Chamu H.S.		73		62	Didāwa H.S.	45	15	25
Amritsar Longitude S.	65		23	Chandaos T.S.		156		6	Diwai H.S.	251	200	53
Amsot H.S.	145		6	Chandipur T.S.		380	189	24	Dōddagunta s.	270		9
Amūa H.S.	326	90	5	Chanduria T.S.		392		16	Dotra s.	222		8
Anandamalalai H.S.	274	232	54	Chānga H.S.		38	19	25	Dūbanli T.S.	361		14
Anandbās T.S.	401	128	16	Chaniāna H.S.		80		26	Dumb h.s.	21		*
Andar H.S.	12	49	32	Charaldānga T.S.		393		16	Dungarpur H.S.	49	222	28
Andhari H.S.	370	195	85	Chaukola H.S.		131	243	11	Etorā T.S.	299		3
Andhīari H.S.	207		2	Chendwār (old) H.S.		371	95	5	Fi Tān H.S.	418	143	52
Ankora H.S.	252		53	Chikalgurki s.		262		9	Fyzabad Longitude S.	317		19
Aramlia S.	89	7	25	Chittagong Longitude S.		412		44	Gandpahar H.S.	11	51	32
Arasakulam S.	281		9	Chūni T.S.		365	120	20	Ganga Choti H.S.	54	72	77
Asu H.S.	40	84	64	Chūlli T.S.		36	23	25	Gangapur T.S.	408	131	48
Ataro Bānki T.S.	394	139	34	Colaba Observatory S.		112	214	7	Garinda S.	90	68	23
Badgaon H.S.	220	177	8	Cuttack H.S.		372	190	24	Gattinārāyantippa h.s.	247		9
Bahak H.S.	158		73	Dadaura T.S.		304		20	Gaus T.S.	321		20
Bajamara H.S.	144		73	Daiādhari H.S.		191		6	Ghorārāo H.S.	100		10
Bandūr s.	263		9	Dājil S.		30	43	32	Godhna T.S.	152		6
Bangalore Base, N.E., S.	268	245	9	Dalea H.S.		332		58	Gogipatri H.S.	454	74	22
Bangalore Base, S.W., S.	269	229	9	Dāmargida Obs. S.		237	180	8	Gora H.S.	325	93	5
Banog H.S.	160	32	6	Dānapa H.S.		271	236	46	Gudali H.S.	347		46
Bāngopāl T.S.	177		2	Dangarvadi H.S.		51		28	Gundak H.S.	6	54	76
Bānskho H.S.	182	62	33	Dāowāla T.S.		24	86	64	Gurāria H.S.	184	5	25
Bisadela T.S.	315	105	20	Dargawa H.S.		211		2	Gūrmi T.S.	204	107	2
Bellary Longitude s.	258		9	Dariāpur T.S.		378		24	Gūru Sikkar H.S.	76	11	25
Bhanar T.S.	25	41	32	Darur H.S.		245	228	9	Gurwāni H.S.	320	92	5
Bhaorāsa H.S.	208		6	Darutippa S.		272		46	Halda s.	232		8
Bhimbhat H.S.	216	175	8	Dasti S.		9		*	Halkāchar T.S.	396	137	56
Bhimsain H.S.	228	199	53	Datāiri T.S.		154	30	6	Harnāsa T.S.	108		18
Bhursu H.S.	369	196	5	Dattaung H.S.		421	144	52	Harpālsid T.S.	173		2
Bichwi H.S.	364	126	27	Daulatpur T.S.		406	130	48	Hātbenā H.S.	339		58
Bihar H.S.	362		14	Dawa H.S.		409	140	44	Hāthria H.S.	48	221	35
Birona S.	78	12	25	Deesa Telegraph Office s.		79	215	26	Hatni h.s.	149		6

Obs. = Observatory, Long. = Longitude. * Old Baluchistan Series.

INDEX TO DEFLECTION STATIONS—(Continued).

Name of Station		Reference Number		No. of Series	Name of Station		Reference Number		No. of Series	Name of Station		Reference Number		No. of Series
		Table XCV	Table XCIV				Table XCV	Table XCIV				Table XCV	Table XCIV	
Hirdepur	T.S.	324	116	15	Kheri	T.S.	141	65	33	Majhār	H.S.	205		2
Hönnavalli	H.S.	137		49	Khimūāna	T.S.	67		23	Mal	H.S.	383		24
Hönūr	H.S.	264		9	Khirsar	H.S.	62		62	Malar	H.S.	41	83	64
Hurilāong	H.S.	356	94	5	Khojak	H.S.	4		†	Male	H.S.	423	159	66
Imlia	T.S.	306		12	Khori	T.S.	37	21	25	Malūncha	H.S.	375	98	5
Indrāwan	T.S.	107	225	18	Khujnaur	s.	147		20	Mandāla	s.	235		8
Injambākam	H.S.	350	233	54	Khundābolo	H.S.	381	191	24	Mandresa	T.S.	59	80	45
Inrogdi	T.S.	52	223	29	Kidarkanta	H.S.	157		22	Māndvi	H.S.	118	212	7
Isanpur	H.S.	139		33	Kisanen Chapper	H.S.	3	57	74	Manēgandi	s.	287	254	63
Jabalpur Longitude	s.	227		53	Kistama	H.S.	273	235	46	Mangalore Longitude	S.	134	230	54
Jalālpur	T.S.	352	124	21	Kodangal	S.	240	227	9	Manichauk	T.S.	313		20
Jalpaiguri Longitude	s.	389	122	34	Koh-i-Malik Siah	H.S.	1	59	74	Marūr	T.S.	319	114	19
Jambo	H.S.	72	77	62	Koramūr	H.S.	133	241	49	Martaban	h.s.	440		52
Jāoli	H.S.	56	47	22	Kudankulam Obs.	S.	284	250	9	Mashelak	H.S.	5	53	76
Jarūra	T.S.	297		3	Kumbhāri	H.S.	132	244	11	Māsi	T.S.	305	104	20
Jetgarh	H.S.	85		23	Kumon Bum	H.S.	430	173	80	Māta-ka-hūra	H.S.	185	4	25
Jhambhera	T.S.	64	81	45	Kumtum Bum	H.S.	431	172	80	Māvinhūnda	H.S.	125		49
Jharipani (IX)	h.s.	162	36	20	Kundgol	H.S.	136		40	Mednipur	T.S.	354	123	21
Jharkil	T.S.	27	45	32	Kunkāvāv	T.S.	50	224	28	Medwāni	H.S.	138	48	22
Kainath	H.S.	93		26	Kurseong	h.s.	387		20	Mergui Base E.,	T.S.	446	149	52
Kaliāna	S.	153	31	6	Kutipārai	S.	289	248	9	Mergui Base W.,	T.S.	447	150	52
Kaliānpur	H.S.	195	1	6	Kyaunggyi	S.	428	146	52	Miāni	T.S.	23	42	32
Kaliānpur	T.S.	180	102	20	Lachkuwa	h.s.	171		6	Mingun	H.S.	425	157	66
Kālingkonda	H.S.	341	185	43	Lādi	H.S.	215		8	Minthangtaung	H.S.	448	152	52
Kallapat	Trestle S.	292	251	63	Lādinslar	T.S.	33	79	45	Mira Donger	H.S.	117		11
Kalsibhānga	T.S.	377	205	17	Lakarwas	H.S.	83		25	Mirya	H.S.	120	242	11
Kalsubai	H.S.	116	220	10	Lakhinagar	T.S.	407	132	48	Mohammadabad	T.S.	210	110	4
Kāmkhera	H.S.	213		25	Lambatach	H.S.	143		22	Mooltan Longitude	S.	32		32
Kānākhera	T.S.	301		3	Letpataung	H.S.	437	155	66	Morali	H.S.	95		26
Kanheri	H.S.	128	208	7	Linganapalle	h.s.	241		9	Moulmein Longitude	S.	441		44
Kanjamaiai	H.S.	286	246	9	Lingmāra	H.S.	330		53	Mugrāla	H.S.	61	78	62
Kankesvar	H.S.	114		11	Lohārgara	T.S.	391		16	Murree	h.s.	55	71	22
Kānkra	H.S.	183	61	33	Loi Hpa Lang	H.S.	433	168	72	Murree Observatory	s.	451		22
Kānnagar	H.S.	81	10	25	Loi Hpatan	H.S.	434	169	72	Mussoores Dome	H.S.	161	37	6
Karabgati	H.S.	126	240	49	Loi Hsam Hsum	H.S.	436	171	72	Myayabeingkyo	H.S.	439	153	66
Karachi Base S,	S.	14	39	32	Loijing	H.S.	413	165	68	Nagūrkhūna	H.S.	411	134	52
Karachi Longitude	S.	16		22	Loi Kiipma	H.S.	435	170	72	Nagarkoil Longitude	S.	283		9
Karachi Observatory	S.	17	25	32	Lora	H.S.	327		5	Nag Tiba	H.S.	159	38	73
Karanja	H.S.	113	213	7	Losalli	S.	197	2	25	Nāharman	H.S.	225		5
Karāra	H.S.	312	91	5	Lūnki	H.S.	42	17	25	Namthabad	s.	261		9
Karaundi	H.S.	226	197	53	Mach	h.s.	8		†	Natkalintaung	H.S.	445	151	52
Kardo	H.S.	92		26	Madhpur	T.S.	376	99	5	Naunangarhi	T.S.	351	119	20
Karia	H.S.	340	204	58	Madhpur	T.S.	400	129	16	Navalūr	H.S.	135		49
Kārothol	H.S.	13	24	25	Madras Observatory	S.	348		54	Nayinipriyān Trestle	S.	294	252	63
Kātpūlaiyam	s.	277		9	Mahabaleswar	H.S.	119		11	Niālamari	H.S.	257		46
Kaulia	H.S.	357		*	Mahadeo Pokra	H.S.	358		*	Nilgarh	H.S.	217	176	8
Kem	H.S.	130		7	Mahar	H.S.	363		14	Nimbāgal	s.	265		9
Kesri	H.S.	189	27	6	Mahesari	T.S.	174		6	Nimkār	T.S.	298	111	3
Khāmor	H.S.	87		23	Mahwari	H.S.	368		5	Nitali	H.S.	230	207	7
Khankharia	S.	47	13	25	Mairūb-ka-Shahar	T.S.	39	20	25	Noh	T.S.	186	29	6
Khānpisura	H.S.	121	210	7	Majala	H.S.	124		49	Nojli	T.S.	151		6

Obs. = Observatory. * Special Triangulation. † Old Baluchistān Triangulation. ‡ Thayetmo and Cape Negrais Series.

INDEX TO DEFLECTION STATIONS—(Continued).

Name of Station	Reference Number		No. of Series	Name of Station		Reference Number		No. of Series	Name of Station	Reference Number		No. of Series
	Table XCV	Table XCVI				Table XCV	Table XCVI			Table XCV	Table XCVI	
Nuon	T.S.	353	21	Rām Thal	S.	70		23	Southern Moscos	H.S.	443	148
Nughallibetta	H.S.	260	281	Rāmnapur (old)	T.S.	296	103	20	Spurpoint (VIII)	h.s.	163	35
Ongole	H.S.	346	46	Rangir (old)	S.	212	89	5	St. Thomas's Mount	Tres. S.	349	284
Orejhar	S.	316	106	Rāngrai	s.	219		8	Sultan-ka-Got	T.S.	22	
Oria	h.s.	77	25	Rangsanobo	H.S.	399	141	44	Sūrāntal	H.S.	193	
Pabbosa	H.S.	311	112	Rānigarh	H.S.	172		20	Takalkhera	s.	218	
Pachapaleiyam	s.	276	247	Ranjitgarh	T.S.	57		23	Talegaon	s.	236	
Pachvad	H.S.	123	239	Rāwal	H.S.	343	192	24	Tamunja	H.S.	415	163
Pahargarh	H.S.	190	26	Retkamauk	H.S.	427	145	52	Tanakarakulam	S.	280	
Pahladpur	T.S.	360		Rewat	H.S.	84		23	Tarhlān	S.	104	
Pāldi	H.S.	97		Robat	S.	449		†	Tāsing	H.S.	181	
Pandalagudi	s.	290	9	Rojhra	H.S.	43	18	25	Taungpila	H.S.	426	
Paphra	T.S.	31	87	Rustamgarhi	h.s.	452	75	22	Taungzun	H.S.	442	147
Parampādi	H.S.	338	238	Sakri	H.S.	223	178	8	Telu	H.S.	60	
Parasāth	H.S.	373	96	Sāler	H.S.	105	218	10	Teona	H.S.	355	
Parewa	T.S.	308		Salighar	H.S.	20		55	Tepri	T.S.	405	135
Parison	T.S.	310		Salimpur	T.S.	198		2	Thikri	H.S.	109	
Pārnera	H.S.	98	219	Salot	H.S.	192	3	25	Thob	H.S.	74	
Patāngdi	H.S.	99	217	Samdari	H.S.	75		62	Thonbinzin	H.S.	417	161
Pathādi	T.S.	333	202	Samenda	T.S.	323	117	15	Thyoliching	H.S.	414	164
Patharankota	S.	295	253	Sānd	H.S.	88	8	25	Tiki	H.S.	82	9
Pathārdi	T.S.	314		Sandawat	H.S.	444		52	Tilabani	H.S.	374	97
Patna	T.S.	379	188	Sangatpur	T.S.	66	70	23	Tinsā	H.S.	196	
Pāvagad	H.S.	101	10	Sānjib	H.S.	342	186	43	Tiruvēndipuram	s.	293	
Pāvagada	H.S.	266		Sankrāo	T.S.	173	108	2	Tonglu	h.s.	385	
Pavia	H.S.	302	3	Sarandi Pat	H.S.	328	198	53	Tōnsalgutta	s.	243	
Piddapād	s.	244	9	Sarey Khan Latitude	S.	329		53	Toungoo	S.	438	154
Peshawar Longitude	S.	18		Sarkāra	T.S.	175		2	Tounsa	T.S.	28	56
Phallut	h.s.	354	20	Sarla	S.	46	14	25	Tuagat	h.s.	240	
Pialmudi	s.	242	9	Saugor	H.S.	224		5	Tuzgi	H.S.	2	58
Piaro	H.S.	10	50	Sawaipur	T.S.	68		23	Ubyetaung	H.S.	422	160
Pirmulo	H.S.	249	182	Seikpa	H.S.	416	162	66	Umarkhel	H.S.	19	46
Poshkar	H.S.	453	73	Semu Tān	H.S.	410	133	52	Usira	H.S.	183	28
Pota	T.S.	359	125	Senchal	h.s.	386		20	Uāmāu	T.S.	307	
Potenda	S.	303	3	Shālpur	T.S.	58		23	Valvādi	H.S.	110	226
Prome Longitude	S.	429	52	Sheinmaga	H.S.	424	158	66	Vānākonda	H.S.	256	183
Punna Observatory	S.	285	9	Shorpur	H.S.	148		6	Vijayāpati	S.	282	
Quetta Telegraph Office	S.	7	60	Shūlakarai	s.	278		9	Vijnot	T.S.	26	85
Rādhāpuram	S.	279	249	Sidhpur	S.	102		10	Virāria	H.S.	44	16
Raitusni	H.S.	398	142	Siliguri	s.	388		20	Vizagapatām Base, N.	S.	344	187
Rājāhari	T.S.	322	118	Sindur	H.S.	335	194	85	Voi	s.	233	
Rājgarh	H.S.	66	67	Singāwāram	H.S.	336	184	43	Waltair Longitude	S.	345	
Rājpur	h.s.	164	34	Sinpitāung	H.S.	432	167	72	Yēponetaung	H.S.	420	166
Rājuli	H.S.	229	53	Sirtānda	T.S.	366	127	13	Yēragunta	h.s.	259	
Rākhi	T.S.	142	64	Sironj Base, N.E.,	S.	194		6	Yēttimalai	S.	275	
Ranai	H.S.	334	203	Sirsa	S.	69	69	23	Yūsuf	S.	34	40
Rānāgh Observatory	s.	15	32	Sirsa	T.S.	176	109	2	Zawa	H.S.	450	52
Rānganj	T.S.	390	121	Sitāpār	H.S.	331		53				
Rāngir	H.S.	254		Somtana	H.S.	234	179	8				
Rānōad	S.	291	53	Sonāda	T.S.	96	216	29				
Rānpūra	H.S.	91	255	Sora	T.S.	309	113	12				

Obs. = Observatory.

Tres. S. = Trestle S.

* Quetta Secondary Series.

† Robat Triangulation.

Deflections of the Plumb-line

in terms of

any Spheroid.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflec- tion†	
1	30 C	Koh-i-Malik Siah H.S.	5393	" " "	" " "	" " "	" " "	" " "	" " "	1
	O	No. 449 <i>ut infra</i>		G 29 51 31'95	G 60 52 19'71	A 321 52 15'7 G 321 52 0'4	Kachakoh E 0 42	+ 26'7		
2	L	Tuzgi H.S.	3131	G 28 53 14'38	G 62 14 58'97	A 216 16 46'3 G 216 16 37'3	Shuri E 0 40	+ 16'3		2
3	34 C	Kisanen Chapper H.S.	4362	G 29 3 54'41	G 64 22 24'93	A 166 46 22'6 G 166 46 15'7	Malik Praji E 1 7	+ 12'4		8
4	J	Khojak H.S.	7851	A 30 51 20'21 G 30 51 24'85	G 66 34 41'08				- 4'6	4
5	J	Mashelak H.S.	7941	G 30 13 30'77	G 66 44 46'79	A 241 55 31'4 G 241 55 27'8	Takatu E 1 20	+ 6'2		5
		No. 450 <i>ut infra</i>								
6	M	Gundak H.S.	8163	G 31 9 49'49	G 67 23 28'55	A 276 29 27'1 G 276 29 25'2	Basha E 0 56	+ 3'1		6
7	N	Quetta Tel Office S.	5500	A 30 11 55'91 G 30 11 57'37	A 67 0 29'27 G 67 0 31'69	A 166 31 12'1 G 166 31 11'8	Takatu E 3 5	+ 0'5	- 1'5	7
8	O	Mach h.s.	3522	A 29 52 20'46 G 29 52 31'51	G 67 18 39'42				- 11'1	8
9	O	Dasti S.	316	A 29 0 27'61 G 29 0 29'93	G 67 53 51'59				- 2'3	9
10	35 J	Piaro H.S.	1438	G 26 3 14'21	G 66 34 7'96	A 159 22 15'3 G 159 22 12'8	Kuliri E 0 16	+ 5'1		10
11	M	Gandpahar H.S.	723	G 27 25 1'26	G 67 30 43'99	A 192 10 55'2 G 192 10 46'4	Kharko D 0 9	+ 17'0		11
12	N	Andar H.S.	4047	G 26 1 22'07	G 67 12 10'58	A 181 7 6'5 G 181 7 1'7	Sulemani D 0 24	+ 9'8		12
13	P	Kārothol H.S.	260	A 24 53 44'78 G 24 53 46'69	G 67 53 32'47	A 121 36 58'0 G 121 36 55'3	Kara E 0 38	+ 5'8	- 1'9	18
14	P	Karachi Base-line S. End S.	46	G 24 52 59'63	G 67 9 24'77	A 205 23 30'5 G 205 23 29'2	Karachi Base-line N. End E 0 10	+ 2'8		14
15	P	Rāmbāgh Obsy. s.	...	A 24 51 20'58 G 24 51 21'44	G 67 0 55'21				- 0'9	15
16	P	Karachi Long. S.		G 24 51 2'44	A 67 0 52'88 G 67 0 53'22				- 0'3	16
17	P	Karachi Obsy. S.	35	A 24 49 50'14 G 24 49 50'25	G 67 1 35'13	A 221 39 9'5 G 221 39 8'4	Mutrani E 0 25	+ 2'4	- 0'1	17
18	38 N	Peshawar Long. S.	...	G 34 0 17'51	A 71 33 14'63 G 71 33 0'27			+ 11'9		18
19	P	Umārkhel H.S.	3036	G 32 25 31'07	G 71 15 20'79	A 73 9 24'0 G 73 9 15'3	Sistarg E 0 18	+ 13'7		19
20	39 A	Saighar H.S.	8284	G 31 29 53'82	G 68 25 5'49	A 1 52 50'1 G 1 52 46'2	Tanispā E 0 38	+ 6'4		20
21	D	Dumb h.s.	183	A 28 15 18'30 G 28 15 21'09	G 68 14 9'96				- 2'8	21
22	D	Sultan-ka-Got T.S.	213	A 28 4 8'05 G 28 4 9'41	G 68 36 32'23				- 1'4	22
23	H	Miāni T.S.	300	G 28 34 15'20	G 69 50 46'91	A 188 2 16'6 G 188 2 4'5	Routi D 0 5	+ 22'2		23
24	H	Dāowāla T.S.	282	G 28 20 12'87	G 69 50 30'68	A 28 49 27'9 G 28 49 21'3	Ghundi D 0 3	+ 12'2		24
25	H	Bhanar T.S.	256	G 28 8 55'00	G 69 17 11'38	A 197 50 8'8 G 197 50 0'7	Khai D 0 5	+ 15'1		25
26	H	Vijnot T.S.	276	G 28 2 3'30	G 69 50 32'77	A 159 35 15'6 G 159 35 10'0	Dewari D 0 5	+ 10'5		26

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
1	"	"	+8.80	"	"	-0.17	"	"	-0.58	"	"	+1.77	"	"	+10.11	+16.4	"	1	
2			+8.23			-0.30			-0.55			+1.83			+9.56	+6.5		2	
3			+7.08			-0.25			-0.48			+1.83			+8.57	+3.6		3	
4	+1.64			-5.19			+0.98			+0.17			-1.78				-2.8	4	
5			+5.71			-0.09			-0.38			+1.78			+7.38	-1.5		5	
6			+5.30			0.00			-0.34			+1.74			+7.04	-4.2		6	
7	+1.54	+6.36	+5.58	-4.74	-0.88	-0.09	+0.98	-0.09	-0.37	+0.17	+0.10	+1.78	-1.54	+5.26	+7.27	-7.5	0.0	7	
8	+1.48			-4.52			+0.98			+0.16			-1.44				-9.7	8	
9	+1.32			-3.89			+0.98			+0.16			-1.12				-1.2	9	
10			+6.30			-0.61			-0.45			+2.05			+7.87	-3.0		10	
11			+5.61			-0.39			-0.39			+1.95			+7.29	+9.5		11	
12			+5.96			-0.57			-0.41			+2.05			+7.60	+2.0		12	
13	+0.42		+5.73	-0.68		-0.69	+0.99		-0.41	+0.16		+2.13	+0.40		+7.40	-1.8	-2.3	13	
14			+6.17			-0.57			-0.43			+2.13			+7.76	-5.2		14	
15	+0.44			-0.64			+0.98			+0.17			+0.45				-1.4	15	
16		+6.37			-0.90			-0.08		+0.01			+5.21			-5.5		16	
17	+0.44		+6.25	-0.63		-0.78	+0.98		-0.43	+0.17		+2.14	+0.40		+7.83	-5.6	-0.6	17	
18		+3.64			-0.49			-0.06		+0.17			+3.20			+8.7		18	
19			+3.21			+0.06			-0.20			+1.70			+5.14	+8.1		19	
20			+4.75			+0.02			-0.31			+1.71			+0.51	-0.4		20	
21	+1.17			-3.33			+0.99			+0.15			-0.89				-1.9	21	
22	+1.12			-3.19			+0.99			+0.14			-0.85				-0.5	22	
23			+4.22			-0.22			-0.29			+1.89			+6.09	+15.7		23	
24			+4.25			-0.22			-0.30			+1.91			+6.14	+5.9		24	
25			+4.56			-0.26			-0.32			+1.91			+6.38	+8.3		25	
26			+4.26			-0.24			-0.28			+1.92			+6.15	+4.1		26	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Fide* p. 2.

Deflections of the Plumb-line

Serial No	Sheet No	Observed at	EVEREST'S SPHEROID.							Serial No
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed Station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflection†	
27	39 I	Jharkil T.S.	532	° ' "	° ' "	° ' "	° ' "	° ' "	° ' "	27
				G 31 21 13.65	G 70 59 44.80	A 208 7 9.2 G 208 7 3.6	Kasain D 0 3	+ 9.2		
28	J	Tounsa T.S.	593	G 30 41 51.59	G 70 39 0.13	A 201 7 57.5 G 201 7 41.4	Lungawala D 0 11	+ 27.1		28
29	J	Dera Dīn Panāh S.	490	A 30 33 59.63 G 30 34 1.87	G 70 56 7.29	A 209 21 14.7 G 209 21 7.3	Sakwala D 0 4	+ 12.5	- 2.2	29
30	K	Dājil S.	412	G 29 33 20.87	G 70 22 52.98	A 239 26 6.7 G 239 25 53.0	Dalura D 0 6	+ 24.2		30
31	K	Paphra T.S.	316	G 29 5 49.37	G 70 49 45.82	A 273 23 2.0 G 273 22 56.8	Ohanikhan D 0 4	+ 9.3		31
32	N	Mooltan Long. S.	420	A 30 10 56.15 G 30 10 58.70	A 71 26 22.19 G 71 26 27.39			- 4.5	- 2.6	32
33	O	Lādimsir T.S.	468	A 29 21 39.83 G 29 21 41.58	G 71 59 19.71	A 195 0 23.1 G 195 0 22.1	Guddan D 0 6	+ 1.8	- 1.8	33
34	40 A	Yūsuf S.	215	G 27 51 8.74	G 68 26 14.75	A 195 51 20.0 G 195 51 16.5	Salar D 0 5	+ 6.6		34
35	1)	Alamkhān T.S.	67	A 24 49 30.50 G 24 49 31.23	G 68 43 47.38	A 174 28 43.3 G 174 28 39.0	Hakimani D 0 4	+ 9.3	- 0.7	35
36	D	Ohūtlī T.S.	72	G 24 46 19.67	G 68 23 40.86	A 141 22 40.1 G 141 22 35.0	R.M.	+ 11.1		36
37	G	Khori T.S.	63	A 25 0 30.60 G 25 0 31.53	G 69 3 5.32	A 247 8 33.5 G 247 8 32.8	Jan Mahamad D 0 5	+ 1.5	- 0.9	37
38	H	Ohānga H.S.	349	A 24 58 47.25 G 24 58 47.00	G 69 51 23.30	A 238 0 7.4 G 238 0 9.1	Sandohar E 0 0	- 3.6	+ 0.3	38
39	H	Mairāb-ka-Shahar T.S.	44	G 24 50 10.79	G 69 20 25.56	A 181 11 36.7 G 181 11 34.6	Amisha D 0 5	+ 4.5		39
40	I	Asu H.S.	479	G 27 10 32.14	G 70 10 59.67	A 201 37 32.6 G 201 37 31.7	Kalu D 0 9	+ 1.8		40
41	J	Malar H.S.	328	G 26 2 25.80	G 70 3 36.19	A 161 26 22.4 G 161 26 23.4	Ramsar D 0 6	- 2.0		41
42	L	Lūnki H.S.	588	A 24 58 18.73 G 24 58 23.15	G 70 39 42.32	A 255 9 1.4 G 255 8 58.2	Karebhit D 0 3	+ 6.9	- 4.4	42
43	L	Rojhura H.S.	518	A 24 57 26.09 G 24 57 26.28	G 70 14 17.90	A 254 1 46.8 G 254 1 44.8	Dharindera D 0 3	+ 4.3	- 0.2	43
44	P	Virāria H.S.	460	A 24 56 32.64 G 24 56 36.13	G 71 2 58.81	A 106 12 49.8 G 106 12 46.3	Karebhit D 0 1	+ 7.5	- 3.5	44
45	P	Didāwa H.S.	212	A 24 51 17.32 G 24 51 19.36	G 71 18 57.69	A 72 32 16.7 G 72 32 14.0	Sohagi D 0 2	+ 5.8	- 2.0	45
46	P	Sarla S.	132	G 24 46 44.68	G 71 34 7.48	A 244 27 47.6 G 244 27 43.2	Dawal D 0 3	+ 9.5		46
47	P	Khankharia S.	362	A 24 36 58.17 G 24 36 56.19	G 71 53 8.91	A 182 0 14.8 G 182 0 15.2	Kosia D 0 6	- 0.9	+ 2.0	47
48	41 E	Hāthria H.S.	696	G 23 27 14.85	G 69 2 45.83	A 154 56 32.9 G 154 56 34.6	Sura Gandara E 0 1	- 3.9		48
49	J	Dungarpur H.S.	404	A 22 48 8.85 G 22 48 13.54	G 70 59 39.44	A 199 56 38.7 G 199 56 32.8	Ohalarwa D 0 15	+ 14.0	- 4.7	49
50	K	Kunkārav T.S.	621	A 21 39 10.31 G 21 39 11.96	G 70 56 8.89	A 161 59 40.0 G 161 59 35.5	Mumaiya D 0 10	+ 11.3	- 1.7	50
51	L	Dangarvadi H.S.	96	A 20 42 52.01 G 20 43 0.53	G 70 56 5.21				- 8.5	51
52	N	Ingrodi T.S.	152	A 22 57 2.50 G 22 57 7.58	G 71 48 34.12	A 198 26 44.5 G 198 26 39.4	Por D 0 8	+ 12.0	- 5.1	52
53	O	Chamardi H.S.	361	A 21 49 23.88 G 21 49 26.65	G 71 55 4.34				- 2.8	53
54	43 F	Ganga Ohoti H.S.	9989	G 34 4 31.37	G 73 44 52.15	A 174 38 11.1 G 174 38 24.0	Kafir Khan E 1 7	- 19.1		54

* A = Astronomical Value.
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a=1$ km			Case II: $\delta b=1$ km			Case III: Latitude $u_0=1''$			Case IV: Azimuth $w_0=1''$			$a=6378200$ metres, $1/e=298.8$						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
27	"	"	+3.92	"	"	0.00	"	"	-0.23	"	"	+1.74	"	"	+5.79	+3.1	"	27	
28			+3.65			-0.03			-0.24			+1.77			+5.56	+21.2		28	
29	+1.43		+3.52	-4.99		-0.03	+0.99		-0.24	+0.11		+1.78	-1.94		+5.46	+6.7	-0.3	29	
30			+3.86			-0.05			-0.26			+1.83			+5.81	+18.1		30	
31			+3.67			-0.14			-0.25			+1.87			+5.62	+3.5		31	
32	+1.37	+2.17		-4.72	-0.30		+0.99	-0.04		+0.10	+0.06		-1.80	+1.86		-6.4	-0.8	32	
33	+1.38		+3.02	-4.14		-0.11	+1.00		-0.20	+0.09		+1.85	-1.37		+5.04	-3.4	-0.4	33	
34			+5.05			-0.30			-0.34			+1.93			+6.83	-0.6		34	
35	+0.37		+5.25	-0.61		-0.05	+0.99		-0.37	+0.14		+2.15	+0.37		+7.03	+2.1	-1.1	35	
36			+5.45			-0.68			-0.38			+2.15			+7.18	+3.7		36	
37	+0.40		+5.03	-0.77		-0.61	+0.99		-0.35	+0.14		+2.14	+0.29		+6.86	-5.6	-1.2	37	
38	+0.37		+4.57	-0.74		-0.55	+0.99		-0.32	+0.12		+2.14	+0.26		+6.47	-10.3	0.0	38	
39			+4.89			-0.61			-0.35			+2.15			+6.73	-2.4		39	
40			+4.16			-0.31			-0.28			+1.98			+6.07	-4.5		40	
41			+4.38			-0.42			-0.30			+2.06			+6.30	-8.5		41	
42	+0.34		+4.10	-0.73		-0.50	+0.99		-0.29	+0.11		+2.14	+0.22		+6.09	+0.6	-4.6	42	
43	+0.35		+4.35	-0.72		-0.53	+0.99		-0.31	+0.12		+2.14	+0.25		+6.28	-2.2	-0.5	43	
44	+0.32		+3.88	-0.70		-0.48	+0.99		-0.27	+0.11		+2.15	+0.22		+5.92	+1.4	-3.7	44	
45	+0.29		+3.73	-0.63		-0.47	+0.99		-0.26	+0.10		+2.16	+0.24		+5.81	-0.2	-2.2	45	
46			+3.59			-0.46			-0.25			+2.16			+5.69	+3.6		46	
47	+0.21		+3.42	-0.43		-0.45	+1.00		-0.24	+0.09		+2.18	+0.31		+5.57	-6.5	+1.7	47	
48			+5.26			-0.82			-0.73			+2.27			+7.06	-11.0		48	
49	-0.29		+4.16	+1.12		-0.75	+0.99		-0.30	+0.11		+2.34	+1.01		+6.21	+8.0	-5.7	49	
50	-0.66		+4.35	+1.14		-0.89	+0.99		-0.31	+0.11		+2.46	+0.59		+6.43	+5.2	-2.3	50	
51	-0.98			+2.79			+1.00			+0.11			+1.61				-10.1	51	
52	-0.27		+4.08	+0.99		-0.63	+1.00		-0.22	+0.08		+2.33	+0.94		+6.23	+6.0	-6.0	52	
53	-0.63			+1.99			+1.00			+0.09			+1.35				-4.2	53	
54			+1.95			+0.07			-0.12			+1.62			+3.90	-23.4		54	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflection†	
	48 G	No. 451 <i>ut infra</i>		° ' "	° ' "	° ' "	° ' "	"	"	
55	G	Murree h.s.	7250	A 33 54 37.35 G 33 54 57.35	G 73 22 50.15				- 20.0	55
56	G	Jāoli H.S.	1918	G 33 16 48.85	G 73 10 26.50	A 214 27 23.4 G 214 27 22.2	Nerh E 0 58	+ 1.8		56
	J	No. 452 <i>ut infra</i>								
	J	No. 453 <i>ut infra</i>								
	K	No. 454 <i>ut infra</i>								
57	L	Ranjitgarh T.S.	900	A 32 35 6.52 G 32 35 12.11	G 74 37 12.48				- 5.6	57
58	P	Shāhpur T.S.	830	A 32 1 34.23 G 32 1 33.77	G 75 5 34.90				+ 0.5	58
59	44 C	Mandresa T.S.	512	G 29 55 9.17	G 72 59 28.42	A 298 34 7.1 G 298 34 5.8	Gajlan D 0 1	+ 2.3		59
60	D	Telu H.S.	470	A 28 56 12.41 G 28 56 11.34	G 72 14 8.80				+ 1.1	60
61	D	Mugrāla H.S.	517	G 28 30 57.06	G 72 22 17.41	A 171 53 31.2 G 171 53 32.0	Habib D 0 8	- 1.5		61
62	D	Khirsar H.S.	603	A 28 29 43.75 G 28 29 40.91	G 72 39 32.34				+ 2.8	62
63	F	Akbar S.	641	A 30 53 38.53 G 30 53 43.27	G 73 17 13.28	A 216 51 25.8 G 216 51 25.4	Firoz D 0 5	+ 0.7	- 4.7	63
64	F	Jhambhera T.S.	630	G 30 5 59.27	G 73 49 16.30	A 185 27 27.5 G 185 27 29.9	Fatehgarh D 0 8	- 4.1		64
65	I	Amritsar Long. S.	770	A 31 38 2.51 G 31 37 58.72	A 74 52 26.46 G 74 52 23.45			+ 2.6	+ 3.8	65
66	M	Sangatpur T.S.	779	A 31 17 35.42 G 31 17 34.43	G 75 2 19.27	A 61 34 52.8 G 61 34 49.1	Rabza D 0 5	+ 6.1	+ 1.0	66
67	N	Khimāna T.S.	731	A 30 22 11.74 G 30 22 14.82	G 75 0 42.52				- 3.1	67
68	O	Sawaipur T.S.	697	A 29 39 13.13 G 29 39 13.96	G 75 3 6.12				- 0.8	68
69	O	Sirsa S.	738	G 29 31 35.39	G 75 1 14.76	A 17 11 0.2 G 17 10 58.1	Banka D 0 6	+ 3.7		69
70	P	Rām Thal S.	951	A 28 29 38.81 G 28 29 39.27	G 75 0 10.60				- 0.5	70
71	45 A	Bithnok H.S.	774	A 27 53 24.97 G 27 53 22.03	G 72 39 54.55				+ 2.9	71
72	A	Jambo H.S.	772	A 27 16 31.94 G 27 16 28.88	G 72 31 5.53	A 153 23 42.9 G 153 23 42.2	Sirad D 0 7	+ 1.4	+ 3.1	72
73	B	Ohamu H.S.	1065	A 26 39 53.44 G 26 39 52.74	G 72 35 26.28				+ 0.7	73
74	B	Thob H.S.	856	A 26 3 2.90 G 26 3 5.85	G 72 22 22.17	A 322 26 25.5 G 322 26 20.4	Samdari D 0 8	+ 10.4	- 3.0	74
75	C	Samdari H.S.	846	A 25 48 59.58 G 25 48 59.55	G 72 34 20.84				0.0	75
76	D	Gūru Sikkar H.S.	5650	G 24 38 58.39	G 72 46 39.73	A 248 53 38.4 G 248 53 36.1	Belka D 1 2	+ 5.0		76
77	D	Oria h.s.	4200	A 24 37 47.63 G 24 37 50.96	G 72 45 39.23				- 3.3	77
78	D	Birona S.	673	G 24 26 38.64	G 72 13 4.51	A 121 43 10.7 G 121 43 10.9	Sitora D 0 8	- 0.4		78

*A = Astronomical Value.

G = Triangulated or Geodetic Value.

†Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	
55	+1'65			-7'15			+1'00			+0'07			-3'39				-16'6	55	
56			+2'25			+0'06			-0'14			+1'66			+4'24	-2'9		56	
57	+1'55			-6'33			+1'00			+0'05			-2'90				-2'7	57	
58	+1'52			-6'12			+1'00			+0'04			-2'78				+3'3	58	
59			+2'46			-0'06			-0'16			+1'82			+4'53	-2'5		59	
60	+1'16			-3'83			+1'00			+0'09			-1'35				+2'5	60	
61			+2'87			-0'13			-0'16			+1'90			+4'95	-6'7		61	
62	+1'07			-3'50			+1'00			+0'08			-1'20				+4'0	62	
63	+1'43		+2'27	-5'21		-0'02	+1'00		-0'15	+0'07		+1'77	-2'06		+4'32	-3'9	-2'6	63	
64			+2'02			-0'04			-0'13			+1'81			+4'13	-8'5		64	
65	+1'46	+1'67		-5'70	-0'23		+1'00	-0'03		+0'04	+0'13		-2'51	+1'53		+1'1	+6'3	65	
66	+1'43		+1'35	-5'48		0'00	+1'00		-0'15	+0'04		+1'76	-2'39		+3'47	+2'3	+3'4	66	
67	+1'31			-4'85			+1'00			+0'04			-2'03				-1'1	67	
68	+1'22			-3'40			+1'00			+0'04			-1'04				+0'2	68	
69			+1'40			-0'02			-0'10			+1'85			+3'64	-0'1		69	
70	+1'01			-3'49			+1'00			+0'04			-1'30				+0'8	70	
71	+0'96			-3'04			+1'00			+0'08			-0'97				+3'9	71	
72	+0'83		+2'86	-2'57		-0'21	+1'00		-0'20	+0'08		+1'90	-0'72		+4'87	-3'7	+3'8	72	
73	+0'70			-2'09			+1'00			+0'08			-0'50				+1'2	73	
74	+0'57		+3'02	-1'59		-0'37	+1'00		-0'21	+0'08		+2'07	-0'18		+5'12	+5'3	-2'8	74	
75	+0'50			-1'41			+1'00			+0'08			-0'18				+0'2	75	
76			+2'89			-0'38			-0'20			+2'17			+5'13	-0'1		76	
77	+0'19			-0'44			+1'00			+0'08			+0'27				-3'6	77	
78			+3'24			-0'44			-0'23			+2'20			+5'43	-5'8		78	

* $\delta a = 0.743$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	
79	45 D	Deesa Tel. Office s.	443	A 24 15 21.15 G 24 15 20.35	A 72 11 1.26 G 72 11 4.85	A 241 16 15.3 G 241 16 15.5	Jairāj E 1 21	- 0.4	- 8.2	79
80	D	Ohaniāna H.S.	953	A 24 6 25.39 G 24 6 36.64	G 72 32 19.66				- 11.3	80
81	H	Kānnagar H.S.	3607	G 24 58 28.78	G 73 18 59.95	A 266 45 16.1 G 266 45 19.0	Māl Niver E 0 2	- 6.2		81
82	H	Tiki H.S.	2369	A 24 55 34.52 G 24 55 38.24	G 73 50 44.41	A 106 4 27.1 G 106 4 23.4	Māl Niver E 0 57	+ 8.0	- 3.7	82
83	H	Lakarwas H.S.	2574	A 24 31 41.05 G 24 31 47.99	G 73 49 43.23				- 6.9	83
84	J	Rewat H.S.	1542	A 26 53 54.74 G 26 53 53.08	G 74 16 53.79				+ 0.8	84
85	J	Jetgarh H.S.	1967	A 26 18 8.02 G 26 18 6.39	G 74 18 36.91				+ 1.6	85
86	J	Rājgarh H.S.	2618	G 26 17 49.31	G 74 35 44.37	A 156 43 41.0 G 156 43 39.9	Kisanpura D 0 8	+ 2.2		86
87	K	Khāmor H.S.	1393	A 25 45 11.00 G 25 45 15.01	G 74 47 29.19				- 4.0	87
88	L	Sāud H.S.	1910	G 24 43 6.11	G 74 32 58.48	A 284 36 7.8 G 284 36 3.9	Mendki D 0 7	+ 8.5		88
89	L	Aramlia s.	1532	A 24 25 2.66 G 24 25 7.27	G 74 59 5.69	A 244 39 1.5 G 244 38 58.9	Nanka Hūāro E 0 5	+ 5.7	- 4.6	89
90	M	Garinda s.	1204	A 27 55 30.05 G 27 55 30.55	G 75 1 18.47	A 115 55 45.3 G 115 55 42.2	Bīramsir D 0 3	+ 5.8	- 0.5	90
91	P	Rāmpūra H.S.	1920	G 24 28 44.16	G 75 26 52.24	A 260 5 35.8 G 260 5 35.0	Nimthūr D 0 16	+ 1.8		91
92	46 A	Kardo H.S.	807	A 23 57 2.27 G 23 57 10.02	G 72 43 52.88				- 7.8	92
93	A	Kainath H.S.	1385	A 23 51 14.99 G 23 51 23.79	G 72 58 51.75				- 8.8	93
94	A	Dhamanva T.S.	397	A 23 32 2.65 G 23 32 8.40	G 72 30 56.82				- 5.8	94
95	A	Morali H.S.	466	A 23 25 17.47 G 23 25 23.18	G 72 57 44.96				- 5.7	95
96	A	Sonāda T.S.	250	A 23 7 15.61 G 23 7 19.89	G 72 46 0.14	A 334 35 18.1 G 334 35 10.2	Mirzāpur D 0 5	+ 18.5	- 4.3	96
97	B	Pāldi H.S.	208	A 22 53 51.60 G 22 53 57.07	G 72 31 30.86				- 5.5	97
98	D	Pārnera H.S.	614	A 20 32 49.83 G 20 32 56.85	G 72 56 56.42	A 349 0 27.3 G 349 0 13.6	Gambhīrgad E 0 17	+ 36.5	- 7.0	98
99	F	Patāngdi H.S.	922	G 22 52 15.70	G 73 53 22.34	A 16 47 30.9 G 16 47 26.6	Bhor D 0 1	+ 10.2		99
100	F	Ghorārāo H.S.	323	A 22 52 8.05 G 22 52 11.17	G 73 21 25.45				- 3.1	100
101	F	Pāvāgad H.S.	2721	A 22 27 39.95 G 22 27 44.33	G 73 31 1.07				- 4.4	101
102	F	Sidhpur s.	169	A 22 4 11.77 G 22 4 15.21	G 73 28 59.81				- 3.4	102
103	G	Alamvādi H.S.	848	A 21 34 30.45 G 21 34 34.13	G 73 30 9.20				- 3.7	103
104	G	Tarbhān s.	140	A 21 0 28.36 G 21 0 34.13	G 73 3 49.79				- 5.8	104
105	H	Sāler H.S.	5140	G 20 43 18.44	G 73 56 21.93	A 151 26 55.7 G 151 26 50.4	Dopāri D 1 29	+ 14.0		105
106	M	Deo Dongri H.S.	1727	A 23 26 43.17 G 23 26 47.79	G 75 32 16.99				- 4.6	106

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $w_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
79	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	79	
80	+0.11	+3.28	+3.27	-0.13	-0.47	-0.46	+1.00	-0.04	-0.23	+0.09	0.00	+2.21	+0.43	+2.67	+5.46	- 5.9	- 8.6	80	
81	+0.07			-0.04			+1.00			+0.09			+0.46				- 11.8	81	
82			+2.55			-0.27			-0.18			+2.16			+4.89	- 11.1		82	
83	+0.25		+2.24	-0.69		-0.29	+1.00		-0.16	+0.06		+2.15	+0.11		+4.57	+ 3.4	- 3.8	83	
84	+0.15			-0.35			+1.00			+0.06			+0.26				- 7.2	84	
85	+0.72			-2.27			+1.00			+0.06			-0.64				+ 1.4	85	
86	+0.59			-1.80			+1.00			+0.05			-0.42				+ 2.0	86	
87			+1.73			-0.17			-0.12			+2.06			+4.09	- 1.9		87	
88	+0.45			-1.36			+1.00			+0.05			-0.24				- 3.8	88	
89			+1.85			-0.24			-0.13			+2.18			+4.30	+ 4.2		89	
90	+0.10		+1.50	-0.26		-0.22	+1.00		-0.11	+0.05		+2.20	+0.26		+4.12	+ 1.6	- 4.9	90	
91	+0.92		+1.46	-3.07		-0.07	+1.00		-0.10	+0.06		+1.95	-1.04		+3.78	+ 1.8	+ 0.5	91	
92			+1.32			-0.18			-0.09			+2.20			+3.90	- 2.1		92	
93	+0.01			+0.13			+1.00			+0.08			+0.52				- 8.3	93	
94	-0.02			+0.22			+1.00			+0.07			+0.54				- 9.3	94	
95	-0.11			+0.49			+1.00			+0.08			+0.68				- 6.5	95	
96	-0.15			+0.59			+1.00			+0.08			+0.70				- 6.4	96	
97	-0.24		+3.03	+0.85		-0.52	+1.00		-0.26	+0.08		+2.32	+0.82		+5.32	+ 13.4	- 5.1	97	
98	-0.30			+1.04			+1.00			+0.08			+0.91				- 6.4	98	
99	-1.09		+3.17	+3.25		-0.75	+1.00		-0.24	+0.08		+2.59	+1.81		+5.63	+ 31.7	- 8.8	99	
100			+2.35			-0.42			-0.17			+2.34			+4.74	+ 6.0		100	
101	-0.32			+1.07			+1.00			+0.07			+0.90				- 4.0	101	
102	-0.46			+1.43			+1.00			+0.07			+1.04				- 5.4	102	
103	-0.58			+1.78			+1.00			+0.07			+1.18				- 4.6	103	
104	-0.48			+2.23			+1.00			+0.07			+1.61				- 5.3	104	
105	-0.59			+2.73			+1.00			+0.07			+1.88				- 7.7	105	
106			+2.49			-0.74			-0.18			+2.57			+5.03	+ 9.5		106	
107	-0.18			+0.68			+1.00			+0.04			+0.69				- 5.3	107	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
107	46N	Indrāwan T.S.	1834	"	"	"	"	"	"	107
				G 22 48 48.54	G 75 10 56.60	A 273 34 2.7 G 273 34 1.9	Harnāsa D 0 9	+ 1.9		
108	N	Harnāsa T.S.	1818	A 22 47 26.71 G 22 47 29.91	G 75 33 10.15				- 3.2	108
109	N	Thikri H.S.	854	A 22 1 3.92 G 22 1 2.77	G 75 24 49.98				+ 1.2	109
110	P	Valvādi H.S.	1128	A 20 44 21.27 G 20 44 27.73	G 75 11 7.12	A 166 52 6.2 G 166 51 59.8	Ajnād D 0 7	+ 16.9	- 6.5	110
111	47B	Bombay, Colaba Long S.	75	A 18 53 39.15 G 18 53 49.48	A 72 48 55.82 G 72 48 49.10			+ 6.4	- 10.3	111
112	B	Bombay, Colaba Obay. S.	63	G 18 53 46.51	G 72 48 47.31	A 288 5 27.7 G 288 5 24.5	Karanja E 1 7	+ 9.3		112
113	B	Karanja H.S.	997	A 18 51 13.79 G 18 51 24.99	G 72 56 21.88	A 173 10 2.5 G 173 9 56.1	Trombay D 0 4	+ 18.7	- 11.2	113
114	B	Kankesvar H.S.	1260	A 18 44 17.89 G 18 44 28.16	G 72 55 34.09				- 10.3	114
115	B	Alibagh Obay. S.	10	A 18 38 26.35 G 18 38 36.60	G 72 52 12.42				- 10.3	115
116	E	Kalsubai H.S.	5400	A 19 35 57.89 G 19 36 1.76	G 73 42 35.26	A 73 2 14.5 G 73 2 11.8	Kāmāndrug D 1 1	+ 7.6	- 3.9	116
117	F	Mira Donger H.S.	1863	A 18 40 55.97 G 18 41 1.68	G 73 9 48.88				- 5.7	117
118	F	Māndvi H.S.	4121	A 18 37 47.94 G 18 37 51.11	G 73 32 21.71	A 271 15 3.9 G 271 15 7.8	Dighi D 0 56	- 11.6	- 3.2	118
119	G	Mahabaleswar H.S.	4719	A 17 55 9.91 G 17 55 15.55	G 73 40 17.41				- 5.6	119
120	G	Mirya H.S.	473	A 17 1 29.65 G 17 1 35.92	G 73 15 39.43	A 167 2 11.4 G 167 2 7.4	Adhūr D 0 12	+ 13.1	- 6.3	120
121	J	Khānpisura H.S.	2751	A 18 45 22.60 G 18 45 30.65	G 74 47 49.81	A 191 14 39.3 G 191 14 43.9	Agargaon D 0 2	- 13.5	- 8.1	121
122	J	Dhauleswar H.S.	2939	A 18 25 41.84 G 18 25 41.64	G 74 9 48.48	A 198 21 22.6 G 198 21 24.4	Bābulsār D 0 31	- 5.4	+ 1.2	122
123	K	Pāchvad H.S.	3138	G 17 31 1.97	G 74 39 43.71	A 331 12 27.4 G 331 12 29.6	Palsi D 0 13	- 7.0		123
124	L	Mājāla H.S.	2613	A 16 46 55.45 G 16 46 56.82	G 74 26 55.57				- 1.4	124
125	L	Māvinhūnda H.S.	2582	A 16 25 4.47 G 16 25 4.19	G 74 47 40.38				+ 0.3	125
126	L	Karabgati H.S.	2544	G 16 7 34.87	G 74 47 56.35	A 179 9 24.9 G 179 9 26.9	Māvinhūnda D 0 7	- 6.9		126
127	M	Dhaigaon S.	1553	A 19 30 30.82 G 19 30 35.04	G 75 12 43.81				- 4.2	127
128	N	Kanheri H.S.	2610	A 18 29 21.84 G 18 29 30.75	G 75 43 16.69	A 311 59 50.6 G 311 59 53.3	Garh Dāud D 0 16	- 8.1	- 8.9	128
129	N	Alsunda H.S.	2165	G 18 26 52.37	G 75 0 35.11	A 227 31 58.4 G 227 32 1.8	Sautāra D 0 4	- 10.2		129
130	N	Kem H.S.	1951	A 18 10 45.68 G 18 10 48.90	G 75 18 23.92				- 3.2	130
131	48 E	Chaukola H.S.	2794	A 15 55 24.94 G 15 55 31.44	G 73 59 21.13	A 166 14 13.4 G 166 14 13.2	Valvan D 0 5	+ 0.7	- 6.5	131
132	I	Kumbhāri H.S.	2898	A 15 9 4.31 G 15 9 1.80	G 74 17 47.20	A 154 15 36.5 G 154 15 33.4	Saili D 0 30	+ 11.4	+ 2.5	132
133	J	Koramūr H.S.	2525	A 14 8 1.71 G 14 8 6.59	G 74 58 24.07	A 235 28 6.8 G 235 28 9.7	Hōnnavalli E 0 3	- 11.5	- 4.9	133
134	L	Mangalore Long S.	186	A 12 52 17.76 G 12 52 14.76	A 74 50 44.70 G 74 50 42.71	A 205 52 50.8 G 205 52 49.6	Mijār E 0 17	+ 5.2	+ 3.0	134

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.8$.						
	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	"	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
107	"	"	+1.55	"	"	-0.31	"	"	-0.12	"	"	+2.35	"	"	+4.20	-2.1	"	107	
108	-0.38			+1.15			+1.00			+0.04			+0.85				-4.1	108	
109	-0.63			+1.83			+1.00			+0.04			+1.14				+0.1	109	
110	-1.06		+1.66	+2.99		-0.39	+1.00		-0.12	+0.04		+2.57	+1.61		+4.53	+12.9	-8.1	110	
111	-1.70	+2.89		+4.70	-0.41		+1.00	-0.03		+0.08	-0.09		+2.33	+2.24		+4.2	-12.6	111	
112			+3.48			-0.99			-0.32			+2.81			+6.00	+4.2		112	
113	-1.72		+3.40	+4.74		-0.97	+1.00		-0.31	+0.08		+2.81	+2.34		+5.95	+13.6	-13.5	113	
114	-1.66			+4.85			+1.00			+0.08			+2.47				-12.8	114	
115	-1.80			+4.95			+1.00			+0.08			+2.42				-12.7	115	
116	-1.45		+2.76	+4.04		-0.73	+1.00		-0.21	+0.06		+2.72	+2.05		+5.44	+3.0	-6.0	116	
117	-1.79			+4.91			+1.00			+0.07			+2.39				-8.1	117	
118	-1.82		+2.83	+4.64		-0.82	+1.00		-0.21	+0.07		+2.69	+2.16		+5.41	-16.1	-5.4	118	
119	-2.10			+5.65			+1.00			+0.06			+2.64				-8.2	119	
120	-2.46		+3.44	+6.52		-1.17	+1.00		-0.26	+0.07		+3.11	+2.97		+6.24	+7.8	-9.3	120	
121	-1.79		+2.07	+4.84		-0.60	+1.00		-0.16	+0.05		+2.84	+2.32		+5.08	-17.7	-10.4	121	
122	-1.91		+2.55	+5.16		-0.76	+1.00		-0.19	+0.06		+2.88	+2.45		+5.45	-10.0	-1.3	122	
123			+2.29			-0.75			-0.17			+3.03			+5.40	-11.1		123	
124	-2.59			+6.77			+1.00			+0.05			+3.01				-4.4	124	
125	-2.75			+7.13			+1.00			+0.05			+3.13				-2.8	125	
126			+2.36			-0.92			-0.18			+3.27			+5.65	-11.2		126	
127	-1.57			+4.29			+1.00			+0.04			+2.10				-6.3	127	
128	-1.92		+1.42	+5.11		-0.20	+1.00		-0.11	+0.03		+2.88	+2.38		+4.84	-12.0	-11.3	128	
129			+1.94			-0.36			-0.15			+2.88			+5.27	-14.6		129	
130	-2.03			+5.40			+1.00			+0.04			+2.69				-5.9	130	
131	-2.95		+3.04	+7.62		-1.15	+1.00		-0.24	+0.06		+3.32	+3.33		+7.08	-5.3	-9.8	131	
132	-3.30		+2.91	+8.41		-1.17	+1.00		-0.23	+0.05		+3.49	+3.58		+6.25	+6.6	-1.1	132	
133	-3.77		+2.47	+9.45		-1.09	+1.00		-0.19	+0.04		+3.74	+3.91		+6.24	-15.7	-8.8	133	
134	-4.36	+1.67	+2.80	+10.77	-0.21	-1.36	+1.00	-0.01	-0.22	+0.05	-0.19	+4.09	+4.34	+1.13	+6.79	+1.0	-1.3	134	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE
Deflections of the Plumb-line

Serial No	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed Station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	
135	48 M	Navalūr H.S.	2445	A 15 25 28.48 G 15 25 31.17	G 75 3 15.42				- 2.7	135
136	M	Kundgol H.S.	2145	A 15 15 14.46 G 15 15 15.28	G 75 14 46.64				- 0.8	136
137	N	Hönnavalli H.S.	2777	A 14 16 30.76 G 14 16 32.46	G 75 10 58.48				- 1.7	137
138	53 A	Medwāni H.S.	1935	G 31 17 40.45 A 30 38 16.03 G 30 38 20.01	G 76 11 58.72	A 64 43 36.5 G 64 43 41.6	Hin D 0 54	- 8.4		138
139	B	Isanpur H.S.	874		G 76 6 40.41				- 4.0	139
140	B	Bowra T.S.	855	G 30 20 50.29	G 76 6 39.00	A 208 37 15.2 G 208 37 12.4	Sudhiwal D 0 4	+ 4.8		140
141	B	Kheri T.S.	822	G 30 5 9.30	G 76 5 54.58	A 212 55 16.6 G 212 55 17.2	Khānpur D 0 3	- 1.0		141
142	C	Rākhi T.S.	785	A 29 17 20.76 G 29 17 21.28	G 76 6 47.49	A 208 30 58.2 G 208 30 55.5	Barowdha D 0 5	+ 4.8	- 0.5	142
143	E	Lambatach H.S.	10474	A 31 0 34.38 G 31 1 8.46	G 76 54 2.93				- 34.1	143
144	F	Bajamara H.S.	9681	A 30 45 27.79 G 30 45 56.20	G 77 54 0.73				- 28.4	144
145	F	Amsot H.S.	3140	A 30 22 16.02 G 30 22 44.86	G 77 41 14.77				- 28.8	145
146	F	Dehra Dun Base-line E. End S.	1967	A 30 16 37.26 G 30 17 7.35	G 77 58 30.74				- 30.1	146
147	F	Khujnaur s.	2576	A 30 15 56.70 G 30 16 23.63	G 77 52 58.67				- 26.9	147
148	F	Shorpur H.S.	2916	A 30 13 15.30 G 30 13 44.43	G 77 57 30.61				- 29.1	148
149	F	Hatni h.s.	3069	A 30 12 31.93 G 30 13 1.52	G 77 52 19.58				- 29.6	149
150	F	Bulāwāla h.s.	2432	A 30 6 22.32 G 30 6 51.29	G 77 59 11.27				- 29.0	150
151	G	Nojli T.S.	929	A 29 53 14.12 G 29 53 27.76	G 77 40 24.59				- 13.6	151
152	G	Godhna T.S.	901	A 29 37 8.73 G 29 37 18.46	G 77 54 2.98				- 9.7	152
153	G	Kaliāna S.	828	A 29 30 47.98 G 29 30 54.70	G 77 39 6.03	A 164 18 46.4 G 164 18 46.9	Dahera E 0 1	- 0.9	- 6.7	153
154	H	Datāiri T.S.	767	A 28 43 58.67 G 28 44 4.49	G 77 38 56.31	A 28 44 34.3 G 28 44 34.2	Bostān D 0 6	+ 6.2	- 5.8	154
155	H	Bostān T.S.	758	A 28 30 54.25 G 28 30 59.64	G 77 30 49.08				- 5.4	155
156	H	Chandaos T.S.	699	A 28 5 0.71 G 28 5 1.59	G 77 51 39.60				- 0.9	156
157	I	Kidarkanta H.S.	12509	A 31 0 51.58 G 31 1 21.71	G 78 10 23.37				- 30.1	157
158	J	Bahak H.S.	9715	A 30 44 37.60 G 30 45 5.22	G 78 13 36.98				- 27.6	158
159	J	Nag Tibba H.S.	9915	A 30 34 41.05 G 30 35 11.57	G 78 9 9.57	A 32 58 41.6 G 32 58 53.9	Eagle's Nest D 8 12	- 20.8	- 30.5	159
160	J	Banog H.S.	7433	A 30 28 4.18 G 30 28 36.91	G 78 0 55.96	A 71 5 55.0 G 71 6 8.7	Amsot D 2 28	- 23.3	- 32.7	160
161	J	Mussooree Dome Obsy. H.S.	6937	A 30 27 4.02 G 30 27 40.55	G 78 4 17.41	A 280 22 46.8 G 280 23 0.5	Top Tibba E 1 7			
						A 6 17 20.1 G 6 17 35.1	Cole's Satellite Station D 5 26	- 25.5	- 36.5	161

* A = Astronomical Value.
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a=1$ km			Case II: $\delta b=1$ km			Case III: Latitude $w_0=1''$			Case IV: Azimuth $w_0=1''$			$a=6378200$ metres, $1/e=298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
135	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	135
136	-3'19			+8'13			+1'00			+0'04			+3'46				- 6'2	136	
137	-3'27			+8'31			+1'00			+0'04			+3'51				- 4'3	137	
138	-3'71			+9'31			+1'00			+0'04			+3'84				- 5'5	138	
139		+0'75				0'00			-0'05			+1'76			+2'93	- 11'8		139	
140	+1'34			-5'03			+1'00			+0'03			-2'16				- 1'8	140	
141		+0'81				-0'02			-0'05			+1'81			+3'05	+ 1'4		141	
142		+0'82				-0'02			-0'05			+1'82			+3'08	- 4'4		142	
143	+1'15	+0'82		-4'08		-0'03	+1'00		-0'06	+0'03		+1'87	-1'63		+3'13	+ 1'3	+ 1'1	143	
144	+1'38			-5'29			+1'00			0'00			-2'35				- 31'7	144	
145	+1'35			-5'12			+1'00			0'00			-2'25				- 26'1	145	
146	+1'30			-4'85			+1'00			0'00			-2'09				- 26'7	146	
147	+1'29			-4'79			+1'00			-0'01			-2'07				- 28'0	147	
148	+1'29			-4'78			+1'00			0'00			-2'06				- 24'8	148	
149	+1'28			-4'75			+1'00			-0'01			-2'05				- 27'0	149	
150	+1'28			-4'74			+1'00			0'00			-2'04				- 27'6	150	
151	+1'27			-4'67			+1'00			-0'01			-2'00				- 27'0	151	
152	+1'24			-4'51			+1'00			0'00			-1'90				- 11'7	152	
153	+1'20			-4'32			+1'00			0'00			-1'81				- 7'9	153	
154	+1'18	0'00		-4'24		0'00	+1'00		0'00	0'00		+1'85	-1'5		+2'39	- 3'8	- 4'9	154	
155	+1'05	0'00		-3'67		0'00	+1'00		0'00	0'00		+1'90	-1'45		+2'45	- 2'7	- 4'3	155	
156	+1'01			-3'51			+1'00			0'00			-1'36				- 4'0	156	
157	+0'94			-3'18			+1'00			0'00			-1'21				+ 0'3	157	
158	+1'38			-5'30			+1'00			-0'01			-2'35				- 27'7	158	
159	+1'35			-5'11			+1'00			-0'01			-2'23				- 25'4	159	
160	+1'33	-0'26		-4'99		+0'01	+1'00		+0'02	-0'01		+1'79	-2'18		+2'08	- 23'4	- 28'3	160	
161	+1'31	-0'19		-4'92		0'00	+1'00		+0'01	-0'01		+1'80	-2'14		+2'15	- 26'0	- 30'6	161	
162	+1'31	-0'22		-4'91		0'00	+1'00		+0'01	-0'01		+1'80	-2'14		+2'12	- 28'1	- 34'4	162	

* $\delta a = 0.924$, $\delta b = 0.743$, $w_0 = 0.31$, $w_0 = 1.29$. *Vide* p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†	
162	53 J	Jharipani (IX) h.s.	5150	A 30 24 17.55 G 30 25 10.05	G 78 5 20.92	A 14 24 59.9 G 14 25 18.0	Dalanwāla D 4 38	- 30.8	- 52.5	162
163	J	Spur point (VIII) h.s.	3850	A 30 23 44.55 G 30 24 37.72	G 78 5 35.96	A 17 59 32.4 G 17 59 49.1	Dalanwāla D 2 48	- 28.5	- 53.2	163
164	J	Rajpur h.s.	3500	A 30 23 9.75 G 30 23 56.83	G 78 5 59.89	A 24 15 4.4 G 24 15 21.8	Dalanwāla D 2 26	- 29.7	- 47.7	164
165	J	VI ...	3050	A 30 22 44.90 G 30 23 30.79	G 78 6 2.00				- 45.9	165
166	J	V ...	2980	A 30 22 7.46 G 30 22 51.83	G 78 5 21.38				- 44.4	166
167	J	IV ...	2780	A 30 21 26.78 G 30 22 8.93	G 78 4 30.87				- 42.2	167
168	J	III ...	2660	A 30 21 5.57 G 30 21 46.61	G 78 4 7.39				- 41.0	168
169	J	Dehra Dun Obsy. (Old) S.	2289	A 30 19 19.56 G 30 19 57.07	G 78 3 34.70	A 165 10 58.8 G 165 11 10.2	Banog E 5 20	- 19.5	- 37.5	169
170	J	Dehra Dun Haig Obsy. S.	2240	A 30 18 51.80 G 30 19 28.73	G 78 2 56.47 G 78 3 22.12			- 22.1	- 36.9	170
171	J	Lachkuwa h.s.	2674	A 30 4 5.34 G 30 4 34.24	G 78 1 41.67				- 28.9	171
172	J	Kānigarh H.S.	7055	A 30 3 34.80 G 30 4 4.47	G 78 42 54.38				- 29.7	172
173	K	Harpālsid T.S.	1000	A 29 39 22.24 G 29 39 50.84	G 78 33 20.81				- 28.6	173
174	K	Mahesari T.S.	821	A 29 30 8.18 G 29 30 18.21	G 78 8 51.70				- 10.0	174
175	K	Sarkāra T.S.	761	A 29 15 35.09 G 29 15 46.91	G 78 32 20.18				- 11.8	175
176	L	Sirsa T.S.	739	A 28 54 30.27 G 28 54 39.64	G 78 32 6.14	A 149 55 17.1 G 149 55 20.5	Milik D 0 4	- 6.2	- 9.4	176
177	L	Bānsgopāl T.S.	677	A 28 33 23.28 G 28 33 28.08	G 78 31 59.71				- 4.8	177
178	L	Sankrāo T.S.	670	A 28 2 28.92 G 28 2 29.00	G 78 32 2.97	A 185 44 20.9 G 185 44 18.8	Sakrora D 0 8	+ 3.9	- 0.1	178
179	O	Birond H.S.	6967	A 29 14 29.73 G 29 15 14.15	G 79 42 57.00				- 44.4	179
180	P	Kaliānpur T.S.	629	G 28 35 11.10	G 79 44 33.75	A 185 30 18.4 G 185 30 17.7	Donao D 0 4	+ 1.3		180
181	54 A	Tāsīng H.S.	2050	A 27 52 59.49 G 27 52 59.47	G 76 12 11.56	A 77 55 36.5 G 77 55 31.6	Jīlo E 0 13	+ 9.3	0.0	181
182	B	Bānskho H.S.	1870	A 26 50 2.37 G 26 50 7.89	G 76 8 20.42	A 148 40 55.6 G 148 40 51.9	Rāmgār E 0 2	+ 7.3	- 5.5	182
183	C	Kānkra H.S.	1652	A 25 37 58.75 G 25 37 59.53	G 76 7 27.15	A 145 33 8.7 G 145 33 6.9	Bhojpur D 0 18	+ 3.8	- 0.8	183
184	D	Gurāria H.S.	1360	A 24 25 31.98 G 24 25 32.46	G 76 5 2.16	A 300 41 56.8 G 300 41 56.2	Kūsalpura D 0 4	+ 1.3	- 0.5	184
185	D	Māta-ka-hūra H.S.	1645	G 24 14 10.67	G 76 36 49.20	A 181 31 35.0 G 181 31 34.3	Sartal D 0 15	+ 1.6		185
186	E	Noh T.S.	710	A 27 50 53.73 G 27 50 53.08	G 77 38 45.56	A 50 22 36.5 G 50 22 33.4	Mānpur D 0 6	+ 5.9	+ 0.1	186
187	E	Agra-group W. Point	550	A 27 9 41.43 G 27 9 45.86	G 77 56 25.26				- 4.4	187
188	F	Usira H.S.	810	A 26 57 0.50 G 26 57 6.22	G 77 37 52.58	A 146 55 27.2 G 146 55 25.9	Madhoni D 0 12	+ 2.6	- 5.7	188
189	G	Kesri H.S.	1487	A 25 46 41.57 G 25 46 35.81	G 77 40 49.02	A 206 41 38.8 G 206 41 40.1	Dīn D 0 10	- 2.7	+ 5.8	189

*A - Astronomical Value.

G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	
162	+1'31	"	-0'23	-4'88	"	0'00	+1'00	"	+0'02	-0'01	"	+1'80	-2'11	"	+2'11	-33'4	-50'4	162	
163	+1'31	"	-0'23	-4'87	"	0'00	+1'00	"	+0'02	-0'01	"	+1'80	-2'11	"	+2'12	-33'1	-51'1	163	
164	+1'30	"	-0'23	-4'86	"	+0'01	+1'00	"	+0'02	-0'01	"	+1'80	-2'11	"	+2'11	-32'3	-45'6	164	
165	+1'30	"	"	-4'86	"	"	+1'00	"	"	-0'01	"	"	-2'11	"	"	"	-43'8	165	
166	+1'30	"	"	-4'85	"	"	+1'00	"	"	-0'01	"	"	-2'10	"	"	"	-42'3	166	
167	+1'30	"	"	-4'84	"	"	+1'00	"	"	-0'01	"	"	-2'10	"	"	"	-40'1	167	
168	+1'30	"	"	-4'84	"	"	+1'00	"	"	-0'01	"	"	-2'10	"	"	"	-38'9	168	
169	+1'30	"	-0'21	-4'82	"	0'00	+1'00	"	+0'01	-0'01	"	+1'81	-2'08	"	+2'13	-22'1	-35'4	169	
170	+1'30	-0'24	"	-4'81	+0'03	"	+1'00	0'00	"	-0'01	+0'11	"	-2'08	-0'05	"	-22'1	-34'8	170	
171	+1'26	"	"	-4'64	"	"	+1'00	"	"	-0'01	"	"	-1'98	"	"	"	-26'9	171	
172	+1'26	"	"	-4'64	"	"	+1'00	"	"	-0'02	"	"	-1'98	"	"	"	-27'7	172	
173	+1'20	"	"	-4'35	"	"	+1'00	"	"	-0'02	"	"	-1'83	"	"	"	-26'8	173	
174	+1'18	"	"	-4'23	"	"	+1'00	"	"	-0'01	"	"	-1'75	"	"	"	-8'2	174	
175	+1'14	"	"	-4'06	"	"	+1'00	"	"	-0'02	"	"	-1'66	"	"	"	-10'1	175	
176	+1'09	"	-0'48	-3'80	"	+0'02	+1'00	"	+0'03	-0'02	"	+1'89	-1'53	"	+2'02	-8'6	-7'9	176	
177	+1'02	"	"	-3'54	"	"	+1'00	"	"	-0'01	"	"	-1'40	"	"	"	-3'4	177	
178	+0'93	"	-0'48	-3'15	"	+0'03	+1'00	"	+0'03	-0'01	"	+1'94	-1'19	"	+2'09	+1'4	+1'1	178	
179	+1'15	"	"	-4'05	"	"	+1'00	"	"	-0'03	"	"	-1'68	"	"	"	-42'7	179	
180	"	"	-1'19	"	"	+0'06	"	"	+0'08	"	"	+1'91	"	"	+1'43	-0'5	"	180	
181	+0'90	"	+0'80	-3'03	"	-0'05	+1'00	"	-0'06	+0'02	"	+1'95	-1'08	"	+3'20	+5'9	+1'1	181	
182	+0'69	"	+0'85	-2'22	"	-0'07	+1'00	"	-0'06	+0'03	"	+2'02	-1'68	"	+3'33	+3'8	-3'8	182	
183	+0'40	"	+0'89	-1'26	"	-0'10	+1'00	"	-0'06	+0'03	"	+2'11	-0'22	"	+3'45	+0'3	-0'6	183	
184	+0'09	"	+0'94	-0'26	"	-0'13	+1'00	"	-0'07	+0'03	"	+2'21	-0'23	"	+3'60	-2'3	-0'3	184	
185	"	"	+0'63	"	"	-0'09	"	"	-0'04	"	"	+2'22	"	"	+3'37	-1'8	"	185	
186	+0'89	"	+0'01	-3'00	"	0'00	+1'00	"	0'00	0'00	"	+1'95	-1'10	"	+2'53	+3'0	+1'2	186	
187	+0'75	"	"	-2'48	"	"	+1'00	"	"	-0'01	"	"	-0'85	"	"	"	-3'5	187	
188	+0'71	"	+0'01	-2'31	"	0'00	+1'00	"	0'00	0'00	"	+2'01	-0'76	"	+2'61	-0'2	-4'9	188	
189	+0'43	"	-0'01	-1'37	"	0'02	+1'00	"	0'00	0'00	"	+2'10	-0'31	"	+2'70	-5'6	+6'1	189	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.								Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflection†		
190	54 H	Pahāgarh H.S.	1641	A 24 56 6.47 G 24 56 6.92	G 77 41 46.32	A 236 19 20.1 G 236 19 17.8	Nimdānt D 0 4	+ 4.9	- 0.5	190	
191	H	Waiādhari H.S.	1867	A 24 38 18.79 G 24 38 17.59	G 77 39 49.31				+ 1.2	191	
192	H	Salot H.S.	1834	G 24 14 52.08	G 77 15 2.41	A 175 58 10.5 G 175 58 10.7	Hatni D 0 8	- 0.4		192	
193	H	Sūrāntal H.S.	1802	A 24 14 21.36 G 24 14 20.42	G 77 40 43.91				+ 0.9	193	
194	H	Sironj Base-line N.E. End S.	1481	A 24 8 55.45 G 24 8 53.57	G 77 50 41.14				+ 1.9	194	
195	H	Kaliānpur H.S.	1765	A 24 7 10.97 G 24 7 11.26	A 77 39 17.57	A 190 27 6.29 G 190 27 5.10	Sūrāntāl D 0 0	+ 2.7	- 0.3	195	
		(Origin)‡	1765	A 27 7 11.57 G 27 7 11.26	A 77 39 17.57	A 190 27 6.39 G 190 27 5.10	Surantal	+ 2.9	+ 0.3		
196	H	Tinsia H.S.	1776	A 24 6 29.05 G 24 6 27.97	G 77 18 30.70				+ 1.1	196	
197	H	Losalli S.	1749	A 24 6 18.19 G 24 6 19.17	G 77 33 14.11	A 149 5 52.1 G 149 5 50.4	Rāmpur D 0 2	+ 3.8	- 1.0	197	
198	I	Salimpur T.S.	645	A 27 46 36.23 G 27 46 36.46	G 78 30 48.70				- 0.2	198	
199	I	Agra-group N. Point	550	A 27 14 10.31 G 27 14 14.10	G 78 1 4.71				- 3.8	199	
200	I	Agra Long. S.	550	A 27 9 34.62 G 27 9 39.93	A 78 1 7.49 G 78 1 1.89			+ 5.0	- 5.3	200	
201	I	Agra-group E. Point	550	A 27 9 16.21 G 27 9 21.00	G 78 6 3.64				- 4.8	201	
202	I	Agra Parade Point	550	A 27 8 52.18 G 27 8 57.47	G 78 1 9.70				- 5.3	202	
203	I	Agra-group S. Point	550	A 27 5 32.95 G 27 5 38.51	G 78 1 2.38				- 5.6	203	
204	J	Gūrmī T.S.	575	A 26 36 5.97 G 26 36 3.63	G 78 30 49.82	A 155 50 8.0 G 155 50 8.8	Panābat D 0 3	- 1.8	+ 2.3	204	
205	J	Majhār H.S.	1028	A 26 6 20.30 G 26 6 17.00	G 78 28 17.73				+ 3.3	205	
206	K	Algi H.S.	854	A 25 29 48.16 G 25 29 46.19	G 78 21 30.98				+ 2.0	206	
207	L	Andhiāri H.S.	1330	A 24 41 11.31 G 24 41 6.78	G 78 13 48.99				+ 4.5	207	
208	L	Bhaorāsa H.S.	1387	A 24 8 5.13 G 24 8 3.74	G 78 0 40.73				+ 1.4	208	
209	L	Budhon H.S.	1867	A 24 5 8.99 G 24 5 8.41	G 78 31 11.89	A 265 22 28.1 G 265 22 27.7	Tinsmal D 0 6	+ 0.9	+ 0.6	209	
210	M	Mohammadabad T.S.	565	G 27 18 24.05	G 79 25 39.80	A 291 59 0.9 G 291 58 51.5	Chandanpur D 0 8	+ 18.2		210	
211	P	Dargawa H.S.	1152	A 24 37 17.32 G 24 37 13.21	G 79 1 24.63				+ 4.1	211	
212	P	Rangār (old) S.	1184	A 24 0 19.28 G 24 0 20.37	G 79 25 59.25	A 106 1 11.0 G 106 1 24.2	Tinsmal E 0 11	- 29.6	- 1.1	212	
213	55 E	Kāmkhara H.S.	1780	A 23 59 42.89 G 23 59 44.93	G 77 43 6.85				- 2.0	213	
214	E	Ahmadpur H.S.	1713	A 23 36 18.42 G 23 36 20.88	G 77 40 48.26	A 185 10 55.0 G 185 10 53.8	Kāmkhara D 0 9	+ 2.7	- 2.5	214	
215	E	Lādi H.S.	1875	A 23 8 39.10 G 23 8 44.13	G 77 42 30.87				- 5.0	215	
216	F	Bhimbhat H.S.	2120	G 22 50 2.06	G 77 37 15.53	A 194 34 0.7 G 194 33 58.6	Lādi D 0 16	+ 5.0		216	

* A - Astronomical Value.

G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

‡ Derived from group of stations surrounding Kaliānpur.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
190	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	190
191	+0.32		-0.02	-0.69		0.00	+1.00		0.00	0.00		+2.16	0.00		+2.77	+2.1	-0.5	191	
192	+0.14			-0.43			+1.00			0.00			+0.12				+1.1	192	
193			+0.24			-0.04			-0.02			+2.22			+3.05	-3.5		193	
194	+0.03			-0.10			+1.00			0.00			+0.27				+0.6	194	
195	+0.01			-0.02			+1.00			0.00			+0.30				+1.6	195	
196	0.00	0.00	0.00	0.00	0.00	0.00	+1.00	0.00	0.00	0.00	0.00	+2.23	+0.30	0.00	+2.88	+0.1	-0.6	196	
(Origin)	0.00	0.00	0.00	0.00	0.00	0.00	+1.00	0.00	0.00	0.00	0.00	+2.23	+0.30	0.00	+2.88	0.0	0.0	(Origin)	
196	0.00			+0.01			+1.00			+0.01			+0.33				+0.8	196	
197	-0.01		+0.06	+0.02		-0.01	+1.00		0.00	0.00		+2.24	+0.32		+2.93	+0.9	-1.3	197	
198	+0.88			-2.95			+1.00			-0.01			-1.09				+0.9	198	
199	+0.77			-2.53			+1.00			-0.01			-0.67				-3.1	199	
200	+0.75	-0.22		-2.48	+0.03		+1.00	0.00		-0.01	+0.05		-0.85	-0.11		+5.1	-4.4	200	
201	+0.75			-2.47			+1.00			-0.01			-0.85				-3.9	201	
202	+0.75			-2.47			+1.00			-0.01			-0.84				-4.5	202	
203	+0.74			-2.42			+1.00			-0.01			-0.82				-4.8	203	
204	+0.63		-0.49	-2.04		+0.04	+1.00		+0.03	-0.01		+2.03	-0.64		+2.20	-4.2	+2.9	204	
205	+0.52			-1.64			+1.00			-0.01			-0.46				+3.8	205	
206	+0.37			-1.14			+1.00			-0.01			-0.22				+2.2	206	
207	+0.16			-1.47			+1.00			-0.01			+0.09				+4.4	207	
208	0.00			-0.01			+1.00			-0.01			+0.29				+1.1	208	
209	-0.01		-0.52	+0.03		+0.08	+1.00		+0.04	-0.01		+2.24	+0.34		+2.48	-1.6	+0.3	209	
210			-0.99			+0.07			+0.07			+1.98			+1.72	+16.3		210	
211	+0.14			-0.42			+1.00			-0.02			+0.10				+4.0	211	
212	-0.02		-1.08	+0.10		+0.16	+1.00		+0.08	-0.03		+2.24	+0.32		+2.04	-31.6	-1.4	212	
213	-0.04			+0.11			+1.00			0.00			+0.36				-2.4	213	
214	-0.15		-0.01	+0.44		0.00	+1.00		0.00	0.00		+2.28	+0.50		+2.93	-0.2	-3.0	214	
215	-0.29			+0.84			+1.00			0.00			+0.67				-5.7	215	
216			+0.02			-0.01			0.00			+2.35			+3.06	+2.1		216	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *vide* p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
217	55 G	Nilgarh H.S.	2533	" " "	" " "	" " "	Sālbaldi E 0 1	- 1'3	"	217
218	G	Takalkhera s.	1094	G 21 45 50'12 A 21 5 50'17 G 21 5 56'76	G 77 39 18'82	A 321 4 43'7 G 321 4 44'2			- 6'6	218
219	H	Rāngrai s.	1046	A 20 48 7'16 G 20 48 14'68	G 77 38 24'94 G 77 35 53'83				- 7'5	219
220	H	Badgaon H.S.	1128	A 20 44 15'54 G 20 44 23'06	G 77 36 31'79	A 183 9 0'9 G 183 8 59'5	Ashti D 0 8	+ 3'7	- 7'5	220
221	H	Dhānura s.	1135	A 20 44 3'35 G 20 44 10'84	G 77 41 43'27				- 7'5	221
222	H	Dotra s.	1140	A 20 41 22'25 G 20 41 28'91	G 77 32 45'66				- 6'7	222
223	H	Sakri H.S.	1810	G 20 0 14'11	G 77 42 7'32	A 175 24 37'7 G 175 24 35'4	Kopdi D 0 20	+ 6'3		223
224	I	Saugor H.S.	2033	A 23 49 48'71 G 23 49 48'07	G 78 46 18'16				+ 0'6	224
225	I	Nāharman H.S.	1940	A 23 30 13'14 G 23 30 18'15	G 78 49 49'13				- 5'0	225
226	M	Karaundi H.S.	1625	A 23 10 45'07 G 23 10 40'03	G 79 59 43'34	A 206 22 35'6 G 206 22 38'4	Lora D 0 1	- 6'5	+ 5'1	226
227	M	Jabalpur Long. s.	...	G 23 10 10'10	A 79 56 52'42 G 79 57 2'61			- 9'4		227
228	P	Bhimsain H.S.	1490	A 20 57 28'54 G 20 57 35'96	G 79 46 7'40	A 297 55 2'8 G 297 55 2'3	Partābgarh E 0 0	+ 1'3	- 7'4	228
229	P	Rājuli H.S.	1070	A 20 12 51'25 G 20 12 55'45	G 79 44 49'27				- 4'2	229
230	56B	Nitali H.S.	2289	A 18 17 2'74 G 18 17 7'16	G 76 16 23'32	A 239 23 1'3 G 239 23 5'7	Harangal D 0 10	- 13'3	- 4'4	230
231	B	Achola H.S.	2274	A 18 14 44'87 G 18 14 48'12	G 76 59 20'51	A 272 47 57'4 G 272 47 58'5	Manganāl D 0 11	- 3'3	- 3'3	231
232	E	Halda s.	1335	A 19 9 24'41 G 19 9 29'38	G 77 41 1'39				- 5'0	232
233	E	Voi s.	1439	A 19 7 14'69 G 19 7 19'89	G 77 34 46'88				- 5'2	233
234	E	Somtana H.S.	1714	G 19 5 0'52	G 77 39 16'29	A 186 51 46'9 G 186 51 48'3	Terbān D 0 5	- 4'0		234
235	E	Mandāla s.	1294	A 19 2 42'84 G 19 2 48'24	G 77 43 35'14				- 5'4	235
236	E	Talegaon s.	1233	A 19 1 21'65 G 19 1 26'64	G 77 37 16'75				- 5'0	236
237	F	Dāmargīda Obsy. S.	1941	A 18 3 14'92 G 18 3 17'35	G 77 40 4'41	A 188 11 59'1 G 188 11 59'8	Burgāpālī D 0 15	- 2'1	- 2'4	237
238	G	Devanūr s.	1593	A 17 10 56'88 G 17 11 0'43	G 77 41 7'41				- 3'6	238
239	G	Akampalle h.s.	1557	A 17 10 50'39 G 17 10 53'96	G 77 34 29'30				- 3'6	239
240	G	Kodangal S.	1906	A 17 7 53'74 G 17 7 57'35	G 77 38 25'73	A 62 29 16'3 G 62 29 17'8	Nēlagat E 0 1	- 4'9	- 3'6	240
241	G	Līnganapalle h.s.	1815	A 17 7 13'40 G 17 7 16'66	G 77 42 28'61				- 3'3	241
242	G	Pīalmudi s.	1869	A 17 4 1'06 G 17 4 6'05	G 77 36 22'06				- 5'0	242
243	H	Tōnsalgutta s.	1133	A 16 18 2'36 G 16 18 6'91	G 77 34 49'44				- 4'6	243
244	H	Pēddapād s.	1090	A 16 17 14'13 G 16 17 20'38	G 77 44 30'54				- 6'3	244

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/\epsilon = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
217	"	"	0'00	"	"	0'00	"	"	0'00	"	"	+2'46	"	"	+3'18	-4'2	"	217	
218	-0'95			+2'66			+1'00			0'00			+1'41				-8'0	218	
219	-1'05			+2'93			+1'00			0'00			+1'52				-9'0	219	
220	-1'08	+0'03		+2'99	-0'01		+1'00		0'00	0'00		+2'58	+1'53		+3'36	+0'6	-9'0	220	
221	-1'08			+3'00			+1'00			0'00			+1'56				-9'1	221	
222	-1'09			+3'04			+1'00			0'00			+1'56				-8'3	222	
223		-0'01				0'00			0'00			+2'67			+3'43	+3'4		223	
224	-0'08			+0'25			+1'00			-0'02			+0'40				+0'2	224	
225	-0'18			+0'53			+1'00			-0'02			+0'52				-5'5	225	
226	-0'26	-1'45	+0'81	+0'25	+1'00	+0'07	+1'00		+0'07	-0'04		+2'32	+0'62	+1'85	-8'3	+4'5		226	
227		-1'38		+0'20				+0'02			-0'02			-1'15		-8'3		227	
228	-0'99	-1'41	+2'79	+0'32	+1'00	+0'10	+1'00		+0'10	-0'03		+2'55	+1'43		+2'26	-0'7	-8'8	228	
229	-1'25		+3'48				+1'00			-0'03			+1'69				-5'9	229	
230	-1'99	+1'03	+5'30	-0'31	+1'00	-0'08	+1'00		-0'08	+0'02		+2'91	+2'44	+4'45	-16'9	-6'8		230	
231	-2'00	+0'52	+5'34	-0'16	+1'00	-0'04	+1'00		-0'04	+0'01		+2'92	+2'44	+4'11	-6'5	-5'7		231	
232	-1'65		+4'47				+1'00			0'00			+2'11				-7'1	232	
233	-1'67		+4'51				+1'00			0'00			+2'12				-7'3	233	
234		+0'01				0'00			0'00			+2'79			+3'61	-7'0		234	
235	-1'70		+4'57				+1'00			0'00			+2'14				-7'5	235	
236	-1'71		+4'60				+1'00			0'00			+2'15				-7'2	236	
237	-2'09	0'00	+5'53	+0'01	+1'00		+1'00		0'00	0'00		+2'95	+2'49	+3'80	-5'0	-4'9		237	
238	-2'45		+6'38				+1'00			0'00			+2'79				-6'4	238	
239	-2'45		+6'30				+1'00			0'00			+2'73				-6'3	239	
240	-2'47	+0'01	+6'43		0'00	+1'00			0'00	0'00		+3'10	+2'81	+4'01	-7'6	-6'4		240	
241	-2'47		+6'44				+1'00			0'00			+2'81				-6'1	241	
242	-2'50		+6'50				+1'00			0'00			+2'83				-7'8	242	
243	-2'89		+7'42				+1'00			0'00			+3'15				-7'8	243	
244	-2'83		+7'27				+1'00			0'00			+3'10				-9'4	244	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
245	56 H	Darur H.S.	1796	G 16 13 35.40	G 77 39 36.51	A 132 35 57.2 G 132 35 59.2	Kōttapalle D O 15	- 6.9		245
246	H	Tuagat h.s.	1450	A 16 9 46.73 G 16 9 51.66	G 77 34 11.59				- 4.9	246
247	H	Gattinārāyantippa h.s.	1225	A 16 7 48.95 G 16 7 54.81	G 77 45 51.00				- 5.9	247
248	H	Devaragat h.s.	1332	A 16 6 31.98 G 16 6 37.27	G 77 41 26.21				- 5.3	248
249	K	Pirmulo H.S.	2093	A 17 52 58.32 G 17 53 2.81	G 78 35 50.98	A 105 0 48.0 G 105 0 49.0	Narsula D O 12	- 3.1	- 4.5	249
250	K	Bolarum P.W.D. Office Long. S.	1971	A 17 30 7.36 G 17 30 13.41	A 78 31 7.84 G 78 31 11.12	A 25 57 35.8 G 25 57 35.8	Hyderābād Naubat- pahar D O 18	0.0	- 6.1	250
251	M	Dīwai H.S.	967	A 19 49 26.87 G 19 49 32.57	G 79 32 28.62	A 154 17 54.2 G 154 17 55.1	Ambāgarh D O 8	- 2.5	- 5.7	251
252	M	Ankora H.S.	1463	A 19 24 26.63 G 19 24 34.75	G 79 36 27.70				- 8.1	252
253	N	Burgpālli H.S.	983	A 18 54 3.48 G 18 54 7.20	G 79 41 36.96	A 142 8 7.5 G 142 8 8.8	Rechni D O 8	- 3.8	- 3.7	253
254	N	Rāmgar H.S.	1772	A 18 35 26.90 G 18 35 26.12	G 79 31 42.36				+ 0.8	254
255	O	Bolikonda H.S.	1363	A 17 42 29.08 G 17 42 35.82	G 79 47 57.98				- 6.7	255
256	O	Vānākonda H.S.	1664	A 17 36 0.22 G 17 36 6.87	G 79 22 20.70	A 180 4 14.5 G 180 4 15.3	Yarābali D O 1	- 2.5	- 6.7	256
257	O	Niālamari H.S.	1144	A 17 1 25.93 G 17 1 33.63	G 79 43 29.78				- 7.7	257
258	57 A	Bellary Long. s.	...	G 15 8 33.06	G 76 55 38.89 G 76 55 39.58			- 0.7		258
259	B	Yēragunta h.s.	1698	A 14 48 27.31 G 14 48 23.26	G 76 58 17.56				+ 4.1	259
260	C	Nughallibētta H.S.	3140	G 13 1 32.95	G 76 28 32.46	A 54 31 39.1 G 54 31 41.7	Sātanhalli E O 8	- 11.2		260
261	E	Namthabad s.	1169	A 15 5 51.75 G 15 5 52.40	G 77 36 26.91				- 0.7	261
262	F	Ohikalgurki s.	1516	A 14 59 5.16 G 14 59 4.53	G 77 11 6.39				+ 0.6	262
263	F	Bandūr s.	1447	A 14 57 44.41 G 14 57 42.32	G 77 0 36.91				+ 2.1	263
264	F	Hōnnūr H.S.	1579	A 14 55 22.20 G 14 55 18.96	G 77 6 2.60				+ 3.2	264
265	F	Nimbāgal s.	1565	A 14 51 56.14 G 14 51 52.43	G 77 11 51.78				+ 3.7	265
266	F	Pāvagada H.S.	3022	A 14 6 18.80 G 14 6 15.39	G 77 16 42.43				+ 3.4	266
267	G	Bōmmasandra s.	2005	A 13 59 42.63 G 13 59 36.34	G 77 28 43.96				+ 6.3	267
268	G	Bangalore Base-line N.E. End S.	3016	A 13 4 53.17 G 13 4 56.05	G 77 39 16.23	A 44 32 19.7 G 44 32 19.3	Bangalore Base-line S.W. End E O 8	+ 1.7	- 2.9	268
269	G	Bangalore Base-line S.W. End S.	3126	A 13 0 36.12 G 13 0 40.91	A 77 34 57.29 G 77 35 0.19	A 224 31 21.7 G 224 31 21.6	Bangalore Base-line N.E. End D O 13	+ 0.4	- 4.8	269
270	H	Dōddagunta s.	3003	A 12 59 51.52 G 12 59 55.76	G 77 37 33.25				- 4.2	270
271	M	Dānapa H.S.	150	A 15 55 59.69 G 15 56 0.14	G 79 56 7.39	A 265 47 36.0 G 265 47 39.4	Babbēpalle D O 22	- 11.9	- 0.5	271
272	M	Darutippa S.	195	A 15 0 33.52 G 15 0 36.47	G 79 55 12.92				- 3.0	272

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $w_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.8$						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
245	"	"	-0.01	"	"	0.00	"	"	0.00	"	"	+3.27	"	"	+4.21	- 9.7	"	245	
246	-2.88			+7.39			+1.00			0.00			+3.14				- 8.0	246	
247	-2.89			+7.43			+1.00			0.00			+3.16				- 9.1	247	
248	-2.90			+7.45			+1.00			0.00			+3.18				- 8.5	248	
249	-2.16		-0.71	+5.70		+0.22	+1.00		+0.05	-0.03		+2.97	+2.51		+3.37	- 5.6	- 7.0	249	
250	-2.31	-0.52	-0.66	+6.06	+0.07	+0.22	+1.00	+0.01	+0.05	-0.01	-0.11	+3.04	+2.67	-0.58	+3.49	- 2.5	- 8.8	250	
251	-1.40		-1.30	+3.83		+0.34	+1.00		+0.10	-0.03		+2.69	+1.83		+2.55	- 4.5	- 7.5	251	
252	-1.55			+4.23			+1.00			-0.03			+1.98				-10.1	252	
253	-1.74		-1.45	+4.71		+0.42	+1.00		+0.11	-0.03		+2.82	+2.16		+2.63	- 5.8	- 5.9	253	
254	-1.87			+5.05			+1.00			-0.03			+2.29				- 1.5	254	
255	-2.22			+5.86			+1.00			-0.04			+2.56				- 9.3	255	
256	-2.27		-1.31	+5.97		+0.43	+1.00		+0.10	-0.03		+3.02	+2.61		+3.04	- 4.6	- 9.3	256	
257	-2.50			+6.53			+1.00			-0.03			+2.81				-10.5	257	
258		+0.43			-0.06			0.00			-0.16			+0.16			- 0.9	258	
259	-3.48			+8.77			+1.00			+0.01			+3.62				+ 0.5	259	
260			+1.17			-0.56			-0.09			+4.05			+5.85	- 14.5		260	
261	-3.35			+8.47			+1.00			0.00			+3.51				- 4.2	261	
262	-3.40			+8.59			+1.00			+0.01			+3.56				- 3.0	262	
263	-3.41			+8.61			+1.00			+0.01			+3.57				- 1.5	263	
264	-3.43			+8.65			+1.00			+0.01			+3.58				- 0.4	264	
265	-3.45			+8.71			+1.00			+0.01			+3.60				+ 0.1	265	
266	-3.80			+9.49			+1.00			+0.01			+3.86				- 0.5	266	
267	-3.85			+9.60			+1.00			0.00			+3.88				+ 2.4	267	
268	-4.28		0.00	+10.55		0.00	+1.00		0.00	0.00		+4.03	+4.19		+5.20	- 0.9	- 7.1	268	
269	-4.32	+0.04	+0.07	+10.63	-0.01	-0.04	+1.00	0.00	-0.01	0.00	-0.19	+4.06	+4.22	-0.21	+5.26	- 2.3	- 9.0	269	
270	-4.32			+10.64			+1.00			0.00			+4.22				- 8.4	270	
271	-2.97		-1.90	+7.62		+0.72	+1.00		+0.15	-0.04		+3.32	+3.18		+3.12	- 13.6	- 3.7	271	
272	-3.38			+8.55			+1.00			-0.04			+3.50				- 6.5	272	

* $\delta a = 0.024$, $\delta b = 0.743$, $w_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Asimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
273	57N	Kistama H.S.	458	A 14 27 12.28 G 14 27 14.56	G 79 45 18.51	A 180 1 54.1 G 80 1 55.6	Pallakōnda E 0 16	- 5.8	- 2.3	273
274	P	Anandalamalai H.S.	923	G 12 55 50.73 A 11 3 52.10 G 11 3 50.00	G 79 23 46.76	A 171 57 36.3 G 171 57 37.6	Pullur E 0 23	- 5.7		274
275	58 E	Yēttimalai S.	617	A 10 39 40.81 G 10 39 39.88	G 77 50 47.10	A 167 34 2.4 G 167 34 1.2	Chennimalai E 0 40	+ 6.2	+ 0.9	275
276	F	Pachapālaiyam s.	970	A 10 56 36.66 G 10 56 35.97	G 77 40 50.63				+ 0.7	276
277	F	Kātpālaiyam s.	878	A 9 32 15.53 G 9 32 13.28	G 77 56 51.22				+ 2.3	277
278	G	Shūlakarai s.	333	A 8 17 1.75 G 8 16 59.44	G 77 42 7.71	A 5 55 25.4 G 5 55 24.1	Kudankulam Obsy. D 0 4	+ 8.9	+ 2.3	278
279	H	Rādhpāram S.	167	A 8 13 57.50 G 8 13 55.39	G 77 38 53.81				+ 2.1	279
280	H	Tanakarakulam S.	176	A 8 13 41.96 G 8 13 39.52	G 77 44 30.98				+ 2.4	280
281	H	Arasākulam S.	55	A 8 12 10.67 G 8 12 8.34	G 77 46 35.58 A 77 26 1.82 G 77 26 3.56				+ 2.3	281
282	H	Vijayāpati S.	90	G 8 11 25.30	A 185 55 18.8 G 185 55 18.6				- 1.7	282
283	H	Nagarkoil Long. S.	110	A 8 10 23.41 G 8 10 21.55	G 77 41 26.26				+ 1.4	283
284	H	Kudankulam Obsy. S.	175	A 8 9 29.92 G 8 9 27.79	G 77 37 35.33	A 38 11 59.1 G 38 12 0.1	Rādhpāram D 0 5		+ 2.1	284
285	H	Punnc Obsy. S.	48	G 9 46 15.13 A 9 31 4.22 G 9 31 1.30	G 78 3 36.52	A 178 0 47.2 G 178 0 50.3	Morur D 1 3	- 4.9		285
286	I	Kanjāmalai H.S.	3236	A 9 28 47.09 G 9 28 44.87	G 78 2 58.77	A 25 17 6.2 G 25 17 6.8	Manikamkota D 0 1		+ 2.9	286
287	K	Manēgandi S.	56	A 9 23 30.55 G 9 23 27.69	G 78 0 37.76				+ 3.6	287
288	K	Black s.	346	G 9 21 51.96	G 78 5 54.11					288
289	K	Kutipārai S.	347	G 9 21 51.96	G 78 49 17.66	A 57 57 54.9 G 57 57 56.2	Koilpati D 0 4	- 7.9	+ 2.2	289
290	K	Pandalagudi s.	217	G 11 57 12.30 A 11 44 43.40 G 11 44 37.64	G 79 33 52.96	A 214 44 19.0 G 214 44 20.0	Uttarakoshamangai E 0 0	- 4.7		290
291	K	Rāmnad S.	48	G 79 42 45.80			Pērumukkal D 0 1		+ 5.8	291
292	M	Kallapat Trestle S.	199	G 10 28 2.31 A 28 22 0.10 G 28 22 11.04	G 79 20 51.19	A 152 57 0.1 G 152 56 57.4				292
293	M	Tiruvēndipuram s.	...	G 79 12 43.59		A 179 40 40.6 G 179 40 42.9	Kachipērumāl E 0 12	+ 13.7		293
294	M	Nayinipiriyan Trestle S.	158	G 80 28 38.33		A 302 56 33.7 G 302 56 30.9	Kakkarakota D 0 4	- 12.4		294
295	N	Pātharankota S.	120	G 80 28 10.95			Rāmnagar D 0 5	+ 5.2	- 10.9	295
296	62 D	Rāmuapur (old) T.S.	541	A 27 59 50.22 G 27 59 55.94	G 80 28 10.95	A 178 58 28.0 G 178 58 20.7			- 5.7	296
297	63 A	Jarūra T.S.	536	A 27 21 8.16 G 27 21 8.09	G 80 29 3.67		Darawal D 0 6	+ 14.1	+ 0.1	297
298	A	Nimkār T.S.	486	A 26 54 22.63 G 26 54 17.85	G 80 39 38.26				+ 4.8	298
299	B	Etora T.S.	429	A 26 15 58.32 G 26 15 52.89	G 80 18 14.46				+ 5.4	299
300	B	Dewarsān T.S.	439							300

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I : $\delta\alpha=1$ km			Case II : $\delta\delta=1$ km			Case III : Latitude $w_0=1''$			Case IV : Azimuth $w_0=1''$			$a=6378200$ metres, $1/e=298.3$						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
273	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	273
	-3.63		-1.90	+9.13		+0.82	+1.00		+0.15	-0.03		+3.65	+3.69		+3.61	-7.5	-6.0		
274			-1.74			+0.84			+0.14			+4.08			+4.31	-7.8		274	
275	-5.27			+12.69			+1.00			0.00			+4.87				-2.8	275	
276	-5.31		+0.04	+12.76		-0.02	+1.00		-0.01	0.00		+4.79	+4.89		+6.19	+3.6	-4.0	276	
277	-5.33			+12.82			+1.00			0.00			+4.90				-4.2	277	
278	-6.05			+14.34			+1.00			-0.01			+5.36				-3.1	278	
279	-6.71		-0.08	+15.70		+0.06	+1.00		+0.01	0.00		+6.34	+5.78		+8.15	+6.2	-3.5	279	
280	-6.74			+15.76			+1.00			0.00			+5.80				-3.7	280	
281	-6.74			+15.76			+1.00			0.00			+5.79				-3.4	281	
282	-6.75			+15.79			+1.00			0.00			+5.80				-3.5	282	
283		+0.13			-0.01			0.00		-0.27				-0.24		-1.5		283	
284	-6.77		-0.06	+15.82		+0.04	+1.00		0.00	0.00		+6.43	+5.81		+8.27	-1.3	-3.9	284	
285	-6.78			+15.84			+1.00			0.00			+5.82				-3.7	285	
286			-0.45			+0.23			+0.03			+4.53			+5.62	-7.1		286	
287			-1.64			+0.98			+0.13			+5.38			+6.20	-20.1		287	
288	-6.06			+14.36			+1.00			-0.01			+5.37				-2.5	288	
289	-6.08		-0.47	+14.40		+0.29	+1.00		+0.04	-0.01		+5.54	+5.38		+6.94	-0.9	-3.2	289	
290	-6.13			+14.49			+1.00			-0.01			+5.41				-2.5	290	
291			-1.57			+0.96			+0.13			+5.61			+6.55	-10.3		291	
292			-2.05			+1.06			+0.16			+4.41			+4.64	-6.9		292	
293	-4.93			+11.96			+1.00			-0.03			+4.61				+1.2	293	
294			-1.94			+1.06			+0.15			+4.73			+5.15	+11.5		294	
295			-1.88			+1.08			+0.15			+5.02			+5.59	-14.8		295	
296	+1.01		-1.53	-3.40		+0.08	+1.00		+0.10	-0.05		+1.92	-1.34		+1.15	+3.6	-9.6	296	
297	+0.94			-3.12			+1.00			-0.05			-1.20				-4.5	297	
298	+0.81		-1.57	-2.62		+0.12	+1.00		+0.11	-0.05		+1.99	-0.95		+1.23	+12.5	+1.1	298	
299	+0.72			-2.28			+1.00			-0.05			-0.78				+5.6	299	
300	+0.57			-1.77			+1.00			-0.04			-0.54				+5.9	300	

* $\delta\alpha = 0.924$, $\delta\delta = 0.743$, $w_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE

Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
801	68 C	Kānākhera T.S.	416	A 25 51 25.97 G 25 51 20.95	G 80 25 31.61				+ 5.0	801
802	C	Pavia H.S.	481	A 25 27 21.18 G 25 27 17.39	G 80 44 12.26				+ 3.8	802
803	D	Potenda S.	993	A 24 37 24.71 G 24 37 23.04	G 80 57 7.18				+ 1.7	803
804	E	Dadaura T.S.	420	A 27 43 3.51 G 27 43 18.33	G 81 42 44.29				- 14.8	804
805	E	Māsi T.S.	406	A 27 38 14.79 G 27 38 25.17	G 81 23 8.97	A 153 5 50.5 G 153 5 52.2	Bela D O 5	- 3.2	- 10.4	805
806	E	Imlia T.S.	428	A 27 19 17.83 G 27 19 18.00	G 81 7 37.37				- 1.1	806
807	F	Utiāmau T.S.	386	A 27 0 1.62 G 26 59 57.08	G 81 12 17.24				+ 4.5	807
808	F	Parewa T.S.	380	A 26 38 11.44 G 26 38 4.00	G 81 12 11.14				+ 7.4	808
809	F	Sora T.S.	400	A 26 17 26.39 G 26 17 18.83	G 81 12 23.12	A 239 43 3.6 G 239 42 55.9	Janki D O 4	+ 15.6	+ 7.6	809
810	G	Pariāou T.S.	346	A 25 50 11.59 G 25 50 5.26	G 81 22 16.31				+ 6.3	810
811	G	Pabhosa H.S.	565	G 25 21 17.32	G 81 19 8.40	A 187 38 4.1 G 187 38 4.7	Karra D O 14	- 1.3		811
812	H	Karāra H.S.	1966	A 24 4 42.20 G 24 4 42.01	G 81 15 47.29	A 269 18 28.7 G 269 18 34.9	Murwās D O 16	- 13.9	+ 0.2	812
813	I	Manichauk T.S.	360	A 27 36 28.91 G 27 36 48.14	G 82 5 3.16				- 19.2	813
814	I	Pathārdi T.S.	320	A 27 25 56.11 G 27 26 14.77	G 82 45 2.97				- 18.7	814
815	I	Bāsadela T.S.	377	A 27 23 50.71 G 27 24 3.24	G 82 16 50.44	A 106 15 8.7 G 106 15 7.9	Saibara D O 8	+ 1.5	- 12.5	815
816	J	Orejhar S.	392	G 26 46 55.54	G 82 12 7.60	A 308 36 18.9 G 308 36 17.7	Bisaul D O 7	+ 2.4		816
817	J	Fyzabad Long. S.	...	G 26 46 40.66	A 82 8 7.60 G 82 8 8.15			- 0.5		817
818	J	Bisaul T.S.	342	G 26 40 37.38	G 82 20 54.43	A 128 40 15.9 G 128 40 14.8	Orejhar D O 1	+ 2.2		818
819	K	Murār T.S.	370	G 25 41 17.20	G 82 14 19.00	A 42 20 13.2 G 42 20 13.1	Buria D O 7	+ 0.2		819
820	L	Gurwāni H.S.	2083	A 24 1 28.93 G 24 1 25.71	G 82 17 28.34	A 210 29 53.8 G 210 29 49.4	Pokra D O 5	+ 9.9	+ 3.2	820
821	M	Ghaus T.S.	296	A 27 20 48.34 G 27 21 5.08	G 83 5 38.81				- 16.7	821
822	N	Rājabāri T.S.	296	G 26 54 3.04	G 83 15 35.49	A 104 47 9.8 G 104 47 9.9	Nandaur D O 5	- 0.2		822
823	N	Samenda T.S.	285	G 26 0 23.97	G 83 13 30.67	A 304 8 50.2 G 304 8 48.7	Ohit Bisram D O 6	+ 3.1		823
824	O	Hirdepur T.S.	289	G 25 24 23.05	G 83 14 15.46	A 304 4 33.1 G 304 4 33.8	Barhāni D O 6	- 1.5		824
825	F	Gora H.S.	1828	G 24 4 55.71	G 83 14 13.47	A 282 48 23.9 G 282 48 27.5	Sewādhi D O 10	- 8.1		825
826	64 A	Amūa H.S.	2113	A 23 59 57.02 G 23 59 56.24	G 80 29 17.26	A 260 4 21.4 G 260 4 19.0	Lakanpura D O 19	+ 5.4	+ 0.8	826
827	A	Lora H.S.	1923	A 23 29 46.30 G 23 29 41.53	G 80 9 56.85				+ 4.8	827
828	B	Sarandi Pat H.S.	1627	G 22 13 18.98	G 80 3 5.98	A 159 45 20.8 G 159 45 20.5	Tālla E O 6	+ 0.7		828

*A - Astronomical Value.

G - Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
801	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	801
801	+0.47			-1.44			+1.00			-0.05			-0.38				+5.4		801
802	+0.38			-1.11			+1.00			-0.05			-0.24				+4.0		802
803	+0.16			-0.43			+1.00			-0.05			+0.08				+1.6		803
804	+0.90			-2.91			+1.00			-0.07			-1.11				-13.7		804
805	+0.88	-2.06		-2.85		+0.14	+1.00		+0.14	-0.06		+1.96	-1.07		+1.78	-5.4	-9.3		805
806	+0.81			-2.60			+1.00			-0.06			-0.95				-0.1		806
807	+0.75			-2.35			+1.00			-0.06			-0.82				+5.3		807
808	+0.66			-2.06			+1.00			-0.06			-0.68				+8.1		808
809	+0.59	-2.02		-1.79		+0.19	+1.00		+0.14	-0.06		+2.06	-0.55		+0.98	+14.4	+8.2		809
810	+0.48			-1.43			+1.00			-0.06			-0.38				+6.7		810
811		-2.13				+0.25			+0.15			+2.13			+1.01	-2.5			811
812	+0.02	-2.18		+0.03		+0.32	+1.00		+0.15	-0.06		+2.23	+0.27		+1.16	-15.1	-0.1		812
813	+0.89			-2.83			+1.00			-0.07			-1.06				-18.1		813
814	+0.87			-2.69			+1.00			-0.08			-1.00				-17.7		814
815	+0.89	-2.57		-2.67		+0.19	+1.00		+0.18	-0.07		+1.98	-0.95		+0.37	+0.9	-11.5		815
816		-2.58				+0.22			+0.18			+2.02			+0.44	+1.7			816
817		-2.69				+0.38			+0.04			+0.05			-2.13	+1.6			817
818		-2.65				+0.23			+0.18			+2.03			+0.40	+1.6			818
819		-2.65				+0.29			+0.19			+2.10			+0.53	-0.5			819
820	+0.02	-2.80		+0.07		+0.41	+1.00		+0.20	-0.08		+2.24	+0.29		+0.66	+9.2	+2.9		820
821	+0.85			-2.63			+1.00			-0.09			-0.97				-15.7		821
822		-3.14				+0.26			+0.22			+2.01			-0.06	-0.3			822
823		-3.19				+0.32			+0.22			+2.07			+0.03	+3.1			823
824		-3.24				+0.37			+0.23			+2.12			+0.08	-1.6			824
825		-3.36				+0.49			+0.24			+2.23			+0.21	-8.3			825
826	-0.01	-1.71		+0.10		+0.25	+1.00		+0.12	-0.05		+2.24	+0.31		+1.54	+3.9	+0.5		826
827	-0.17			+0.53			+1.00			-0.04			+0.50				+4.3		827
828		-1.53				+0.30			+0.11			+2.41			+1.96	-1.1			828

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observation†	Meridian Deflection†	
329	64 B	Sarey Khan Lat. S.	1409	A 22 12 50.66 G 22 12 55.61	G 80 2 49.79				- 5.0	329
330	C	Lingmāra H.S.	1400	A 21 42 55.36 G 21 43 3.07	G 80 7 36.30				- 7.7	330
331	C	Sitāpār H.S.	1237	A 21 24 43.83 G 21 24 50.54	G 80 19 26.36				- 6.7	331
332	J	Dalea H.S.	1622	A 22 19 30.25 G 22 19 33.62	G 82 1 31.25				- 3.4	332
333	K	Pathāidi T.S.	879	A 21 48 43.06 G 21 48 45.96	G 82 16 46.96	A 198 23 42.8 G 198 23 43.3	Konārgarh D 0 5	- 1.2	- 2.9	333
334	L	Ramai H.S.	1313	A 20 56 50.31 G 20 56 51.47	G 82 8 18.55	A 223 15 23.2 G 223 15 23.8	Khalāri E 0 9	- 1.6	- 1.2	334
335	P	Sindur H.S.	2918	G 20 15 33.64	G 83 39 42.83	A 201 20 4.4 G 201 20 10.3	Lakh Parbat D 0 85	- 16.0		335
336	65 C	Singāwāram H.S.	714	A 17 45 8.71 G 17 45 10.38	G 80 56 9.04	A 249 3 4.9 G 249 3 6.5	Nārākonda E 0 29	- 5.0	- 1.7	336
337	D	Dhūlipalla S.	245	A 16 25 53.47 G 16 25 56.75	G 80 5 29.59	A 125 53 37.6 G 125 53 39.9	Kachalboru E 0 47	- 7.8	- 3.3	337
338	G	Parampudi H.S.	684	A 17 12 32.63 G 17 12 38.28	G 81 12 10.06	A 114 12 9.2 G 114 12 13.2	Nāgaldurgam E 0 15	- 12.9	- 5.7	338
339	I	Hātbenā H.S.	2600	A 19 51 42.60 G 19 51 42.34	G 82 1 25.96				+ 0.3	339
340	I	Karīa H.S.	2014	A 19 12 2.67 G 19 12 5.98	G 82 7 7.97	A 201 43 17.4 G 201 43 17.9	Motiaon E 0 4	- 1.4	- 3.3	340
341	K	Kālingkonda H.S.	4634	G 17 49 42.44	G 82 18 40.67	A 189 41 25.0 G 189 41 24.8	Kaurālbiding D 0 25	+ 0.6		341
342	K	Sānjib H.S.	2142	A 17 31 12.32 G 17 31 18.68	G 82 41 24.30	A 135 38 16.0 G 135 38 15.9	Dhār E 0 65	+ 0.3	- 6.4	342
343	N	Rāwal H.S.	874	A 18 32 4.73 G 18 32 9.22	G 83 33 11.63	A 317 29 5.0 G 317 29 5.0	Piudi D 0 11	0.0	- 4.5	343
344	N	Vizagapatām Base-line N. End S.	181	A 18 0 56.66 G 18 1 2.93	G 83 13 43.36	A 203 44 24.5 G 203 44 24.5	Bor E 0 12	0.0	- 6.3	344
345	O	Waltair Long. S.	200	A 17 43 20.44 G 17 43 29.31	A 83 19 0.17 G 83 19 3.52			- 3.2	- 8.9	345
346	66 A	Ongole H.S.	250	A 15 29 52.87 G 15 29 56.85	G 80 2 27.72				- 4.0	346
347	B	Gudali H.S.	292	A 14 1 10.65 G 14 1 9.45	G 80 1 13.36				+ 1.2	347
348	C	Madras Observatory Long. S.	54	A 13 4 8.97 G 13 4 4.17	A 80 14 47.06 G 80 14 54.33			- 7.1	+ 4.8	348
349	C	St. Thomas's Mount Trestle S.	250	A 13 0 20.64 G 13 0 14.79	G 80 11 41.38	A 12 30 5.3 G 12 30 6.2	Nannangalam D 0 7	- 3.9	+ 5.9	349
350	D	Injambākam H.S.	29	G 12 54 51.18	G 80 15 11.23	A 99 4 39.1 G 99 4 40.6	Nannangalam E 0 28	- 6.5		350
351	72 B	Nannangarhi T.S.	344	G 26 59 10.19	G 84 23 46.86	A 107 52 43.1 G 107 52 47.8	Bakwa D 0 7	- 9.2		351
352	B	Jalāipur T.S.	232	A 26 3 45.56 G 26 3 39.42	G 84 23 9.46	A 111 52 41.5 G 111 52 40.0	Katwārpur D 0 3	+ 3.1	+ 6.1	352
353	C	Nūāon T.S.	251	A 25 34 45.64 G 25 34 37.91	G 84 14 15.86				+ 7.7	353
354	O	Medinipur T.S.	335	A 25 5 22.35 G 25 5 14.02	G 84 22 6.95	A 215 46 30.0 G 215 46 33.5	Bisunpur D 0 6	- 7.5	+ 8.3	354
355	D	Teona H.S.	740	A 24 34 49.76 G 24 34 38.94	G 84 10 26.42				+ 10.8	355
356	D	Hurilāong H.S.	1378	A 24 2 16.74 G 24 2 5.99	G 84 21 50.58	A 128 18 18.3 G 128 18 24.0	Khaira Pāndu D 0 4	- 12.8	+ 10.8	356

* A = Astronomical Value.
G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
829	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	829	
830	-0.56			+1.65			+1.00			-0.04			+0.97				- 6.0	830	
831	-0.72			+2.10			+1.00			-0.04			+1.15				- 8.9	831	
832	-0.83			+2.37			+1.00			-0.04			+1.25				- 8.0	832	
833	-0.50			+1.55			+1.00			-0.07			+0.91				- 4.3	833	
834	-0.60		-2.98	+2.01		+0.60	+1.00		+0.22	-0.07		+2.45	+1.10		+0.92	- 1.9	- 4.0	834	
835	-0.95		-2.98	+2.79		+0.67	+1.00		+0.22	-0.07		+2.55	+1.41		+1.10	- 2.4	- 2.6	835	
836			-4.09			+1.00			+0.30			+2.62			+0.44	- 15.9		836	
837	-2.19		-2.48	+5.81		+0.79	+1.00		+0.19	-0.05		+2.99	+2.54		+2.21	- 6.3	- 4.2	837	
838	-2.75		-1.97	+7.12		+0.71	+1.00		+0.15	-0.04		+3.22	+3.01		+2.92	- 9.3	- 6.3	838	
839	-2.40		-2.75	+6.34		+0.92	+1.00		+0.21	-0.06		+3.08	+2.73		+2.19	- 13.8	- 8.4	839	
840	-1.35			+3.79			+1.00			-0.07			+1.80				- 1.5	840	
841	-1.59		-3.17	+4.42		+0.87	+1.00		+0.24	-0.07		+2.77	+2.03		+1.36	- 2.4	- 5.3	841	
842			-3.51			+1.11			+0.27			+2.97			+1.50	0.0		842	
843	-2.25		-3.85	+6.03		+1.25	+1.00		+0.29	-0.08		+3.02	+2.61		+1.36	- 0.2	- 9.0	843	
844	-1.82		-4.31	+5.04		+1.27	+1.00		+0.33	-0.09		+2.86	+2.26		+0.75	+ 0.1	- 6.8	844	
845	-2.03		-4.16	+5.54		+1.29	+1.00		+0.31	-0.09		+2.94	+2.44		+1.00	- 0.1	- 8.7	845	
846	-2.15	-3.29		+5.83	+0.46		+1.00	+0.03		-0.09	-0.11		+2.53	-2.83		- 0.4	- 11.4	846	
847	-3.16			+8.06			+1.00			-0.04			+3.33				- 7.3	847	
848	-3.83			+9.57			+1.00			-0.04			+3.83				- 2.6	848	
849	-4.28	-1.54		+10.56	+0.20		+1.00	+0.01		-0.04	-0.19		+4.16	-1.52		- 5.6	+ 0.6	849	
850	-4.31		-2.52	+10.63		+1.20	+1.00		+0.20	-0.04		+4.05	+4.18		+3.85	- 5.6	+ 1.7	850	
851			-2.60			+1.25			+0.21			+4.08			+3.86	- 8.2		851	
852			-3.77			+0.30			+0.26			+2.00			-0.60	- 8.8		852	
853	+0.61		-3.84	-1.62		+0.38	+0.99		+0.27	-0.11		+2.06	-0.41		-0.53	+ 3.6	+ 6.5	853	
854	+0.48			-1.22			+0.99			-0.11			-0.30				+ 8.0	854	
855	+0.36		-3.93	-0.82		+0.49	+0.99		+0.28	-0.11		+2.14	-0.11		-0.42		+ 8.4	855	
856	+0.22			-0.40			+0.99			-0.10			+0.09				+ 10.7	856	
857	+0.08		-4.04	+0.06		+0.58	+0.99		+0.29	-0.11		+2.22	+0.28		-0.34	- 12.3	+ 10.5	857	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
857	72 E	Kaulia H.S.	7051	A 27 48 25.5 G 27 48 58.6	G 85 14 20.7				- 33.1	857
858	E	Mahadeo Pokra H.S.	7095	A 27 40 53.6 G 27 41 31.5	G 85 31 19.9				- 37.9	858
859	F	Pota T.S.	222	G 26 22 40.13	G 85 26 20.33	A 180 4 5.0 G 180 4 8.3	Madanpur D O 4	- 6.7		859
860	F	Fahladpur T.S.	175	A 26 4 27.24 G 26 4 21.01	G 85 27 13.16				+ 6.2	860
861	G	Dābauli T.S.	189	A 25 40 22.99 G 25 40 16.23	G 85 20 16.53				+ 6.8	861
862	G	Bihar H.S.	391	A 25 12 39.27 G 25 12 26.05	G 85 30 31.33				+ 13.2	862
863	H	Mahar H.S.	1606	A 24 44 31.12 G 24 44 20.88	G 85 9 55.13				+ 10.2	863
864	K	Biohwi H.S.	321	G 25 9 49.15	G 86 8 3.99	A 357 49 29.7 G 357 40 32.4	Bkgora E O 40	- 5.7		864
865	N	Chūni T.S.	197	G 26 11 4.72	G 87 2 52.80	A 185 49 39.4 G 185 49 49.0	Minai D O 4	- 19.5		865
866	O	Sirkanda T.S.	132	G 25 27 47.43	G 87 8 23.57	A 145 34 17.2 G 145 34 21.8	Pureni D O 4	- 9.7		866
867	73 A	Bulbul H.S.	3352	A 23 37 53.44 G 23 37 44.63	G 84 26 13.94				+ 8.8	867
868	A	Mahwari H.S.	3153	A 23 26 9.28 G 23 26 4.96	G 84 54 1.75				+ 4.3	868
869	A	Dhursu H.S.	2271	G 23 15 57.13	G 84 44 19.21	A 149 58 57.7 G 149 59 0.8	Bagru E O 38	- 7.2		869
870	C	Andhari H.S.	1442	G 21 57 39.56 A 23 57 16.82 G 23 57 13.75	G 84 14 57.17	A 17 40 39.6 G 17 40 43.1	Garpati D O 1	- 8.7		870
871	E	Ohendwār (old) H.S.	2817	G 23 57 13.75 A 20 28 52.05 G 20 29 0.68	G 85 26 9.29	A 92 35 20.3 G 92 35 20.5	Kasātu D O 16	- 0.5	+ 3.1	871
872	H	Outtack H.S.	133	A 20 28 52.05 G 20 29 0.68	G 85 52 1.43	A 155 35 54.6 G 155 35 54.3	Kaplās E 1 21	+ 0.8	- 8.6	872
873	I	Pārasnāth H.S.	4481	G 23 57 34.89	G 86 8 10.86	A 145 7 21.0 G 145 7 22.9	Bāmāni D 1 2	- 4.3		873
874	I	Tilabani H.S.	1329	G 23 24 59.87	G 86 33 14.64	A 272 58 23.5 G 272 58 23.1	Sūsiniā D O 8	+ 0.9		874
875	M	Malūncha H.S.	970	A 23 54 29.64 G 23 54 29.02	G 87 5 41.86	A 74 46 32.3 G 74 46 35.7	Durgapur D O 1	- 7.7	+ 0.6	875
876	M	Madhpur T.S.	180	G 23 9 53.06	G 87 44 37.29	A 206 49 9.1 G 206 49 5.5	Farhat D O 8	+ 8.4		876
877	N	Kalsibhānga T.S.	303	G 22 20 23.80 A 21 47 28.82 G 21 47 27.95	G 87 8 19.19	A 115 7 20.2 G 115 7 17.7	Kalābani D O 1	+ 6.1		877
878	O	Dariāpur T.S.	63	G 21 47 27.95 A 21 47 17.28 G 21 47 20.83	G 87 52 3.32	A 207 38 56.0 G 207 38 58.5	Dāntūn D O 8	- 6.3	- 3.6	878
879	O	Patna T.S.	80	A 21 47 17.28 G 21 47 20.83	G 87 11 45.53	A 96 49 54.7 G 96 49 55.1	Nilgiri E O 59	- 1.0	- 3.0	879
880	O	Ohandipur T.S.	53	A 21 26 34.03 G 21 26 36.99	G 87 2 3.66	A 196 41 21.2 G 196 41 22.9	Ohiklikhāi D O 27	- 4.7	- 5.9	880
881	74 A	Khundābolo H.S.	3115	A 19 51 7.03 G 19 51 12.90	G 84 58 17.43	A 146 26 29.1 G 146 26 33.0	Thaladi D 1 21	- 11.4		881
882	B	Deodonger H.S.	4534	G 18 54 32.37 A 18 47 6.75 G 18 47 16.07	G 84 3 35.84				- 10.2	882
883	B	Mal H.S.	483	G 18 47 6.75 G 18 47 16.07	G 84 30 44.31				- 36.6	883
884	74 A	Phallut h.s.	11815	A 27 12 4.30 G 27 12 40.86	G 88 1 0.96					884

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	v cos λ	w cot λ	u	v cos λ	w cot λ	u	v cos λ	w cot λ	u	v cos λ	w cot λ	u	v cos λ	w cot λ	Deflection in Prime Vertical	Deflection in Meridian		
357	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	357	
	+1.01			-2.99			+0.99			-0.12			-1.13				-32.0	358	
358	+1.00			-2.90			+0.99			-0.13			-1.08				-36.8	359	
359			-4.41			+0.40			+0.30			+2.04			-1.05	-5.6		360	
360	+0.64			-1.63			+0.99			-0.12			-0.47				+6.7	361	
361	+0.54			-1.30			+0.99			-0.12			-0.32				+7.1	362	
362	+0.43			-0.92			+0.99			-0.13			-0.14				+13.3	363	
363	+0.29			-0.54			+0.99			-0.12			+0.03				+10.2	364	
364			-4.94			+0.58			+0.35			+2.12			-1.28	-4.2		365	
365			-5.33			+0.50			+0.37			+2.04			-1.82	-17.9		366	
366			-5.48			+0.60			+0.38			+2.10			-1.79	-7.9		367	
367	-0.04			+0.41			+0.99			-0.11			+0.43				+8.4	368	
368	-0.08			+0.57			+0.99			-0.12			+0.50				+3.8	369	
369			-4.36			+0.71			+0.31			+2.29			-0.43	-6.6		370	
370			-4.23			+0.83			+0.31			+2.43			-0.06	-8.4		371	
371	+0.09		-4.69	+0.12		+0.68	+0.99		+0.33	-0.12		+2.23	+0.32		-0.85	+0.6	+2.8	372	
372	-1.02		-5.54	+3.20		+1.30	+0.99		+0.41	-0.13		+2.58	+1.58		-0.69	+2.0	-10.2	373	
373			-5.10			+0.74			+0.37			+2.23			-1.19	-2.9		374	
374			-5.44			+0.86			+0.39			+2.27			-1.35	+2.5		375	
375	+0.13		-5.69	+0.15		+0.83	+0.99		+0.41	-0.15		+2.22	+0.35		-1.63	-5.9	+0.2	376	
376			-6.21			+1.02			+0.44			+2.28			-1.88	+10.5		377	
377			-5.99			+1.11			+0.43			+2.37			-1.53	+8.1		378	
378	-0.49			+2.00			+0.98			-0.16			+1.13				-0.2	379	
379	-0.52		-6.13	+2.00		+1.22	+0.99		+0.45	-0.15		+2.43	+1.12		-1.49	-4.3	-4.7	380	
380	-0.64		-6.10	+2.32		+1.27	+0.99		+0.45	-0.15		+2.47	+1.24		-1.37	+0.9	-4.2	381	
381	-1.28		-5.05	+3.79		+1.29	+0.99		+0.37	-0.12		+2.66	+1.79		-0.16	-3.9	-7.7	382	
382			-4.60			+1.30			+0.34			+2.80			+0.44	-10.9		383	
383	-1.70			+4.80			+0.99			-0.11			+2.16				-12.4	384	
384	+1.00			-2.53			+0.98			-0.17			-0.86				-35.7	384	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Azimuth*	Name and angular Elevation or Depression of observed station	(A - G) cot λ for azimuth or (A - G) cos λ for longitude observations†	Meridian Deflec- tion†	
385	78 A	Tonglu h.s.	10073	A 27 1 11'30 G 27 1 53'54	G 88 5 2'93				-42'2	385
386	B	Senchal h.s.	8600	A 26 58 33'01 G 26 59 8'25	G 88 17 44'78				-35'2	386
387	B	Kurseong h.s.	4428	A 26 51 15'05 G 26 52 5'56	G 88 15 54'68				-50'5	387
388	B	Siliguri s.	401	A 26 41 18'10 G 26 41 40'37	G 88 24 49'54				-22'3	388
389	R	Jalpaiguri Long. s.	280	A 26 31 11'44 G 26 31 17'30	A 88 43 52'42 G 88 44 12'77	A 321 33 25'3 G 321 33 33'0	Dharampur D O 2	- 15'4	- 6'0	389
390	B	Rāmganj T.S.	249	G 26 18 55'51	G 88 17 30'43	A 218 51 56'2 G 218 52 8'5	Kanchābāri D O 2	- 24'9		390
391	H	Lohārgara T.S.	205	A 26 2 14'17 G 26 2 12'02	G 88 21 56'60				+ 2'2	391
392	O	Chanduria T.S.	160	A 25 44 31'93 G 25 44 27'47	G 88 22 17'15				+ 4'5	392
393	D	Charaldānga T.S.	149	A 24 52 45'36 G 24 52 43'95	G 88 23 4'21				+ 1'4	393
394	F	Ataro Bānki T.S.	133	G 26 4 50'62	G 89 28 3'10	A 70 52 20'4 G 70 52 32'5	Ohandrapur D O 3	- 24'7		394
395	G	Alaṇājāni T.S.	143	G 25 59 6'81	G 89 45 41'19	A 293 0 46'2 G 293 0 57'3	Sāmding E O 1	- 22'8		395
396	G	Halkāchar T.S.	103	G 25 9 55'94	G 89 42 48'42	A 145 54 38'0 G 145 54 48'9	Kānchāpāra D O 5	- 23'2		396
397	H	Aloākāndi T.S.	88	G 24 45 29'80	G 89 38 42'19	A 205 17 22'4 G 205 17 28'5	Gaborgram D O 4	- 13'2		397
398	J	Itakueni H.S.	803	G 26 8 11'37	G 90 39 47'24	A 136 38 12'9 G 136 38 24'2	Bhairaber Chura E O 34	- 23'0		398
399	O	Rangsanobo H.S.	4455	G 25 15 19'60 A 23 56 42'82 G 23 56 38'97	G 91 43 20'86	A 125 49 11'9 G 125 49 18'5	Mosingi E I 21	- 14'0		399
400	79 A	Madhupur T.S.	92	G 23 56 38'97	G 88 29 7'66	A 172 57 25'5 G 172 57 31'7	Imāmnagar D O 3	- 14'0	+ 3'9	400
401	A	Anandbās T.S.	67	G 23 21 19'24	G 88 22 40'36	A 6 58 55'2 G 6 58 58'6	Jeodhāra D O 3	- 7'9		401
402	B	Aknāpur T.S.	98	G 22 54 22'85	G 88 3 6'66	A 147 41 14'5 G 147 41 15'7	Hākistāpur D O 4	- 2'8		402
403	B	Calcutta Base-line S. End T.S.	13	G 22 36 55'68	G 88 22 54'43	A 177 10 27'3 G 177 10 30'3	Calcutta Base-line N. End D O 2	- 7'2		403
404	B	Calcutta Long. s.	18	A 22 32 55'58 G 22 32 54'67	A 88 21 17'84 G 88 21 29'10			- 10'4	+ 0'9	404
405	K	Tepri T.S.	67	G 23 57 24'45	G 89 53 11'99	A 156 35 52'8 G 156 35 55'7	Bangaon D O 4	- 6'5		405
406	E	Daulatpur T.S.	60	G 23 8 43'76	G 89 42 57'76	A 202 38 51'3 G 202 38 49'3	Maheshpur D O 5	+ 4'7		406
407	I	Lakhinagar T.S.	51	G 23 0 39'73	G 90 45 43'08	A 85 27 44'7 G 85 27 39'6	Kānchāpur D O 4	+ 12'0		407
408	J	Gangapur T.S.	54	G 22 59 34'77	G 90 27 28'63	A 151 19 48'9 G 151 19 49'0	Malgaon D O 5	- 0'2		408
409	M	Dawa H.S.	205	G 23 45 17'63	G 91 20 16'63	A 173 18 51'3 G 173 18 53'0	Lambusara D O 5	- 3'9		409
410	N	Semu Tān H.S.	226	G 22 48 38'48 A 22 22 57'08 G 22 22 56'40	G 91 47 38'49	A 272 20 55'4 G 272 20 57'1	Bhatti Moin E I 22	- 4'0		410
411	N	Nagarkhāna H.S.	290	G 22 22 56'40	G 91 48 30'42	A 155 47 13'3 G 155 47 16'6	Ohandranath E O 22	- 8'0	+ 0'7	411
412	N	Chittagong Long. S.	...	G 22 20 18'43	A 91 50 4'91 G 91 50 16'68			- 10'9		412

* A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
885	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	- 41.4	385
886	+0.96			-2.40			+0.98			-0.17			-0.79					- 34.4	386
387	+0.96			-2.36			+0.98			-0.17			-0.78					- 49.8	387
388	+0.94			-2.27			+0.98			-0.17			-0.74					- 21.6	388
388	+0.90			-2.13			+0.98			-0.17			-0.67					- 5.4	389
389	+0.88	-6.63	-6.23	-2.00	+0.93	+0.54	+0.98	+0.09	+0.43	-0.18	+0.03	+2.01	-0.60	-5.37	-2.64	- 13.0			390
890			-6.02			+0.55			+0.42			+2.02			-2.41	- 22.7			391
391	+0.75			-1.61			+0.98			-0.17			-0.43				+ 2.6		392
392	+0.68			-1.37			+0.98			-0.17			-0.32				+ 4.8		393
393	+0.46			-0.67			+0.98			-0.17			0.00				+ 1.4		394
394			-6.71			+0.64			+0.47			+2.03			-2.97	- 21.9			395
395			-6.89			+0.67			+0.48			+2.04			-3.09	- 19.7			396
396			-7.00			+0.81			+0.49			+2.10			-3.01	- 20.2			397
397			-7.04			+0.88			+0.50			+2.13			-2.95	- 10.2			398
398			-7.37			+0.69			+0.51			+2.02			-3.54	- 19.5			399
399			-8.13			+0.91			+0.57			+2.07			-4.00	- 10.0			400
400	+0.20		-6.51	+0.12		+0.94	+0.98		+0.46	-0.17		+2.21	+0.36		-2.32	- 11.7	+ 3.5		401
401			-6.56			+1.04			+0.47			+2.26			-2.22	- 5.5			402
402			-6.44			+1.10			+0.46			+2.31			-2.02	- 0.6			403
403			-6.71			+1.19			+0.49			+2.33			-2.16	- 4.8			404
404	-0.23	-6.41		+1.32	+0.90		+0.98	+0.07		-0.17	-0.03		+0.85	-5.27		- 5.1	0.0		405
405			-7.40			+1.05			+0.52			+2.20			-3.06	- 3.2			406
406			-7.51			+1.23			+0.54			+2.27			-2.93	+ 7.8			407
407			-8.08			+1.34			+0.58			+2.28			-3.36	+ 15.4			408
408			-7.90			+1.32			+0.57			+2.28			-3.20	+ 3.0			409
409			-8.26			+1.19			+0.59			+2.20			-3.73	- 0.2			410
410			-8.86			+1.28			+0.63			+2.29			-4.08	+ 0.1			411
411	-0.10		-9.02	+1.45		+1.32	+0.97		+0.64	-0.22		+2.33	+1.00		-4.15	- 3.8	- 0.3		412
412		-8.49			+1.20			+0.09		-0.04				-6.97		- 3.9			

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. *Vide p. 2.*

TABLE
Deflections of the Plumb-line

Serial No.	Sheet No.	Observed at	EVEREST'S SPHEROID.							Serial No.
			Height in feet	Latitude*	Longitude*	Asimuth*	Name and angular Elevation or Depression of observed station	(A-G) cot λ for azimuth or (A-G) cos λ for longitude observations†	Meridian Deflec- tion†	
413	83 H	Loijing H.S.	6610	"	"	"	Pangkibot D 0 48	- 8.7	"	413
414	K	Thyolloching H.S.	6566	G 24 44 28.17	G 93 43 44.79	A 120 58 12.0 G 120 58 16.0	Sirohifurar E 1 15	- 7.3		414
415	L	Tamunja H.S.	3387	G 25 0 5.76	G 94 43 46.98	A 113 3 3.9 G 113 3 7.3	Khambiching E 1 84	- 0.4		415
416	P	Seikpa H.S.	3857	G 24 39 9.33	G 94 36 48.01	A 116 36 26.7 G 116 36 26.0	Madhun E 0 17	- 27.3		416
417	P	Thonbinzin H.S.	1932	G 24 35 37.84	G 95 45 33.72	A 258 32 25.1 G 258 32 37.6	Katha E 0 37	- 18.9		417
418	84 C	Pi Tan H.S.	563	G 24 14 3.03	G 95 58 8.16	A 277 46 13.1 G 277 46 21.6	Luraintong E 1 84	- 16.2		418
419	D	Akyab Long. S.	20	G 21 49 20.78 A 20 8 14.87 G 20 8 13.10	G 92 8 16.02 A 92 53 38.63 G 92 53 49.63	A 256 23 22.7 G 256 23 20.2		- 10.3	+ 1.8	419
420	H	Yaponetaung H.S.	2819	A 20 14 51.83 G 20 14 55.91	G 93 41 49.34	A 74 17 19.6 G 74 17 23.3	Rongdong D 1 1	- 10.0	- 4.1	420
421	H	Dattaung H.S.	455	G 20 13 14.44	G 93 1 9.09	A 171 27 28.3 G 171 27 31.0	Bengara E 0 37	- 7.3		421
422	M	Ubyetaung H.S.	2766	G 23 40 52.06	G 95 57 42.75	A 303 38 45.7 G 303 38 50.1	Tagaungtaung D 0 17	- 10.0		422
423	M	Male H.S.	848	G 23 2 53.30	G 95 57 18.09	A 316 31 54.4 G 316 32 0.8	Wapyadaung E 1 12	- 15.0		423
424	N	Sheinmaga H.S.	456	G 22 16 33.89	G 95 58 15.79	A 354 23 23.8 G 354 23 31.5	Mingun E 0 31	- 18.8		424
425	N	Mingun H.S.	1343	G 22 3 0.71	G 95 59 41.41	A 174 23 57.7 G 174 24 5.6	Sheinmaga D 0 48	- 19.5		425
426	P	Taungpila H.S.	1012	G 20 41 52.71	G 95 53 4.50	A 240 23 15.5 G 240 23 17.0	Yuba E 0 50	- 4.0		426
427	85 E	Betkamauk H.S.	1582	A 19 47 37.32 G 19 47 38.55	G 93 28 13.32	A 229 44 59.4 G 229 44 59.2	Ingrauntaung D 0 14	+ 0.6	- 1.2	427
428	N	Kyaunggyi S.	...	G 18 49 20.95	G 95 12 55.40	A 109 26 42.1 G 109 26 46.0	Prome E 3 45	- 11.4		428
429	N	Prome Long. S.	100	A 18 49 18.62 G 18 49 14.28	A 95 12 42.20 G 95 12 57.44			- 14.4	+ 4.3	429
430	92 G	Kumon Bum H.S.	7970	G 25 38 13.48	G 97 3 34.06	A 308 54 46.0 G 308 54 46.8	Maran Bum D 2 7	- 14.2		430
431	H	Kumtum Bum H.S.	1833	G 24 46 44.32	G 97 9 17.40	A 210 24 26.5 G 210 24 32.9	Maran Bum D 0 17	- 13.9		431
432	93 A	Sinpitaung H.S.	2649	G 23 29 48.36	G 96 45 45.79	A 162 2 20.3 G 162 2 25.8	Taungkalat D 0 32	- 12.7		432
433	E	Loi Hpa Lang H.S.	3591	G 23 14 13.07	G 97 37 24.38	A 188 35 42.8 G 188 35 49.4	Loi Song E 1 11	- 15.4		433
434	J	Loi Hpattan H.S.	6419	G 22 55 35.51	G 98 0 52.58	A 210 28 36.0 G 210 28 43.6	Loi Taow E 0 18	- 18.0		434
435	O	Loi Kiipma H.S.	3792	G 22 1 56.87	G 98 4 30.30	A 154 30 44.7 G 154 30 53.0	Loi Hsimu E 0 2	- 20.5		435
436	94 A	Loi Hsam Hsum H.S.	6472	G 21 41 45.04	G 99 53 57.51	A 115 53 15.6 G 115 53 22.4	R. M.	- 17.1		436
437	B	Letpataung H.S.	3975	G 19 34 7.27	G 96 28 38.76	A 174 46 24.5 G 174 46 31.5	Byingye E 0 36	- 19.7		437
438	B	Toungoo S.	186	G 18 56 1.54	G 96 25 55.21	A 30 46 36.9 G 30 46 42.5	Bhondan E 0 49	- 16.3		438
439	B	Myayabeingkyo H.S.	1411	G 18 21 33.93 A 16 31 39.6 G 16 31 33.08	G 96 22 53.46	A 169 33 42.4 G 169 33 43.8	Khengdan E 0 32	- 4.2		439
440	H	Maraban h.s.	273		G 97 36 59.57				+ 6.5	440

*A = Astronomical Value.

G = Triangulated or Geodetic Value.

† Minus sign indicates Easterly or Northerly Deflection of Plumb-line.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres. $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
413	"	"	-9'40	"	"	+1'14	"	"	+0'66	"	"	+2'09	"	"	-4'93	-4'0		418	
414			-9'91			+1'14			+0'70			+2'06			-5'43	-2'1		414	
415			-9'93			+1'23			+0'70			+2'09			-5'34	+4'7		415	
416			-10'00			+1'32			+0'75			+2'08			-5'90	-21'6		416	
417			-10'82			+1'45			+0'76			+2'11			-5'96	-13'1		417	
418			-9'48			+1'33			+0'67			+2'37			-4'49	-11'9		418	
419	-0'79	-9'11		+3'48	+1'29		+0'97	+0'09		-0'24	-0'08		+1'85	-7'52		-2'8	-0'1	419	
420	-0'69		-11'20	+3'37		+1'70	+0'96		+0'80	-0'25		+2'54	+1'83		-5'57	-5'2	-5'9	420	
421			-10'77			+1'59			+0'77			+2'55			-5'24	-2'9		421	
422			-11'01			+1'56			+0'78			+2'16			-5'99	-4'2		422	
423			-11'28			+1'65			+0'80			+2'21			-6'09	-9'1		423	
424			-11'63			+1'75			+0'83			+2'29			-6'24	-13'1		424	
425			-11'75			+1'78			+0'84			+2'31			-6'31	-13'7		425	
426			-12'38			+1'98			+0'88			+2'45			-6'54	+2'0		426	
427	-0'88		-11'29	+3'80		+1'71	+0'96		+0'81	-0'25		+2'60	+1'99		-5'56	+5'4	-3'2	427	
428			-13'04			+2'18			+0'94			+2'70			-6'66	-5'6		428	
429	-1'12	-10'48		+4'70	+1'45		+0'95	+9'10		-0'28	-0'11		+2'40	-8'71		-5'7	+1'9	429	
430			-11'05			+1'09			+0'77			+1'99			-6'59	-7'8		430	
431			-11'34			+1'35			+0'80			+2'05			-6'58	-7'5		431	
432			-11'55			+1'66			+0'82			+2'16			-6'43	-6'5		432	
433			-12'17			+1'78			+0'87			+2'18			-6'84	-8'8		433	
434			-12'55			+1'87			+0'90			+2'20			-7'09	-11'1		434	
435			-13'01			+2'03			+0'93			+2'28			-7'28	-13'4		435	
436			-14'28			+2'31			+1'02			+2'03			-8'54	-8'9		436	
437			-13'41			+2'27			+0'97			+2'58			-7'07	-13'2		437	
438			-13'79			+2'41			+0'99			+2'66			-7'20	-9'7		438	
439			-14'15			+2'51			+1'02			+2'75			-7'35	+2'3		439	
440	-1'87			+6'90			+0'94			-0'31			+3'29				+3'2	440	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

XCV.

in terms of any Spheroid.

Serial No.	FOR CHANGES OF AXES.						FOR CHANGES OF ORIGIN.						HELMERT'S SPHEROID*						Serial No.
	Case I: $\delta a = 1$ km			Case II: $\delta b = 1$ km			Case III: Latitude $u_0 = 1''$			Case IV: Azimuth $w_0 = 1''$			$a = 6378200$ metres, $1/e = 298.3$.						
	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	u	$v \cos \lambda$	$w \cot \lambda$	Deflection in Prime Vertical	Deflection in Meridian		
	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	"	
441	-1.88	-11.86		+6.93	+1.66		+0.94	+0.10		-0.31	-0.15		+3.37	-9.89		-6.0	+4.8	441	
442	-1.90		-16.69	+7.00		+3.38	+0.94		+1.21	-0.31		+3.04	+3.33		-8.61	-6.0	+4.4	442	
443			-19.85			+4.56			+1.45			+3.58			-9.88	+3.4		443	
444	-3.64			+11.05			+0.93			+0.33			+4.72				+3.5	444	
445			-22.77			+5.76			+1.67			+3.96			-11.13	+1.1		445	
446	-3.68		-22.91	+11.14		+5.82	+0.93		+1.68	-0.33		+3.98	+4.74		-11.19	+2.5	+3.6	446	
447	-3.69		-22.90	+11.16		+5.81	+0.93		+1.68	-0.33		+3.98	+4.75		-11.18	+1.1	+3.9	447	
448	-3.70		-23.01	+11.19		+5.87	+0.93		+1.69	-0.33		+3.99	+4.76		-11.23	-1.6	+4.3	448	
A d d e n d a .																			
449	+1.85			-4.49			+0.96			+0.26			-0.98				+11.4	449	
450			+5.92			-0.22			-0.40			+1.85			+7.58	+2.5		450	
451			+2.12			+0.08			-0.13			+1.63			+4.08	-10.0		451	
452			+1.91			+0.05			-0.09			+1.63			+3.87	-34.3		452	
453	+1.63		+2.07	-7.22		+0.06	+1.00		-0.10	+0.05		+1.63	-3.48		+4.03	+5.2	+18.3	453	
454	+1.62		+1.49	-7.12		+0.05	+1.00		-0.09	+0.05		+1.64	-3.42		+3.50	+4.4	+6.4	454	

* $\delta a = 0.924$, $\delta b = 0.743$, $u_0 = 0.31$, $w_0 = 1.29$. Vide p. 2.

To complete the statement of data concerning gravity values of the residuals, observed *minus* theoretical values are now given. These have been taken from Professional Paper No. 15 and for a full explanation reference should be made thereto. It is only necessary to explain that γ_a , γ_b , γ_c are the theoretical values of gravity based on Helmert's formula:— $\gamma_0 = 978.030 (1 + 0.005302 \sin^2 \phi - 0.000007 \sin^2 2\phi)$ and assuming the corrections according to the Free Air, Bouguer and Hayford hypotheses respectively.

TABLE XCVI.

No.	Name	Latitude	Longitude	Height	$g - \gamma_a - .011$	$g - \gamma_b + .030$	$g - \gamma_c - .011$
		° ' "	° ' "	feet	dynes	dynes	dynes
1	Agra ...	27 10	78 1	585	— .012	+ .011	+ .006
2	Aligarh ...	27 54	78 1	612	— .039	— .019	— .018
3	Allahābād ...	25 26	81 55	288	— .023	+ .008	— .002
4	Amgaon ...	21 22	80 28	1032	— .015	— .009	— .014
5	Amraoti ...	20 56	77 46	1123	+ .014	+ .017	+ .015
6	Arrah ...	25 34	84 39	188	— .067	— .032	— .039
7	Asarori ...	30 14	77 58	2467	— .061	— .101	...
8	Asirgarh ...	21 28	76 18	2077	+ .046	+ .023	+ .019
9	Badnūr ...	21 54	77 54	2103	+ .045	+ .015	+ .027
10	Bangalore ...	13 1	77 35	3118	+ .014	— .050	...
11	Bassein ...	16 47	94 44	23	+ .006	+ .046	...
12	Bhopāl ...	23 16	77 25	1630	+ .018	+ .004	+ .011
13	Bilāspur ...	22 4	82 12	878	— .006	+ .005	+ .002
14	Bina ...	24 11	78 12	1355	+ .015	+ .010	+ .015
15	Buxar ...	25 35	83 59	207	— .051	— .017	— .025
16	Chātra ...	24 13	88 23	64	— .025	+ .014	— .006
17	Colāba ...	18 54	72 49	34	+ .052	+ .092	+ .052
18	Cuttack ...	20 29	85 52	92	— .005	+ .033	— .005
19	Daltonganj ...	24 2	84 4	707	— .004	+ .013	+ .014
20	Damoh ...	23 50	79 26	1213	— .012	— .012	— .007
21	Darjeeling ...	27 3	88 16	6966	+ .044	— .124	...
22	Dehra Dūn ...	30 19	78 3	2239	— .085	— .115	— .005
23	Dera Ghāzi Khān ...	30 4	70 46	397	— .108	— .080	...
24	Dholpur ...	26 42	77 55	577	— .030	— .008	— .016
25	Edgar Shaft (Surface) ...	12 56	78 16	2945	+ .053	— .005	...
26	Ellichpur ...	21 18	77 31	1314	+ .019	+ .016	+ .020
27	Fatehpur ...	30 26	77 44	1434	— .085	— .089	...
28	Ferozepur ...	30 56	74 37	647	— .004	+ .015	...
29	Gaya ...	24 48	85 0	361	— .031	— .002	— .008
30	Gesupur ...	28 33	77 42	691	— .031	— .013	— .006
31	Goona ...	24 39	77 19	1569	+ .015	+ .003	+ .008
32	Gorakhpur ...	26 45	83 23	257	— .127	— .095	— .081
33	Gwalior ...	26 14	78 13	658	— .030	— .011	— .018

TABLE XCVI.—(Continued).

No.	Name	Latitude	Longitude	Height	$g-\gamma_s-.011$	$g-\gamma_s+.030$	$g-\gamma_s-.011$
		°	'	feet	dynes	dynes	dynes
34	Hardwār ...	29 56	78 9	949	-.117	-.106	...
35	Hāthras ...	27 37	78 3	587	-.020	+.001	000
36	Henzada ...	17 39	95 27	46	-.031	+.008	...
37	Hoshangābād ...	22 45	77 44	1002	.000	+.007	+.010
38	Jacobābād ...	28 17	68 27	183	+.003	+.038	+.027
39	Jalgaon ...	21 0	75 34	760	.000	+.015	+.009
40	Jalpaiguri ..	26 31	88 44	268	-.124	-.091	-.031
41	Japla ...	24 32	84 0	474	-.031	-.006	-.009
42	Jhānsi ...	25 27	78 34	858	-.004	+.008	+.003
43	Jubbulpore ..	23 9	79 59	1467	+.017	+.009	+.019
44	Kaliāna ...	29 31	77 39	810	-.065	-.051	-.018
45	Kaliānpur ...	24 7	77 39	1763	+.039	+.021	+.028
46	Kālka ...	30 50	76 56	2202	-.045	-.074	...
47	Kālsi ...	30 31	77 50	1684	-.084	-.090	...
48	Katni ...	23 50	80 26	1254	-.010	-.011	-.004
49	Kesarbāri ...	26 8	88 31	204	-.071	-.037	...
50	Khandwa ...	21 50	76 22	1014	+.033	+.010	+.036
51	Khurja ...	28 14	77 52	649	-.054	-.035	-.030
52	Kisnapur ...	25 2	88 28	113	+.001	+.038	+.028
53	Kodaikānal ...	10 14	77 28	7665	+.156	-.049	...
54	Kurseong ...	26 53	88 17	4913	-.011	-.117	...
55	Lalitpur ...	24 41	78 24	1199	-.016	-.015	-.013
56	Ludhiāna ...	30 55	75 51	835	-.053	-.040	...
57	Mach ...	29 52	67 18	3522	-.032	-.103	...
58	Madras ...	13 4	80 15	20	-.024	+.016	-.064
59	Maihar ...	24 16	80 48	1161	-.020	-.018	-.014
60	Majhāuli Rāj ...	26 18	83 58	219	-.105	-.071	-.068
61	Mandalay ...	22 0	96 6	244	-.028	+.006	...
62	Maymyo ...	22 1	96 28	3495	+.050	-.026	...
63	Meiktila ...	20 51	95 52	799	-.003	+.011	...
64	Meerut ...	29 0	77 42	734	-.035	-.019	...
65	Mhow ...	22 33	75 46	1903	-.002	-.025	-.026
66	Miān Mir ..	31 32	74 23	708	-.004	+.013	+.029
67	Moghal Sarai ...	25 17	83 6	257	-.040	-.008	-.016
68	Mogok ...	22 55	96 30	3685	+.063	-.020	...
69	Mohan ...	30 11	77 55	1660	-.081	-.093	...
70	Monghyr ...	25 23	86 28	154	-.067	-.031	-.036
71	Montgomery ...	30 40	73 6	557	-.011	+.011	+.008
72	Mortakka ...	22 13	76 3	576	-.022	.000	-.006

TABLE XCVI—(Continued).

No.	Name	Latitude	Longitude	Height	$g-\gamma_0-.011$	$g-\gamma_0+.030$	$g-\gamma_0-.011$
		° ' "	° ' "	feet	dynes	dynes	dynes
73	Mukhtiāra ...	22 24	75 59	926	-.039	-.029	-.030
74	Multān ...	30 11	71 26	404	-.066	-.039	...
75	Mussooree (Camel's Back)	30 28	78 5	6924	+.074	-.093	+.042
76	Mussooree (Dunseverick)	30 27	78 4	7129	+.076	-.098	...
77	Muttra ...	27 28	77 42	562	-.015	+.007	+.004
78	Muzaffarpur ...	26 7	85 25	179	-.091	-.056	-.053
79	Myingyan ...	21 29	95 24	248	-.020	+.013	...
80	Mysore ...	12 19	76 40	2501	+.003	-.040	...
81	Nojli ...	29 53	77 40	879	-.099	-.087	...
82	Ootacamund ...	11 25	76 42	7395	+.184	-.016	+.001
83	Pathānkot ...	32 17	75 39	1088	-.175	-.169	-.087
84	Pendra ...	22 47	82 0	1996	+.010	-.016	-.003
85	Prome ...	18 50	95 14	101	-.027	+.011	...
86	Pyinmana ...	19 44	96 12	409	-.014	+.014	...
87	Quetta ...	30 12	67 1	5520	+.020	-.123	-.004
88	Raipur ...	21 14	81 41	996	-.013	-.005	-.014
89	Rājpur ...	30 24	78 6	3321	-.051	-.113	+.015
90	Rāmchāndpur ...	25 41	88 33	132	-.031	+.006	...
91	Rānchi ...	23 23	85 19	2167	+.040	+.008	+.019
92	Rangoon ...	16 48	96 9	164	+.010	+.045	...
93	Roorkee ...	29 52	77 54	867	-.112	-.099	-.055
94	Salem ...	11 40	78 9	948	-.047	-.038	-.059
95	Sandakphu ...	27 6	88 0	11766	+.178	-.125	+.037
96	Sasarām ...	24 57	83 59	340	-.025	+.005	-.002
97	Saugor ...	23 52	78 48	1757	+.010	-.008	.000
98	Seoni ...	22 5	77 29	2032	+.041	+.014	+.025
99	Shāhpur ...	22 12	77 54	1286	+.006	+.004	+.012
100	Sibi ...	29 33	67 53	434	-.137	-.109	-.070
101	Siliguri ...	26 42	88 25	387	-.160	-.130	-.050
102	Simla ...	31 6	77 10	7043	+.080	-.100	...
103	Sipri ...	25 26	77 39	1533	+.027	+.016	+.018
104	Sultānpur ...	26 16	82 5	314	-.064	-.034	...
105	Toungoo ...	18 56	96 27	159	-.011	+.025	...
106	Ujjain ...	23 11	75 47	1612	-.013	-.026	-.022
107	Umaria ...	23 32	80 54	1499	+.016	+.007	+.018
108	Yercaud ...	11 47	78 12	4493	+.072	-.027	-.044

CHAPTER X.

Deflections of the Plumb-line and values of "g" derived in Turkistan (Ferghana) by the Russian observations.

1. The problem of the origin of the Himalayas has already been attacked from the geodetic point of view making use of the deflection results and gravity anomalies obtained in India. All these results relate to points south of the main chain. It is now possible to put certain results obtained by the Russian surveyors into the same terms: and it has accordingly been considered suitable to do this, so that all geodetic data closely related to the Himalayas will be conveniently available in one volume.

2. Values of deflection can be found for Osh base, N.W. end, the starting point of the Russian triangulation which links up with the Indian Pamir triangulation. The Russian triangulation emanates from a base near Osh latitude $40^{\circ} 31'$, longitude $42^{\circ} 30'$ E. of Pulkowa. Latitude and azimuth were observed astronomically while the longitude of Osh was found by electric telegraph from Tashkent, and transferred from Osh to the north-west end of the base by chronometer. Presumably the longitude of Tashkent is in terms of astronomical longitude measured from Pulkowa (Poulcovo) which is $2^h 1^m 18^s \cdot 57$ of time E. of Greenwich (*vide* Nautical Almanac). This converted becomes $30^{\circ} 19' 38'' \cdot 55$ which must be added to the values of longitudes expressed in Russian terms. Calculations of the Russian triangulation were performed on Bessel's spheroid. Suppose the deflections at Osh are ξ, η in prime vertical and meridian (positive if S. or W.). Then the geodetic elements are found by deducting $\eta = -u, \xi \sec \lambda = -v, \xi \tan \lambda = -w$ from the astronomic values of the latitude, longitude and azimuth respectively. The Astronomic elements of the N.W. end Osh base are:—

Latitude	$40^{\circ} 37' 16'' \cdot 67$
Longitude	$72^{\circ} 56' 11'' \cdot 17^*$
Azimuth	not stated.

From this astronomic origin and a measured base triangulation was carried to a station Kukhtek of the Indian triangulation, the deduced latitude being $37^{\circ} 17' 43'' \cdot 94$ and the longitude $74^{\circ} 59' 55'' \cdot 53^*$.

It is necessary to express these in terms of the geodetic value of Osh and the Helmert spheroid. This can be done approximately by means of the tables already prepared with reference to Kalianpur as origin. Kukhtek is $2^{\circ} 3' 44'' \cdot 36$ east of the Osh origin, and as the tables prepared for Kalianpur are shown in absolute longitude the corresponding longitude required is that of Kalianpur increased by this i.e. $77^{\circ} 39' 17'' \cdot 57 + 2^{\circ} 3' 44'' \cdot 36 = 79^{\circ} 43' 2'' \div 79^{\circ} \cdot 7$.

* This includes $30^{\circ} 19' 38'' \cdot 55$, the difference of longitude of Greenwich and Pulkowa.

3. Suppose corresponding to changes u, v, w at Osh, changes u', v', w' occur at Kukhtek and that both are due to imaginary changes u_0, v_0, w_0 at an origin O on the longitude of Osh and at the latitude of Kalianpur together with a change of axes from those of Bessel to those of Helmer: for which $\delta a = +.803$ and $\delta b = +.739$ (*vide* Appendix).

From tables XVII-XX a change at O of u_0, v_0 and w_0 causes changes of $u_0, v_0 + .370 w_0, 1.20 w_0$ at Osh and from the axes change the further changes (from tables XXIX-XXXIV) at Osh are $+.803 \times 1.24 - .739 \times 10.56, .803 \times 0 + .739 \times 0, .803 \times 0 + .739 \times 0$ i.e. $-6.808, 0, 0$. The total changes at Osh accordingly are $u_0 - 6.808, v_0 + .370 w_0, 1.20 w_0$. The changes at Kukhtek, $2^\circ 3' 44''$ E. of meridian origin and latitude $37^\circ.3$ may be found in the same tables under longitude $79^\circ.7$.

$$\begin{array}{l} \text{They are} \\ \left. \begin{array}{l} u_0 - .032 w_0 \\ + .803 \times 1.626 - .739 \times 9.033 \\ 1.806 - 6.675 \end{array} \right| \left. \begin{array}{l} v_0 + .027 u_0 + .284 w_0 \\ -.803 \times 1.525 + .739 \times .191 \\ -1.225 + .141 \end{array} \right| \left. \begin{array}{l} w_0 + .045 u_0 + 1.145 w_0 \\ -.803 \times .752 - .739 \times .056 \\ -.604 - .041 \end{array} \right| \end{array}$$

which reduce to $u' = u_0 - .032 w_0 - 5.369$ $v' = v_0 + .027 u_0 + .284 w_0 - 1.084$ $w' = w_0 + .045 u_0 + 1.145 w_0 - .645$

Now at Osh

$$u = u_0 - 6.808, v = v_0 + .370 w_0, w = 1.2 w_0$$

In terms of vertical deflection $u = -\eta, v = -\xi \sec \lambda, w = -\xi \tan \lambda$ which determine u_0, v_0, w_0 in terms of ξ and η as follows:—

$$u_0 = -\eta + 6.808, w_0 = -\frac{1}{1.2} \xi \tan \lambda, v_0 = -\xi \sec \lambda + .308 \xi \tan \lambda \quad \text{when } \lambda = 40^\circ 37'$$

$$u_0 = -\eta + 6.808, w_0 = -.715 \xi, \quad v_0 = -1.317 \xi + .264 \xi = -1.053 \xi$$

Hence $u' = -\eta + 6.808 + .023 \xi - 5.369, v' = -1.053 \xi + .027 (-\eta + 6.808) - .203 \xi - 1.084$
and $w' = -.715 \xi + .045 (-\eta + 6.808) - .818 \xi - .645$
 $= u' = -\eta + -.023 \xi + 1.439, v' = -1.256 \xi - .027 \eta - .900, w' = -1.533 \xi - .045 \eta - .339.$

These are the corrections which have to be applied to the Russian values of coordinates of Kukhtek, to bring them into terms of the Indian triangulation.

The Indian values have to be corrected to bring into terms of the Helmert spheroid and the observed values of the elements at Kalianpur, corresponding to $u_0 = 31, v_0 = 0^*, w_0 = 1.29, \delta a = .924, \delta b = .743$. Denoting the corrections by u'', v'', w'' the corrections at Kukhtek latitude $37^\circ.3$ and longitude $75^\circ.0$ are:—

$$\begin{array}{l} u'' = .31 \times .999 + 1.29 \times .043 \\ + .924 \times 1.590 - .743 \times 9.032 \\ = .310 + .055 + 1.470 - 6.711 \\ = -4.877 \end{array} \left| \begin{array}{l} v'' = -.31 \times .036 + 1.29 \times .282 \\ + .924 \times 1.980 - .743 \times .247 \\ = -.011 + .364 + 1.830 - .184 \\ = +1.999 \end{array} \right| \left| \begin{array}{l} w'' = -.31 \times .058 + 1.29 \times 1.144 \\ + .924 \times .979 + .743 \times .072 \\ = -.018 + 1.476 + .905 + .053 \\ = +2.416 \end{array} \right|$$

The values obtained by the Indian triangulation for latitude and longitude of Kukhtek are:—

	37° 17' 32".97		75° 0' 12".19
u''	-4.88	v''	+2.00
Values in terms of Helmert spheroid	37 17 28.09		75 0 14.19
Values obtained by Russian triangulation are	37° 17' 43".94		74° 59' 55".53

and to bring these into accord with the Indian values it is necessary to apply $-15''.85$ to the latitude and $+18''.66$ to the longitude.

* This is zero because the deflection at Kalianpur in prime vertical had never been taken account of, although it was implied by the values of azimuths adopted.

The following equations are formed giving the quantities ξ and η at Osh:—

$$\begin{aligned} -\eta + \cdot 023 \xi + 1\cdot439 &= -15\cdot85 \\ -\cdot 027 \eta - 1\cdot256 \xi - \cdot 900 &= +18\cdot66 \end{aligned}$$

whence

$$\begin{aligned} \eta &= + \cdot 023 \xi + 17\cdot29 \\ 1\cdot256 \xi &= - \cdot 027 \eta - 19\cdot56 \\ &= - \cdot 001 \xi - \cdot 467 - 19\cdot56 \\ \xi &= -\frac{20\cdot03}{1\cdot257} = -15\cdot94 \\ \eta &= 17\cdot29 - \cdot 37 = +16\cdot92 \end{aligned} \quad \left. \vphantom{\begin{aligned} \eta \\ \xi \\ \eta \end{aligned}} \right\} \text{ at Osh.}$$

The values found by the Russian observers in terms of Tashkent vertical are $\xi = -6\cdot04$ and $\eta = 23\cdot43$, *vide* table *XCVIII*.

4. It is possible to bring these results into agreement with the values deduced from the Indian side by supposing that the vertical at Tashkent, latitude $41^\circ 21'$ and longitude $69^\circ 18'$, is deflected with reference to Kalianpur.

Consider the effect of changes u_0, v_0, w_0 at Tashkent longitude and Kalianpur latitude, and from Bessel to Helmert spheroid for which $\delta a = +803$, $\delta b = +739$.

From tables *XVII-XX* a change at origin of u_0, v_0, w_0 causes changes of $u_0, v_0 + \cdot 372 w_0, 1\cdot204 w_0$ at Tashkent and from axes changes the further changes from tables *XXIX-XXXIV* at Tashkent are $-7\cdot962, 0, 0$. The total changes at Tashkent accordingly are $u_0 - 7\cdot962, v_0 + \cdot 372 w_0, 1\cdot204 w_0$.

The total changes at Osh $\lambda 40^\circ 37', L 81^\circ 17', \{ = 77^\circ 39' + (72^\circ 56' - 69^\circ 18') \}$ calculated from the same tables are:—

$$+ \cdot 998 u_0 - \cdot 059 w_0 - 6\cdot791 \mid v_0 + \cdot 052 u_0 + \cdot 357 w_0 - 2\cdot015 \mid + \cdot 081 u_0 + 1\cdot197 w_0 - 1\cdot283$$

The following equations are formed:—

$$+ \cdot 998 u_0 - \cdot 059 w_0 - 6\cdot791 = u = 6\cdot5 \quad (1)$$

$$v_0 + \cdot 052 u_0 + \cdot 357 w_0 - 2\cdot015 = v = 9\cdot9 \sec \lambda = 13\cdot042 \quad (2)$$

$$+ \cdot 081 u_0 + 1\cdot197 w_0 - 1\cdot283 = w = 9\cdot9 \tan \lambda = 8\cdot490 \quad (3)$$

in which the numerical quantities 6·5 and 9·9 are the corrections necessary to ξ and η as determined by the Russian observers, to bring them into agreement with the Kalianpur terms.

Hence

$$u_0 = \cdot 059 w_0 + 13\cdot318 \quad (1)$$

$$1\cdot197 w_0 = - \cdot 081 u_0 + 9\cdot773 \quad (3)$$

$$= - \cdot 005 w_0 - 1\cdot079 + 9\cdot773$$

$$\text{or } 1\cdot202 w_0 = 8\cdot694$$

$$\text{or } w_0 = 7\cdot23$$

$$\therefore u_0 = \cdot 427 + 13\cdot318 = 13\cdot75 \quad (1)$$

$$\text{and } v_0 = - \cdot 715 - 2\cdot581 + 2\cdot015 + 13\cdot042 = 11\cdot76 \quad (2)$$

5. With these origin changes the corrections at certain degree squares have been computed from the tables and the results are exhibited in table *XCVII*.

TABLE *XCVII*.

	$\lambda \backslash L$	69°	70°	71°	72°	73°
u	40°	7·20	7·09	6·98	6·87	6·76
v		14·41	14·05	13·70	13·34	12·99
w		8·63	8·59	8·55	8·52	8·48
u	41°	6·78	6·67	6·56	6·45	6·34
v		14·59	14·24	13·89	13·54	13·20
w		8·75	8·71	8·66	8·62	8·57

6. The results of the Russian observations* are now given first in terms of Tashkent and Bessel's spheroid and then corrected to the Kalianpur vertical and Helmert's spheroid. The stations are shown in chart No. VI.

TABLE XCVIII.

Serial No.	Station	Latitude	Longitude† in Greenwich Terms	Bessel's Spheroid and Tashkent Vertical		Corrections		Helmert's Spheroid and Kalianpur Vertical	
				Deflection in Prime Vertical	Deflection in Meridian	-u	-u cot λ	Deflection in Prime Vertical	Deflection in Meridian
	Tashkent ...	41° 21'	69° 18'	0"00	0"00	"	"	- 9"9	- 6"5
1	Ohodschent ...	40 17	37	+ 13.06	+ 3.83	-6.94	-10.12	+ 2.9	- 3.1
2	Karatschekum ...	16	70 5	+ 3.71	+ 22.52	-6.89	-10.09	- 6.4	+ 15.6
3	Kanybadam ...	19	26	- 4.63	+ 22.66	-6.83	-10.05	-14.7	+ 15.8
4	Tschil-Machram ...	33	33	...	+ 2.36	-6.72	- 4.4
5	Begowat (south) ...	19	43	...	+ 42.82	-6.80	+ 36.0
6	Puntan ...	44	43	...	- 8.13	-6.62	-14.8
7	Sary-Kurgan ...	20	71 2	+ 9.47	+ 30.34	-6.77	-10.08	- 0.6	+ 23.6
8	Pap ...	54	4	...	-20.89	-6.53	-27.4
9	Begowat (north) ...	38	14	...	+ 8.92	-6.60	+ 2.3
10	Karail-tjube ...	31	15	+ 18.14	+ 16.64	-6.65	- 9.97	+ 8.2	+ 10.0
11	Waryk ...	41 7	16	...	-26.87	-6.40	-33.3
12	Katput ...	40 16	20	...	+ 41.13	-6.76	+ 34.4
13	Gurt-tjube ...	50	28	...	- 4.13	-6.51	-10.6
14	Kassan ...	41 15	36	...	-20.30	-6.31	-26.6
15	Chalmion ...	40 11	39	...	+ 49.41	-6.76	+ 42.7
16	Namangan ...	41 0	41	- 2.94	- 9.73	-6.41	- 9.83	-12.8	-16.1
17	Martelan ...	40 23	47	+ 3.53	+ 32.56	-6.65	- 9.99	- 6.5	+ 25.9
18	Bjälaja ...	41 4	50	...	-13.08	-6.37	-19.5
19	Kara-tjube ...	40 37	50	...	+ 17.17	-6.57	+ 10.6
20	Balyktschi ...	53	52	...	- 1.87	-6.46	- 8.3
21	Utsch-Kurgan I ...	41 6	72 3	...	-10.63	-6.33	-17.0
22	Utsch-Kurgan II ...	40 14	4	...	+ 41.63	-6.70	+ 34.9
23	Kuwa ...	31	4	+ 0.42	+ 28.40	-6.58	- 9.94	- 9.5	+ 21.8
24	Tasch-tjube ...	40	10	...	+ 18.97	-6.51	+ 12.5
25	Tschumbagysch ...	54	13	...	- 0.38	-6.41	- 6.8
26	Isbasken ...	41 2	21	...	-12.36	-6.34	-13.7
27	Min-tjube ...	40 29	22	- 6.84	+ 33.74	-6.55	- 9.94	-16.8	+ 27.2
28	Andischan ...	47	24	...	+ 12.63	-6.43	+ 6.2
29	Kisyl-Kurgan ...	20	24	...	+ 32.39	-6.63	+ 25.8
30	Salb-tjube ...	38	34	...	+ 25.85	-6.50	+ 19.4
31	Massy ...	41 5	39	...	-15.71	-6.27	-22.0
32	Ohodscha-Syrjan ...	40 43	43	...	+ 11.49	-6.38	+ 5.1
33	Tjulka-tjube ...	18	46	...	+ 32.71	-6.59	+ 26.1
34	Osh ...	31	49	- 6.04	+ 23.43	-6.50	- 9.90	-15.9	+ 16.9
35	Chasret Ujunys ...	46	56	...	+ 13.70	-6.38	+ 7.3
36	Mady ...	34	56	...	+ 22.80	-6.47	+ 16.3
37	Dschalabad ...	55	73 1	...	+ 4.38	-6.32	- 1.9

* c. f. Comptes-Rendus de L' Association Geodesique Internationale for 1898 (Annexe A IIc, p. 268).

† Converted from Pulkowa Longitude by applying +30° 20' (more accurately 30° 19' 38".55).

The following description is extracted from *Comptes-Rendus de L' Association Geodesique Internationale* for 1896 (Annexe B XI p. 309):—

“The researches recently completed on the deviation of the plumb-line in Ferghana (Turkistan) are of special interest. This valley lying between $40^{\circ} 15'$ and $41^{\circ} 15'$ in latitude and $39^{\circ} 30'$ and $42^{\circ} 45'$ in longitude, east of Pulkowa is a deep depression the walls of which are pierced in their western part by the narrow bed of the Syr Daria. The bottom of the valley has an approximately elliptic figure with its major axis 250 km. long following the direction of the parallel, and its minor axis 110 km. that of the meridian. On the north the valley is enclosed by chains of mountains of an average height of 2500 to 3500 m. and on the south are the Alai, the Trans Alai, the Pamirs and the Hindu Kush. Such a position leads one to expect considerable deviation particularly in latitude and explains the investigations which have been made in order to verify this supposition. To this end 37 determinations of latitude have been made of points equally distributed over the district and 10 of longitude at points nearly on the same parallel. Taking Bessel's Ellipsoid, and the point Balyktschi as zero, we obtain for the deviation in latitude values for A—G from $-25''$ on the north up to $+51''$ on the south of the valley, in longitude the deviation of opposite sign amounts to $25''$ ”.

7. A table of gravity residuals in the same district* is also given for the sake of completeness.

TABLE XCIX.

	No.	Station	Latitude	Longitude	Height in Metres	$\gamma_0 - \gamma_0$ Cm 10-3 x
TURKISTAN	1	Pamir Post ...	$38^{\circ} 10' 0''$	$73^{\circ} 58' 2''$	3700	- 30
	2	Kala-i-Wanj ...	$38^{\circ} 22' 2''$	$71^{\circ} 27' 0''$	1795	- 177
	3	Sar-i-pul ...	$38^{\circ} 24' 5''$	$70^{\circ} 5' 5''$	1500	-100
	4	Kala-i-Chamb ...	$38^{\circ} 27' 3''$	$70^{\circ} 46' 5''$	1345	-152
	5	Rabat Ak Baital ...	$38^{\circ} 29' 7''$	$73^{\circ} 51' 5''$	4100	+ 74
	6	Rabat Maskol ...	$38^{\circ} 42' 0''$	$73^{\circ} 31' 7''$	4200	+169
	7	Kara Kul Lake ...	$39^{\circ} 6' 4''$	$73^{\circ} 31' 2''$	3920	+ 35
	8	Irkeshtan Fort ...	$39^{\circ} 41' 9''$	$73^{\circ} 55' 5''$	2850	- 56
	9	Ak-bossaga ...	$39^{\circ} 48' 6''$	$73^{\circ} 13' 7''$	2875	-128
	10	Sufi Kurgan ...	$40^{\circ} 1' 5''$	$73^{\circ} 30' 0''$	2115	- 91
	11	Karaul Kishlak ...	$40^{\circ} 2' 2''$	$72^{\circ} 6' 0''$	1800	-157
	12	Gultsha ...	$40^{\circ} 19' 0''$	$73^{\circ} 25' 7''$	1583	-126
	13	New Marghilan ...	$40^{\circ} 23' 7''$	$71^{\circ} 46' 7''$	581	-159
	14	Langar ...	$40^{\circ} 24' 6''$	$73^{\circ} 5' 7''$	1685	- 67
	15	Osh ...	$40^{\circ} 31' 4''$	$72^{\circ} 46' 6''$	1021	-106
	16	Andijan ...	$40^{\circ} 45' 8''$	$72^{\circ} 20' 6''$	530	-185
	17	Tashkent ...	$41^{\circ} 19' 5''$	$69^{\circ} 17' 7''$	478	- 50
	18	Wysokoji Khojand ...	$42^{\circ} 30' 9''$	$70^{\circ} 33' 9''$	1060	- 1
	19	Chodient ...	$40^{\circ} 17' 1''$	$69^{\circ} 34' 7''$	320	-140
	20	Namangan ...	$40^{\circ} 59' 7''$	$71^{\circ} 38' 7''$	440	-178

* c. f. C. R. 1911—Volume III pp. 156-158.

APPENDIX.

Various Determinations of the Axes of the Earth.

For convenience of reference the principal values of the elements of the figure of the earth obtained from time to time are given below, expressed in units of 1000 feet and kilometers. It is to be observed that the datum of height in different continents is only in the same terms on the assumption that the geoid is identical with the spheroid. The quantities determined really refer to the several concentric spheroids through these sea level datum points.

1. RADIUS OF THE EARTH CONSIDERED AS A SPHERE.

Reference No.	Authority	Date	RADIUS		Data used
			in 1,000 feet	in kilometres	
1	Eratosthenes ...	250 B. C.	24370	7428	Arc from Syene (Upper Egypt) to Alexandria.
2	Posidonius ...	80 "	23190	7069	
3	Richard Norwood...	1637 A. D.	21038	6412	Arc from London to York. Mean Lat. 52° 42' 30".
4	Jean Picard ...	1689 "	20906	6372	Arc from Paris to Amiens. Mean Lat. 49° 30'.

2. AXES OF THE EARTH CONSIDERED AS A SPHEROID.

Jean Richer (d. 1696) pointed out that the Earth was not a sphere.

Reference No.	Authority	Date	RADIUS OF CURVATURE IN MEAN LATITUDE OF ARC		DEDUCED		Data used	
			in 1,000 feet	in kilometres	$\frac{a}{\sin^2 \phi}$ in 1,000 feet	$\frac{1}{e}$		
5	J. and D. Cassini ...	1684-1718	20880.2	6360.8			Arc from Paris to Dunkirk. Mean Lat. 49° 56' 9".	The immediate inference was that the degree diminishing with the increasing latitude, the Earth must be a prolate spheroid. Bouguer, De la Condamine, Maupertuis, Clairault, De Thury and De Lacaille proved that the Earth was an oblate and not a prolate spheroid.
6	J. and D. Cassini	20919.4	6376.1			Arc from Paris to Collioure. Mean Lat. 45° 40' 42".	
7	Bouguer and De la Condamine	1735-1751	20795.4	6338.3	20988.5 (= 6397.3 km)	216.82	Arc in Peru. Mean Lat. 1° 31' 0" S.	
8	Maupertuis and Clairault	1736	21038.4	6412.4			Arc in Finland. Mean Lat. 66° 18' 35".	
9	De Lacaille ...	1752	20697.4	6369.4			Arc at Cape of Good Hope. Mean Lat. 33° 18' 30" S.	

8. AXES OF THE EARTH CONSIDERED AS A SPHEROID DETERMINED FROM A GROUP OF ARCS, GRAVITY ETC.

Quantities given by authority named are shown in roman figures; deduced quantities are in italics.

Reference No.	Authority	Date	SEMI MAJOR AXIS (=a)		SEMI MINOR AXIS (=b)		$\frac{1}{e}$	Data used etc.
			in 1,000 feet	in kilometres	in 1,000 feet	in kilometres		
10	Laplace	1799	20919.768	6376.840	20862.222	6355.935	812.20	Arcs in Peru, India, France, England, and Sweden.
11	Everest	1830	22.84095	77.876	58.28403	56.075	800.8017	Arc from Damargida to Kalianpur Mean Lat. $21^{\circ} 5' 13''$.
			(22.93180)	...	(58.37458)	As expressed by Everest in terms of Indian 10-foot bar A (=9.99985658 feet).
12	Airy	1830	23.713	76.842	58.810	56.286	299.33	14 meridian arcs and 4 arcs of parallel.
13	Bessel	1841	23.237 \pm .709	77.397 \pm .214	58.296	56.079	299.15	From 10 meridian arcs.
14	Clarke	1857	26.348	78.845	55.233	56.669	294.26	Arcs: Anglo-Gallic, Russian, Indian, Prussian, Peruvian, Hanoverian, Danish.
15	Pratt	1863	26.189	78.297	55.316	56.695	295.26	Semi axes and ellipticity of the Mean Figure of the Earth. From a comparison of the Anglo-Gallic, Russian and Indian arcs &c.
16	Clarke	1866	26.062	78.253	55.121	56.635	294.98	Arcs: Anglo-Gallic (rejecting 21 lat. stations), 2nd Indian, Russian, Peruvian, Cape.
17	Clarke	1880	26.202	78.801	54.695	56.871	293.47	
18	Clarke-Bessel	...	26.202	78.801	56.282	[56.980]	299.15	
19	Clarke-Bessel	...	26.239	78.190	55.888	[56.869]	299.15	From Clarke's value of 1866 with an old conversion factor.
20	Darwin	1889	296.4	From consideration of precession.
21	25.329	78.085	55.381	56.715	299.15	(1.0001) \times Bessel's value; adopted by the Central Bureau, International Geodetic Association.
22	Hayford (O.and G.S.)	1906	26.144 \pm .112	78.283 \pm .034	55.885	56.868	297.8 \pm 0.9	
23	Helmert	1907	26.271	78.200	55.721	56.818	296.3	From gravity determinations.
24	Hayford (O.and G.S.)	1909	26.488 \pm .059	78.388 \pm .018	56.019	56.909	297.0 \pm 0.5	Adopted for International $\frac{1}{M}$ map.
25	Helmert-Hayford	...	26.436	78.372	55.976	56.896	297	Obtained by S. Wallisch taking Helmert's value with weight=unity and the modified Hayford values with weight=4.
26	International Map Committee	1909	26.002	78.24	54.874	56.56	[294.2]	
27	E.W. Brown	1914	298.7 \pm 0.3	From Lunar theory. The value will make the observed motions of perigee and node agree with the theoretical values.
28	Nautical Almanac	1911	297	Adopted in the conference of Nautical Almanac directors.
29	Crommelin	294.4 \pm 1.5	From Moon's parallax at Greenwich and Cape. Obtained by a hundred pairs of simultaneous observations at the Cape and Greenwich Observatories by a comparison between theoretical and observed values of the Moon's parallax.

4. AXES OF THE EARTH CONSIDERED AS AN ELLIPSOID.

Reference No.	Authority	Date	Equatorial Axes				Polar Axis		Longitude of major axis E. of Greenwich	Data used
			a		b		c			
			in 1,000 feet	in kilometres	in 1,000 feet	in kilometres	in 1,000 feet	in kilometres		
80	Schubert ...	About 1860	20927.397 (3272671 toises)	6378.665	20925.044 (3272308 toises)	6377.948	20855.759 (3261468 toises)	6356.880	41° 4'	Arcs: French, English, Russian, Indian, Cape, Prussian and Peruvian.
81	Clarke ...	1880	„ 26.629	„ 78.481	„ 25.105	„ 77.966	„ 53.477	„ 56.499	15° 34'	Arcs: French, English, Indian, Russian, Prussian, Peruvian, Cape.

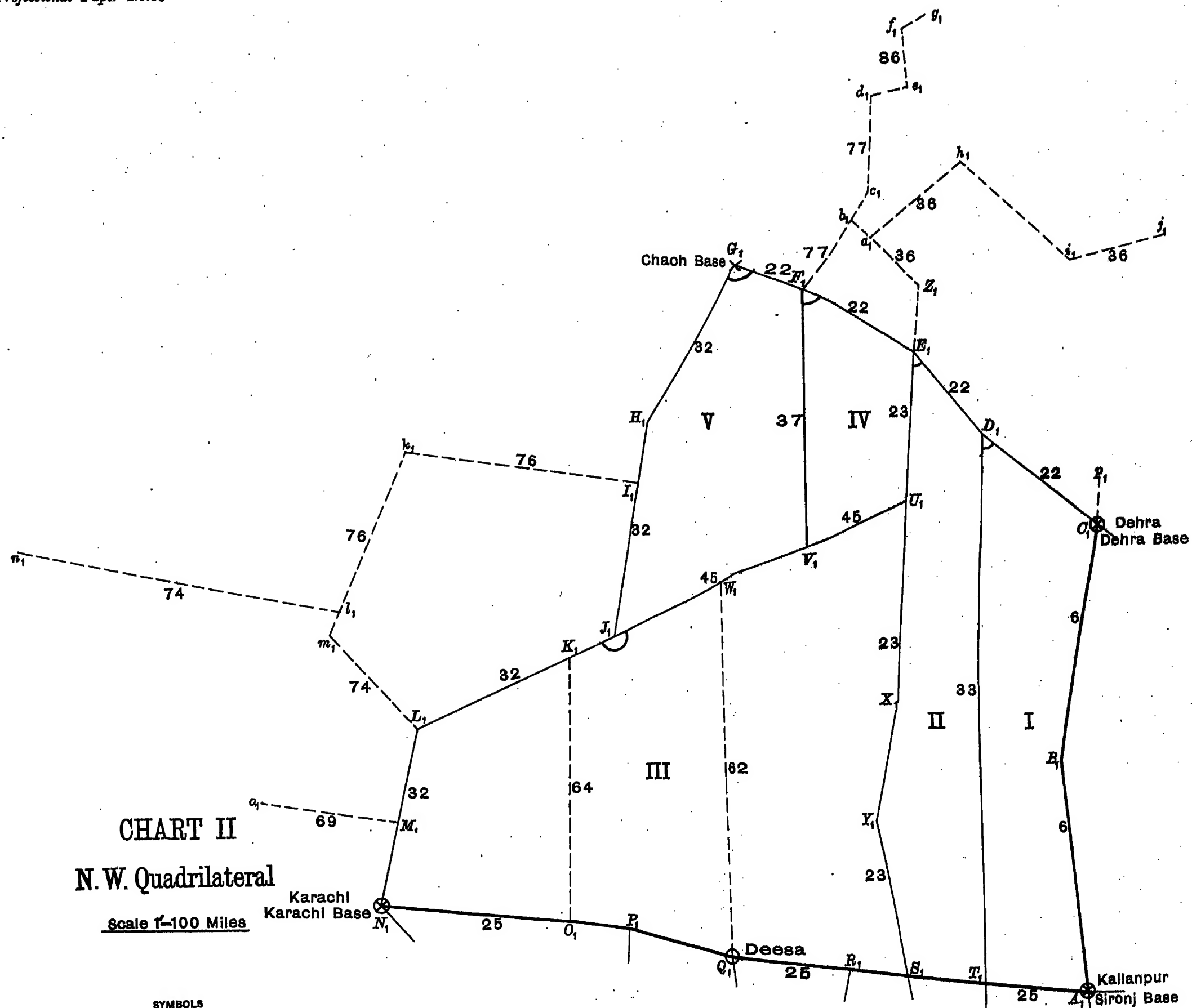
REFERENCES.

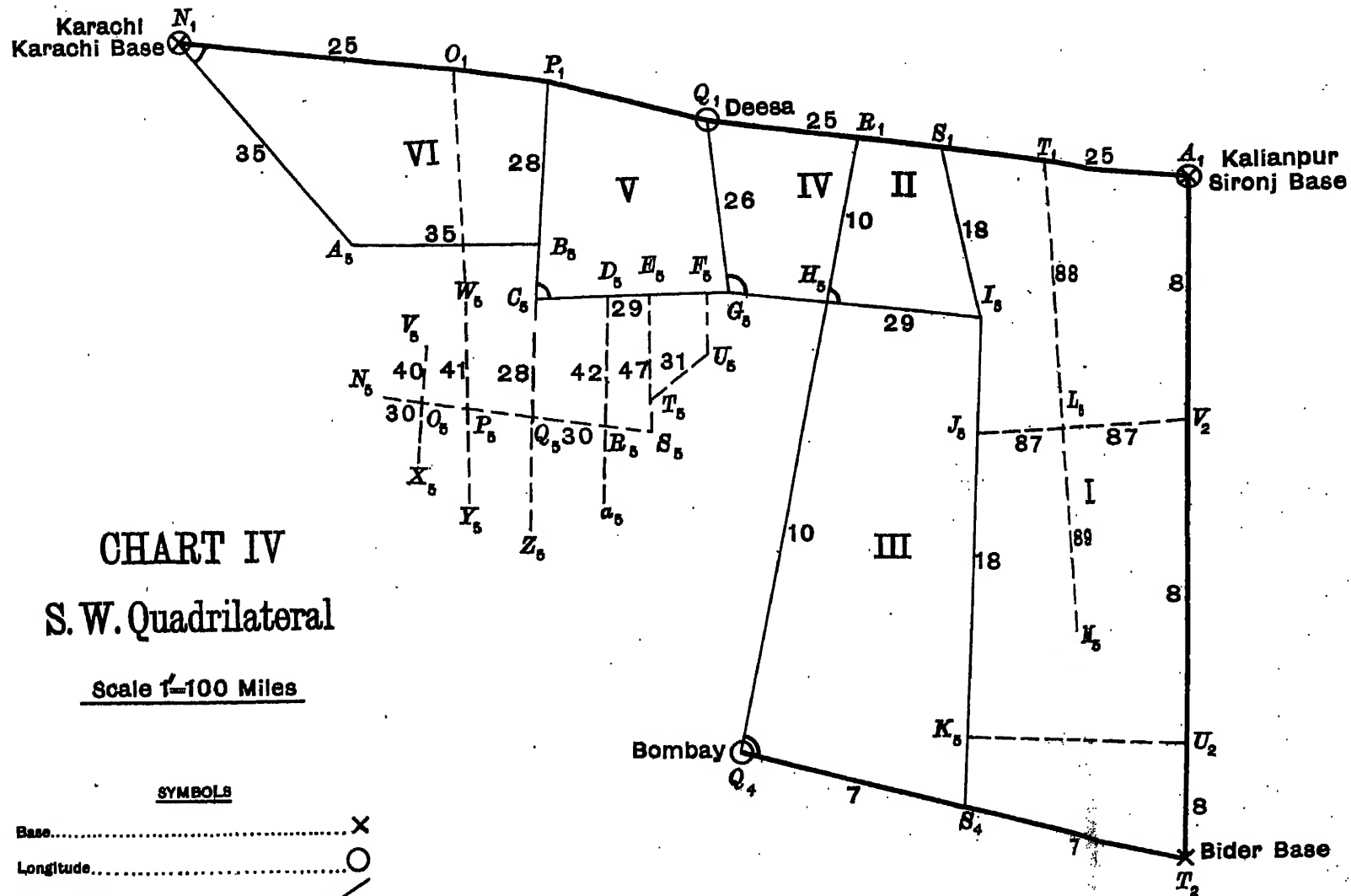
1. Earth's circumference was reckoned to be 250,000 stadia (1 stadia = 50,563 feet) *Vide Encyclopædia Metropolitana* 1848, Art. *Figure of the Earth*; also *Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*.
2. Earth's circumference = 240,000 stadia.—*Ibid*.
3. 1° = 367176 feet.—*Ibid*.
4. 1° = 57060 toises.—*Ibid*.
5. 1° = 56960 „ *Ibid*.
6. 1° = 57097 „ *Ibid*.
7. 1° = 56767 „ *Ibid*.
8. 1° = 57183 „ *Vide Encyclopædia Metropolitana* 1848, Art. *Figure of the Earth*.
9. 1° = 57037 „ *Vide Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*.
10. *Account of the Observations and Calculations of the Principal Triangulations &c.* p. 733 by Capt. A. R. Clarke, R. E., F. R. A. S.
11. *Vide* pp. 1, 2
12. *Account of the Observations and Calculations of the Principal Triangulations &c.* p. 734 by Capt. A. R. Clarke, R. E., F. R. A. S. also *Geodesy* p. 303 by Col. A. R. Clarke, C. B.
13. *Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*, p. 807; also *The Figure of the Earth and Isostasy from measurements in the United States* 1909, p. 172 by John F. Hayford.
14. *Geodesy* 1880, p. 305 by Col. A. R. Clarke, C. B.; also *Account of the Observations and Calculations of the Principal Triangulations &c.* p. 771 by Capt. A. R. Clarke, R. E., F. R. A. S.
15. *A Treatise on Attraction, Laplace's Function, and the Figure of the Earth*, 1868, p. 134 by John H. Pratt, M.A.
16. *Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*, p. 808; also *Smithsonian Geographical Tables* (1906) p. XLIII.
17. *Geodesy* 1880, p. 319 by Col. A. R. Clarke, C. B.; also *Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*, p. 808.
18. Semi major axis of Clarke, 1880 (*Vide* 17) and ellipticity of Bessel (*Vide* 13).
19. *Philosophical Transactions*, Series A, Vol. 205 pp. 289-318—*On the Intensity and Direction of the Force of gravity in India* by Lieut-Colonel S. G. Burrard, R. E., F. R. S.
20. *Proceedings of the Royal Society*, Vol. LX, p. 119—*The Theory of the Figure of the Earth*.
21. *Encyclopædia Britannica* 11th Edition, Vol. 8, Art. *Figure of the Earth*, p. 813.
22. *The Figure of the Earth and Isostasy from measurements in the United States* 1909, p. 172 by John F. Hayford.
23. *Vide* p. 3; also *The Figure of the Earth and Isostasy from measurements in the United States* 1909, p. 173 by John F. Hayford.
24. *Supplementary Investigation in 1909 of the Figure of the Earth and Isostasy* 1910, p. 77 by John F. Hayford.
25. *Vide Nature* No. 2411, Vol. 96, January 13, 1916.
26. *Vide Resolutions and Proceedings of the International Map committee assembled in London*, November, 1909.
- 27, 28 & 29. *Vide Opening Address by Prof. E. W. Brown, F. R. S. at the Australian Meeting of the British Association*,—*Nature* No. 2346, Vol. 94, Oct. 15, 1914.
30. *Memoirs of the Royal Astronomical Society* 1859-60 Vol. XXIX, p. 25; also *Monthly Notices of Royal Astronomical Society*, Vol. XX, pp. 104-107.
31. *Geodesy* 1880, p. 308 by Col. A. R. Clarke, C. B.; also *Philosophical Magazine*, August 1878.

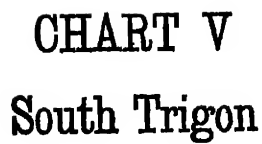
For some other values see *La Figure de La Terre*, Paris, 1901, by Capt. G. Perrier.

Conversion factors used— *1 metre = 3.28084275 feet. log = 0.5159854152
or 1 foot = 0.30479973 metre. „ = 1.4840145780
1 toise = 6.39459252 feet „ = 0.8058128753

* *Vide "Determination du Rapport du yard au metre"* by M. Benoit, Paris, 1896; also *Text Book of Topographical and Geographical Surveying*, p. 359 by Colonel C. F. Close, C. M. G., R. E.







SYMBOLS

Baro.....X

Longitude 0

Closing Point.....

Rm.No.7 D.D.1917 (2 D.O.)-8.Ir420